

# Cosmic Rays in IceCube

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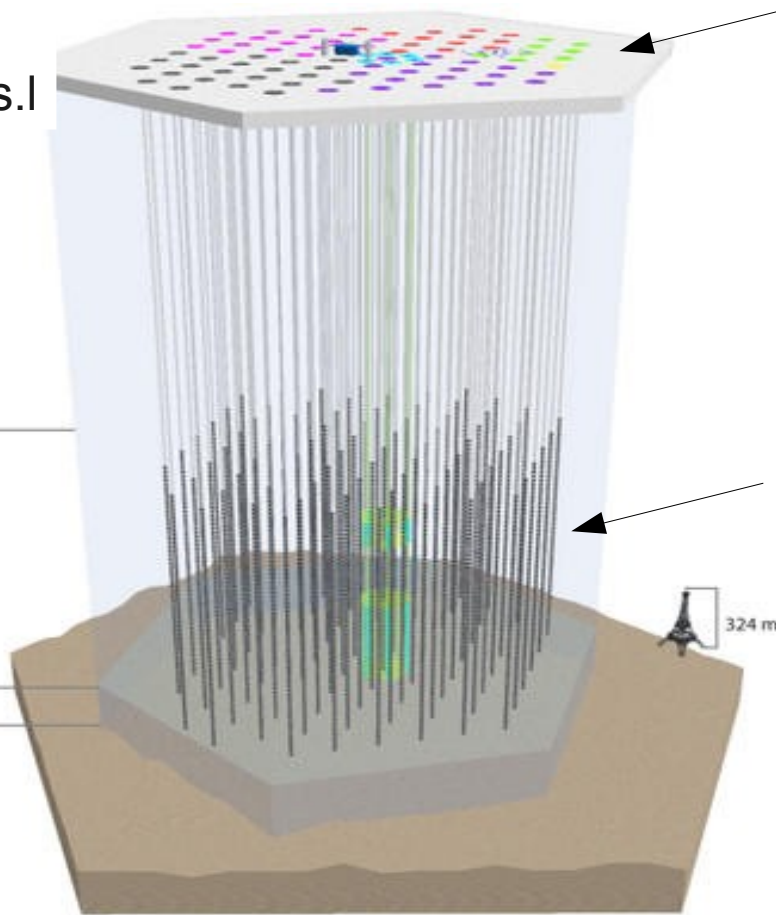
# IceCube Components

Surface:  
2835 m a.s.l

-1450 m

-2450 m

-2620 m



**IceTop** Surface Array:  
1km<sup>2</sup> area

**LE** electromagnetic

**InIce** Volume Detector:  
1km<sup>3</sup> volume

**HE** muons

Electromagnetic Particles  
(10s-100s of MeV)

IceTop  
Surface Array

LE Muons  
(1-10 GeV)



Primary Energy:  
1-100s PeV

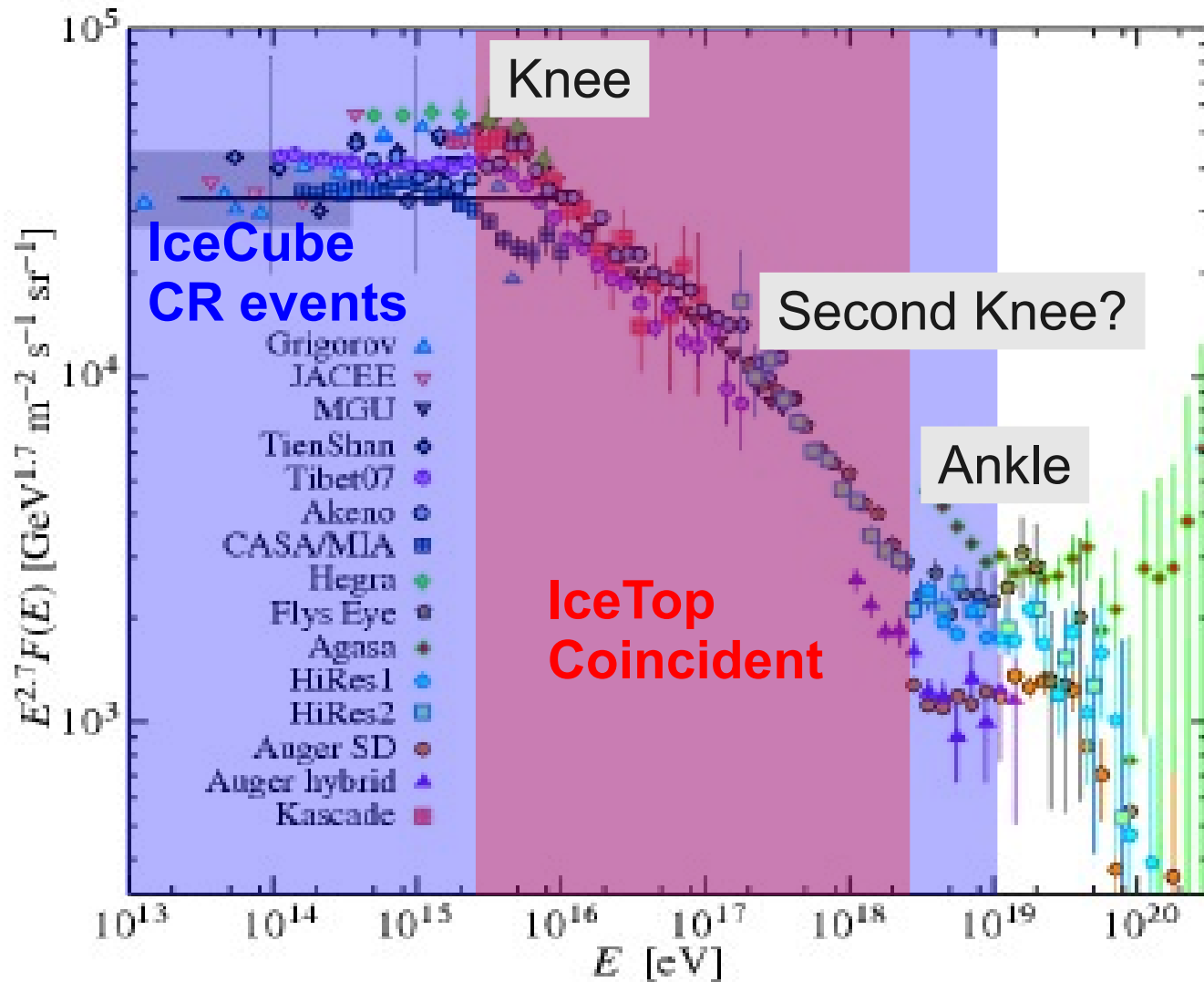
InIce Muon Multiplicity:  
10s-1000s

HE Muons  
(TeV)

InIce

Shower Axis

# CR Energy range of IceCube/IceTop



## Main Science Goals:

Primary composition change around Knee

Transition to Extragalactic CR at “Second Knee”?



Temporary Cover

2 Tanks  
per Station

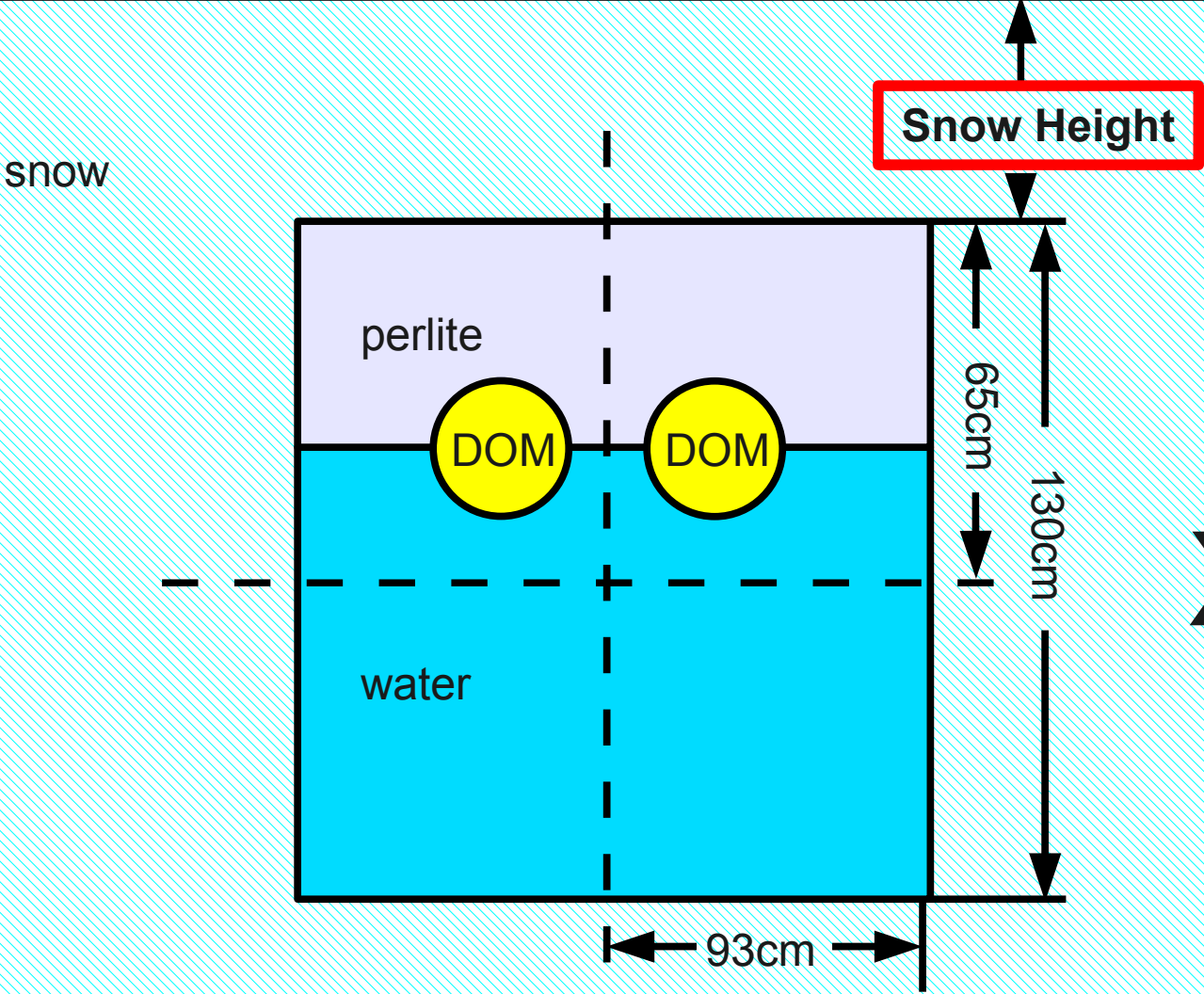
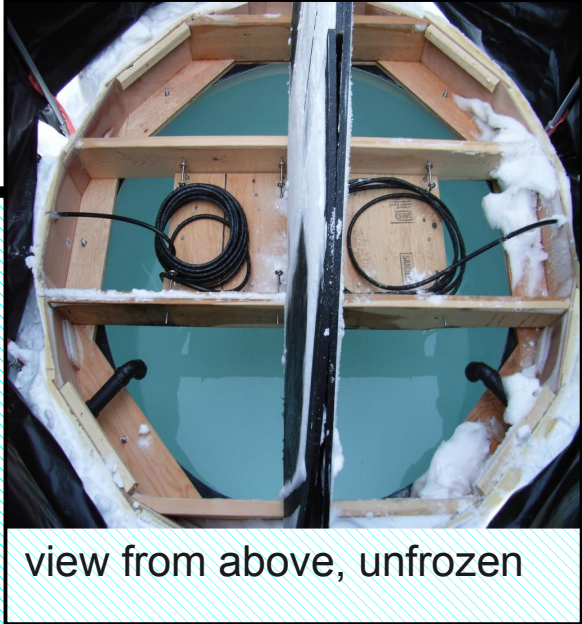
Freezing Unit

IceTop  
Deployment

# IceTop Station



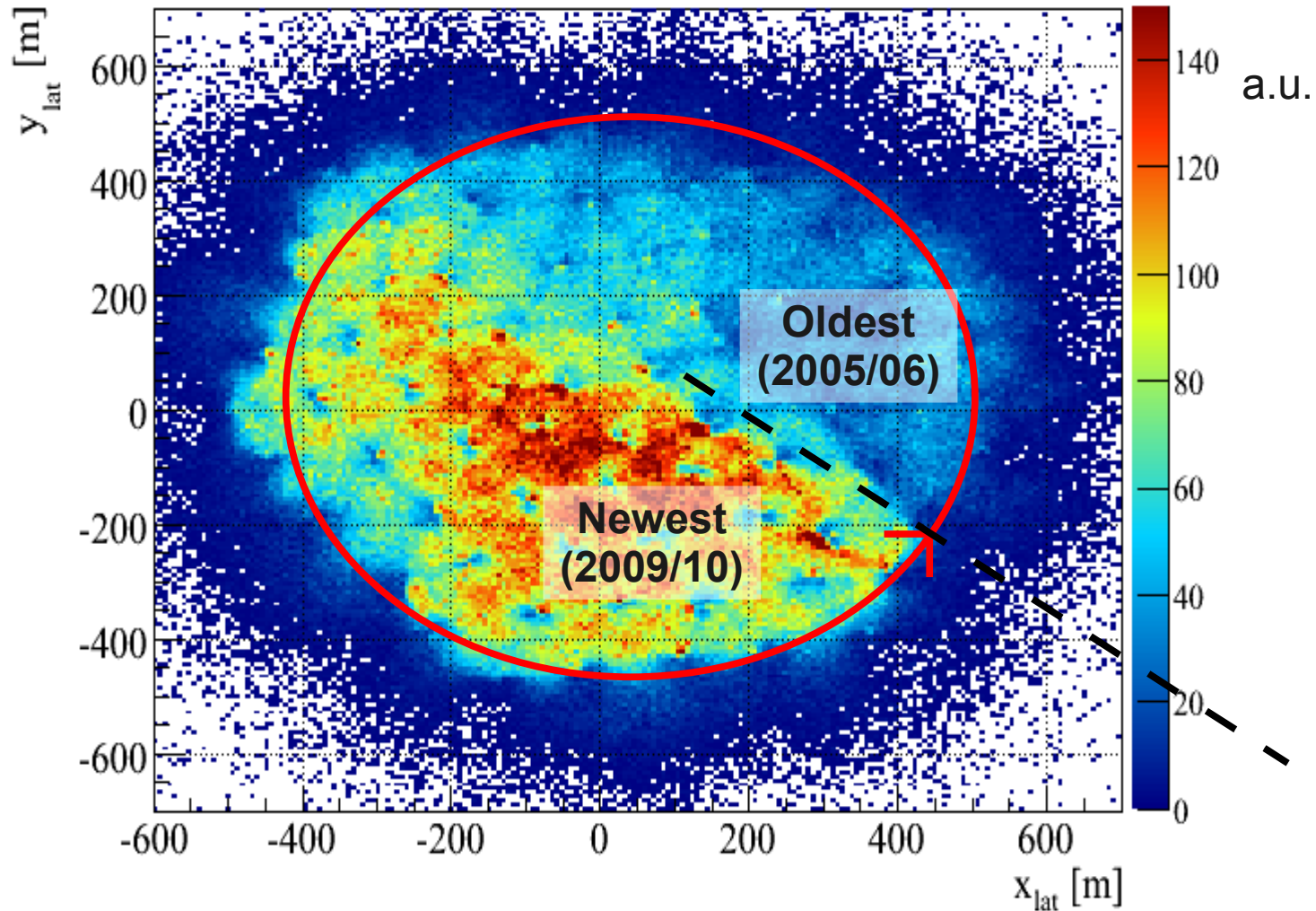
# IceTop Tank



x2 x78

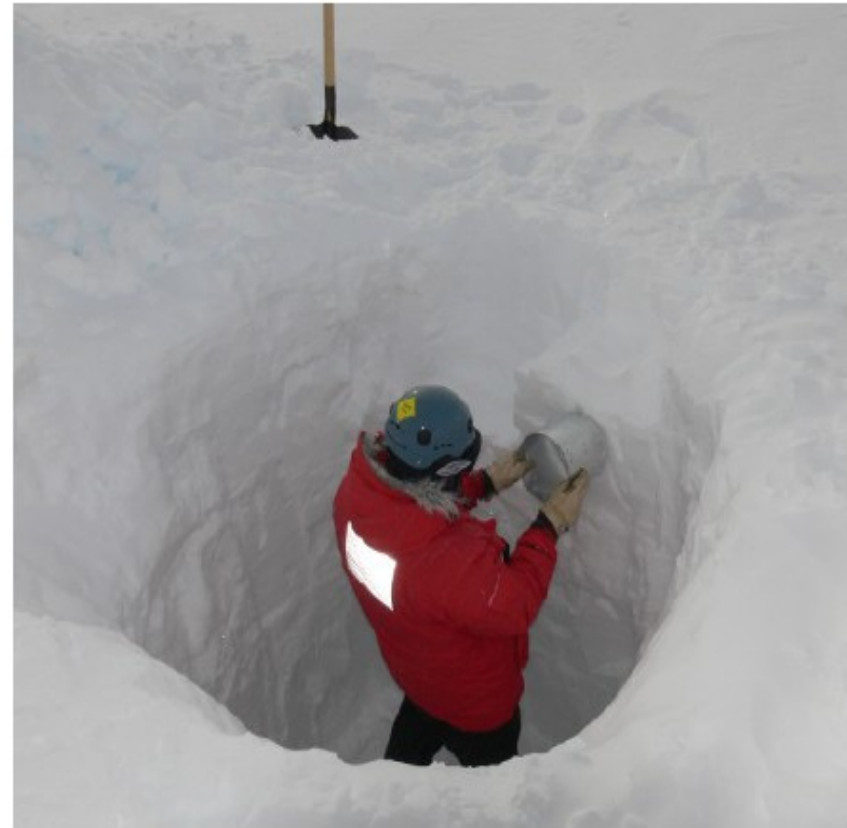
# Effect of Snow Coverage

(Reconstructed Shower Core Position)





# IceTop Calibration Procedure



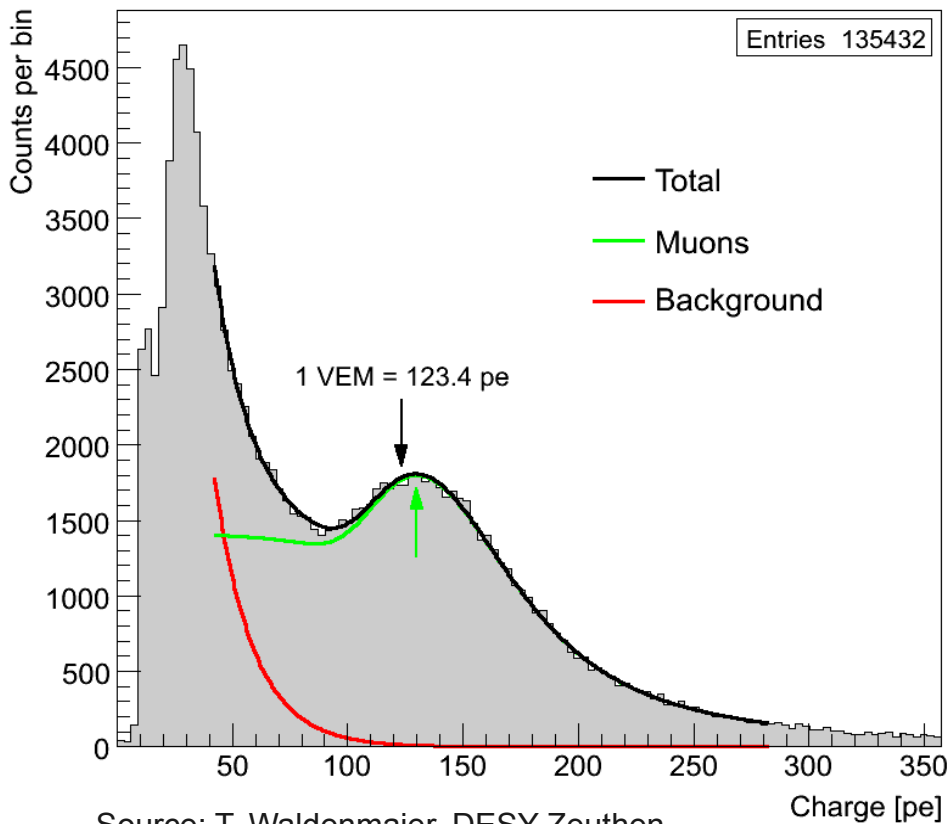
Natural snow density:  $\sim 0.35 \text{ g/cm}^3$

Snow density in trenches:  $\sim 0.4 \text{ g/cm}^3$

# Tank Calibration

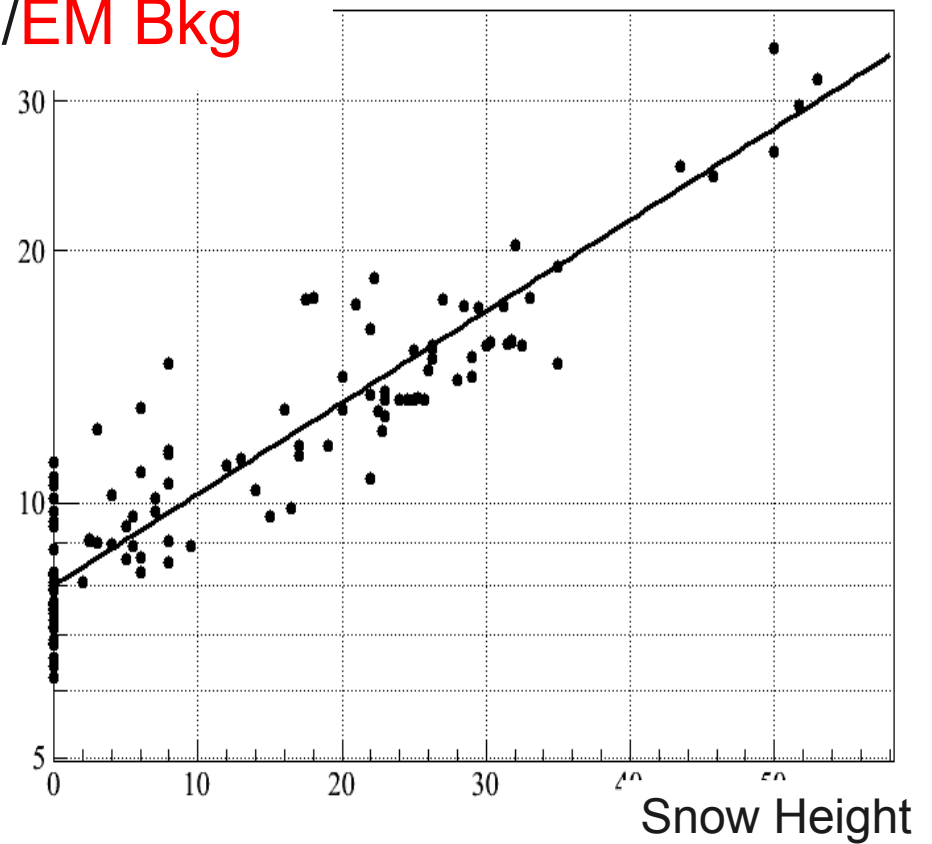
Snow reduces EM signal,  
muons unaffected

Muon Spectrum of DOM(19, 61)



Source: T. Waldenmaier, DESY-Zeuthen

Muon Peak  
/EM Bkg

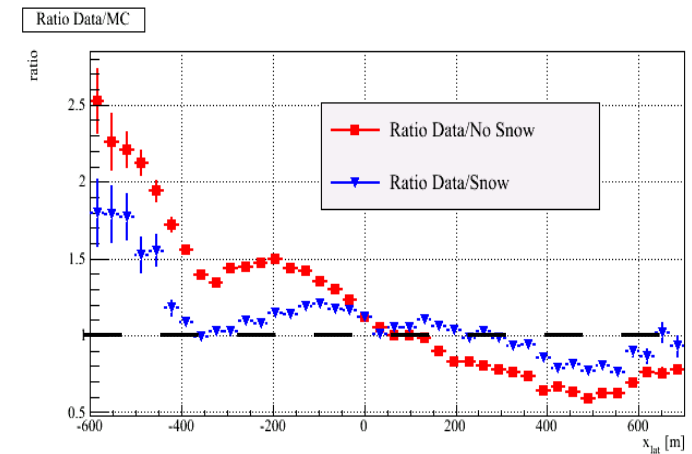
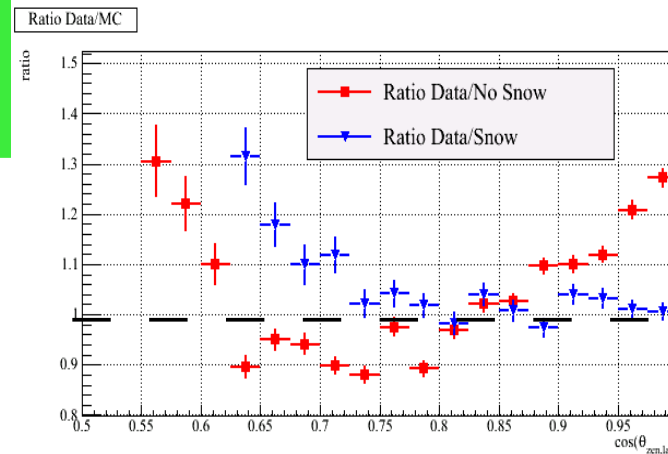
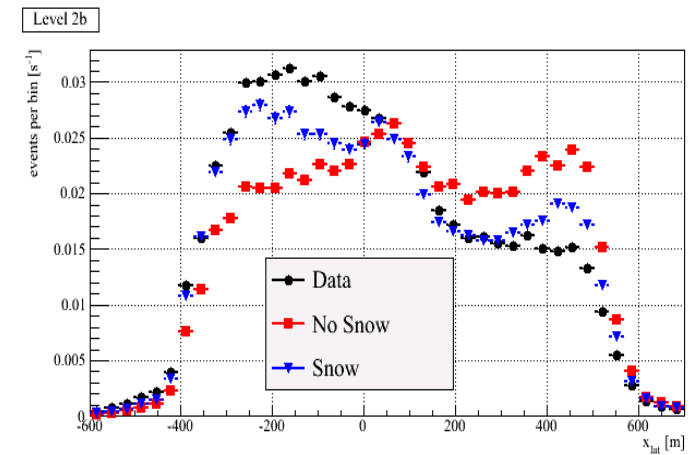
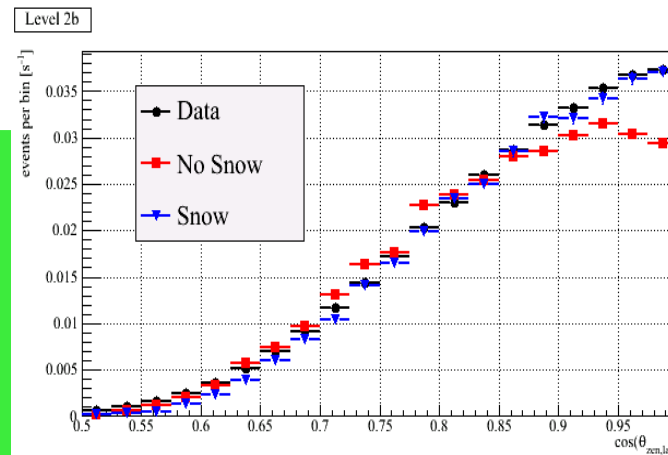


$$S_{\mu}/B_{EM} \simeq \exp\left(\frac{h_{snow}}{40.0''} + 2.08\right)$$

# Snow Effects in Detector Simulation

Here:  
 Parametrization of  
 detector response  
 in dependence of  
 particle type and snow  
 height

Soon:  
 Full GEANT-4 simulation



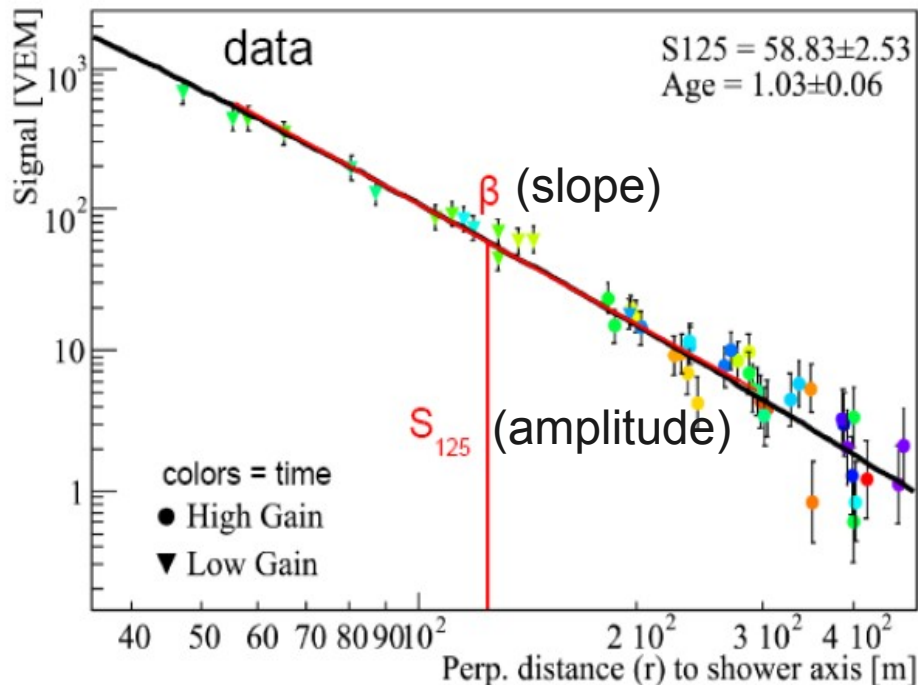
**Red:**  
 Default Response  
 Parametrization

**Blue:**  
 Snow Attenuation and  
 individual Snow Height  
 for each tank

Zenith Angle

Shower Core Position

# IceTop Shower Reconstruction



Lateral shower profile at 125m

$$S(r) = S_{125} \left( \frac{r}{125m} \right)^{-\beta - \kappa \log_{10} \left( \frac{r}{125m} \right)}$$

$S_{125}$ : signal at  $r = 125m$

$\beta$ : slope at  $r = 125m$

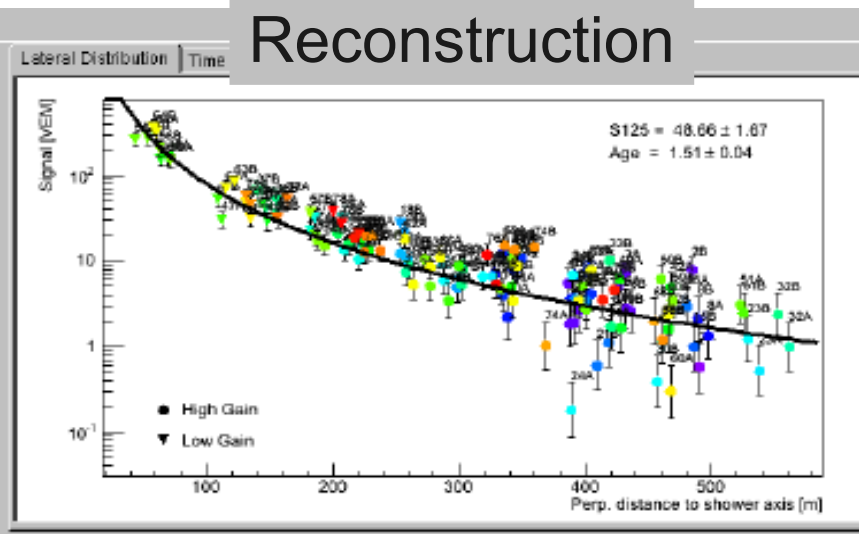
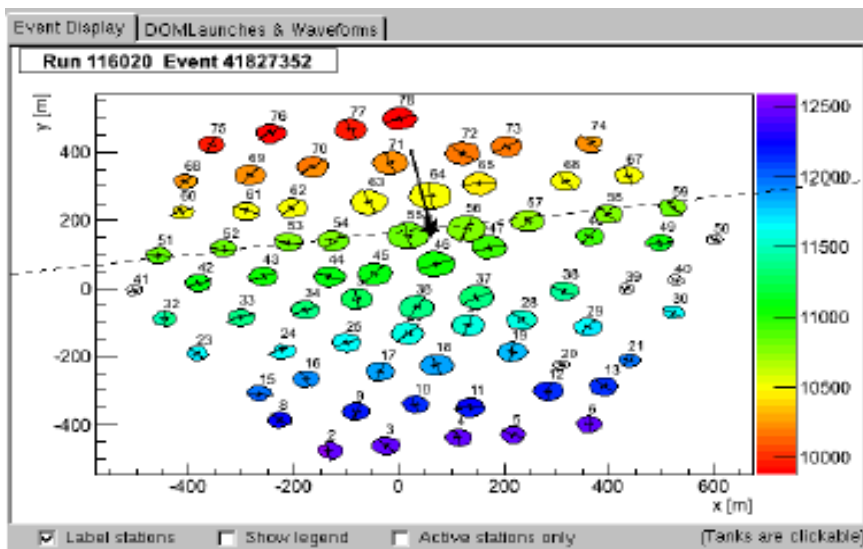
$\kappa = 0.303$  fixed

arXiv: 0711.0353

$$E_{\text{prim}} = f(S_{125}, \theta_{\text{zen}})$$

“Double-Logarithmic Parabola”:  
MC-derived empirical description

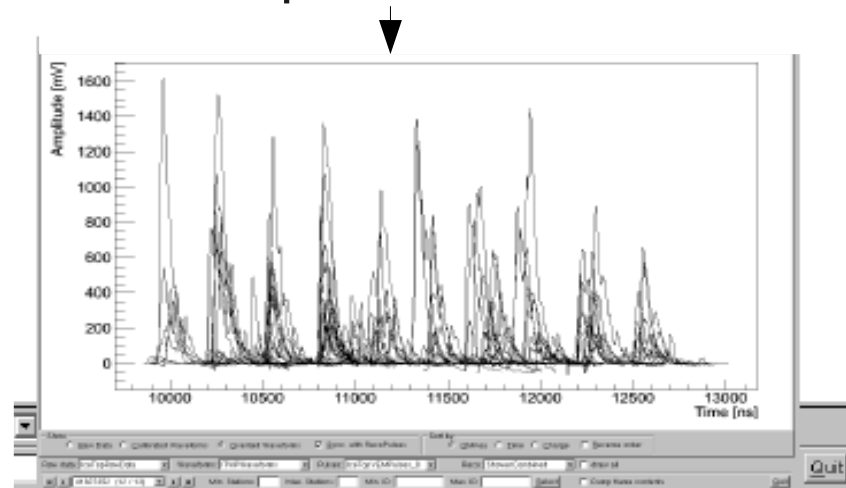
# IceTop Event



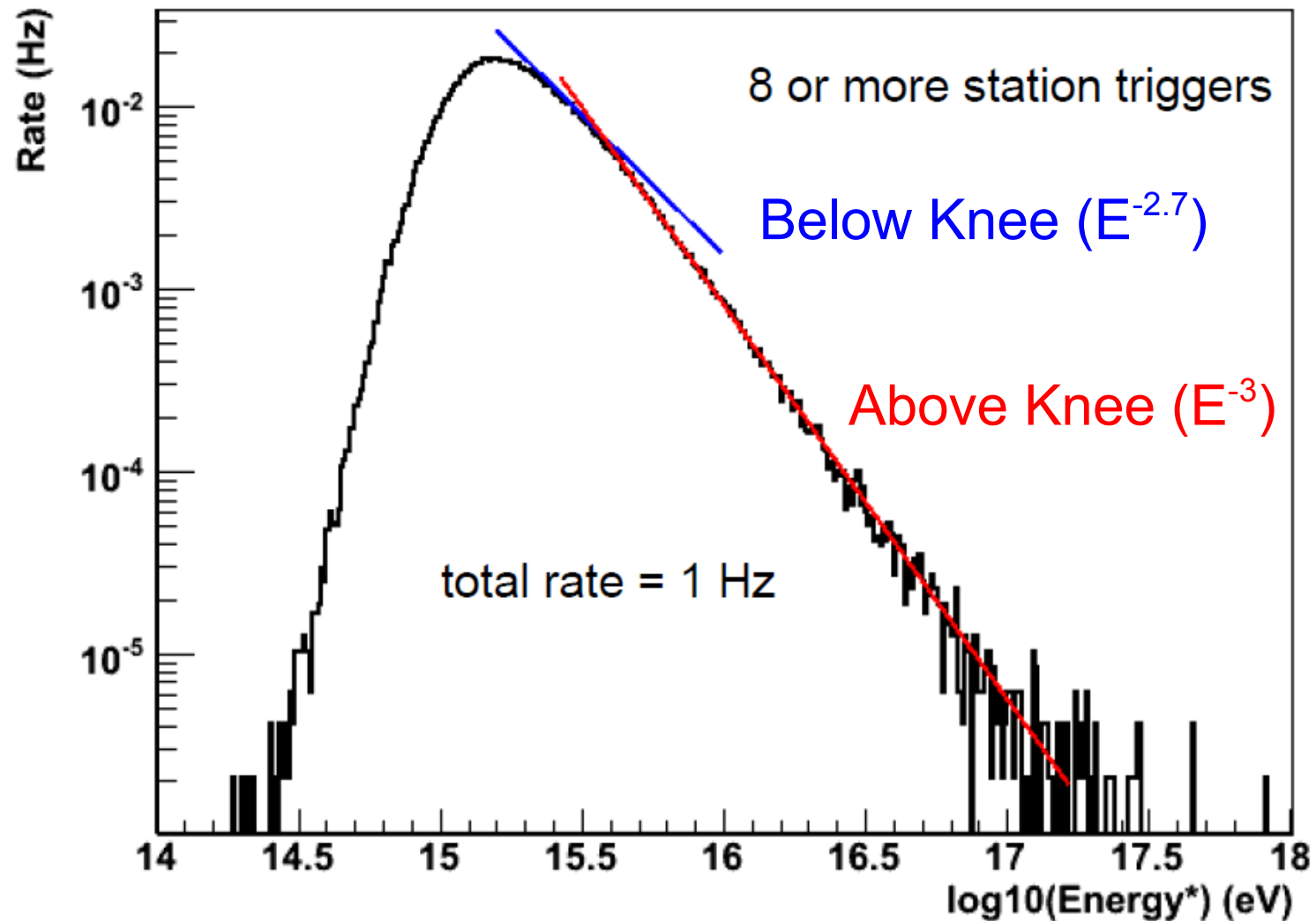
color: time  
size: signal

Estimated Energy: 183 PeV  
Zenith Angle:  $50^\circ$

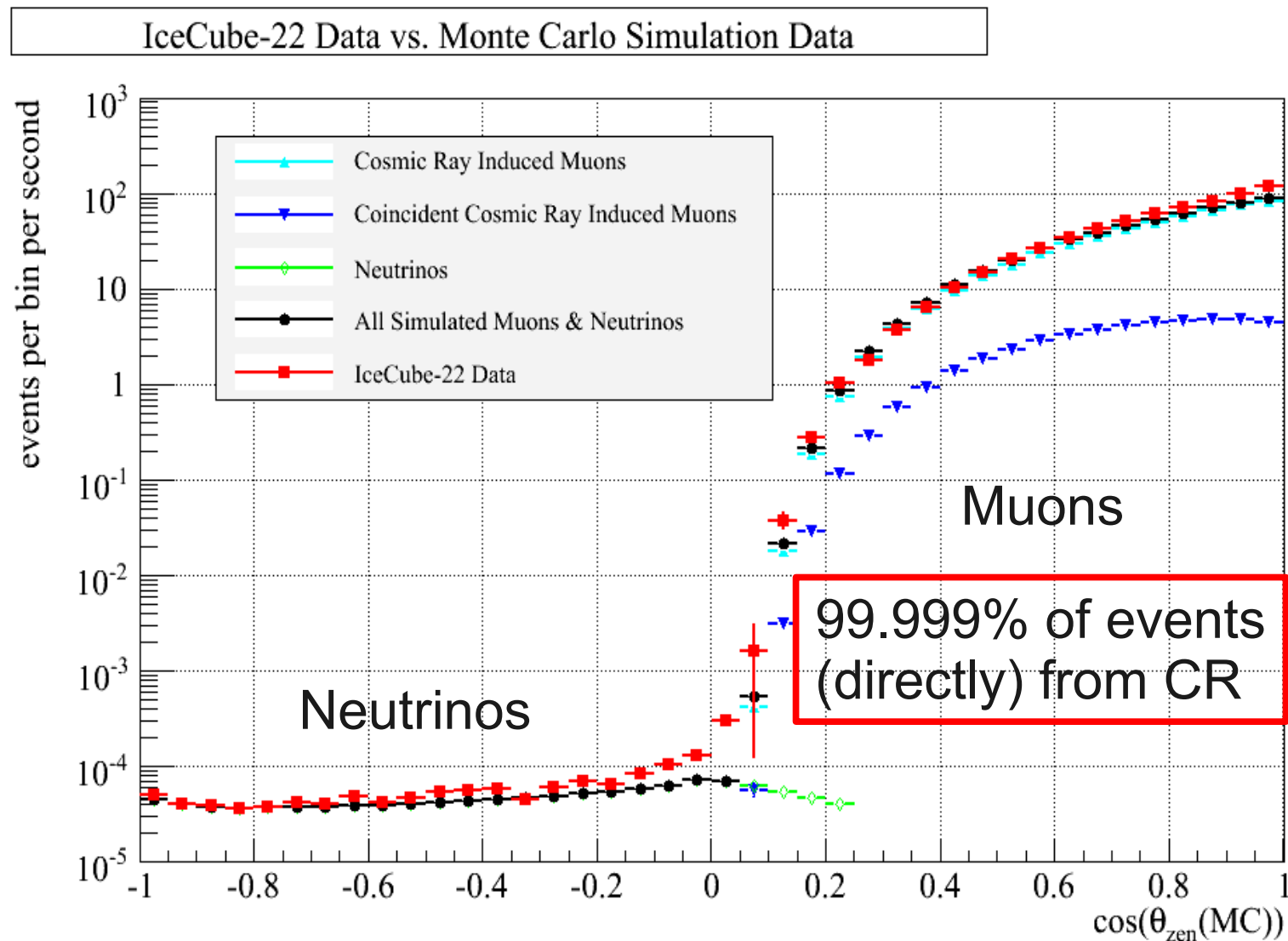
All IceTop Waveforms



# Raw IceTop-only CR Spectrum



# InIce Muon Flux (IC22, 2008)

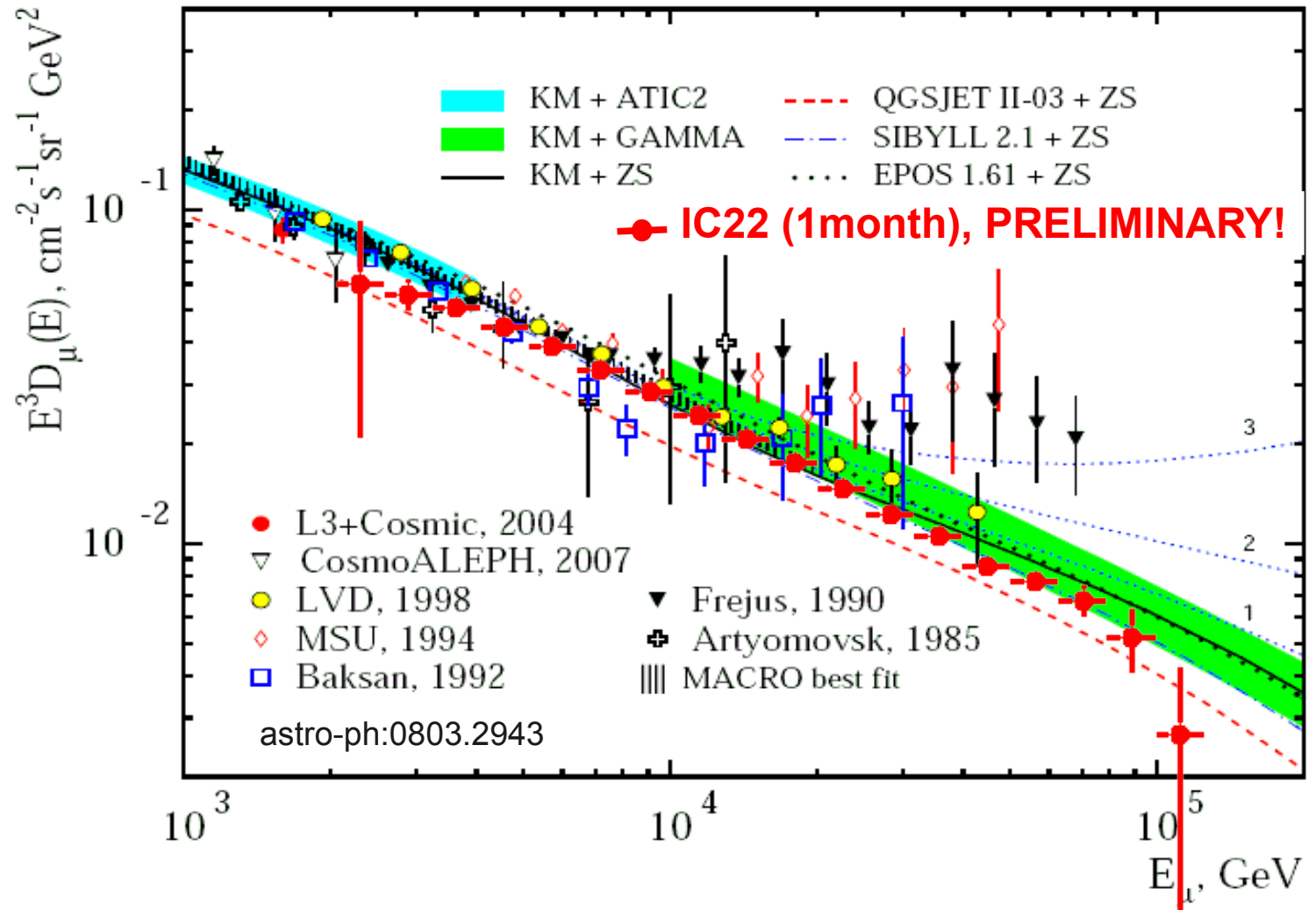


arXiv:0902.0021

# Muon Spectrum

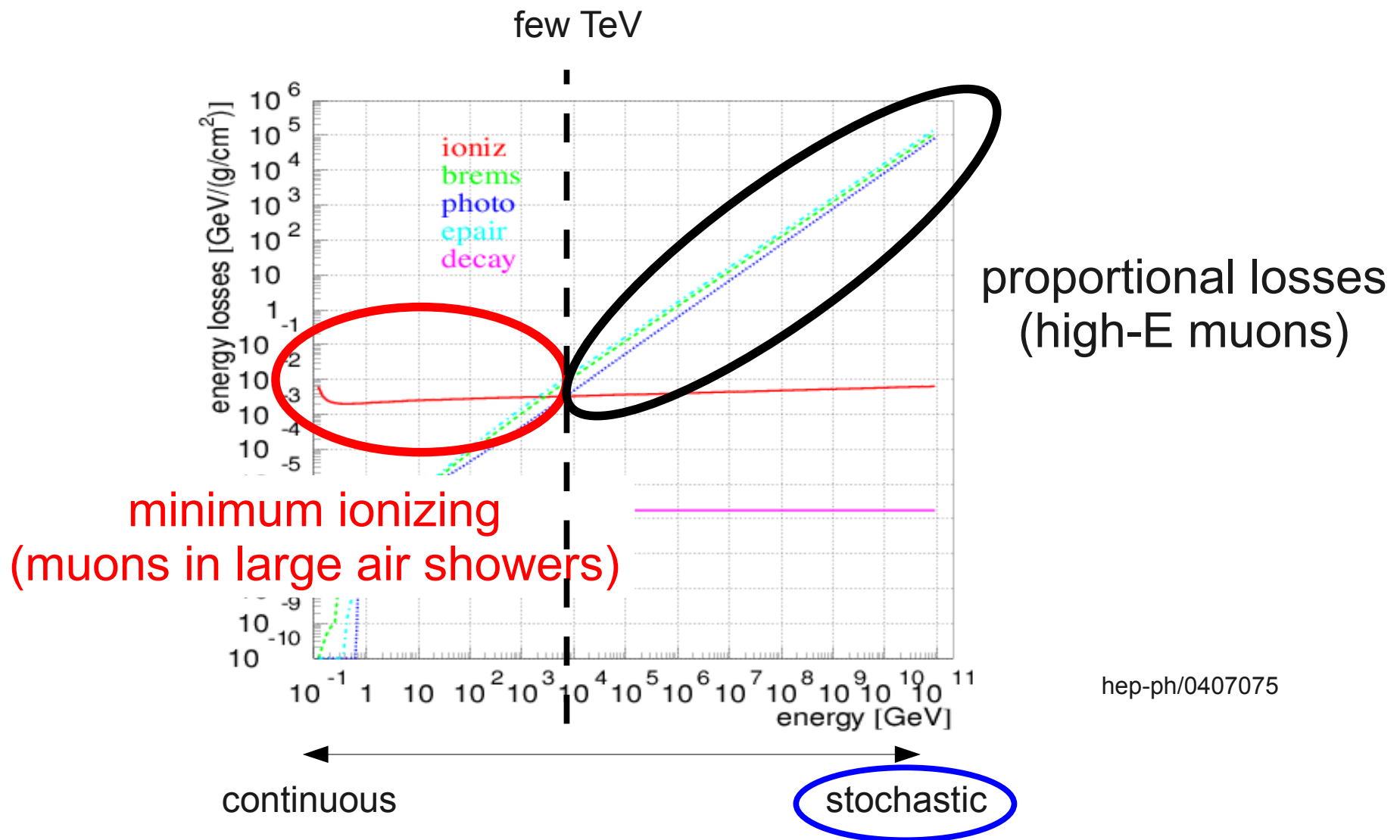
Derived from  
muon tracks  
near the horizon

Higher angles:  
High-multiplicity  
muon bundles  
dominate over  
single HE muons

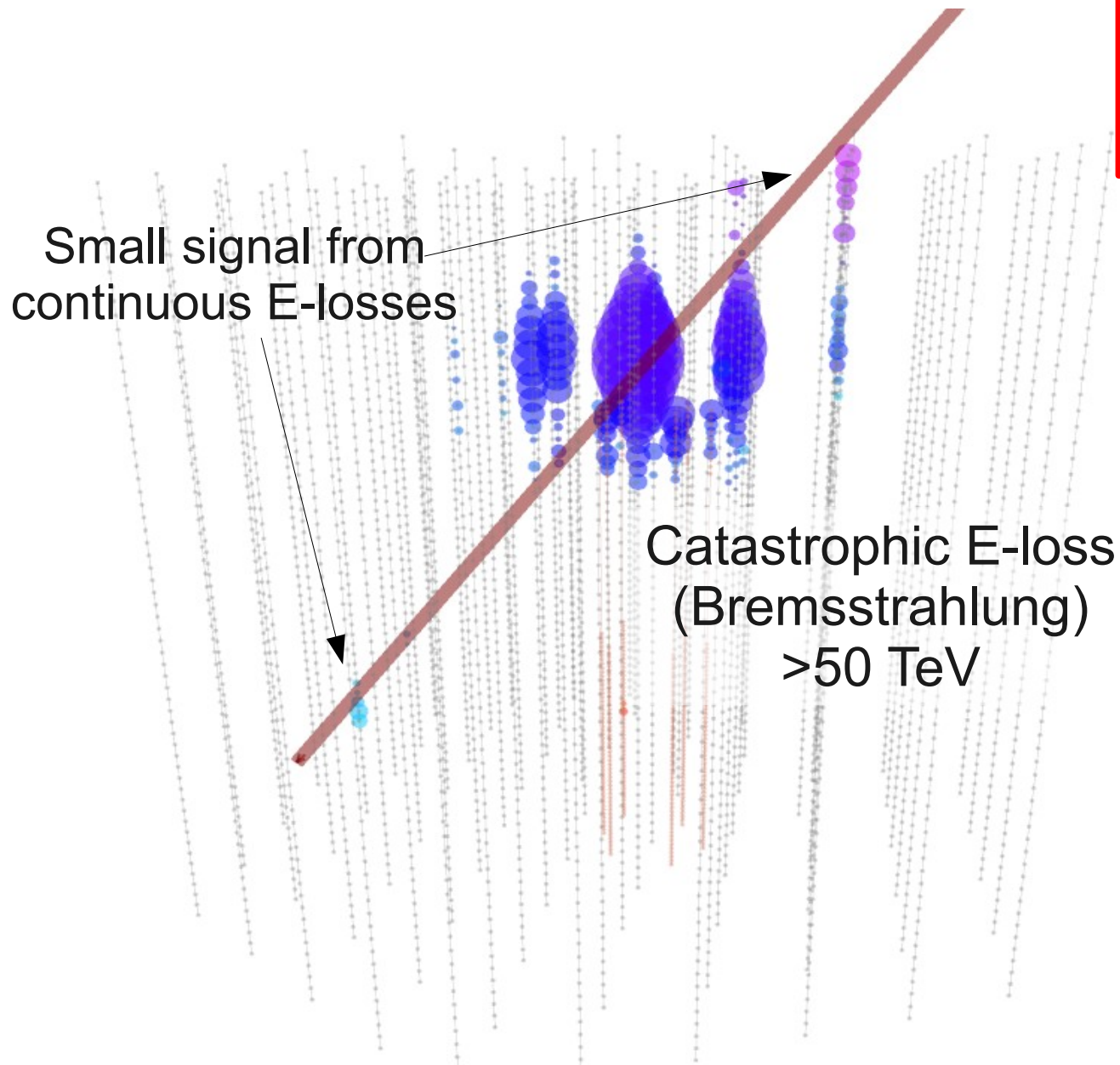




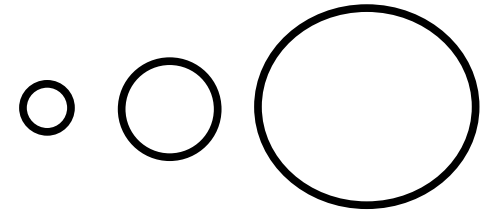
# Muon Energy Losses in Matter (Ice)



# High-Energy Muon



Identify stochastic losses to distinguish HE muons from high-multiplicity muon bundles



Size: PMT Signal

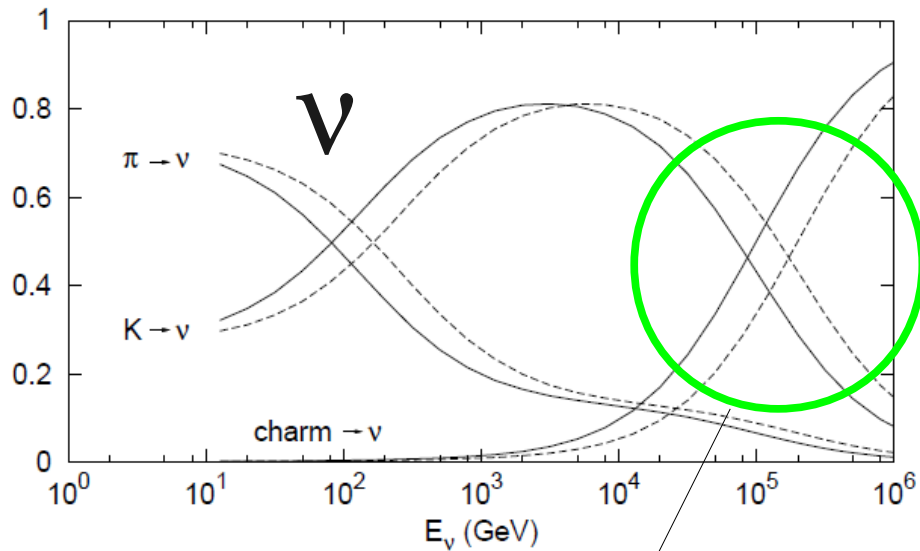


Color: Time

# Prompt Leptons from Charm Decay

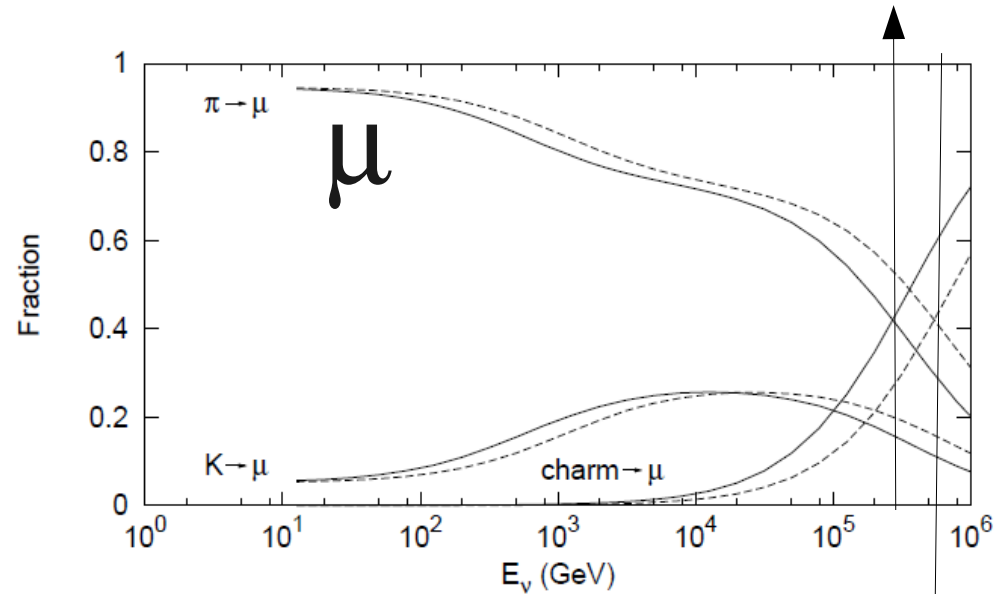
Light mesons: Flux  $\propto 1/\cos\theta_{zen}$   
 Prompt: constant angular Flux

For **vertical** muons, prompt component becomes dominant at  $\approx 200$  TeV



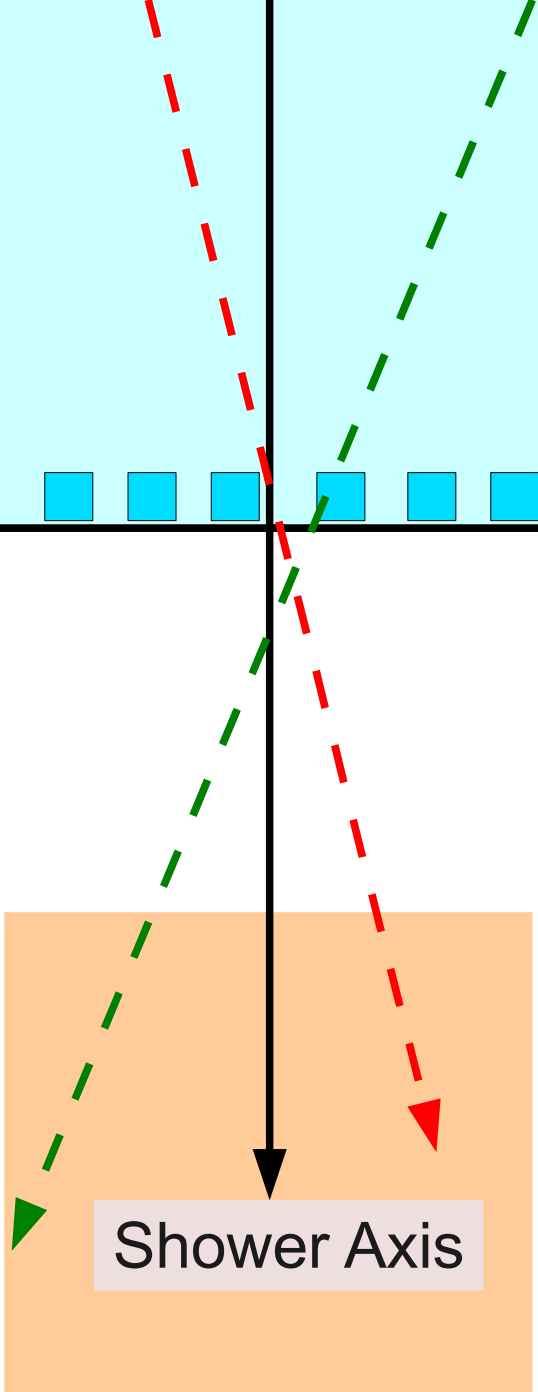
Source: T. Gaisser

Charm represents major systematic uncertainty for neutrinos above 100 TeV



At  $60^\circ$ , light meson decay dominates up to  $\approx 500$  TeV

**IceTop Reco:**  
 $\Delta \log_{10} E \simeq 0.1$  (prim)  
 $\Delta \theta \simeq 1^\circ$   
 $\Delta r \simeq 20$  m

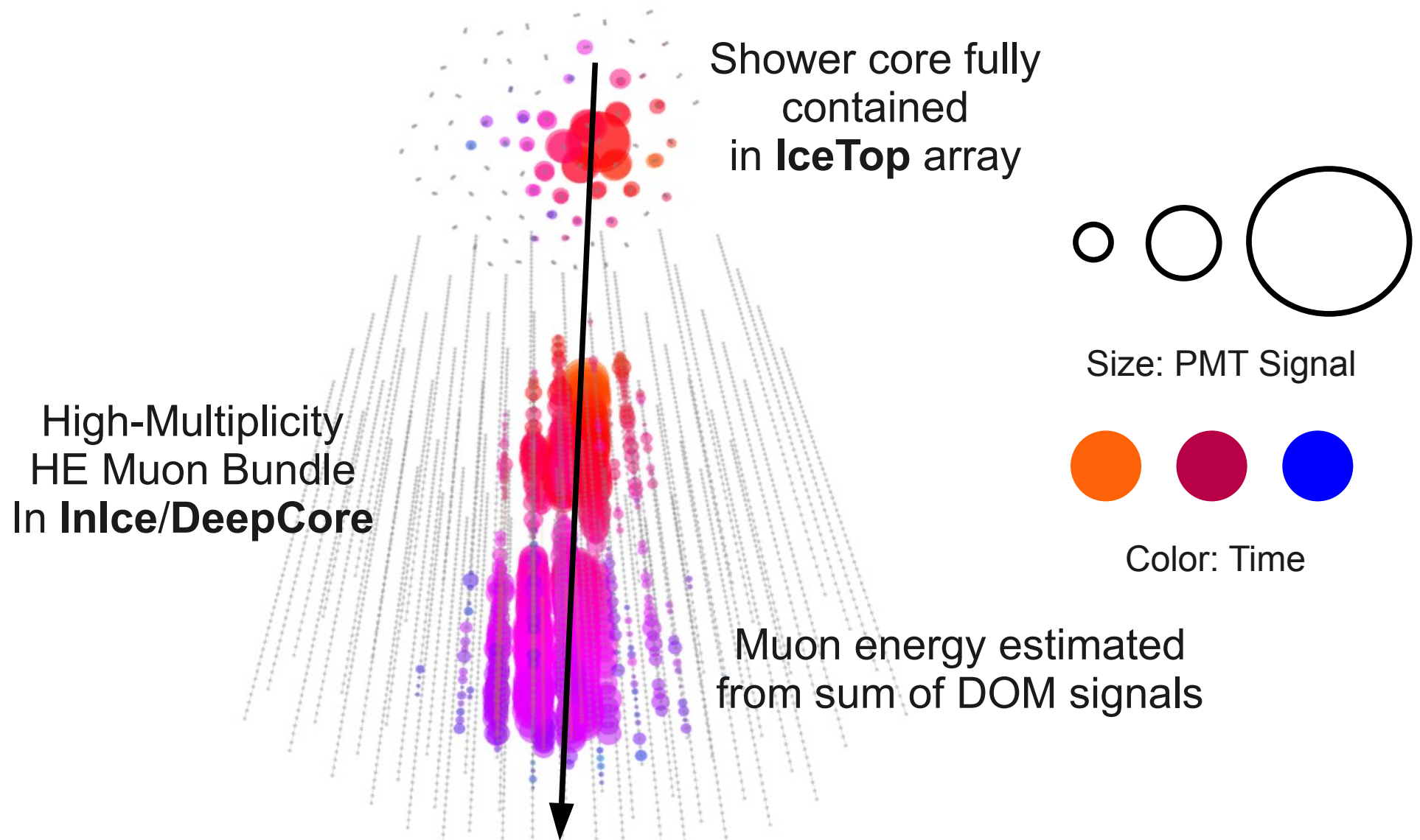


**InIce Reco:**  
 $\Delta \log_{10} E \simeq 0.3$  ( $\mu$ )  
 $\Delta \theta \simeq 0.7^\circ$   
 $\Delta r \simeq 30$  m

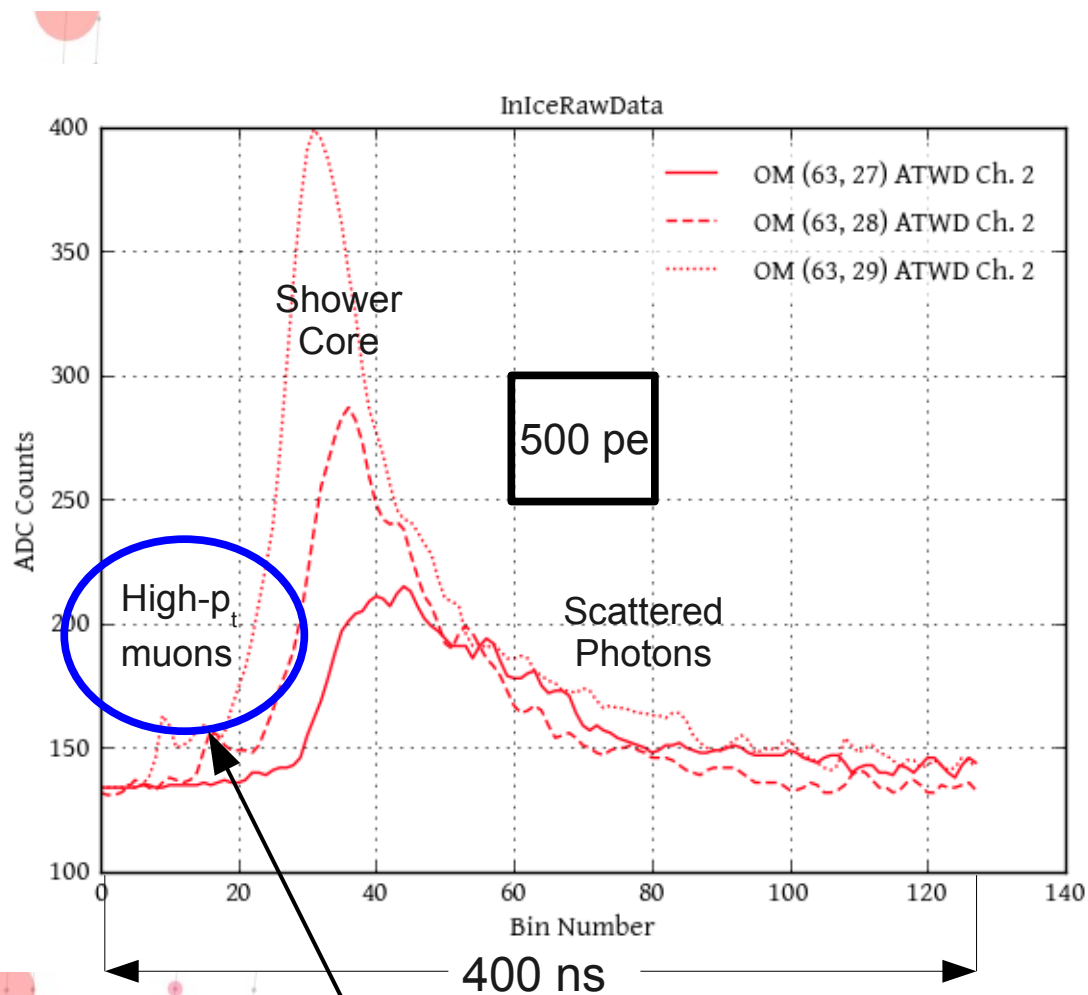
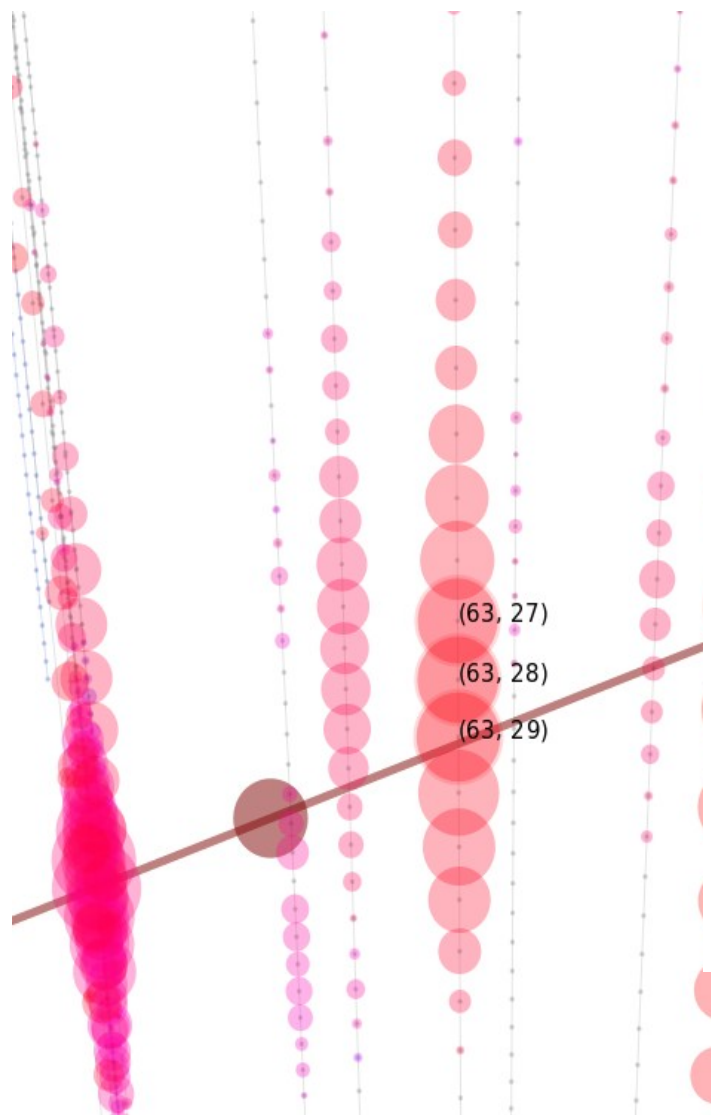
Reconstructions  
are completely  
Independent

Relation between  
energies depends  
On CR primary type

# IceTop/InIce Coincident Event

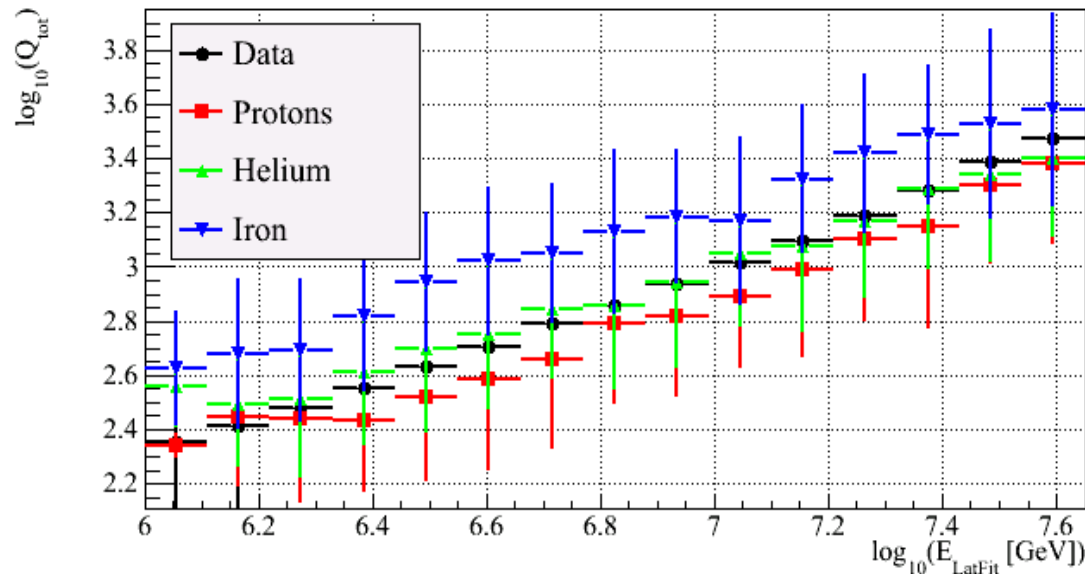


# DOM Waveforms



DOM readout triggered by outlying muons:  
Characteristic for large CR bundles

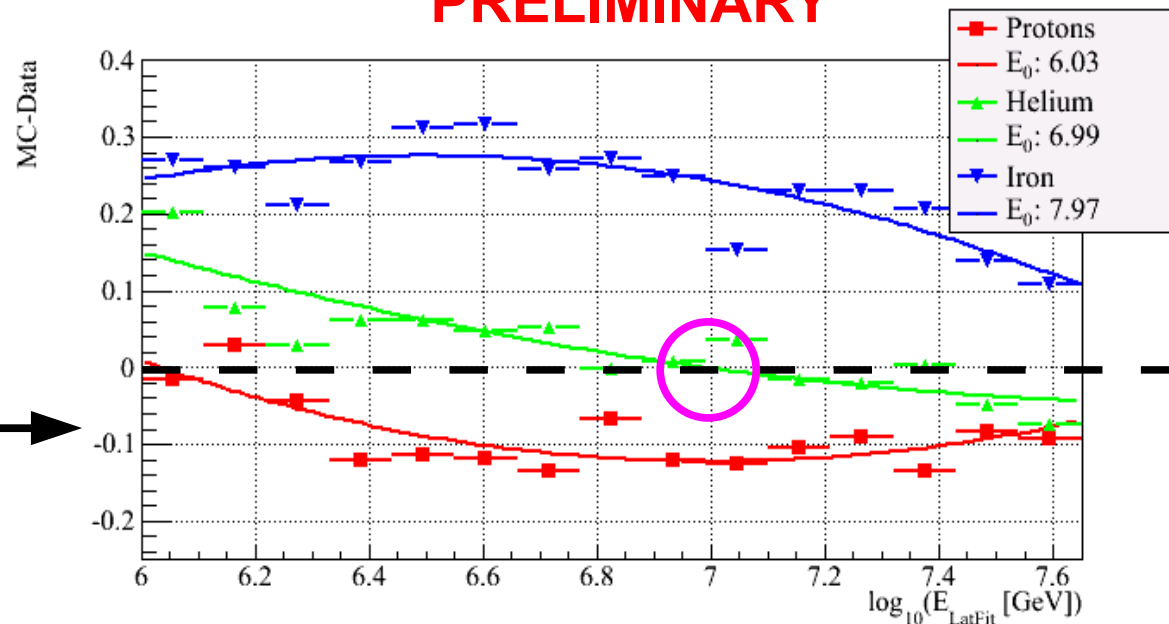
# IceTop/InIce Energy Correlation



**PRELIMINARY**

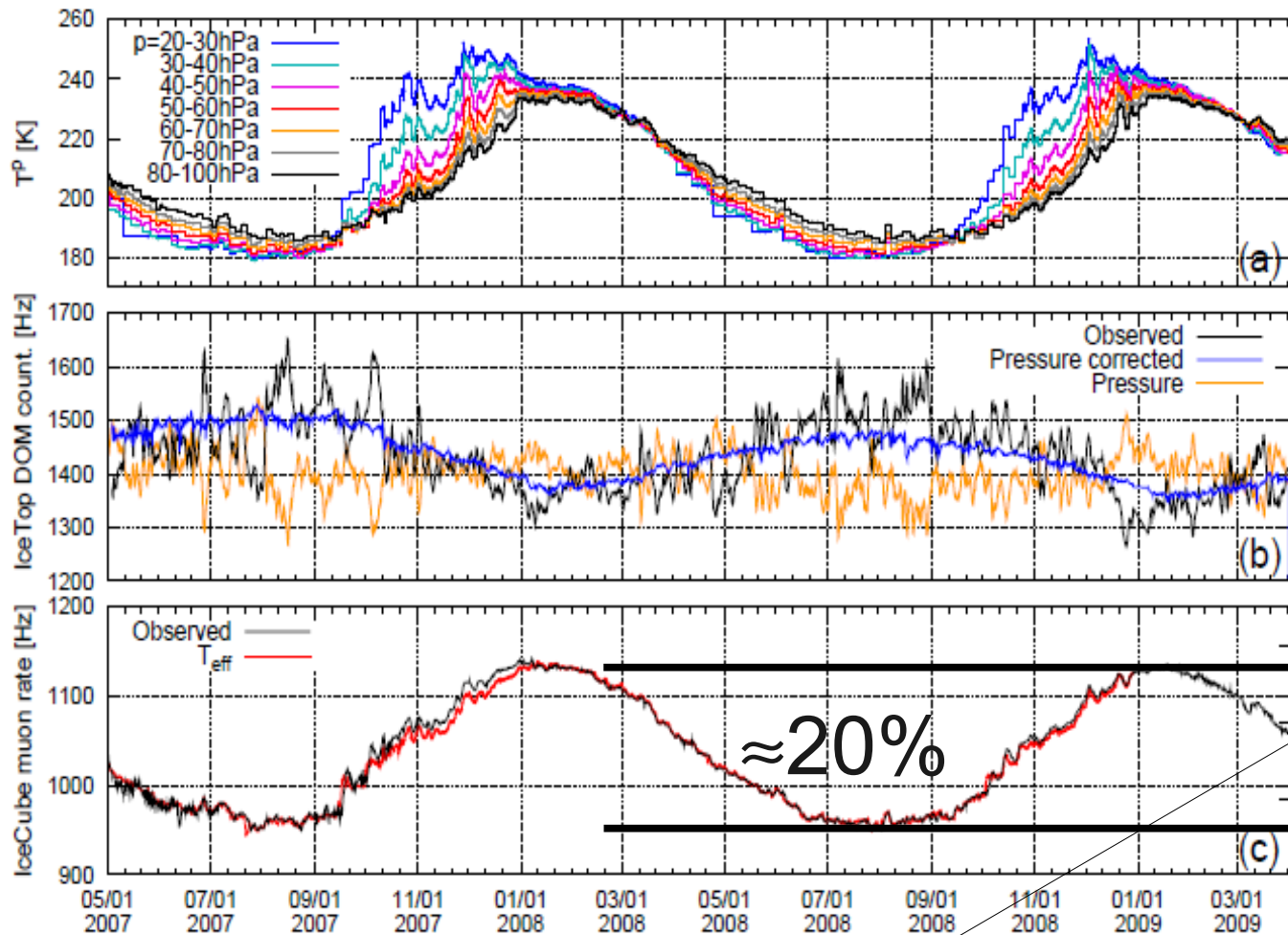
$\langle \log A \rangle = \text{He}$   
 at  
 $E_{\text{prim}} \approx 10 \text{PeV}$

↑ InIce (HE muons)  
 → IceTop (EM)



$$\begin{aligned}
 \langle Q_{\text{MC}}(A) \rangle \\
 = \\
 \langle Q_{\text{data}} \rangle
 \end{aligned}$$

# Seasonal Variations



Atmospheric  
Temperature

IceTop  
(barometric)

InIce  
(temperature)

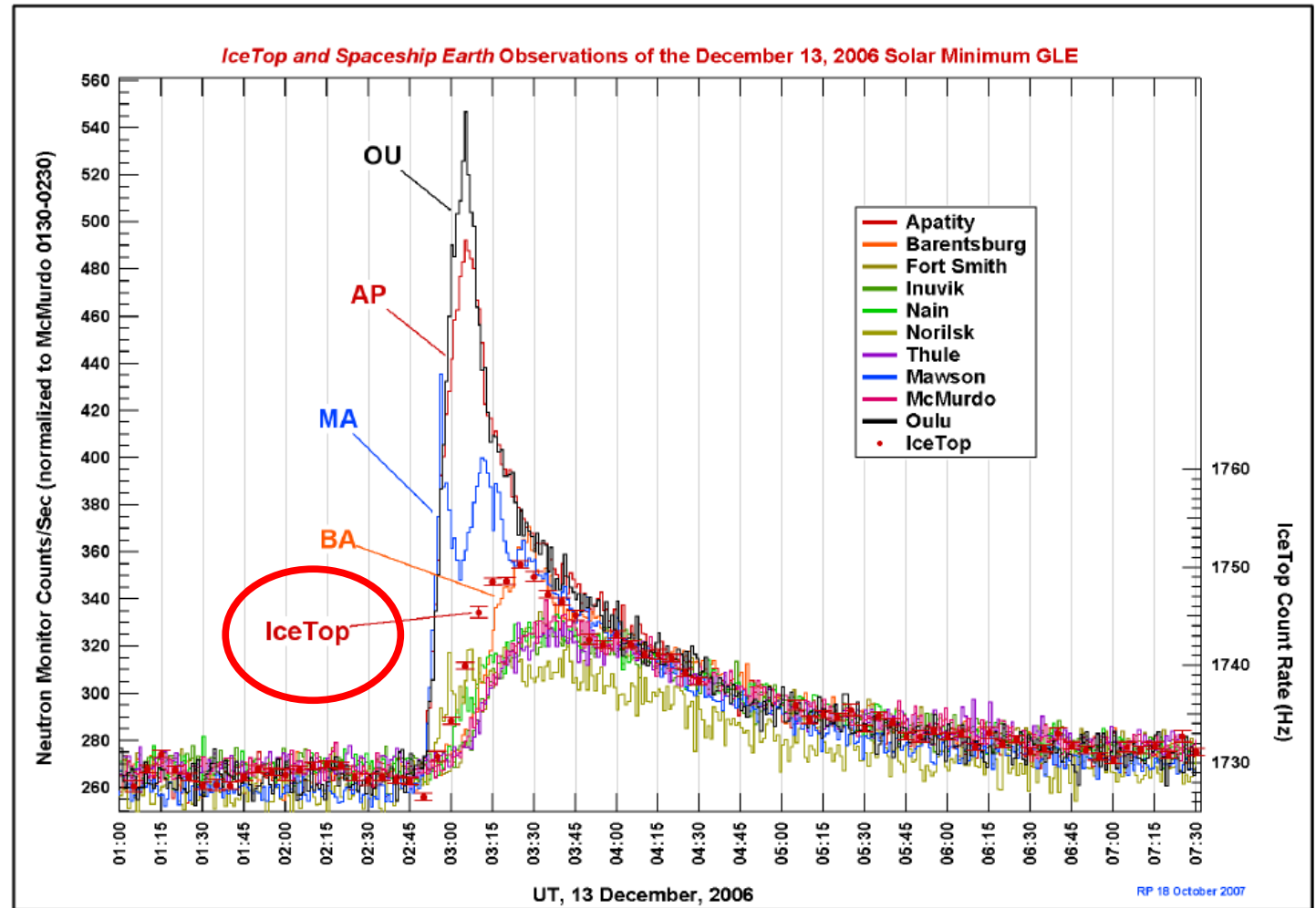
$T_{\text{eff}}$ : Temperature weighted by muon production probability

$$T_{\text{eff}} = \frac{\int_0^\infty \frac{dX}{X} T(X) (e^{-X/\Lambda_\pi} - e^{-X/\Lambda_N})}{\int_0^\infty \frac{dX}{X} (e^{-X/\Lambda_\pi} - e^{-X/\Lambda_N})}$$



# First Extraterrestrial IceCube Signal

13 Dec. 2006:  
Solar Flare seen  
in IceTop  
Count Rate





Thank you  
for your  
attention!