



University of Zurich



# The XENON100 direct Dark Matter search Experiment

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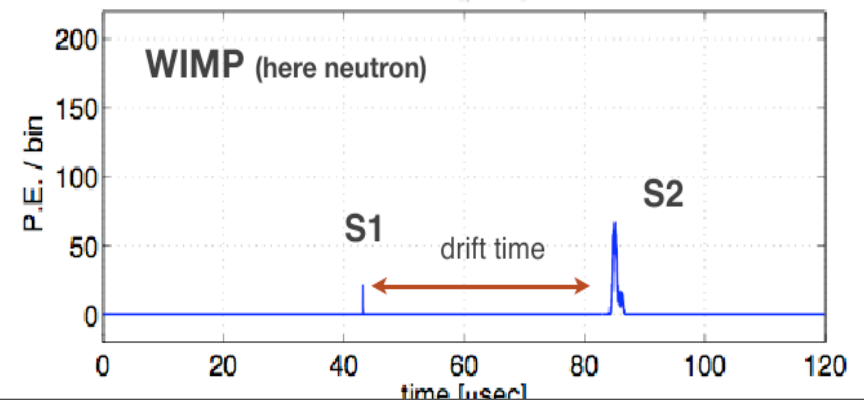
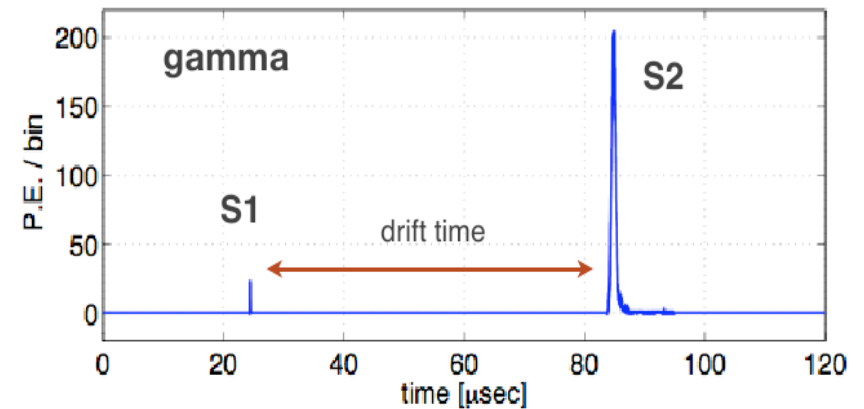
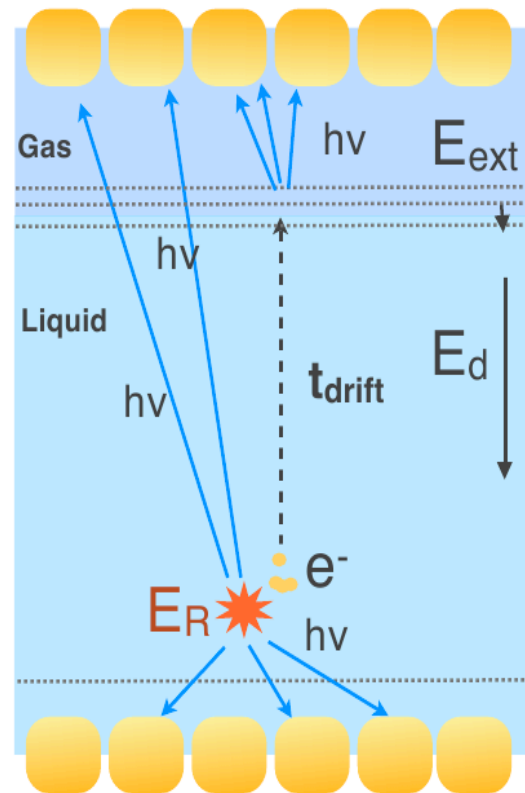
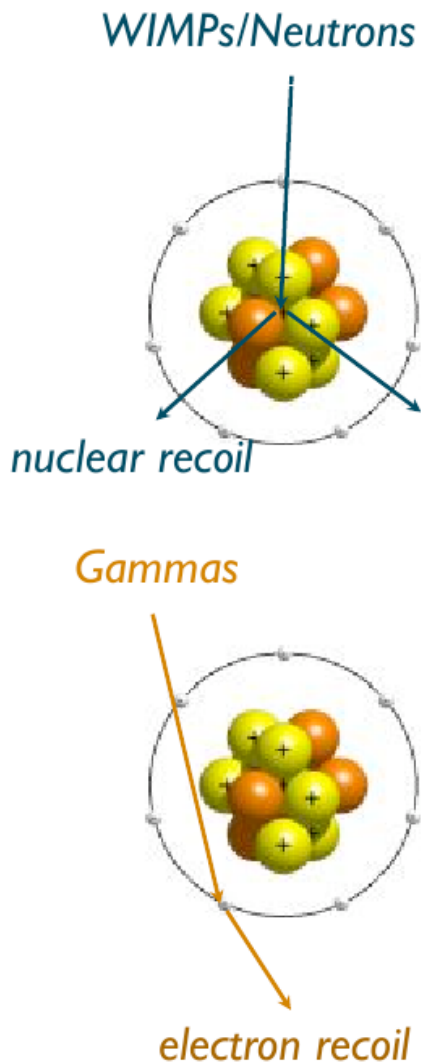
Alfredo Davide Ferella  
University of Zurich (UZH)

On Behalf of the XENON Collaboration

TeVPA  
19 - 23 July 2010

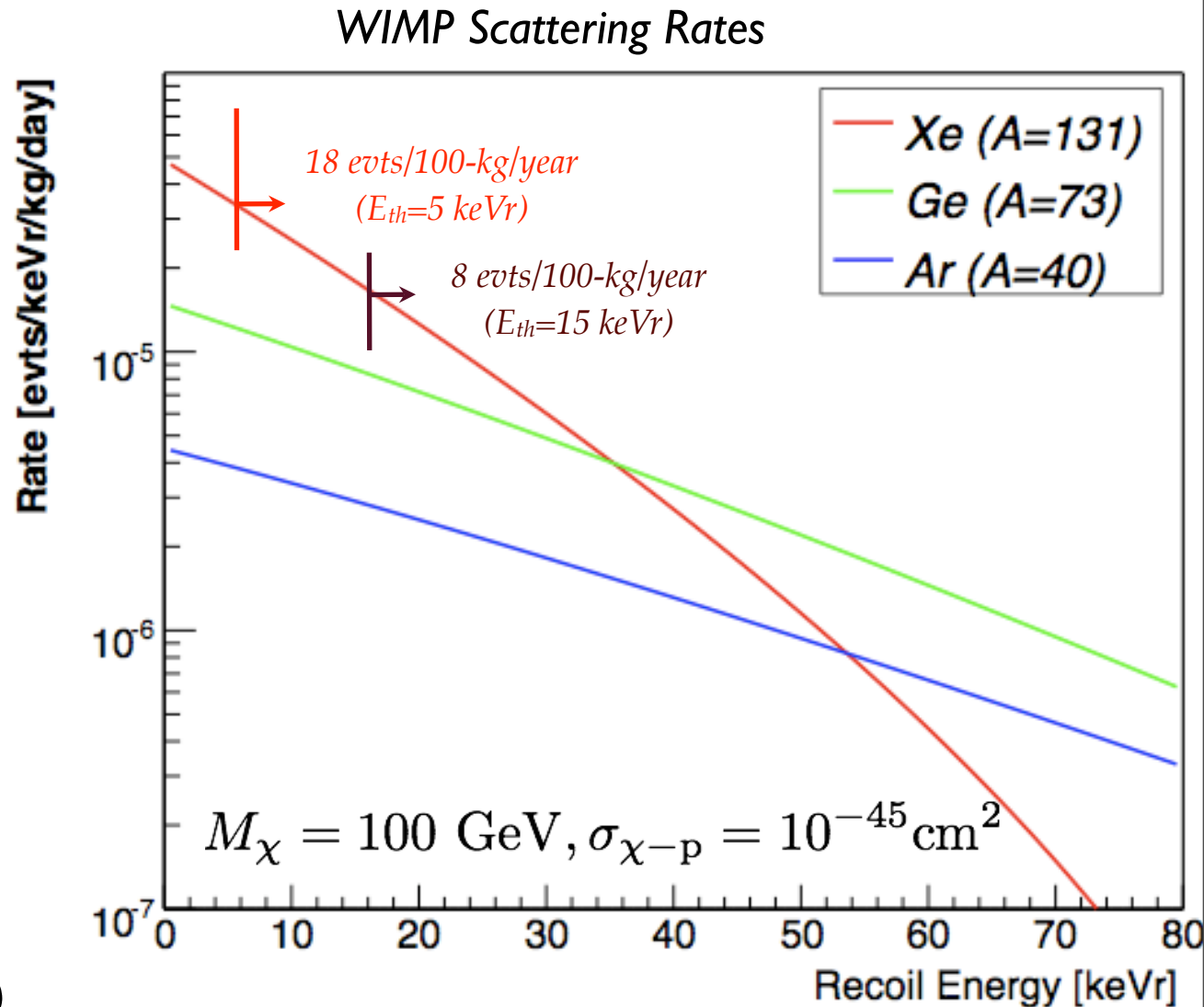
# Double phase TPC

- Primary scintillation signal (S1)
- Electrons drift over 30 cm max distance
- Electrons are extracted and accelerated generating secondary scintillation signal
- The time difference between the two signals gives information on event position in z



# Why Liquid Xenon?

- ✓ large mass (ton scale)
- ✓ easy cryogenics
- ✓ low energy threshold (a few keV)
- ✓  $A \sim 131$  (good for SI)
- ✓  $\sim 50\%$  odd isotopes (SD)
- ✓ background suppression
  - good self shielding features ( $\sim 3 \text{ g/cm}^3$ )
  - low intrinsic radioactivity
  - gamma background discrimination
  - position sensitive (TPC mode)





# Collaboration

USA, Switzerland, Portugal, Italy, Germany, France, China, Netherlands



COLUMBIA



RICE



UCLA



ZURICH



COIMBRA



LNGS



MPIK



BOLOGNA



SHANGHAI



MUENSTER



SUBATECH



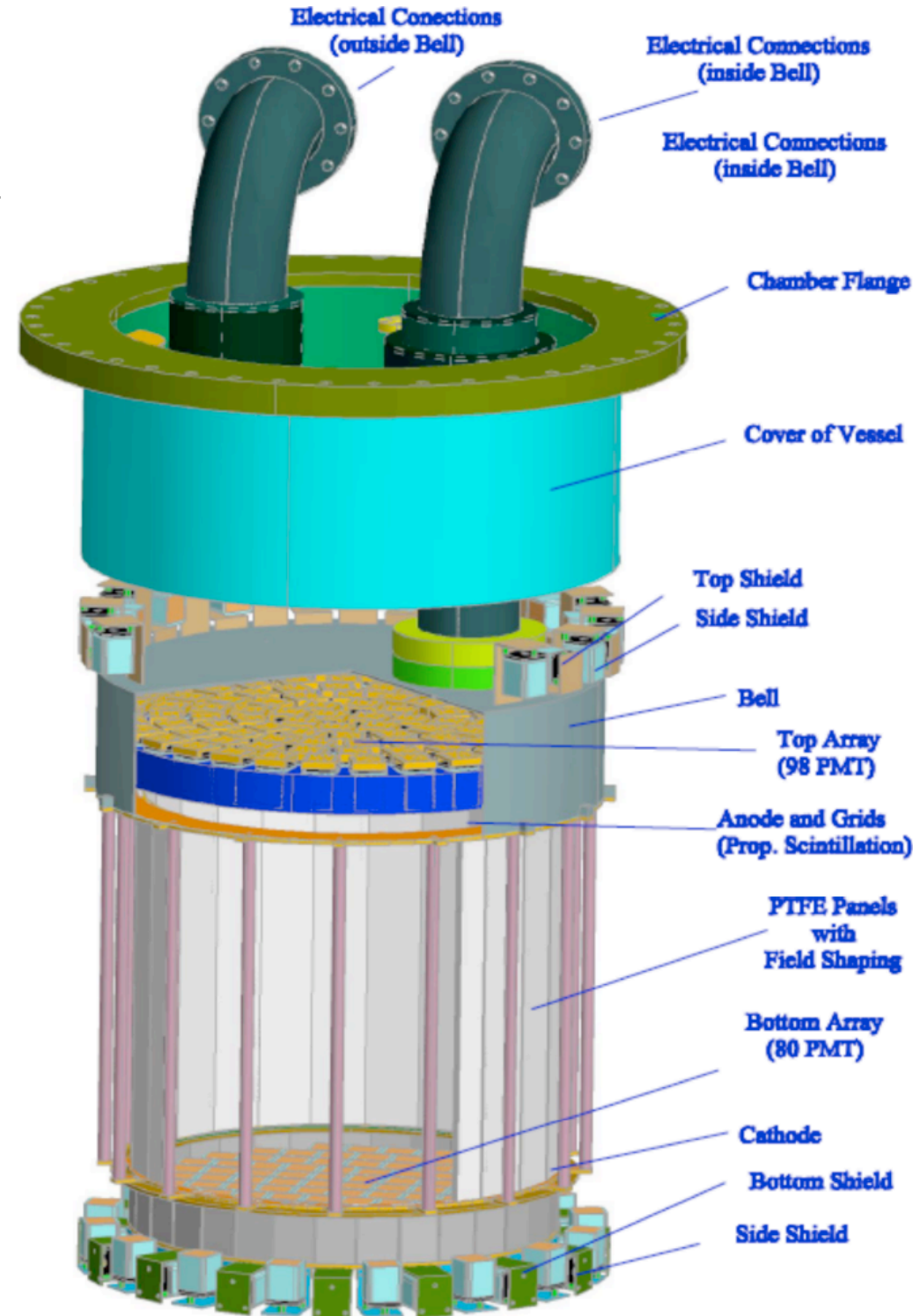
NIKHEF



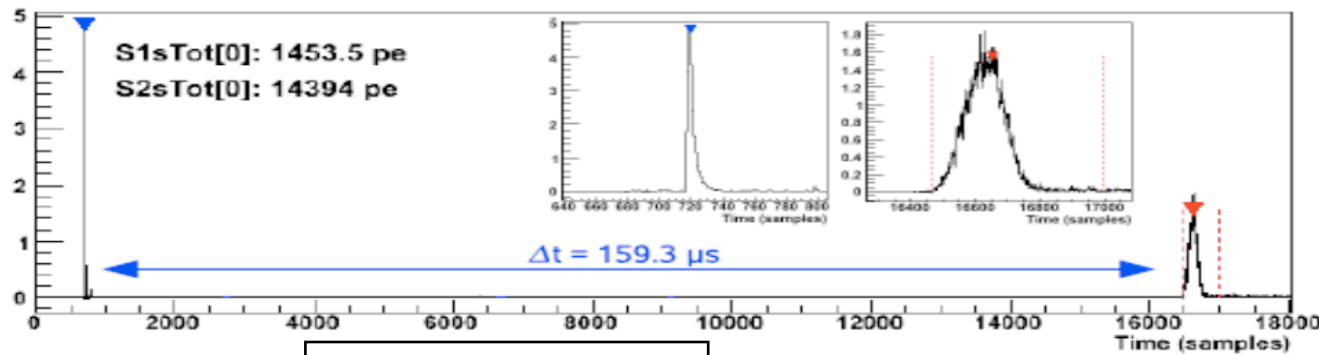
JOHANNES GUTENBERG  
UNIVERSITÄT MAINZ

# Xenon100 design: TPC

- ~161 kg total / ~62 kg target LXe (15 cm radius, 30 cm drift)
- Active LXe veto (64 PMTs)
- 70 new high QE (>32%@175nm) low activity 1" R8520 PMTs (total 242 PMTs)

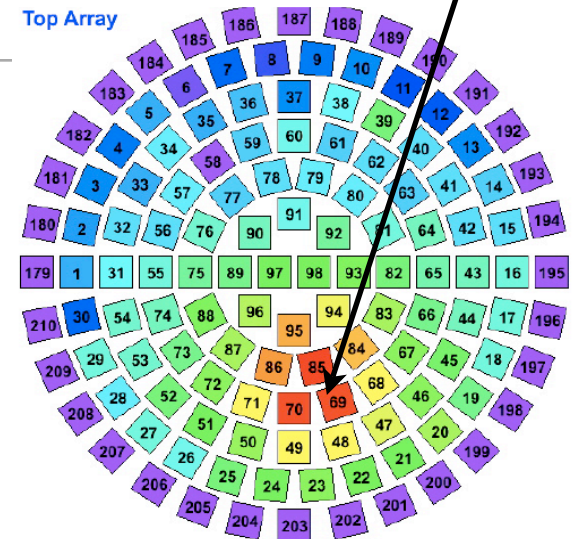


# Xenon100: Position reconstruction



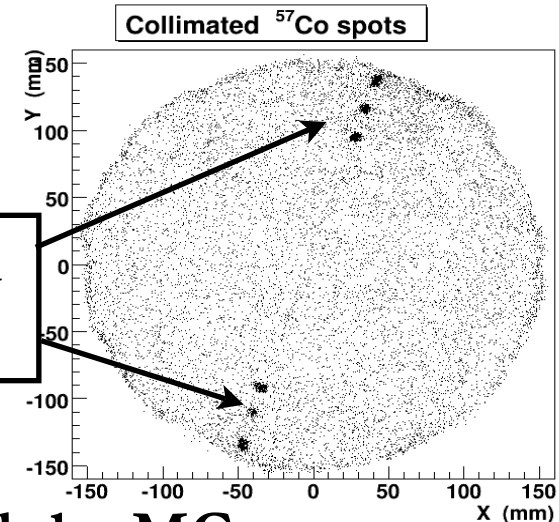
drift time  $\rightarrow z$

Very localized S2 hit pattern (xy position information)



3 different methods for xy position reconstruction:  
neural network  
support vector machine  
Least squares minimization

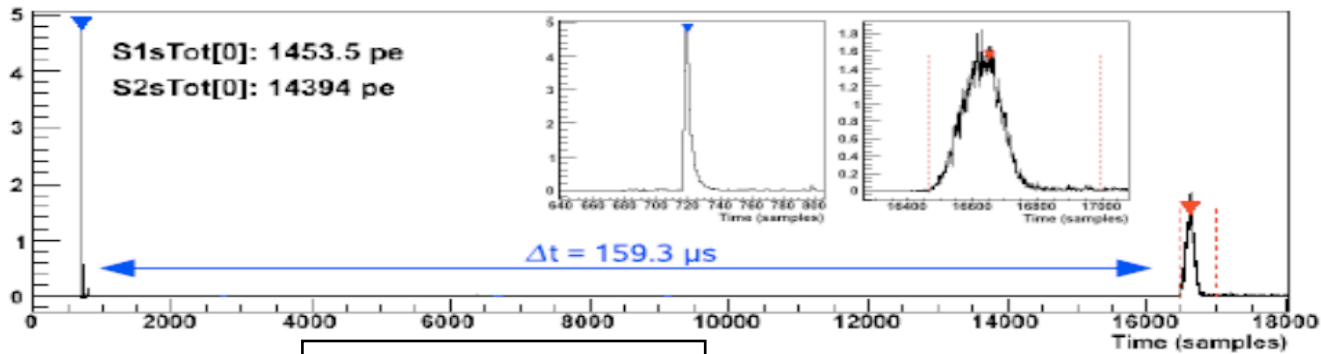
position resolution measured  
with collimated source



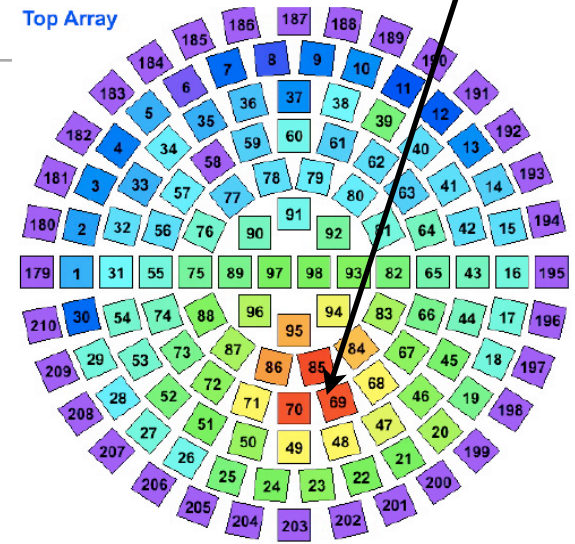
Agreement between the results and the MC  
yields a resolution  $\leq 3$  mm

# Xenon100: Position reconstruction

Very localized S2 hit pattern (xy position information)

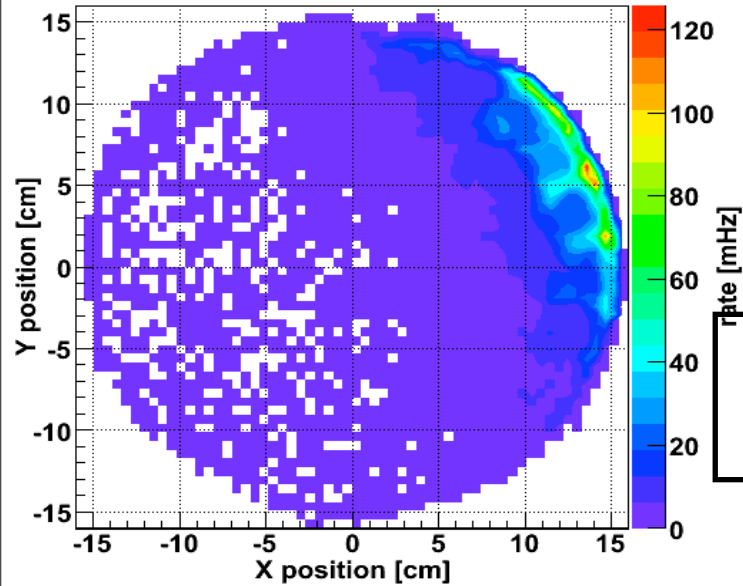


drift time  $\rightarrow z$

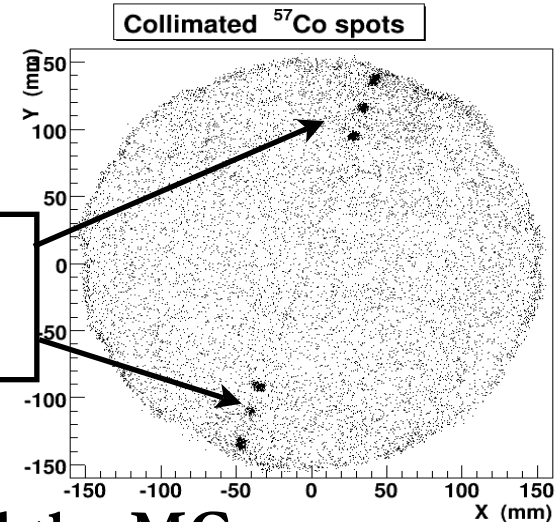


3 different methods for xy position reconstruction:  
 neural network  
 support vector machine  
 Least squares minimization

position resolution measured with collimated source



Cs137 from the side



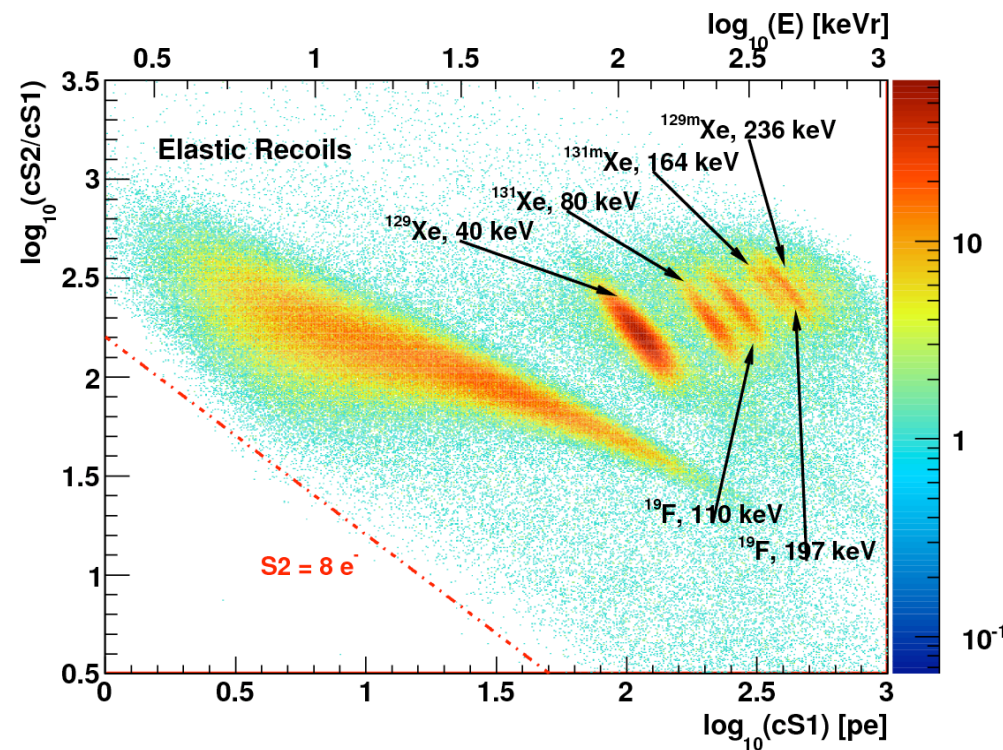
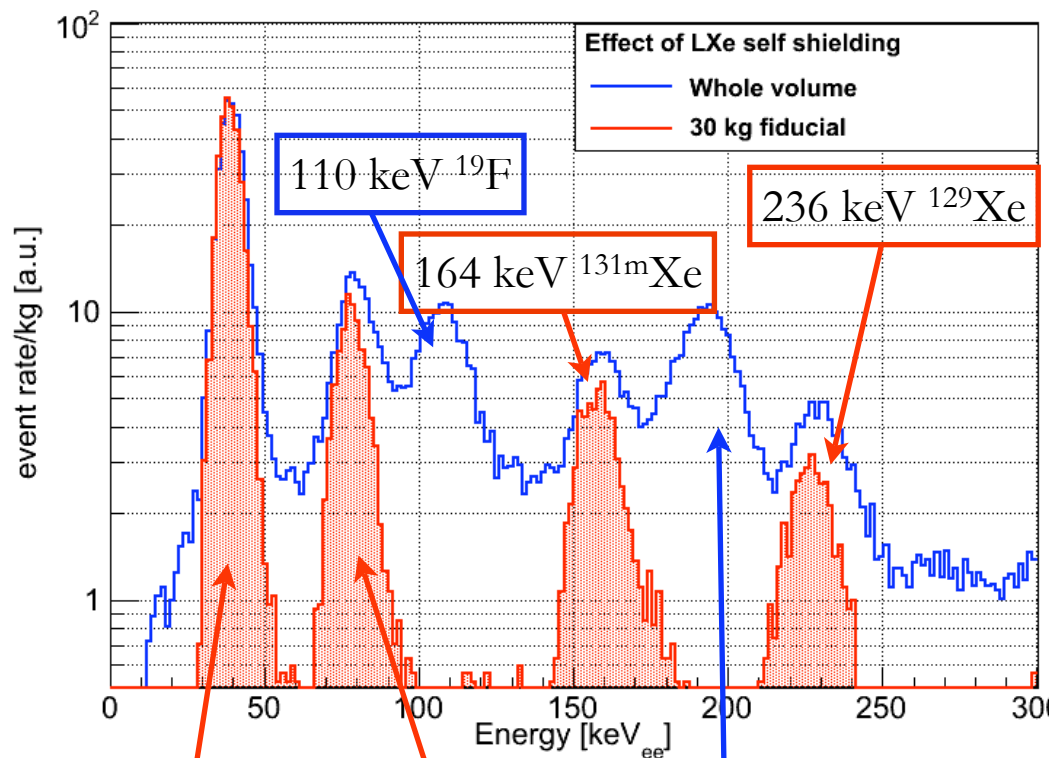
Agreement between the results and the MC yields a resolution  $\leq 3$  mm

# Xenon100: calibration

Gamma sources:

- $^{137}\text{Cs}$  for regular detector checks and calibration
- $^{60}\text{Co}$  electron recoil response determination
- Xenon inelastic and activation lines from AmBe run

Neutron source:  $^{241}\text{AmBe}$

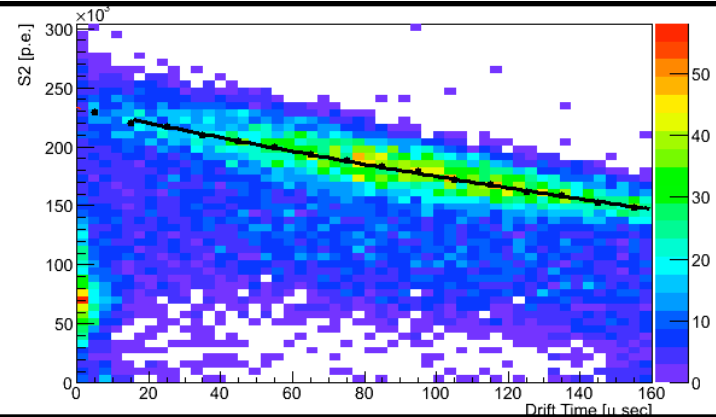




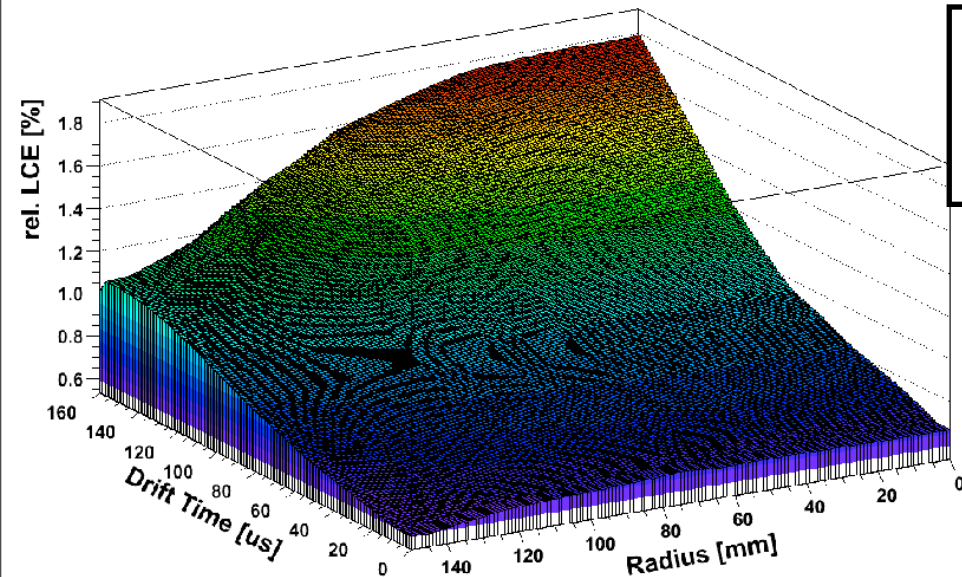
# Xenon100: signal position dependence

- Light yield from different positions in the detector changes due to solid angle, absorption length and teflon reflectivity
- Several sources distributed in the active volume have been used to measure the collection efficiency of the detector
- The results from these sources (40 keV inelastic,  $^{131}\text{mXe}$ , and  $^{137}\text{Cs}$ ) agree within each other

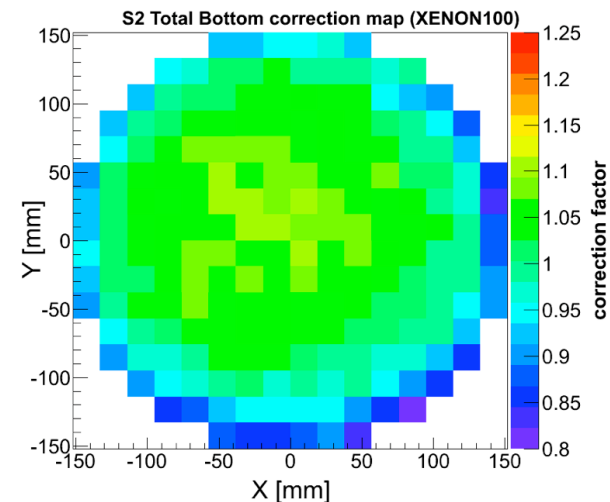
Signal corrected by the electron lifetime:  $Q_o \sim Q e^{dt/T}$



Differences in the signal due to the different solid angles in different XY positions are also corrected. No inhomogeneity is observed



Average light yield with electric field  
2.2 pe/keV @ 122 keV

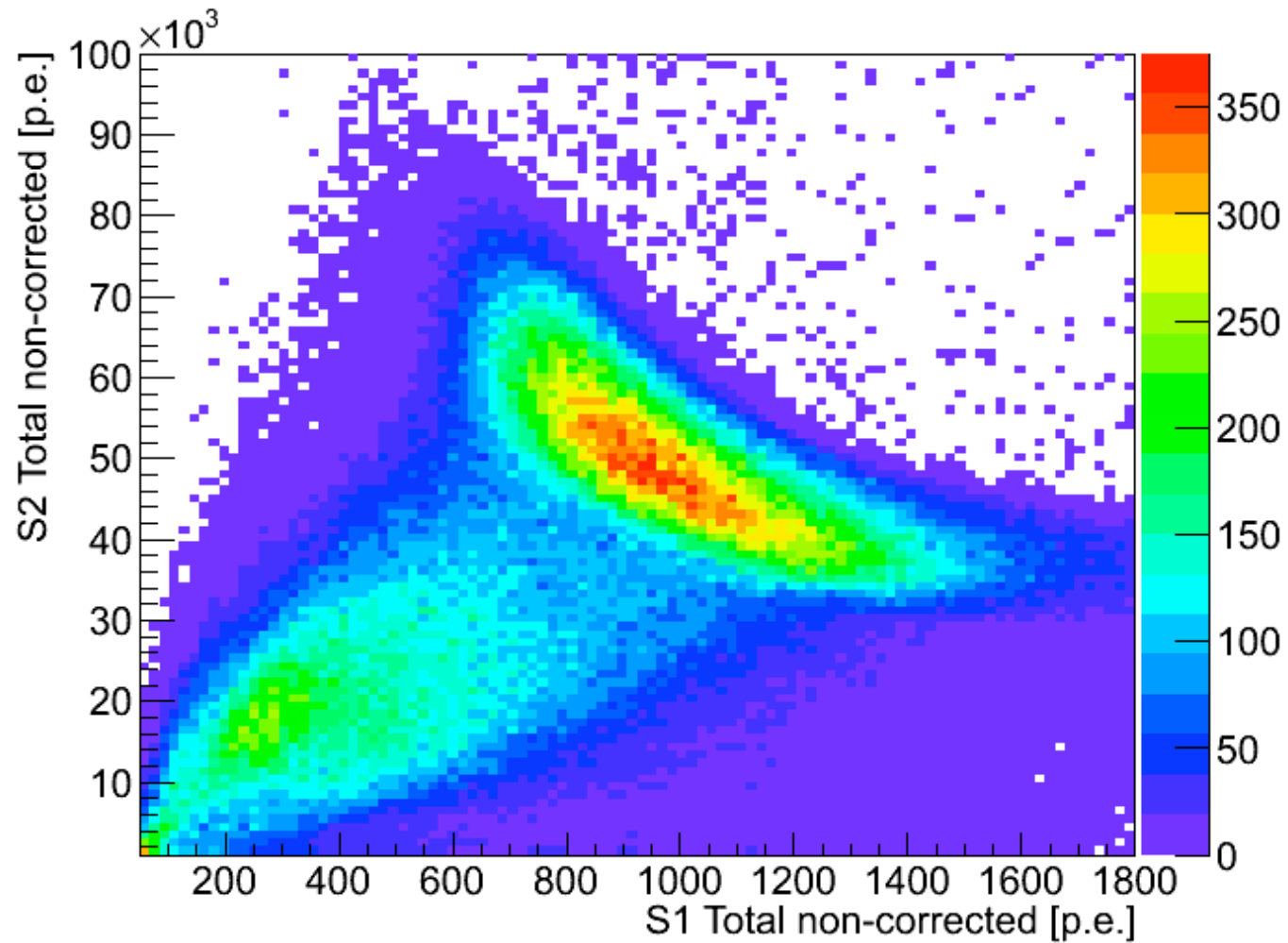


# Xenon100: calibration

Effect of the corrections:

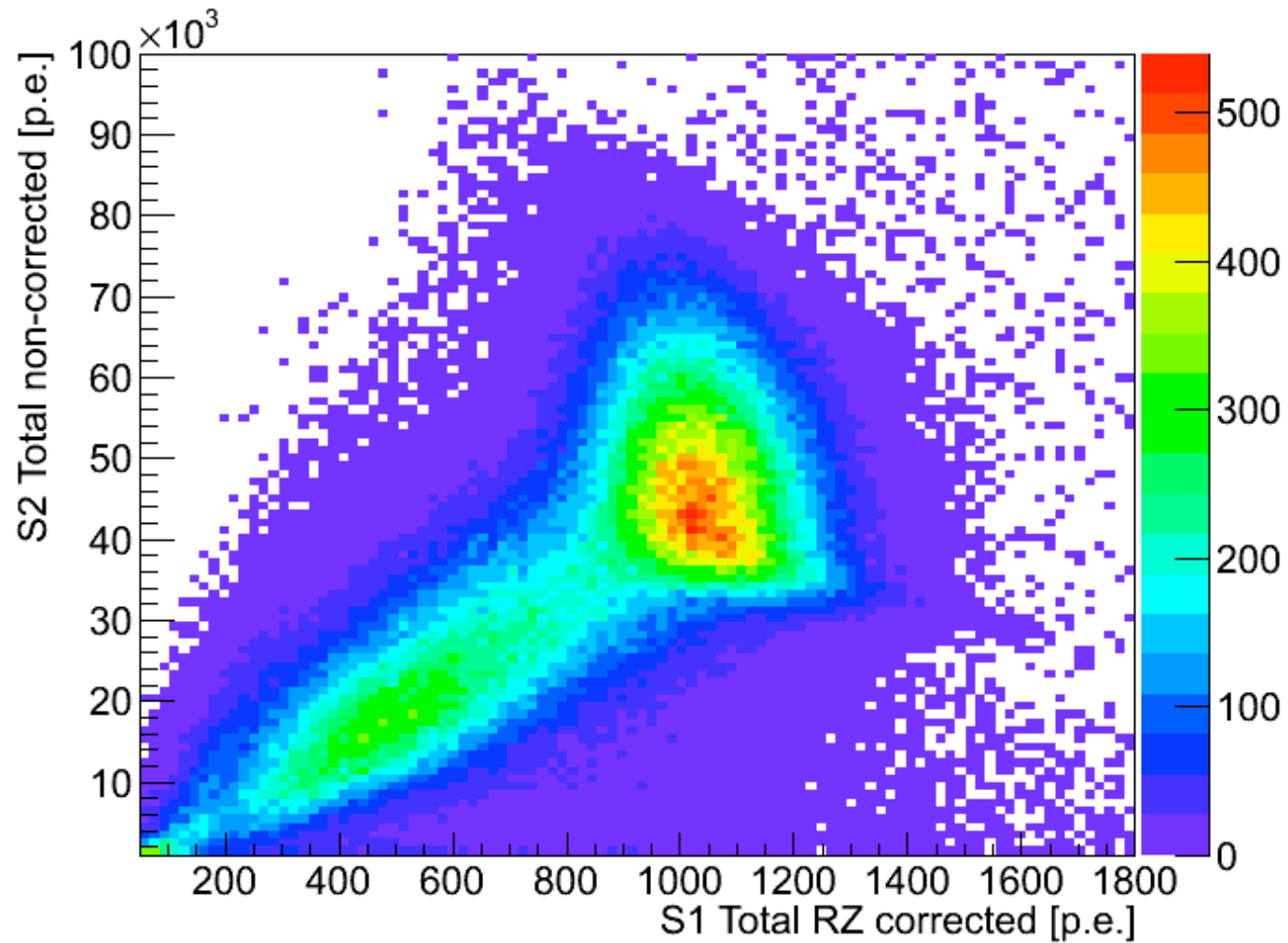
# Xenon100: calibration

Effect of the corrections:



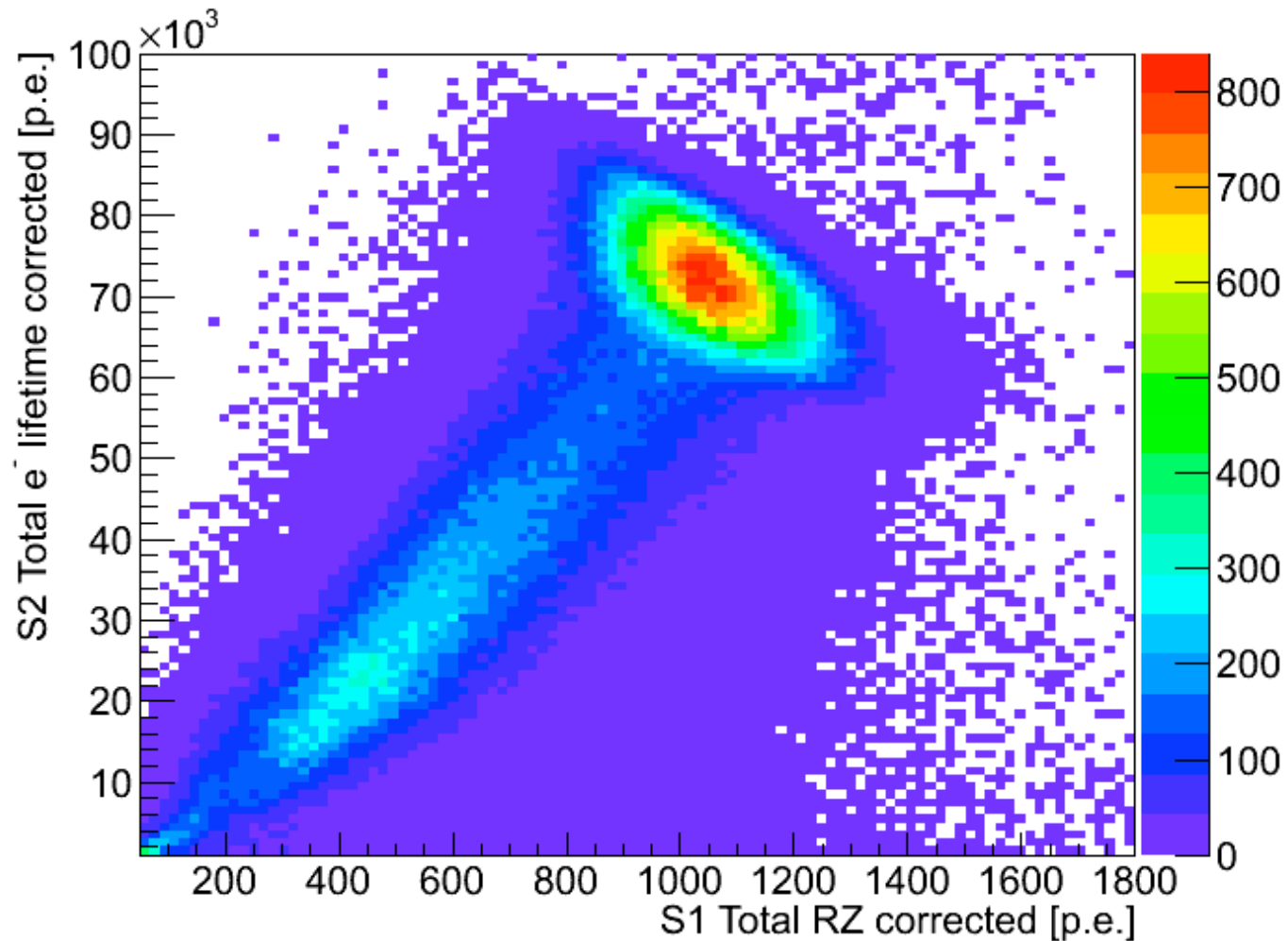
# Xenon100: calibration

Effect of the corrections:



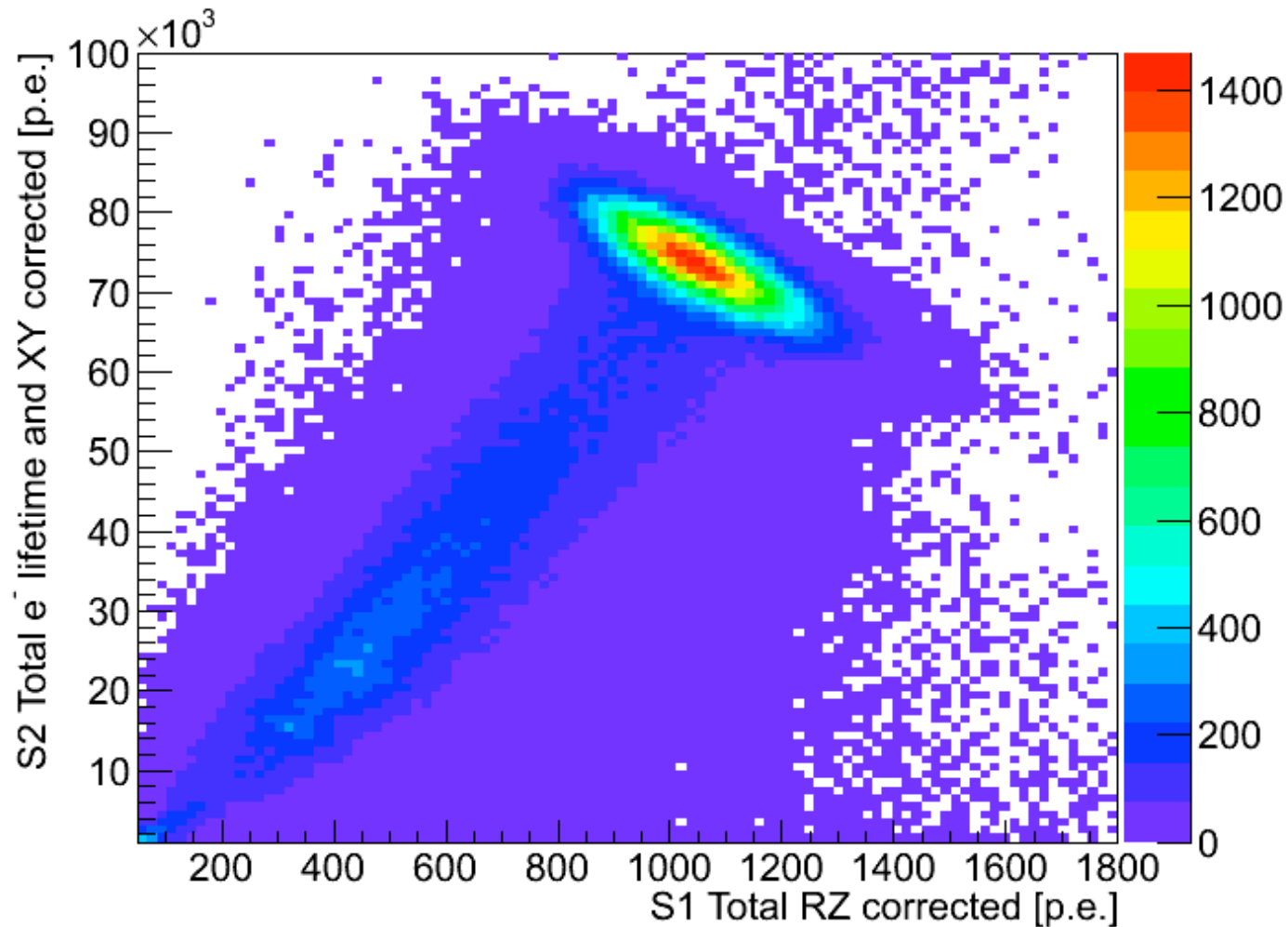
# Xenon100: calibration

Effect of the corrections:

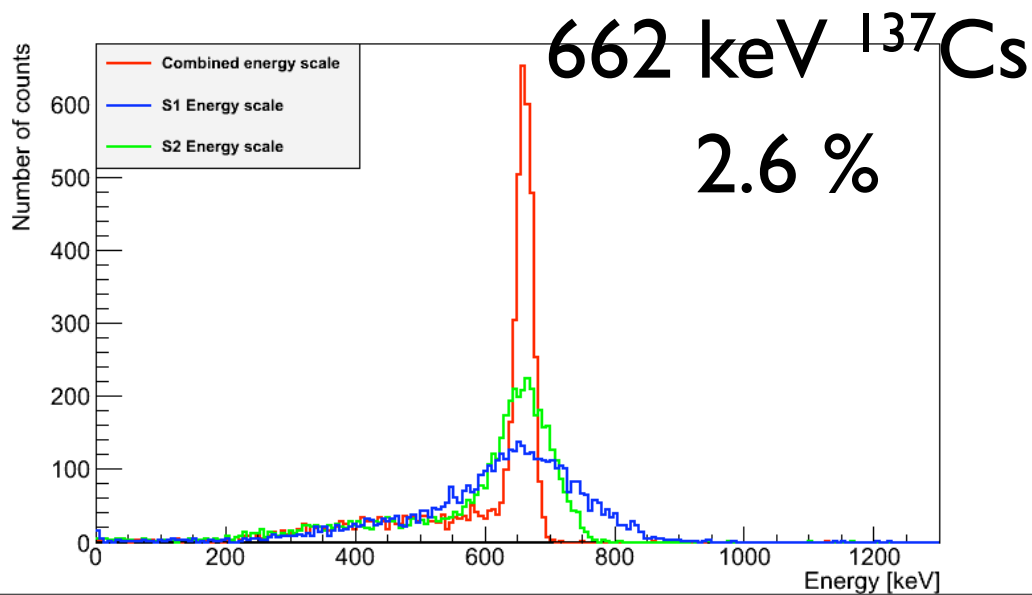
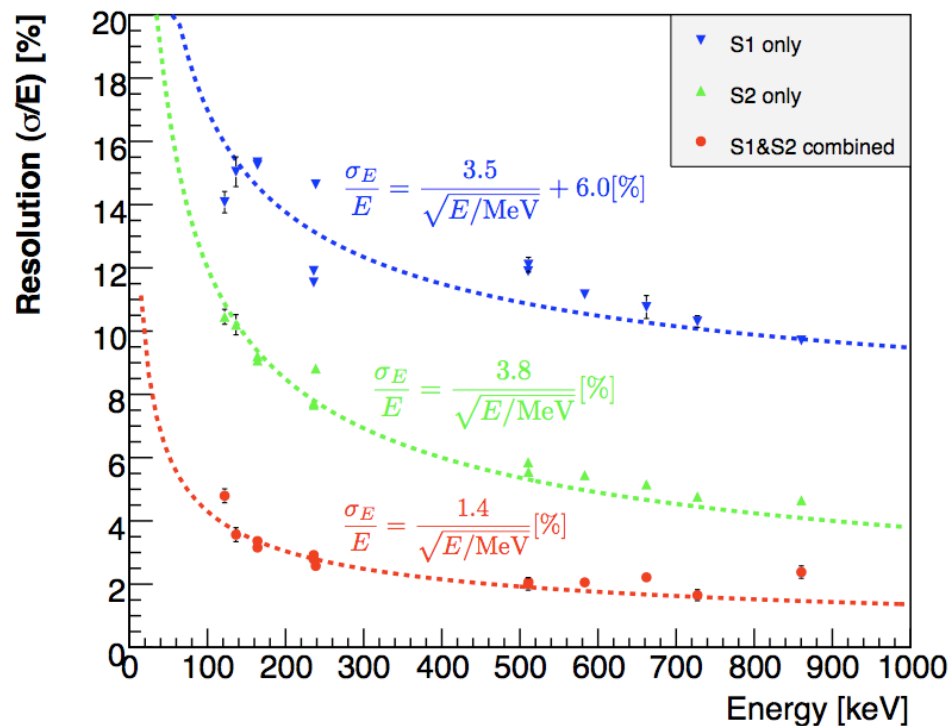
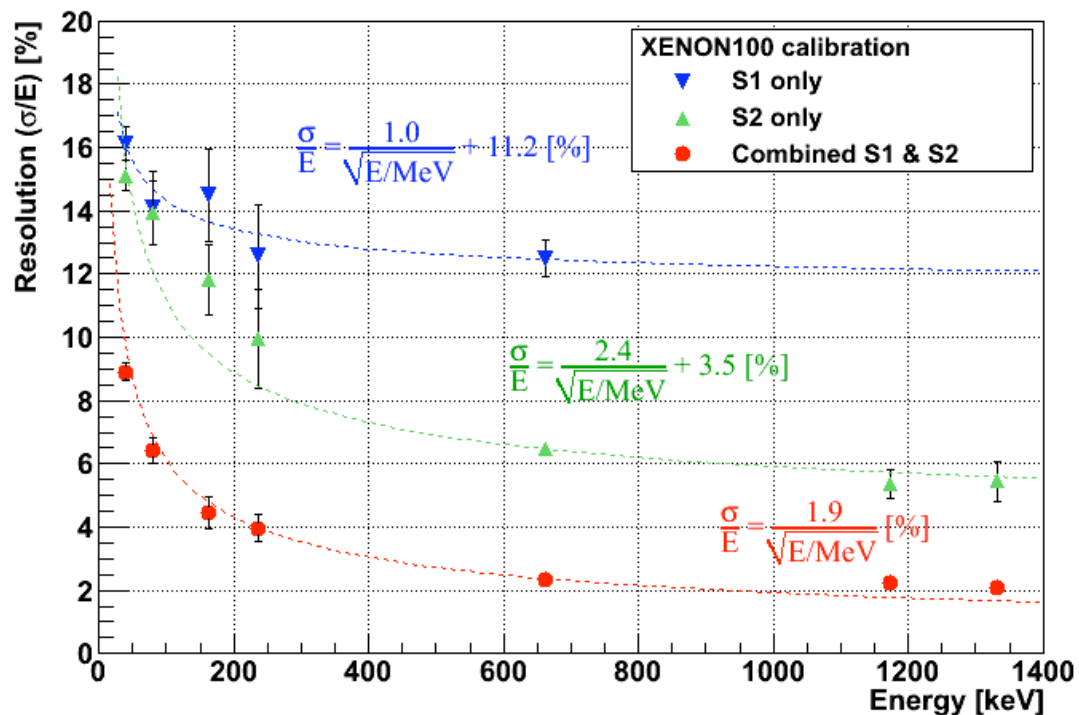


# Xenon100: calibration

Effect of the corrections:

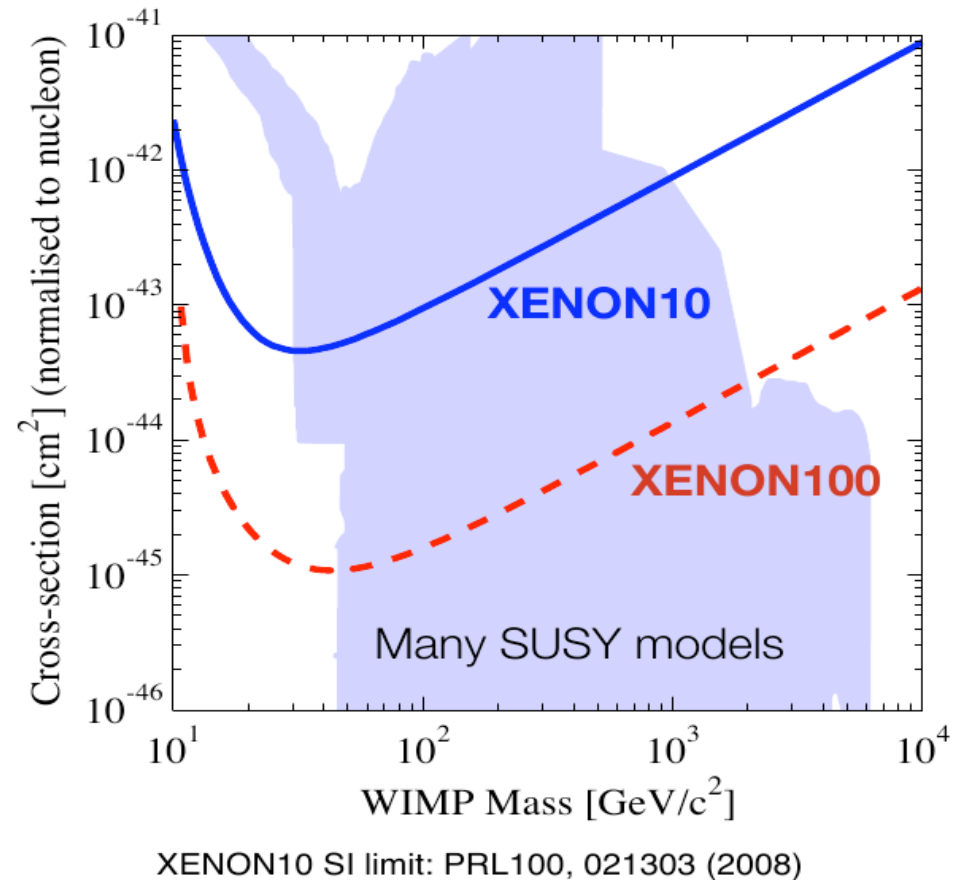


# Xenon100: calibration



# Xenon100: goals

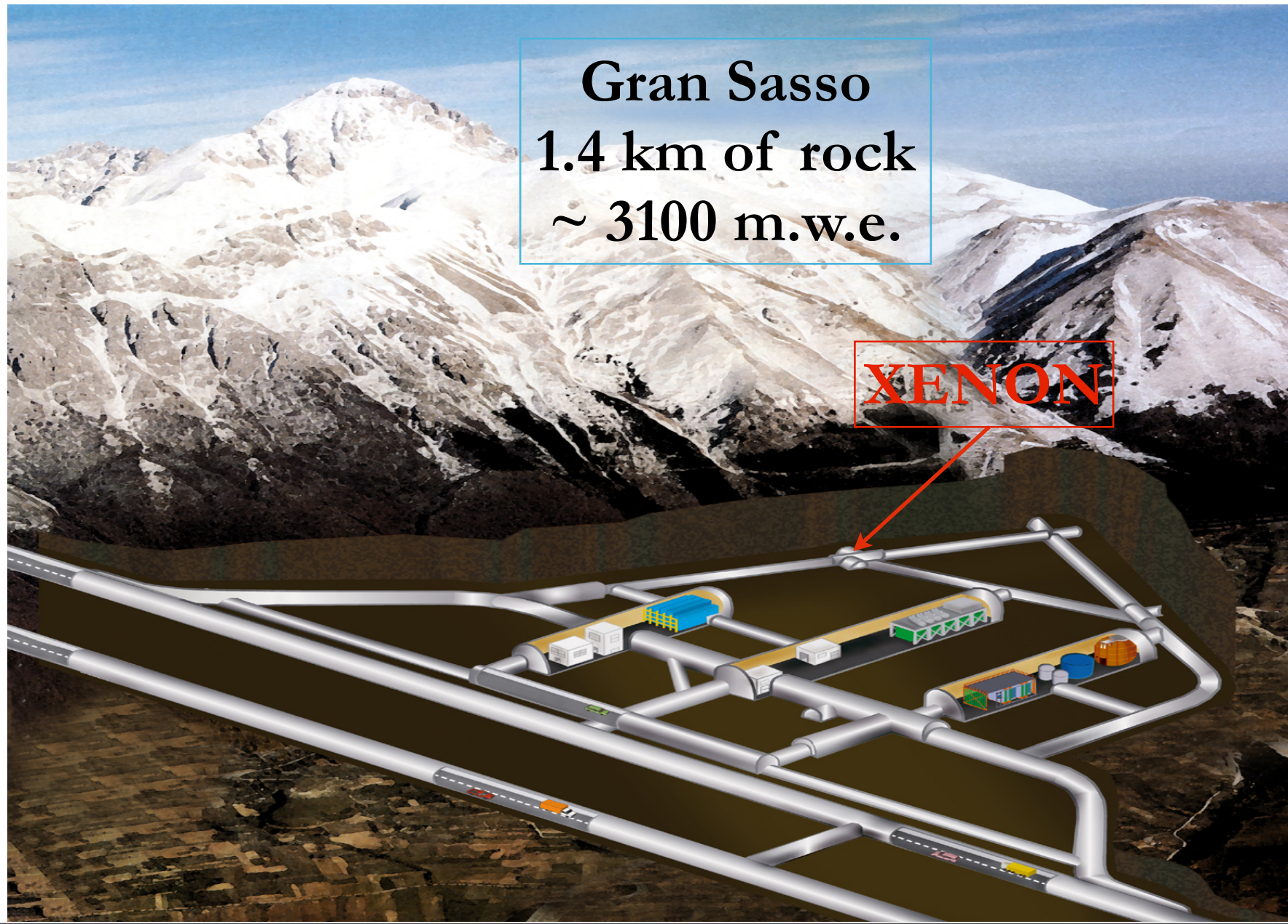
- Improve the sensitivity  $\sim 50$  times over XENON10.
- Assuming same energy threshold and same discrimination power as XENON10, the required background in the fiducial volume needs to be 100 times lower with a mass increase of a factor 10.



What was done in order to reach the goal?



# Install the detector underground...



Most of the stuff goes outside of the shield (improved)...

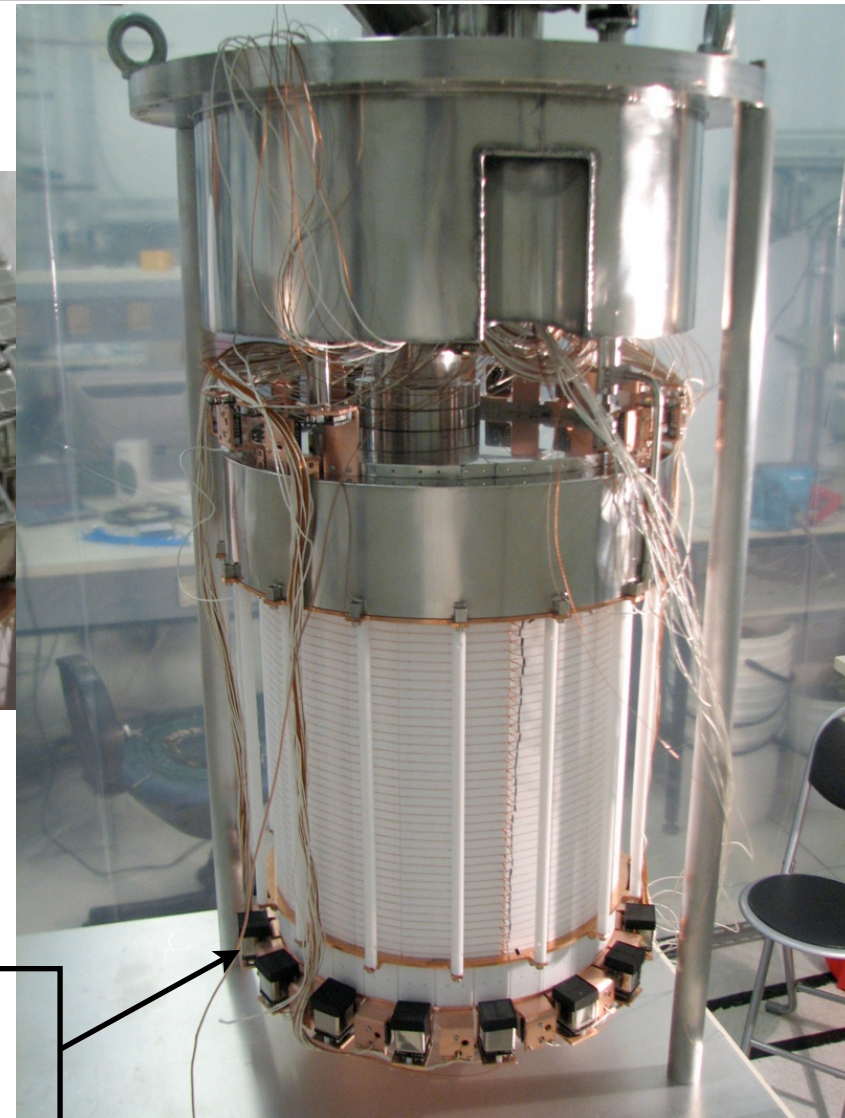
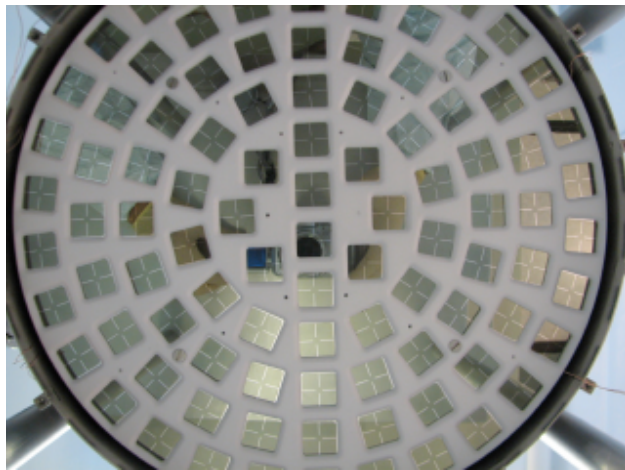
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# What is inside has to be carefully selected



242 (Hamamatsu R8520) 1"x1"  
low radioactivity PMTs



SS  
PTFE  
Copper  
Cables  
Screws

100 kg LXe Active veto  
(side, top and bottom)

# Material screening results (selection)

## Stainless Steel

Material	$^{238}\text{U}$ [mBq/kg]	$^{232}\text{Th}$ [mBq/kg]	$^{60}\text{Co}$ [mBq/kg]	$^{40}\text{K}$ [mBq/kg]
25 mm SS Nironit (flange and bars)	< 1.3	$2.9 \pm 0.7$	$1.4 \pm 0.3$	< 7.1
2.5 mm SS Nironit (bottom cryo)	< 2.7	< 1.5	$13 \pm 1$	< 12

## Inner detector materials

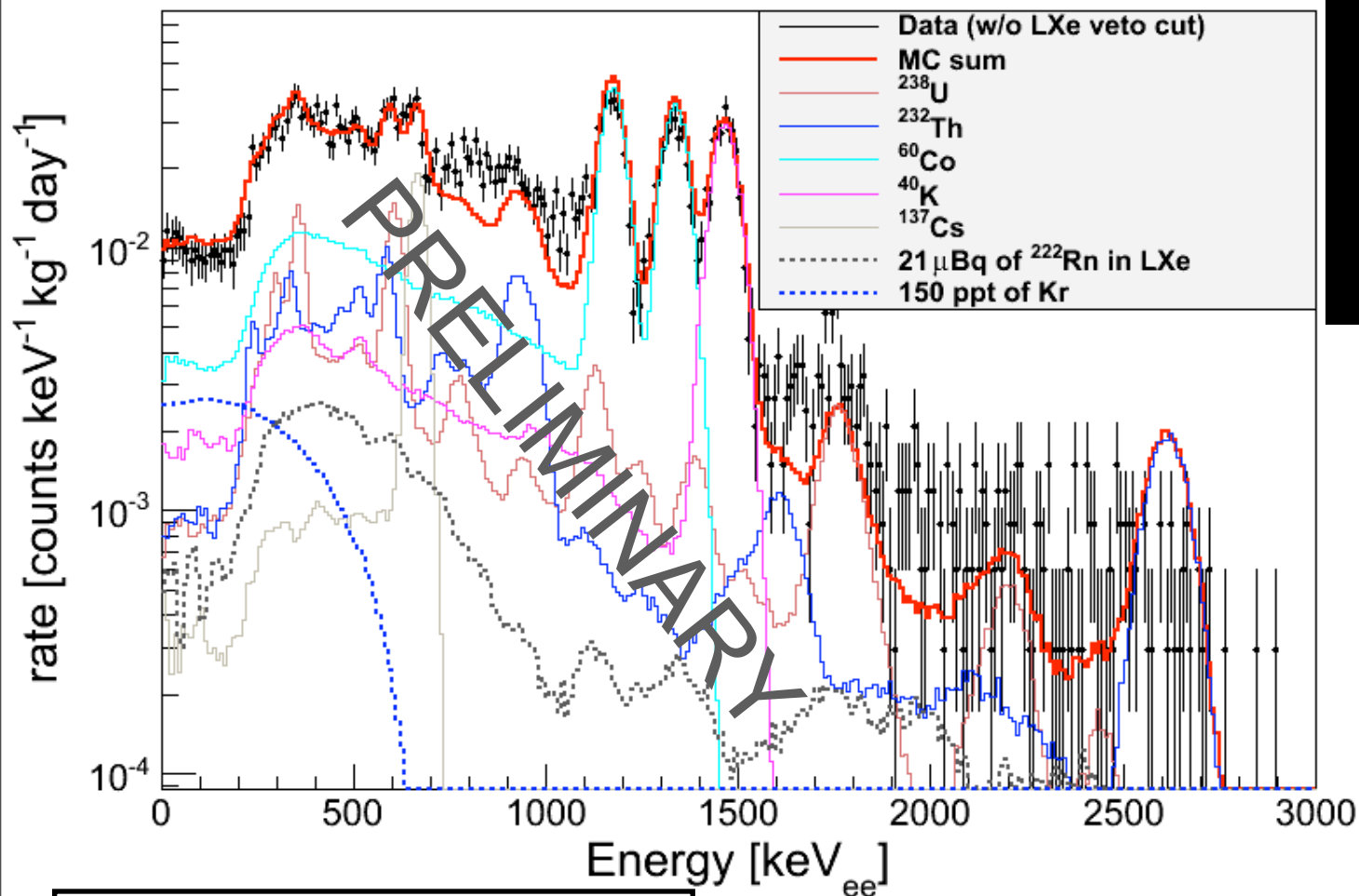
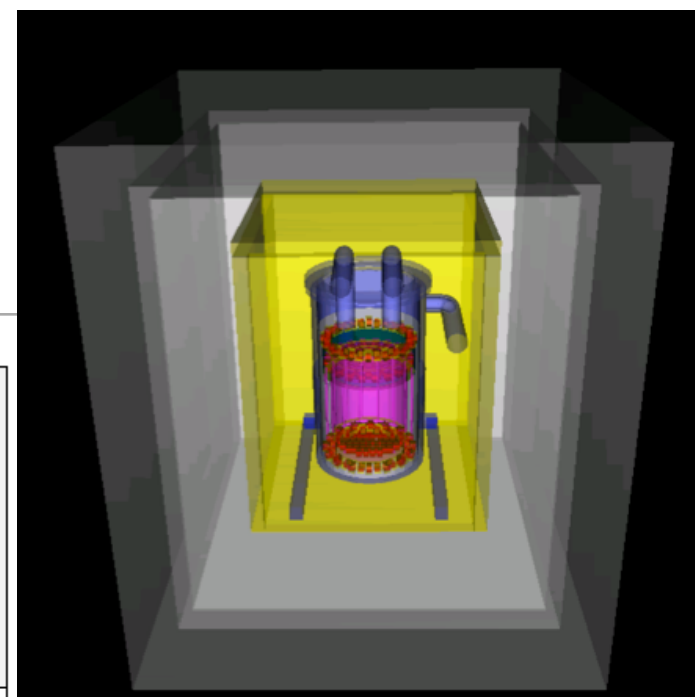
PMT Bases (Cirlex)	$65 \pm 8$	$31 \pm 10$	< 3.6	< 66
Teflon (in use)	< 0.31	< 0.16	< 0.11	< 2.25
Copper (TPC inner structure)	< 0.22	< 0.21	$0.21 \pm 0.07$	< 1.34
Small Screws (SS)	< 9.2	$16 \pm 4$	$9 \pm 3$	< 46.4

## PMTs

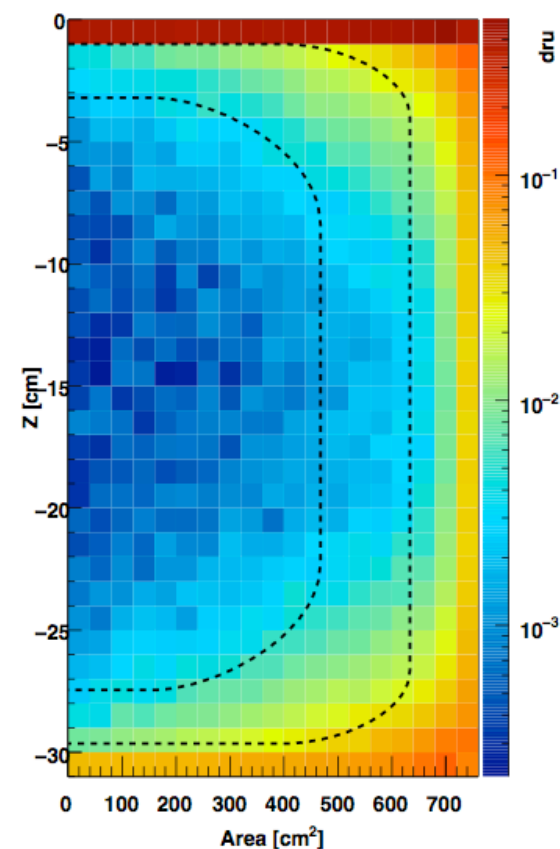
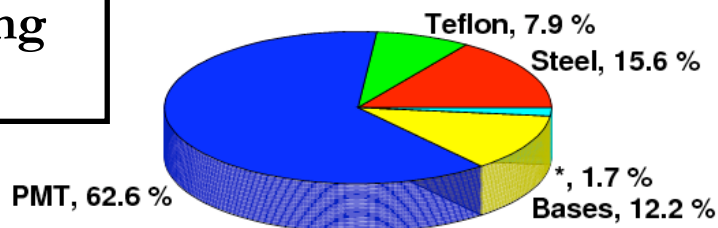
	$^{238}\text{U}$ [mBq/PMT]	$^{232}\text{Th}$ [mBq/PMT]	$^{60}\text{Co}$ [mBq/PMT]	$^{40}\text{K}$ [mBq/PMT]
39 PMTs	$0.12 \pm 0.01$	$0.11 \pm 0.01$	$1.5 \pm 0.1$	$6.9 \pm 0.7$
48 PMTs	$0.11 \pm 0.01$	$0.12 \pm 0.01$	$0.56 \pm 0.04$	$7.7 \pm 0.8$
22 HQE PMTs	< 0.64	$0.18 \pm 0.06$	$0.6 \pm 0.1$	$12 \pm 2$
23 HQE PMTs	$0.16 \pm 0.05$	$0.46 \pm 0.16$	$0.73 \pm 0.07$	$14 \pm 2$

**Special thanks to Matthias Laubenstein (LNGS screening facility)**

# Gamma background



Only input from Screening  
NO TUNING

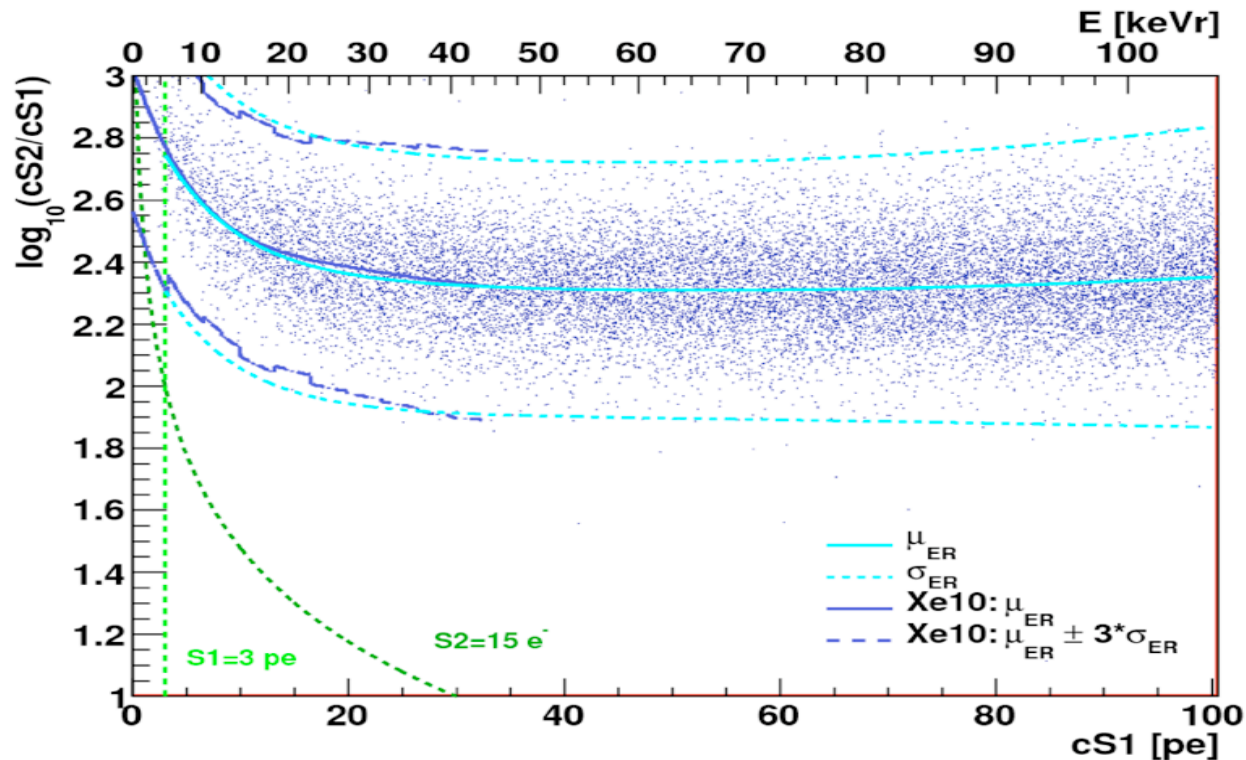


# Xenon100: gamma band

Multiple calibrations with  $^{60}\text{Co}$  to study the response of the detector to low energy electron recoils

Statistics achieved are more than 10 times the expected background

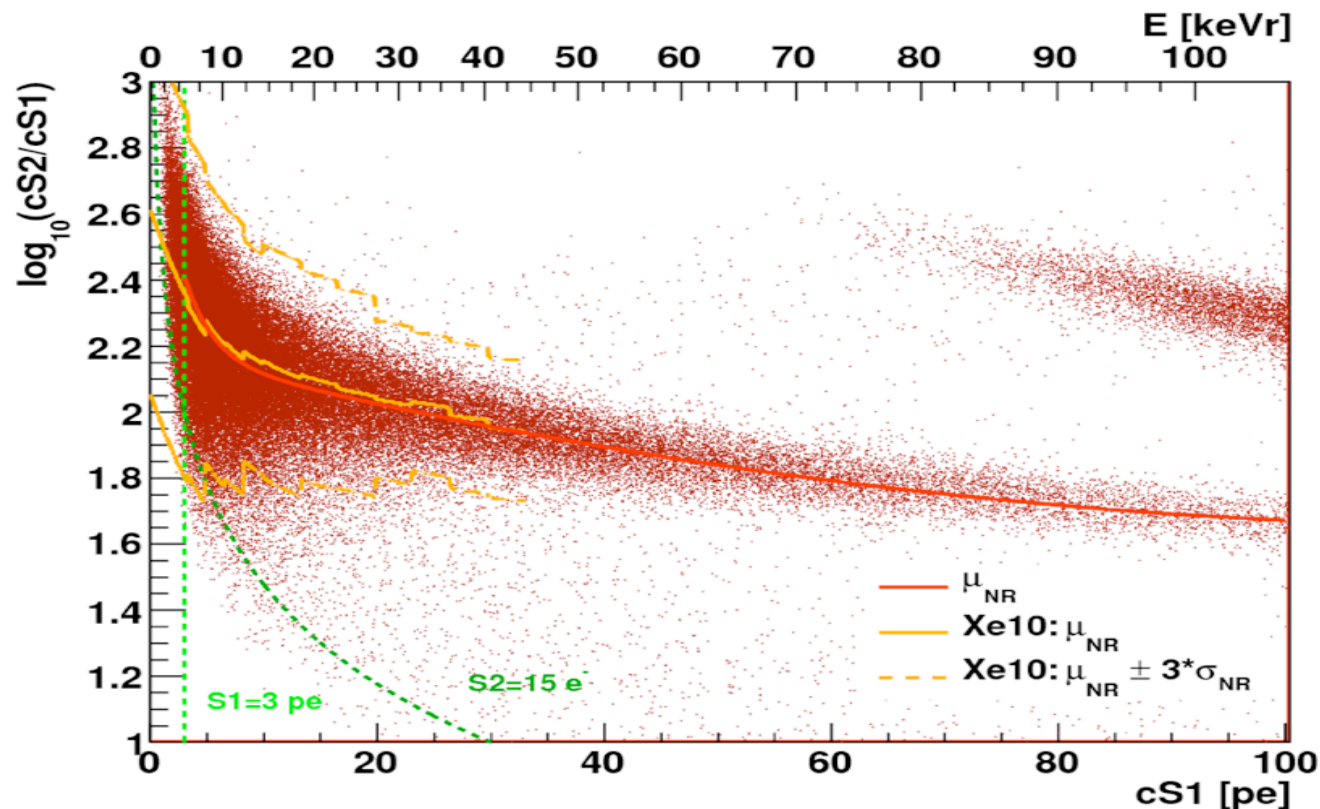
Results in good agreement with XENON10



# Xenon100: neutron band

Calibration of the detector using an AmBe source has been performed during December 2009

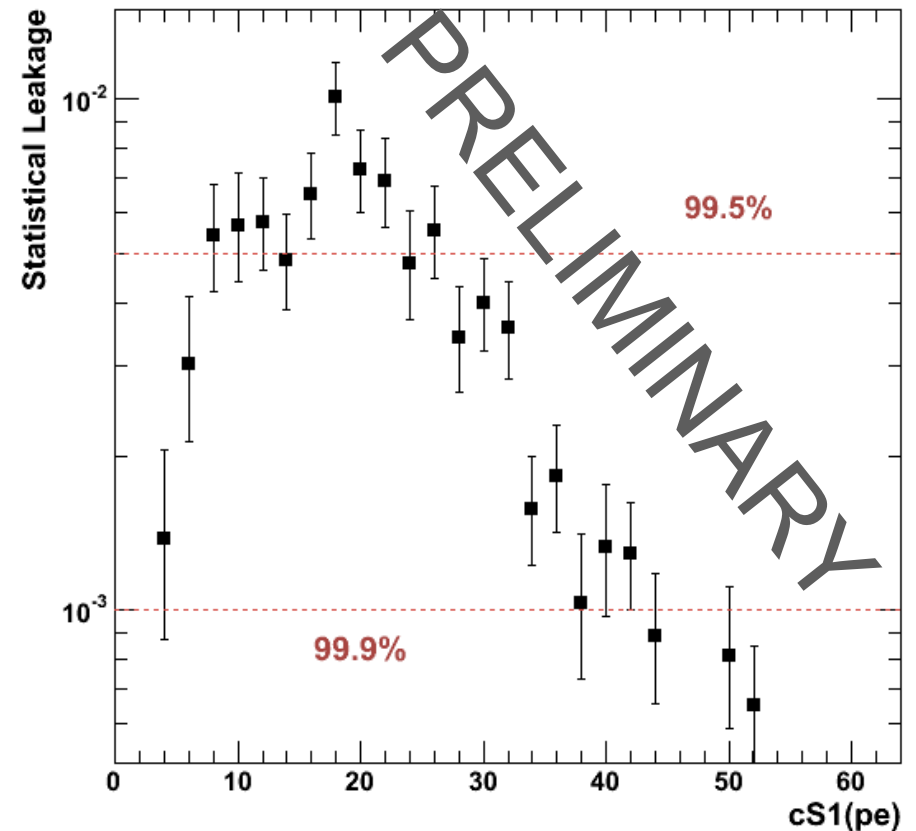
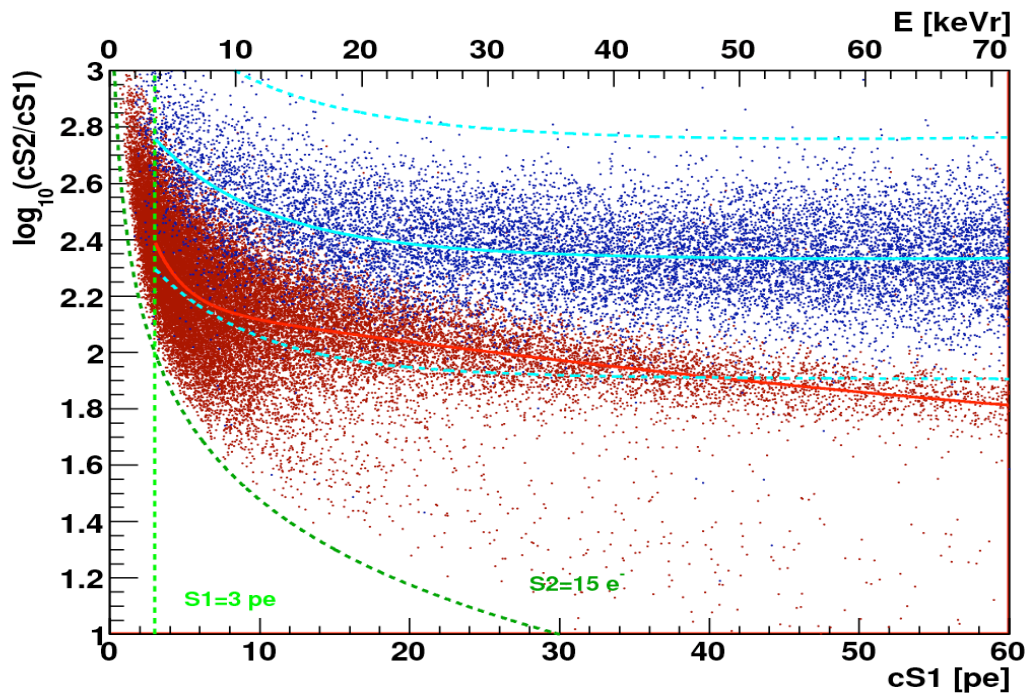
In addition to multiple gamma lines above 40keV, the detector response to low energy nuclear recoils has been studied  
Results are in good agreement with XENON10



# Xenon100: rejection power

It is possible to distinguish between nuclear recoils and electron recoils due to their different charge/light ratio

The rejection efficiency is  $\sim 99\%$  in the range from 4 to 20 pe





# Background analysis

11.2 days of non blinded data were taken in the period Oct-Nov 2009

Applied cuts are only optimized in calibration data

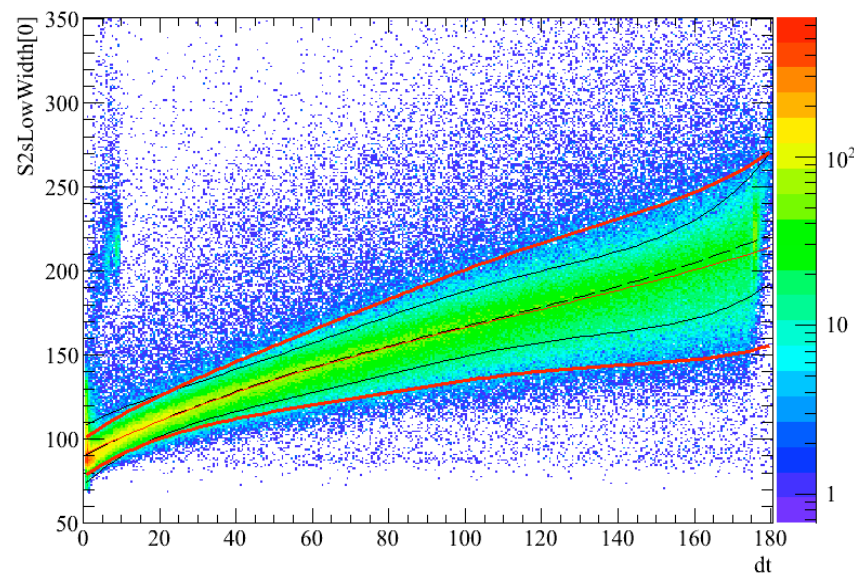
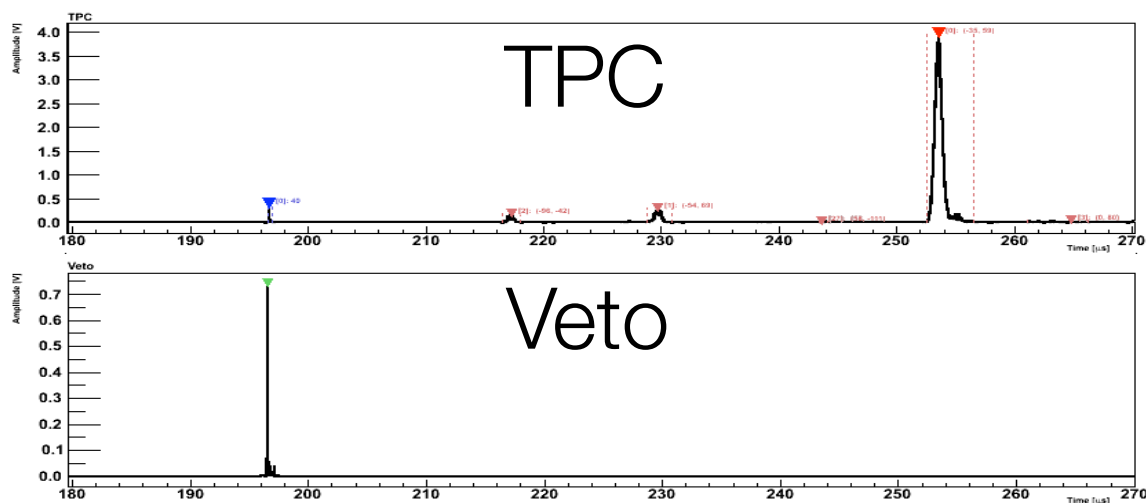
Only very basic cuts are used:

Single scatterers

Reasonable signal to noise ratio

Width and drift time of the event compatible(remove gas events)

Veto anticoincidence

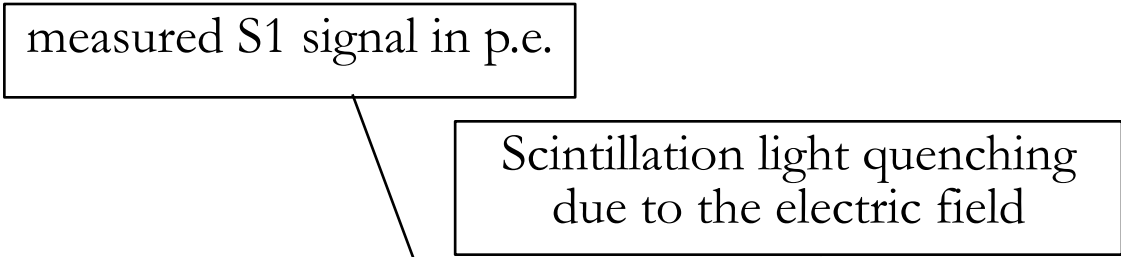


# Energy scale for nuclear recoils

We use a global fit of the available data to compute the quenching factor for nuclear recoils

Ongoing efforts to measure this quantity with a better precision

In XENON100 [4-20] pe  $\sim$  [7-27]keVr

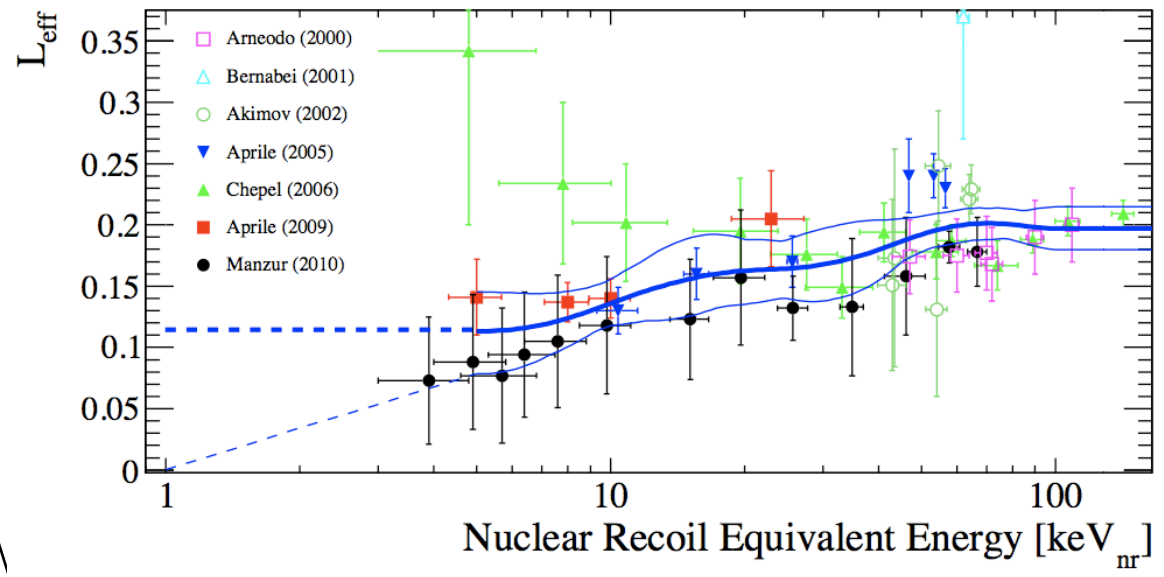


$$E_{nr} = \frac{S1}{L_y L_{eff}} \cdot \frac{S_e}{S_r}$$

Light yield @ 122 keV

Scintillation efficiency at 0 field

Scintillation light quenching due to the electric field

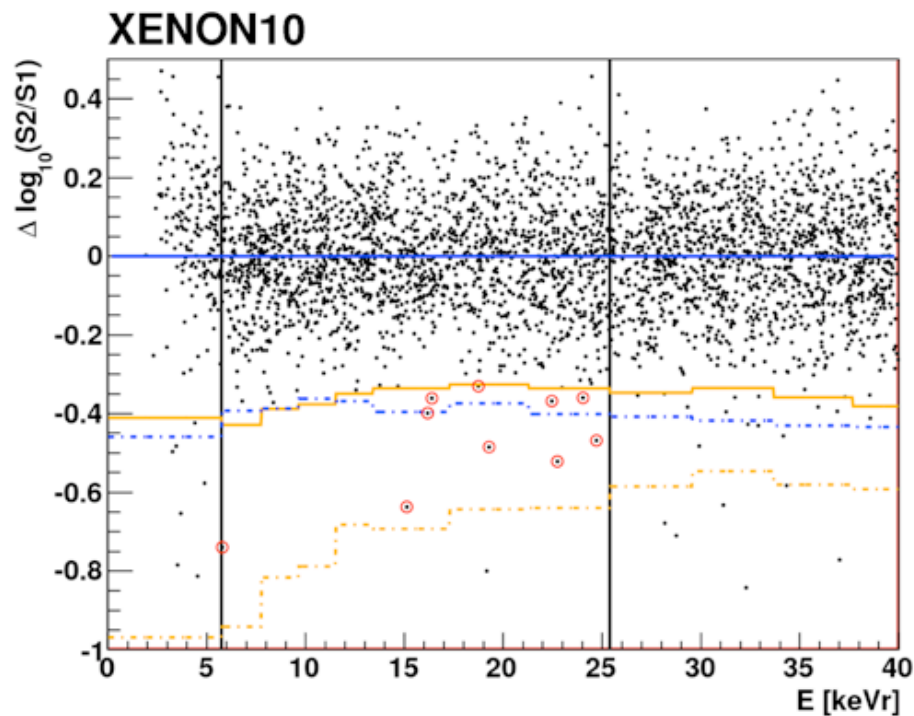


# Background analysis

## XENON10 PRL 100, 021303 (2008)

136 kg-days Exposure =

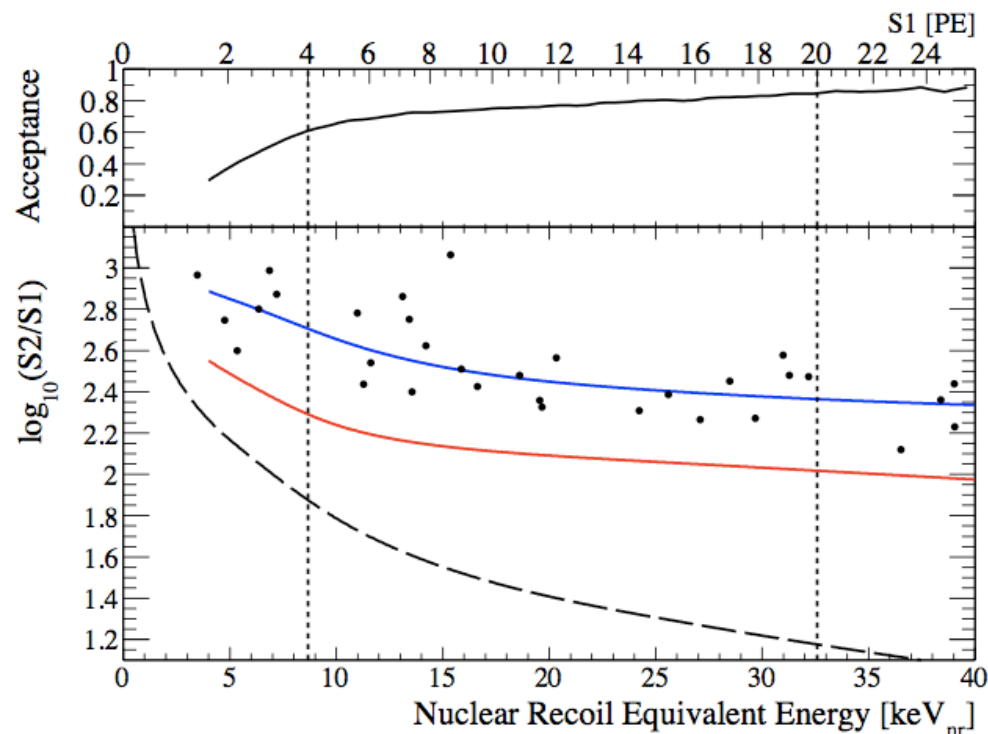
58.6 live days x 5.4 kg x 0.86 ( $\epsilon$ ) x 0.50 (50% NR)  
(data collected between Oct.2006 and Feb.2007)



## XENON100 PRL in preparation

161 kg-days Exposure =

11.2 live days x 40 kg x  $\epsilon$  x 0.50 (50% NR)  
(data collected between Oct.2009 and Nov.2009)



**0 events with a bigger exposure than XENON10!!**

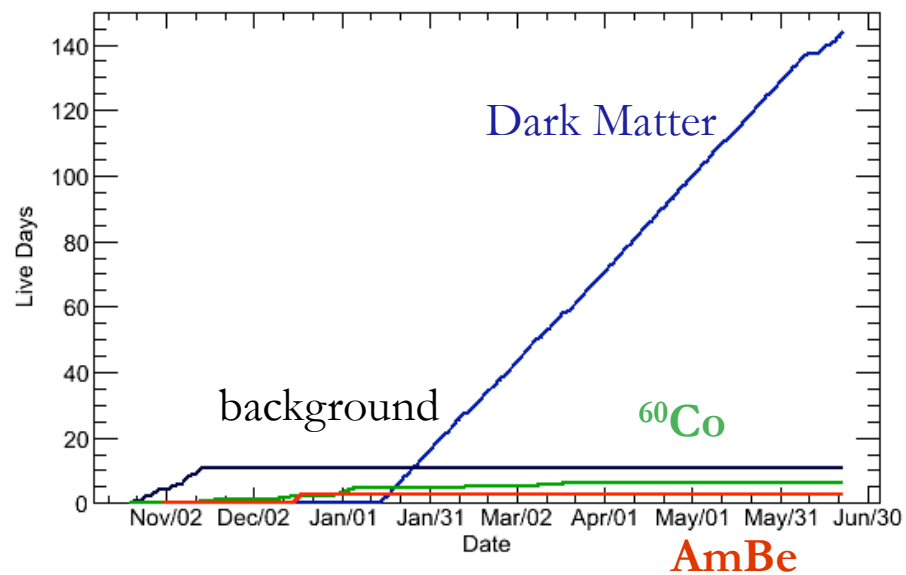
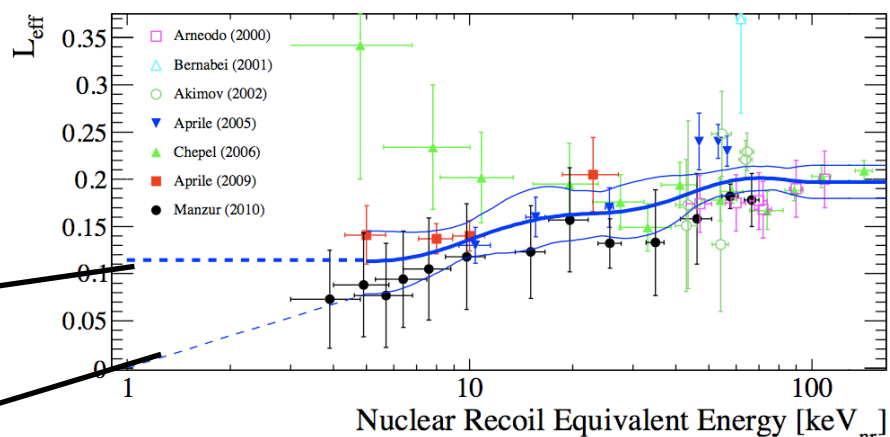
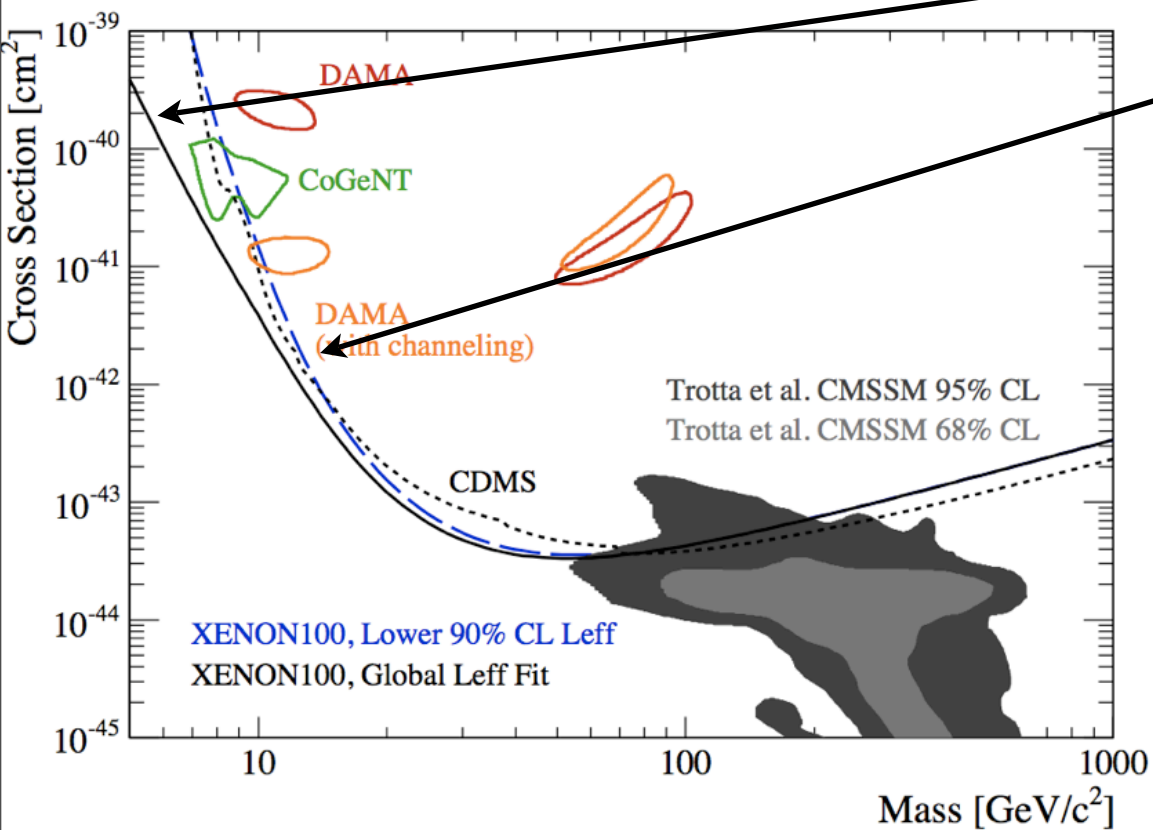
# Background analysis

Standard astrophysical assumptions:

$$v_o = 220 \text{ km/s}$$

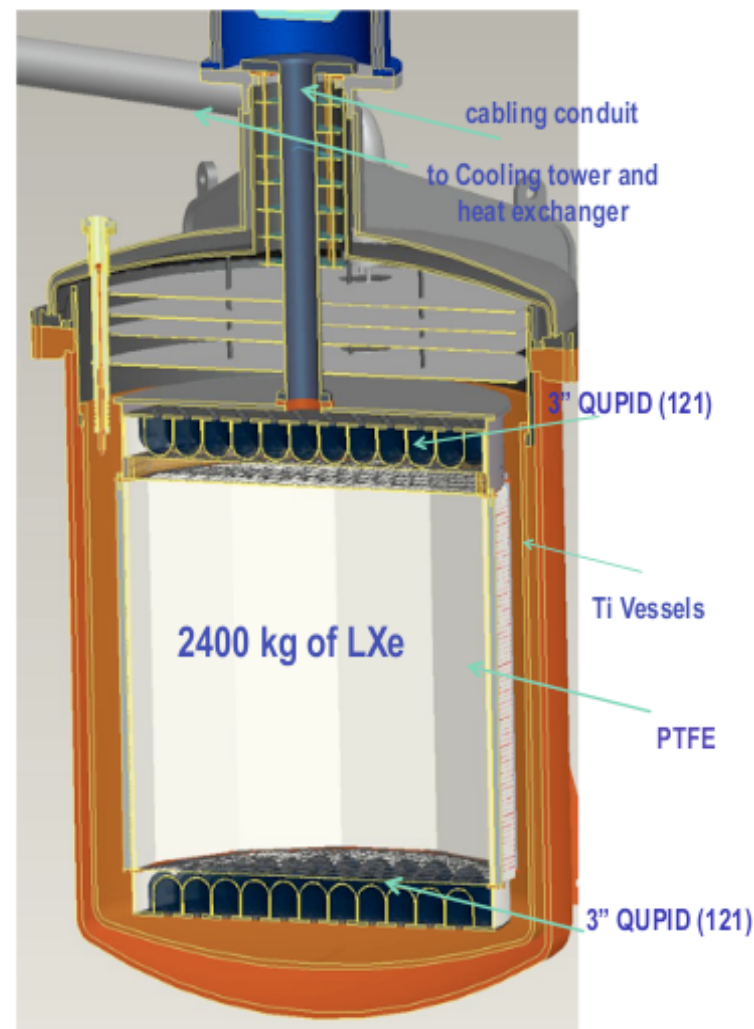
$$\rho = 0.3 \text{ GeV}/c^2$$

$$v_{esc} = 544 \text{ km/s}$$



# Future: XENON1T

- ➔ The Xenon100 detector has been successfully calibrated and is already taking science data, with a performance as good as expected
- ➔ Within this year, it will either see a signal or constrain significantly the models for WIMP SI or SD interactions
- ➔ In both cases, larger experiments with reduced backgrounds are needed
- ➔ Critical technologies developed within the XENON10/100 programs can be directly applied to the next scale. Risks and the costs are fully understood.
- ➔ A strong international collaboration, with valuable expertise and resources, is in place.
- ➔ A technical design proposal for a XENON1T is in preparation. With 50 - 50 share of resources between US and other groups, we plan to realize the experiment before 2015.

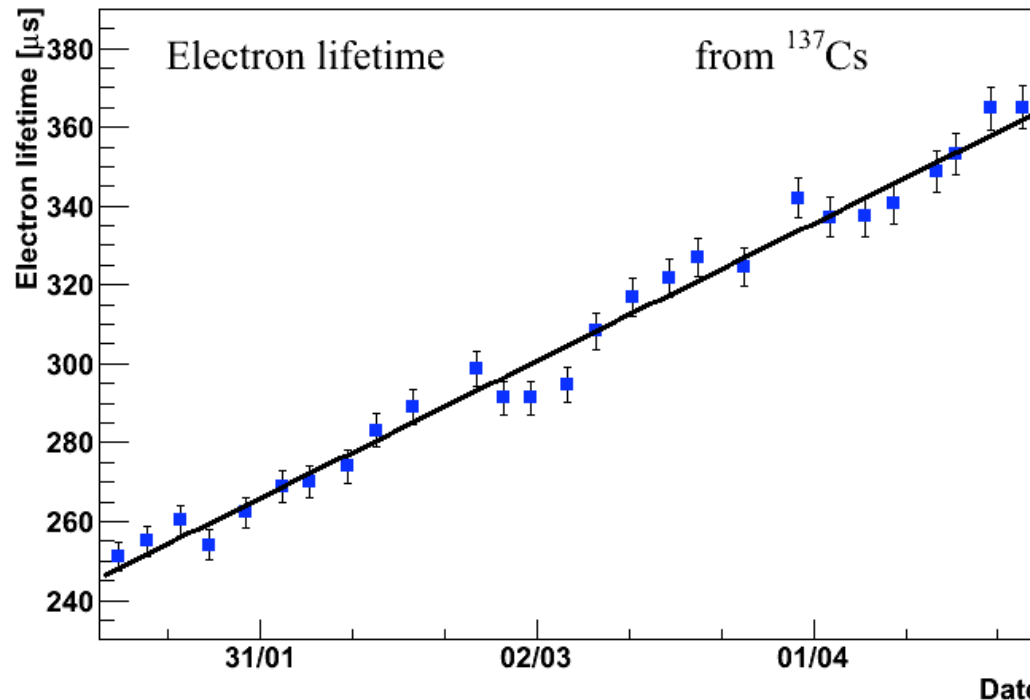
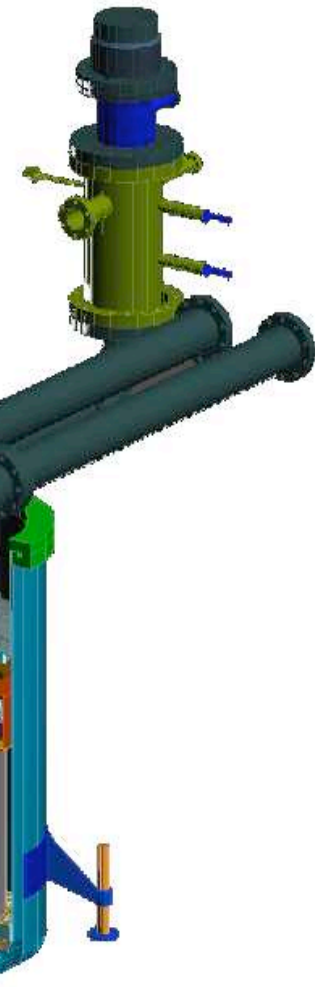


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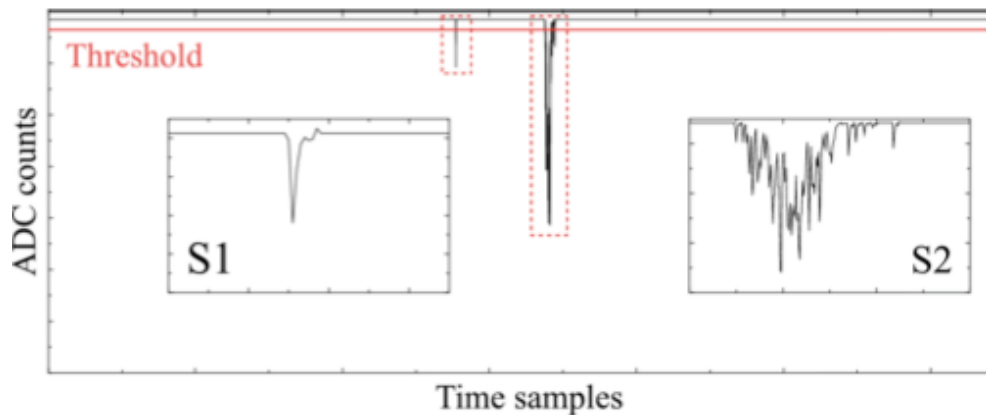
# Xenon100 design: Cooling system

- ➔ The Xenon is continuously recirculated and purified through a hot getter (SAES)
- ➔ Cooling power is provided by a Pulse Tube Refrigerator (160W)
- ➔ Vacuum cryostat extends outside the shield to surround the cooling tower
- ➔ Recirculation in gas phase 10 SLPM

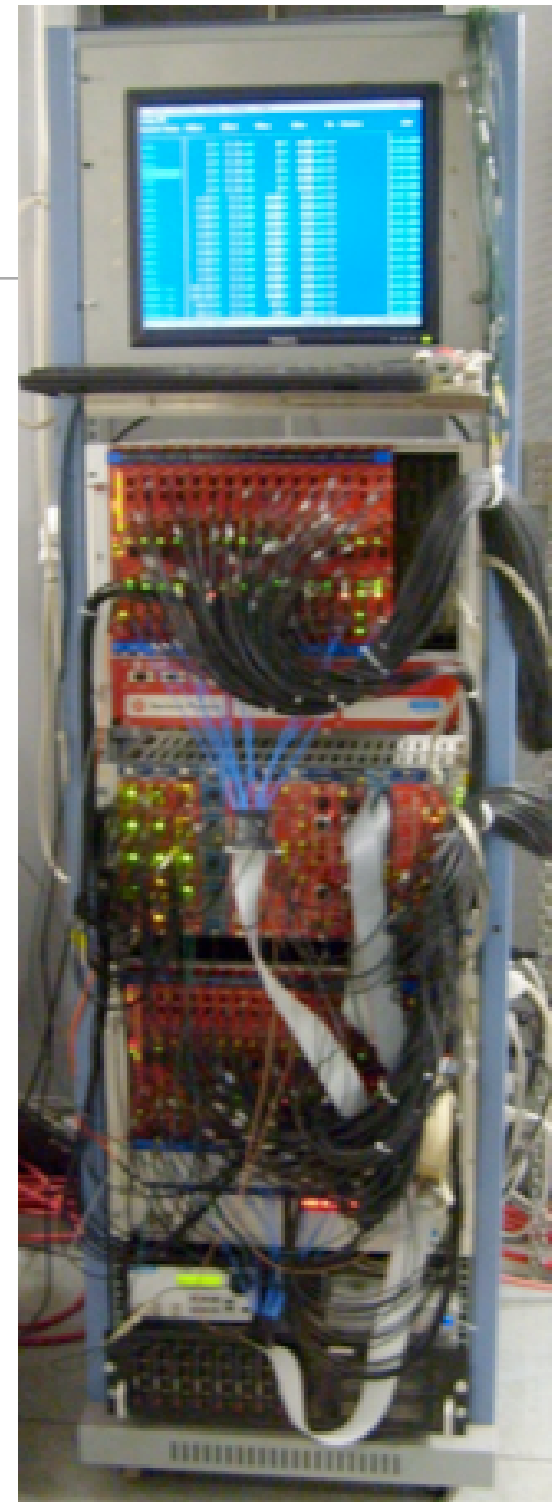


# Xenon100: Data Acquisition

- CAEN V1724 100 MHz digitizer (14 bit resolution)
- Circular buffer -> dead time free
- Integrated FPGA for zero length encoding



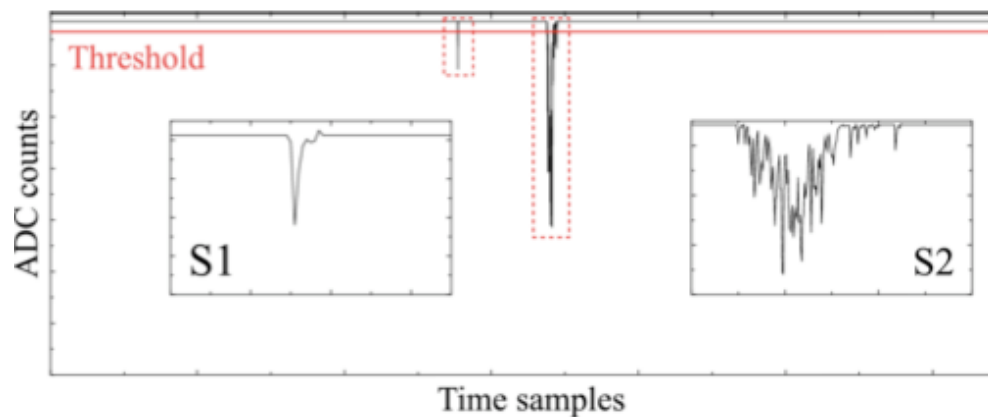
- Slow control to monitor the detector crucial parameters
- sms alarms are sent to people on shift in case of emergency



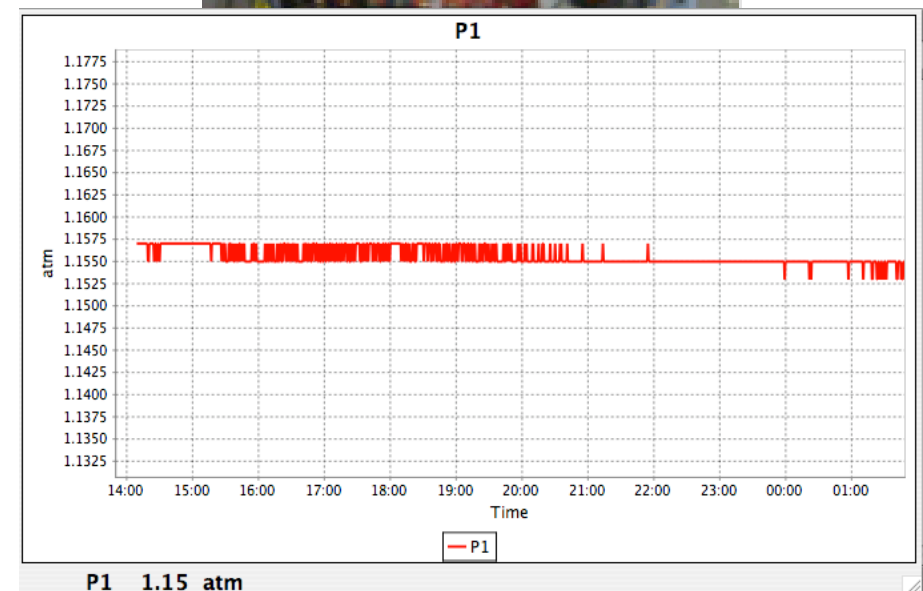
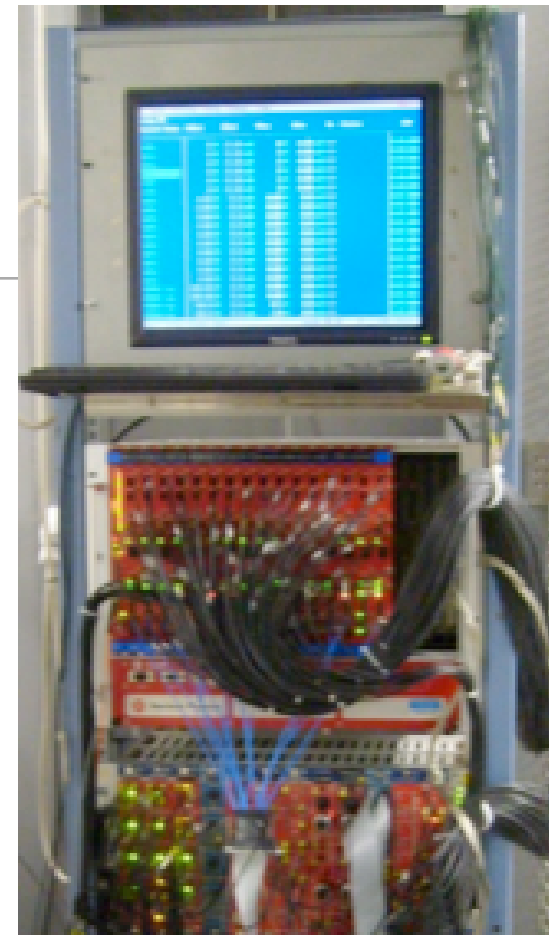


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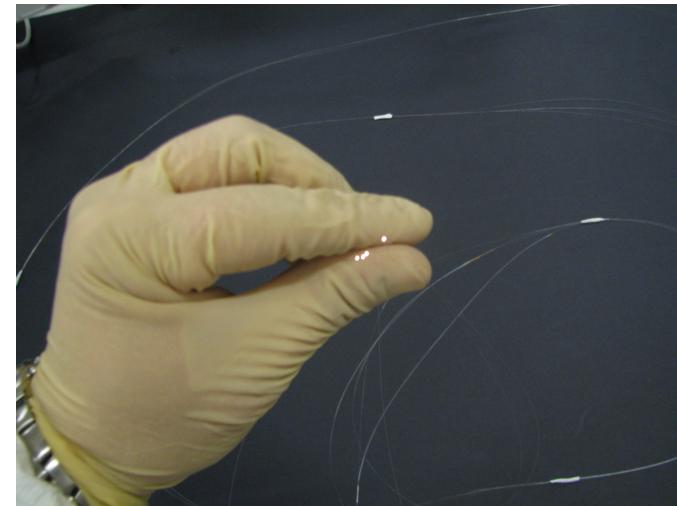
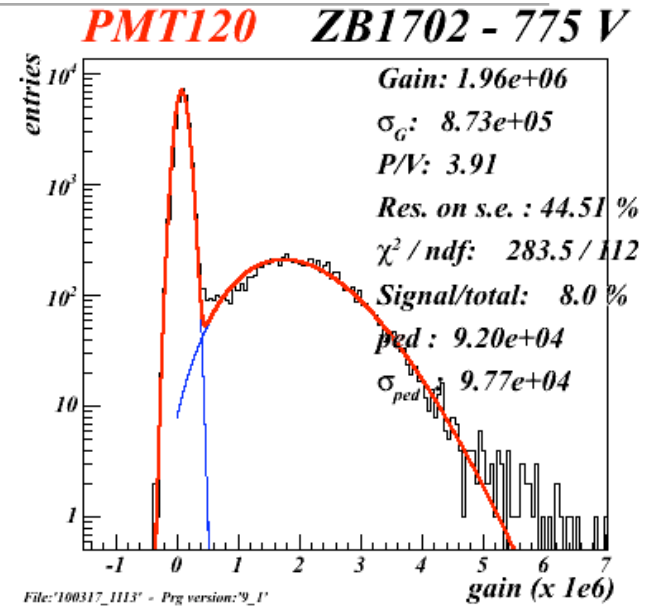
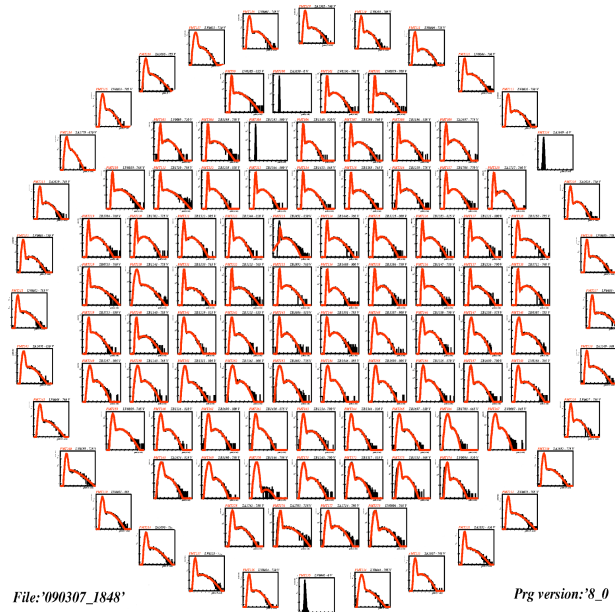
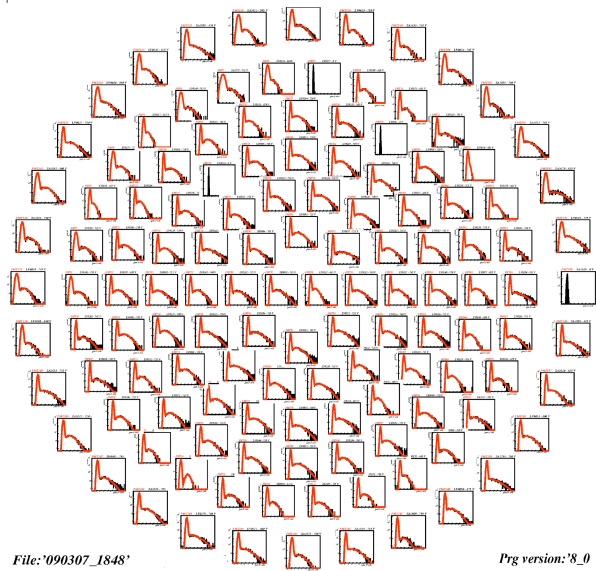
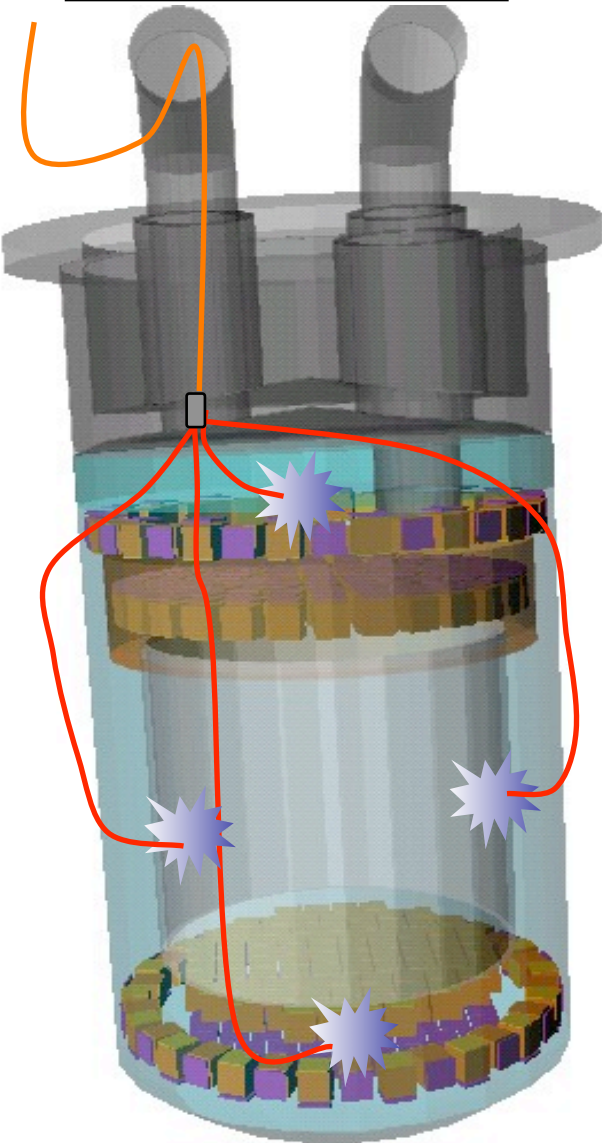


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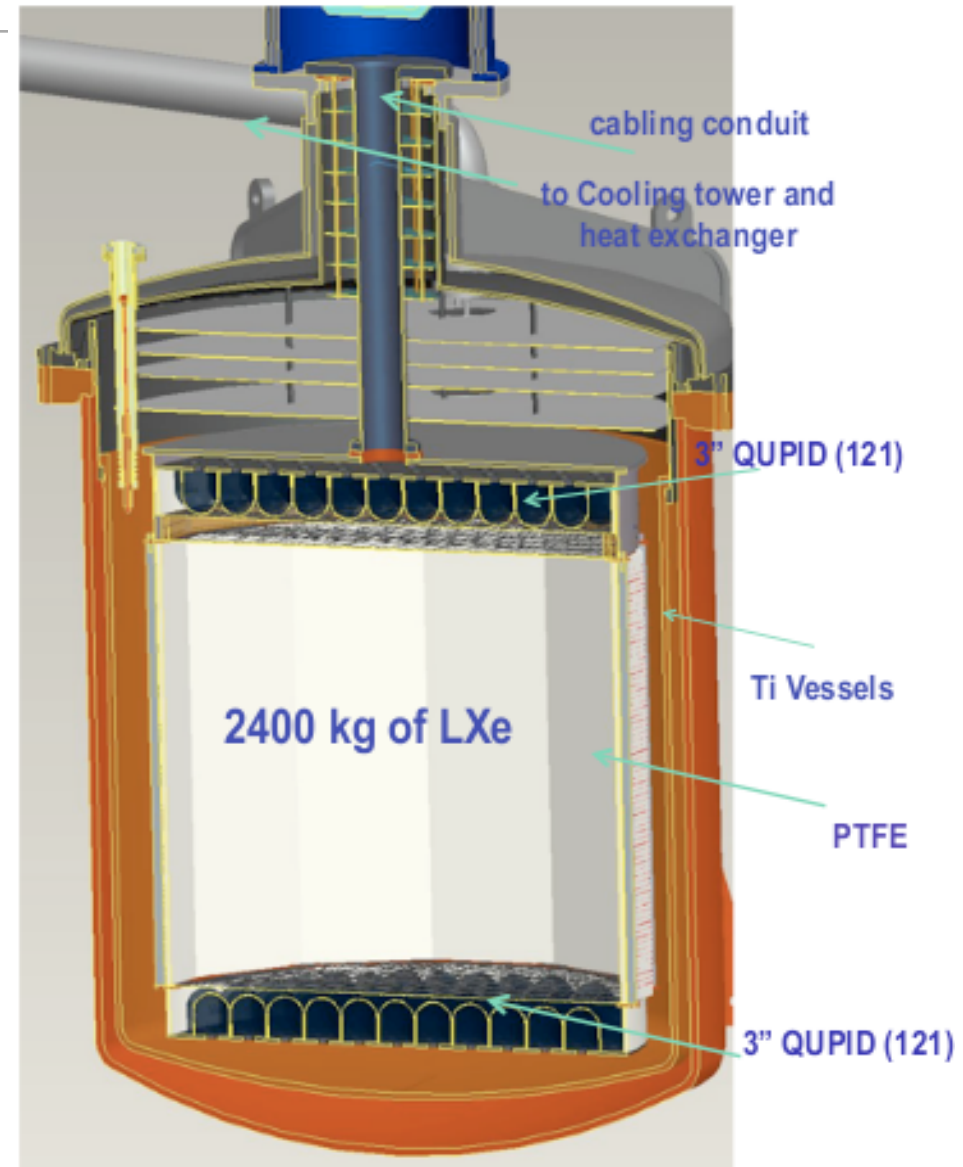
# Xenon100: PMT light calibration

4 optical fibers



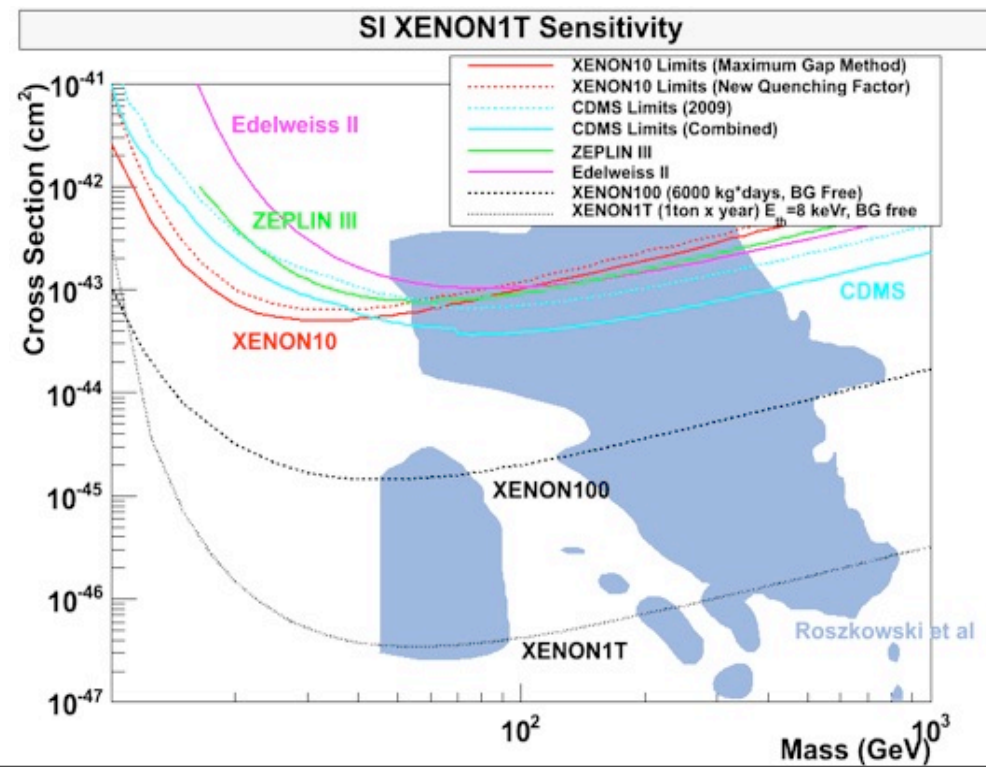
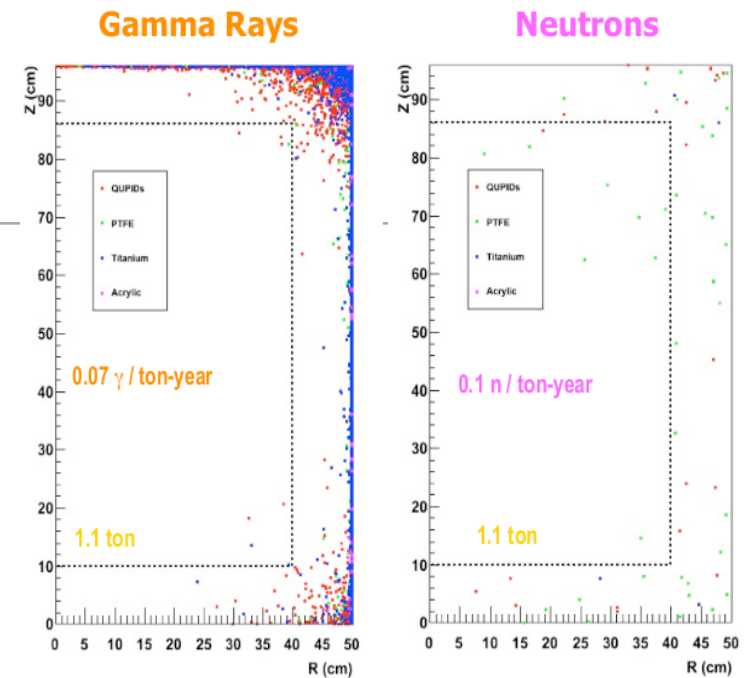
# XENON1T: Detector design

- ➔ Baseline design similar to XENON100 with improvements in different areas
  - lower radioactivity cryostat (Ti and Cu)
  - lower radioactivity PMTs (QUPIDs)
  - high efficiency heat exchanger
  - filling & recovery in liquid phase
- ➔ Design has been validated with detailed MC studies of internal/external background sources
- ➔ Capital cost ~ 8M\$ shared equally between US and foreign groups

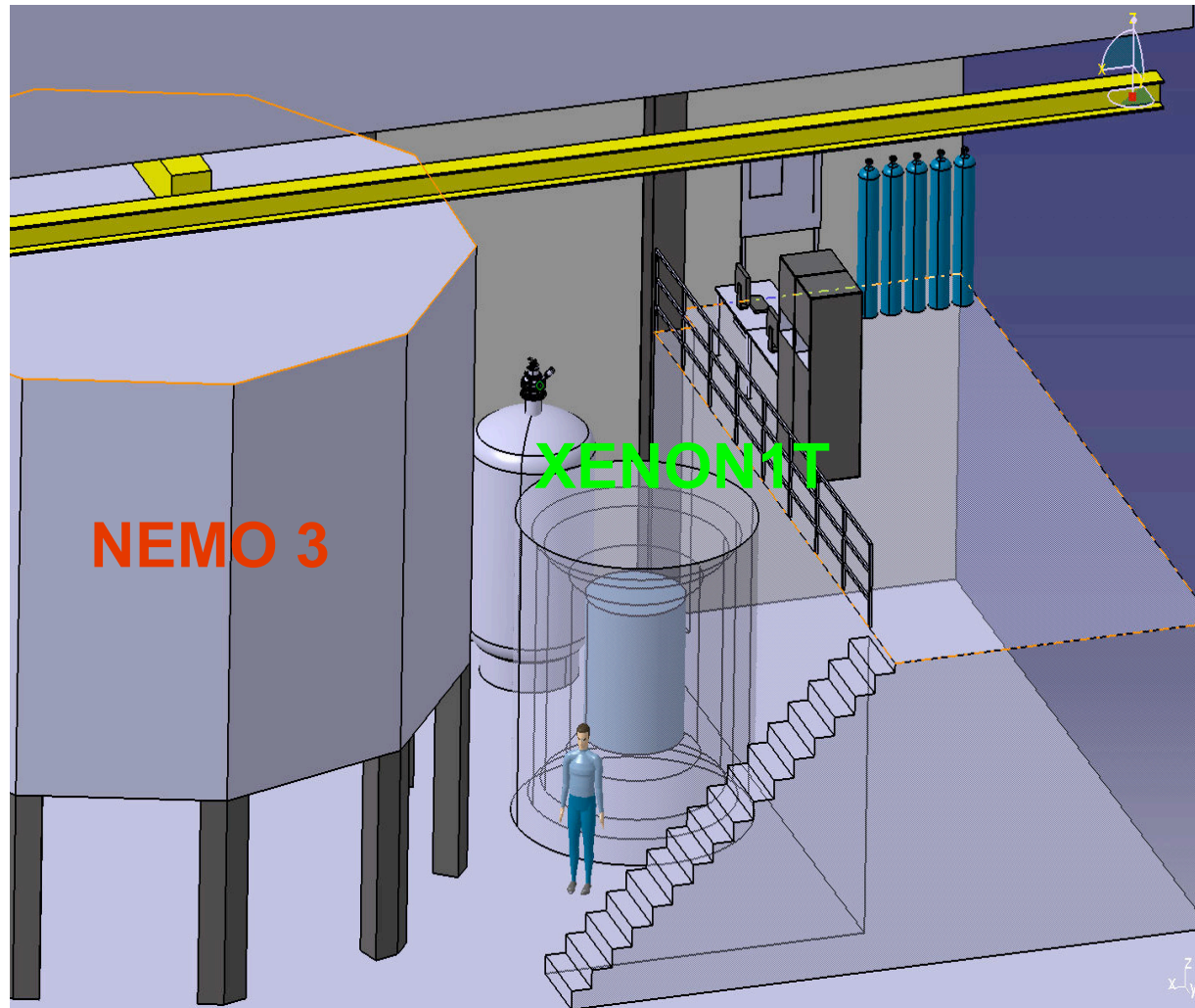


# XENON1T: Scientific goal

- ➔ The detector will have a fiducial mass of  $\sim 1$  ton of LXe
- ➔ QUPID sensors will measure the light from the interactions
- ➔ Simulations of the radioactivity from the material components show a background of less than 1 event/ton·year
- ➔ Extensive simulations in the proposed sites and with the proposed shield configurations are being carried out to show a similar level from external components
- ➔ After one year of background free measurement, the sensitivity will be  $\sim 5 \cdot 10^{-47} \text{cm}^2$ , covering most of the CMSSM predicted region for SI interactions

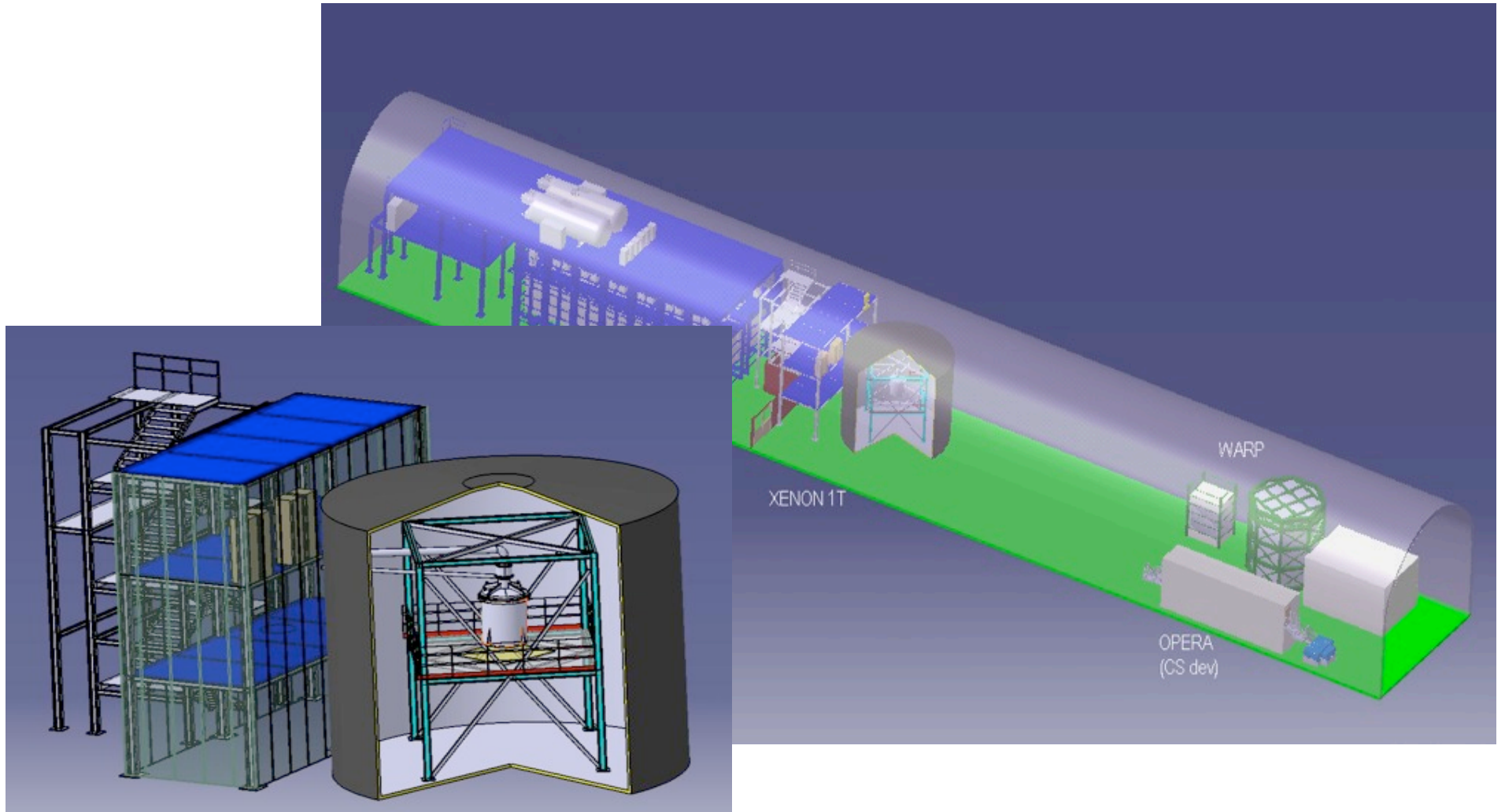


# XENON1T where? @ LSM



**Solid shield (55 cm Poly, 20 cm Pb, 15 cm Poly, 2 cm ancient Pb) plus >99 % muon veto**

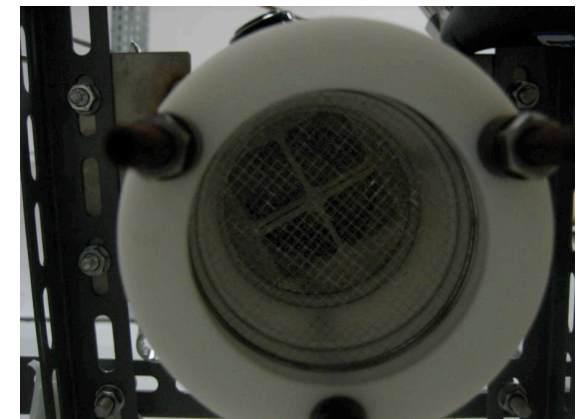
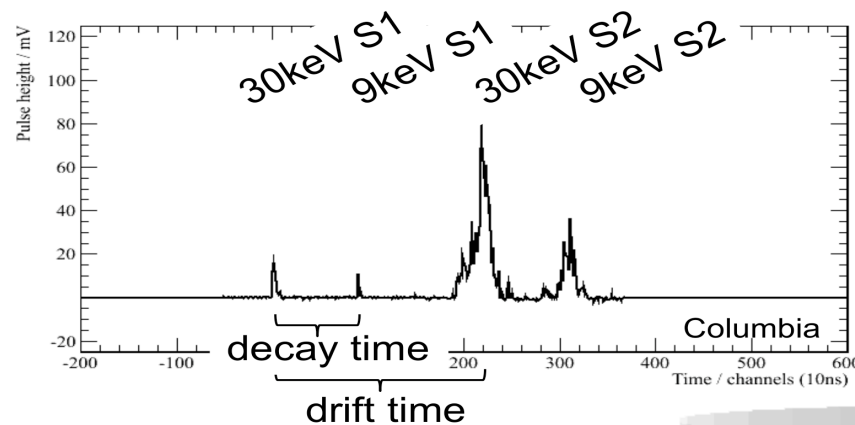
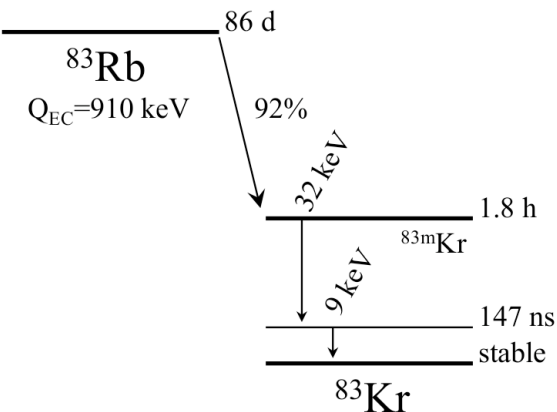
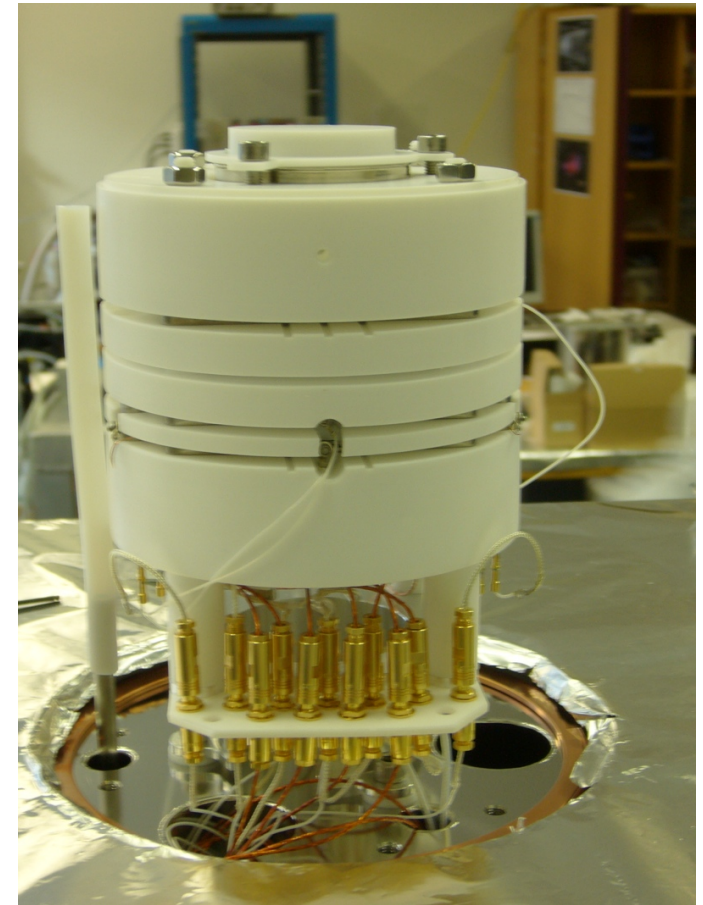
# XENON1T where? @ LNGS



**5 m-thick water shield**

# Xenon100: calibration

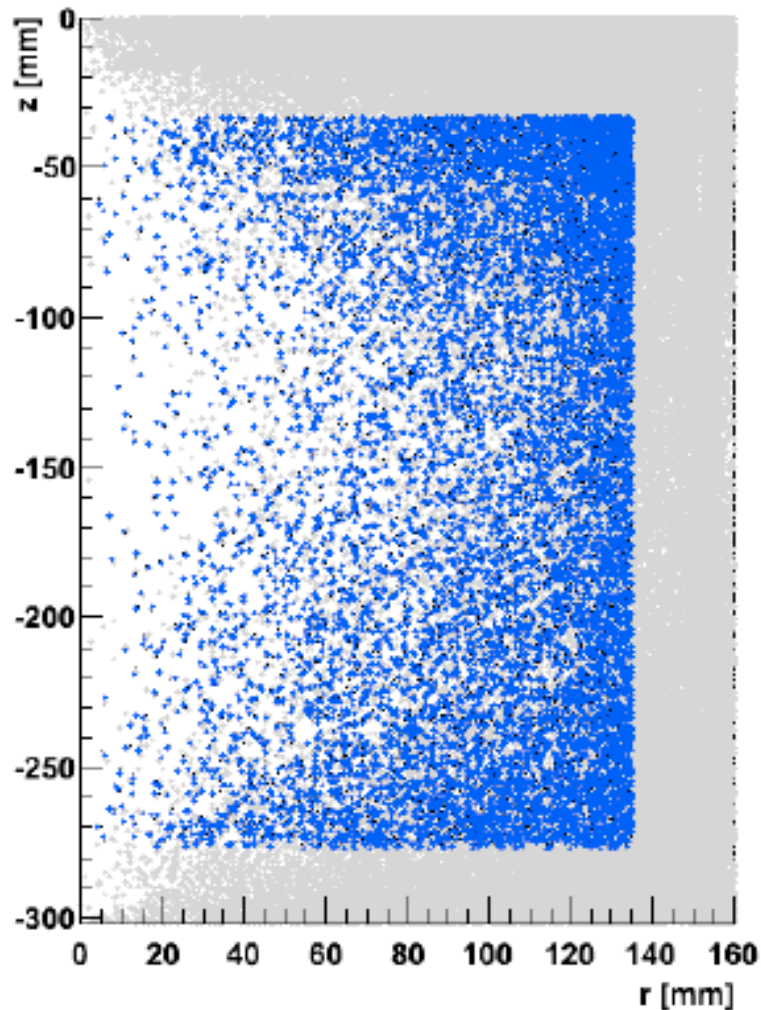
- $^{83}\text{Kr}$  is an ideal candidate for homogeneous calibration of the detector:
  - Not electronegative: no effect for electron attachment
  - Fast decay time  $\sim 2\text{h}$
  - Provides 2 lines at low energies (32keV and 9keV) with a 147ns delay
- Principle demonstrated in two small setups at Zurich and Columbia
- Extensive R&D already done
- A calibration with  $^{83\text{m}}\text{Kr}$  is planned



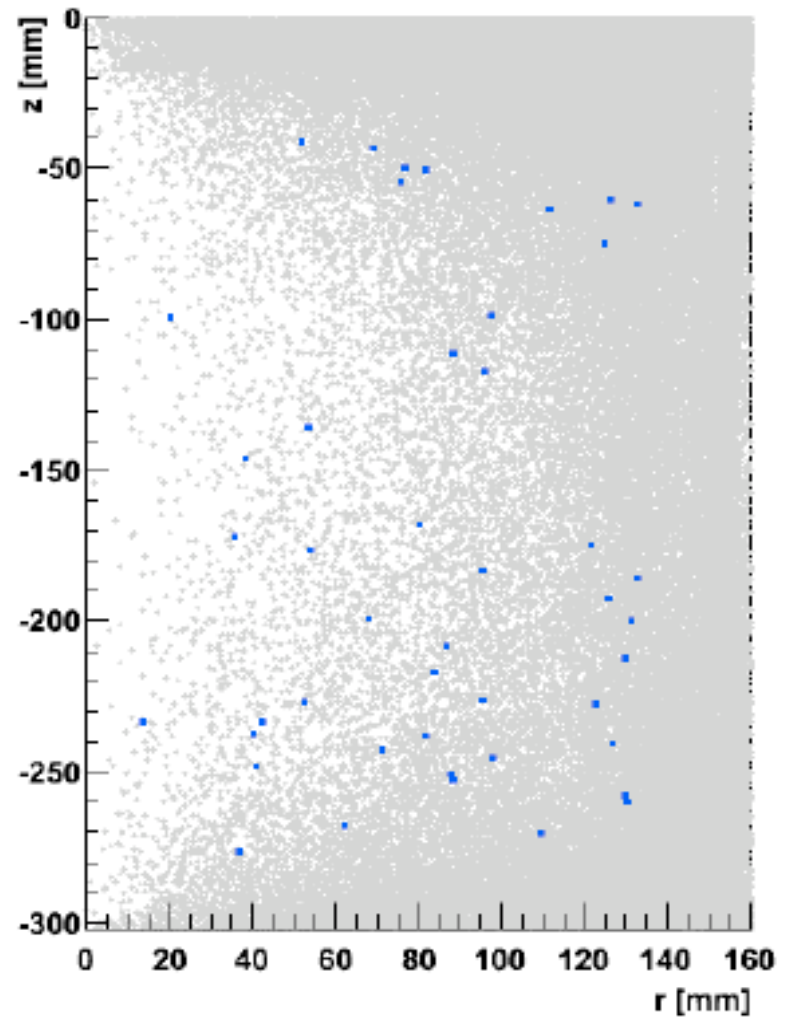
# Background analysis

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Selection of a 40kg  
cylindrical fiducial volume

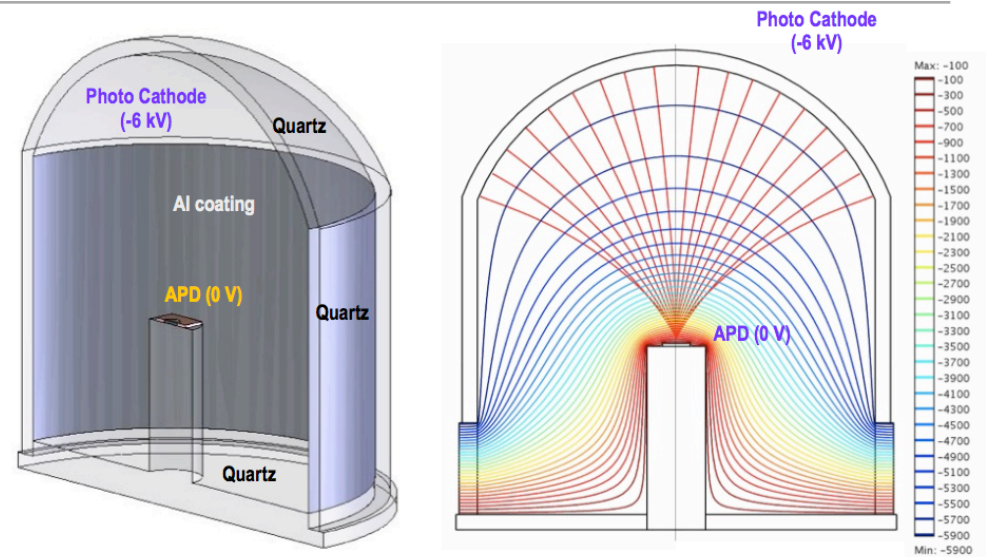


Energy range selection  
 $< 28 \text{ keV}_r$





# XENON1T: QUPIDs (QUartz Photons Intensifying Detector)

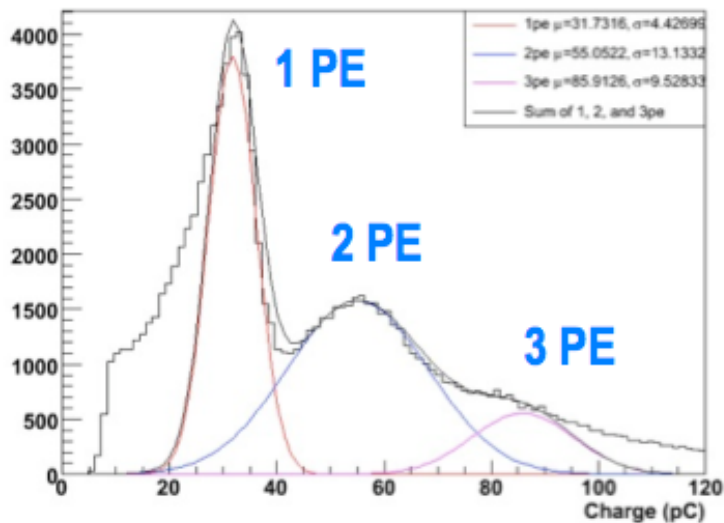


New concept of Light sensors

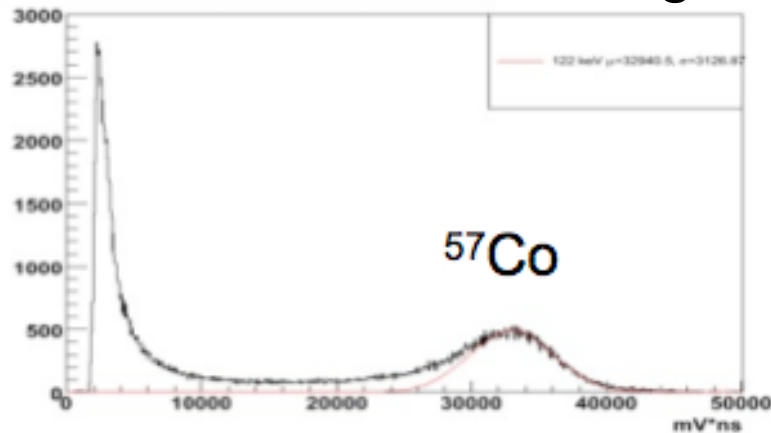
Very low radioactivity ( $<0.1$  mBq  $^{238}\text{U}/^{232}\text{Th}$ )

High QE photocathode

1, 2, and 3 Photoelectron Peaks

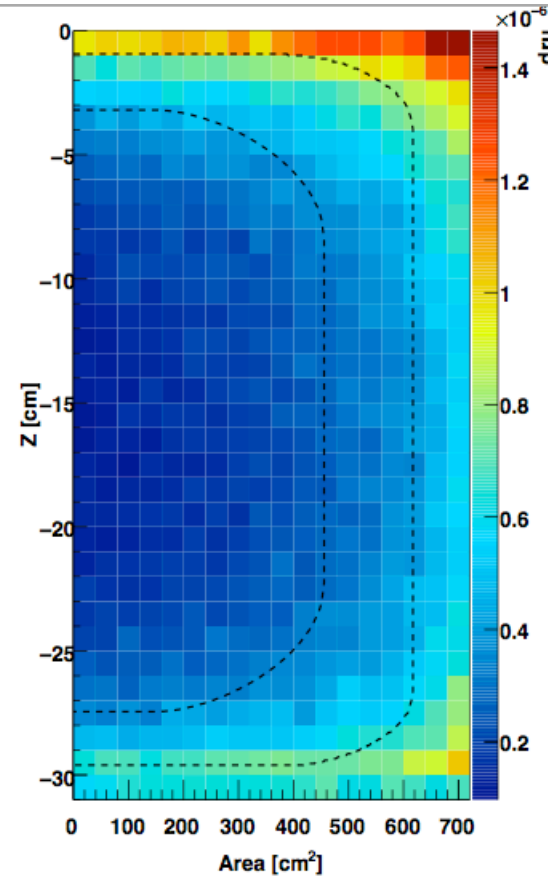
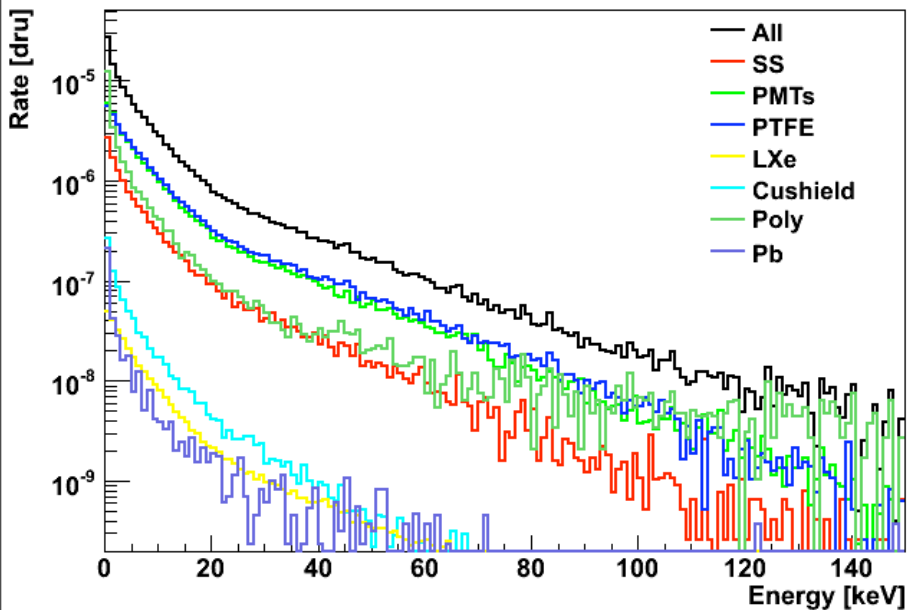


Spectrum 4.5kV, 305V Bias, 122keV

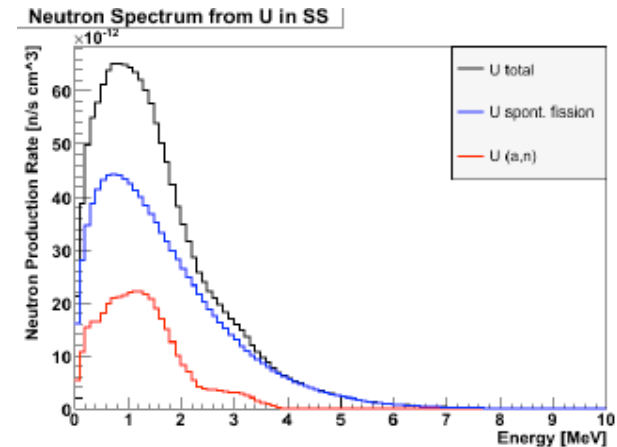


# Neutron background

Single nuclear recoils in the whole active volume from materials



SOURCES4A code



Total single nuclear recoil rate  $[5,27] \text{ keV}_r$   
 (including rock and muons)  
 1.62 n/year (50 kg)      0.60 n/year (30 kg)