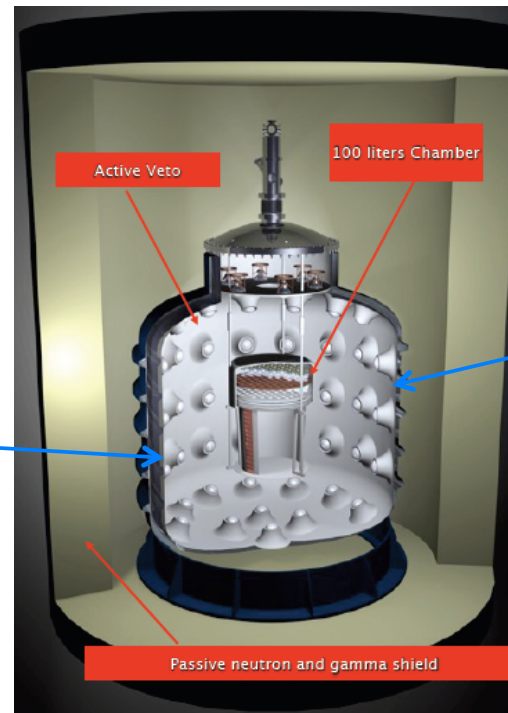


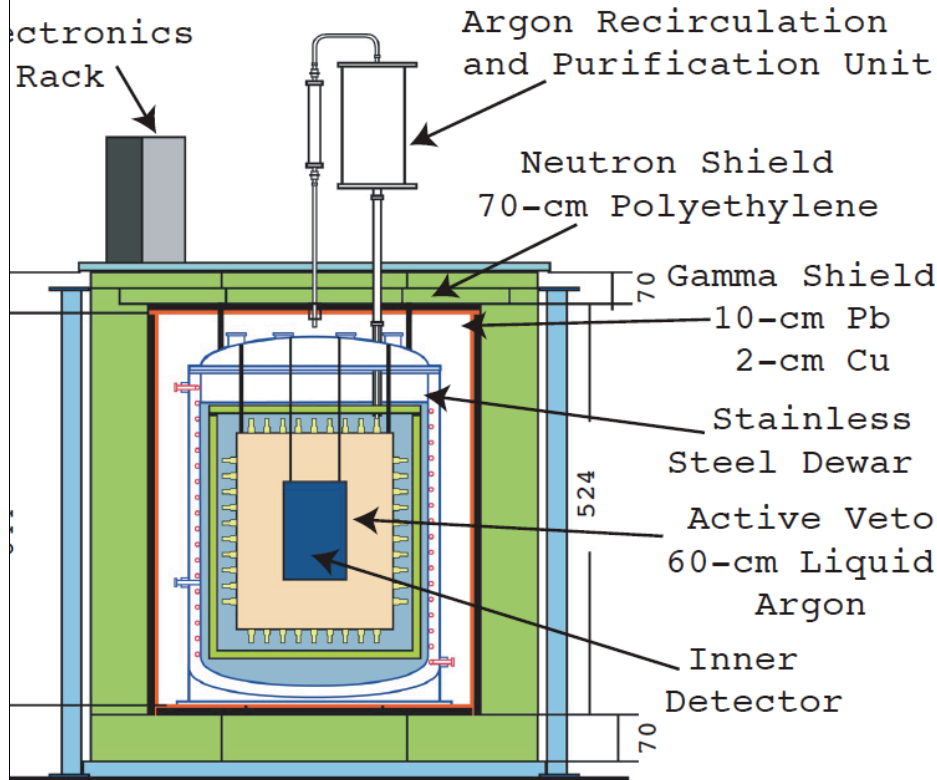
# The WArP Experiment @ LNGS

- The WArP detector is under commissioning at LNGS
- 140 kg active target, with active LAr shield (9 tons and 300 PMTs)
- 3D Event localization and definition of fiducial volume for surface background rejection
- Cryostat designed to allocate a possible 1400 kg detector
- Ar gas with **low level of radioactive  $^{39}\text{Ar}$**  from US u.g. reservoir promising for multi-ton expt

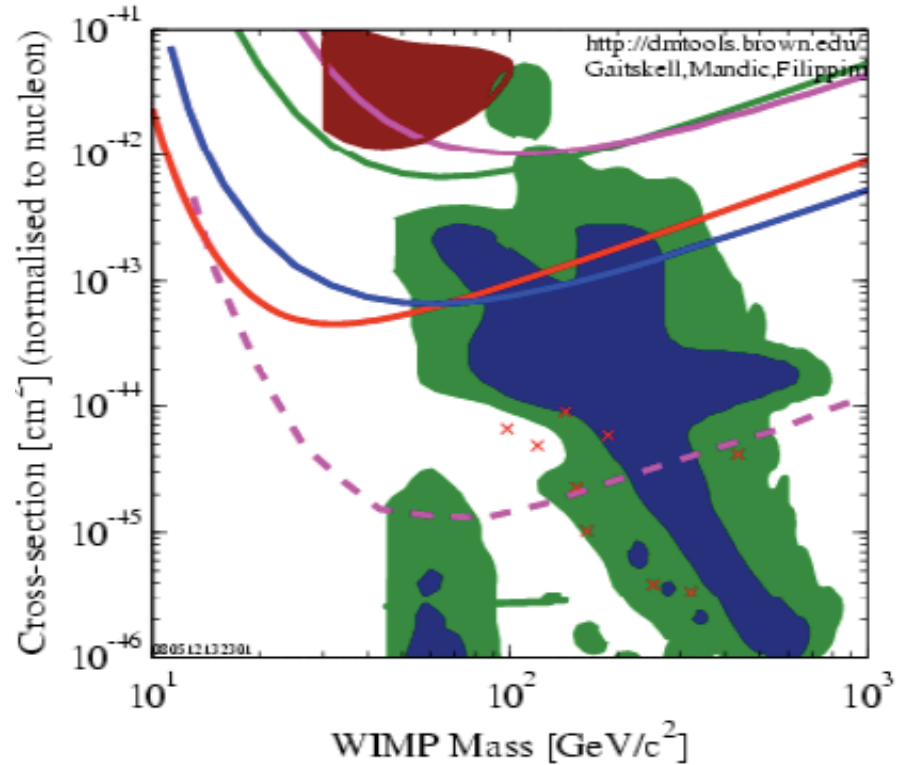
## Main Cryostat

after positioning in Hall-B (Jul.'07) inside sustaining structure





- We expect to complete the detector assembly and commissioning and to be able to run with the complete detector within the end of the year.



	<p>DATA listed top to bottom on plot</p> <p>DAMA 2000 58k kg-days NaI Ann. Mod. 3sigma w/DAMA 1996</p> <p>WARP 2.3L, 96.5 kg-days 55 keV threshold</p> <p>ZEPPLIN II (Jan 2007) result</p> <p>CDMS 2008 Ge</p> <p>XENON10 2007 (Net 136 kg-d)</p> <p>WARP 140kg (proj)</p> <p>Roszkowski/Ruiz de Austri/Trotta 2007, CMSSM Markov Chain Monte Carlos (1)</p> <p>Roszkowski/Ruiz de Austri/Trotta 2007, CMSSM Markov Chain Monte Carlos (1)</p> <p>Ellis et. al Theory region post-LEP benchmark points</p> <p>08051213230L</p>
--	--

WARP Update

Cryostat f

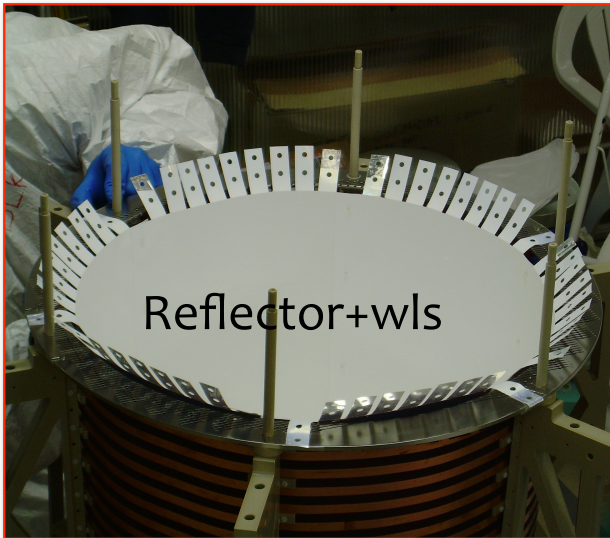
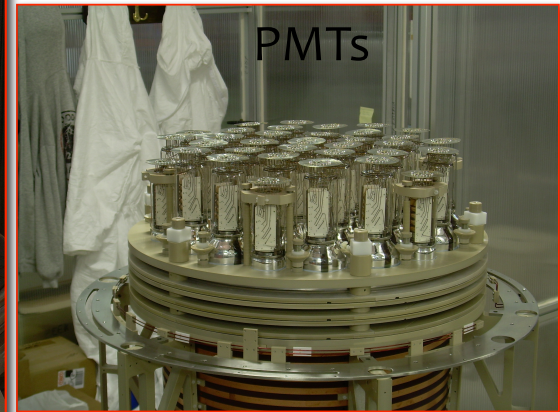
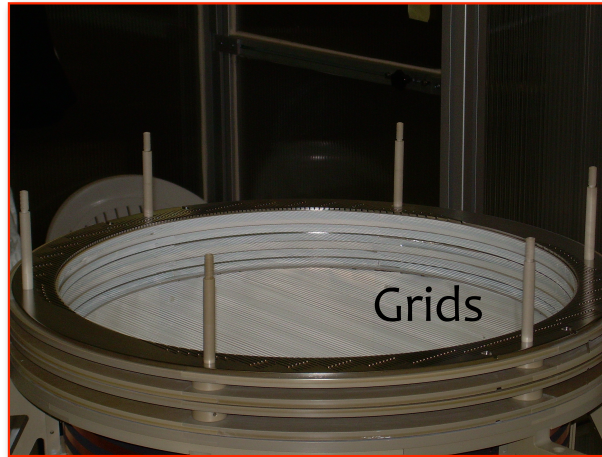
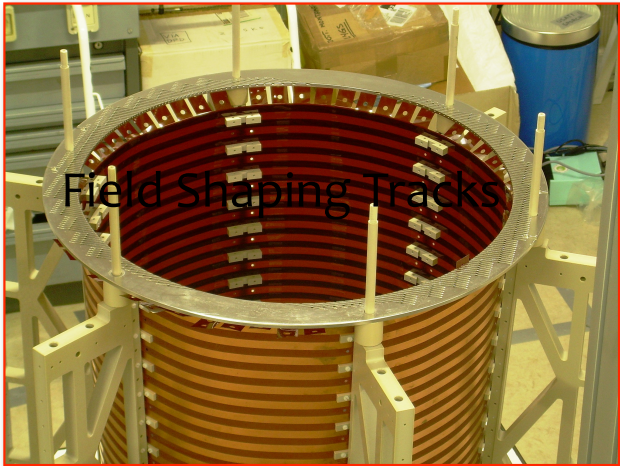
140-k

U

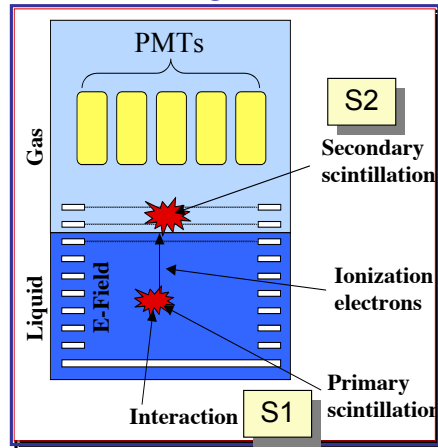
December 2008



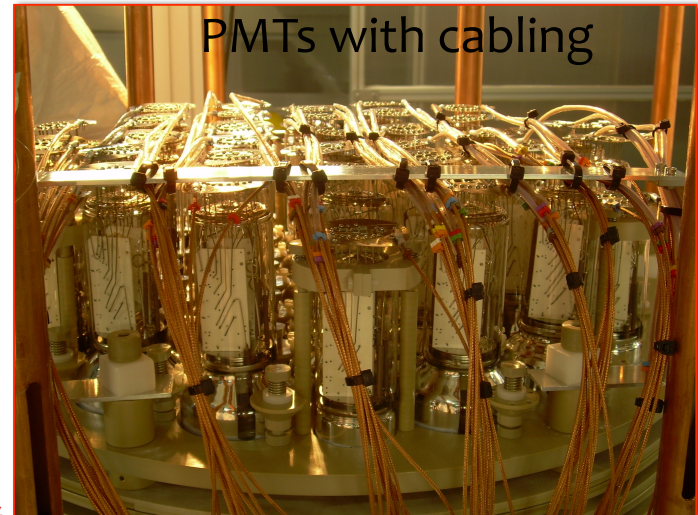
# WArP Inner Detector Components



Two-Phase Argon Drift Chamber

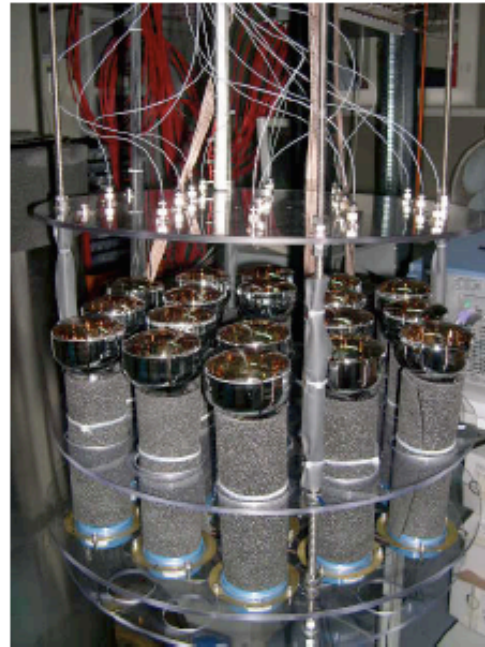
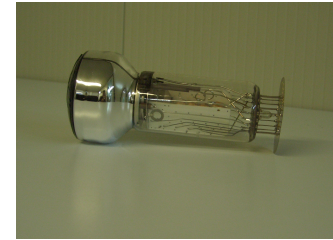


Recoil-like event       $\gamma$ -like event  
 Slow component  $\approx$  10 %      Slow component  $\approx$  70 %

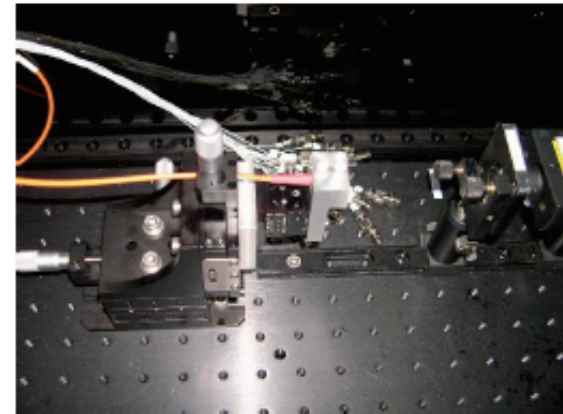


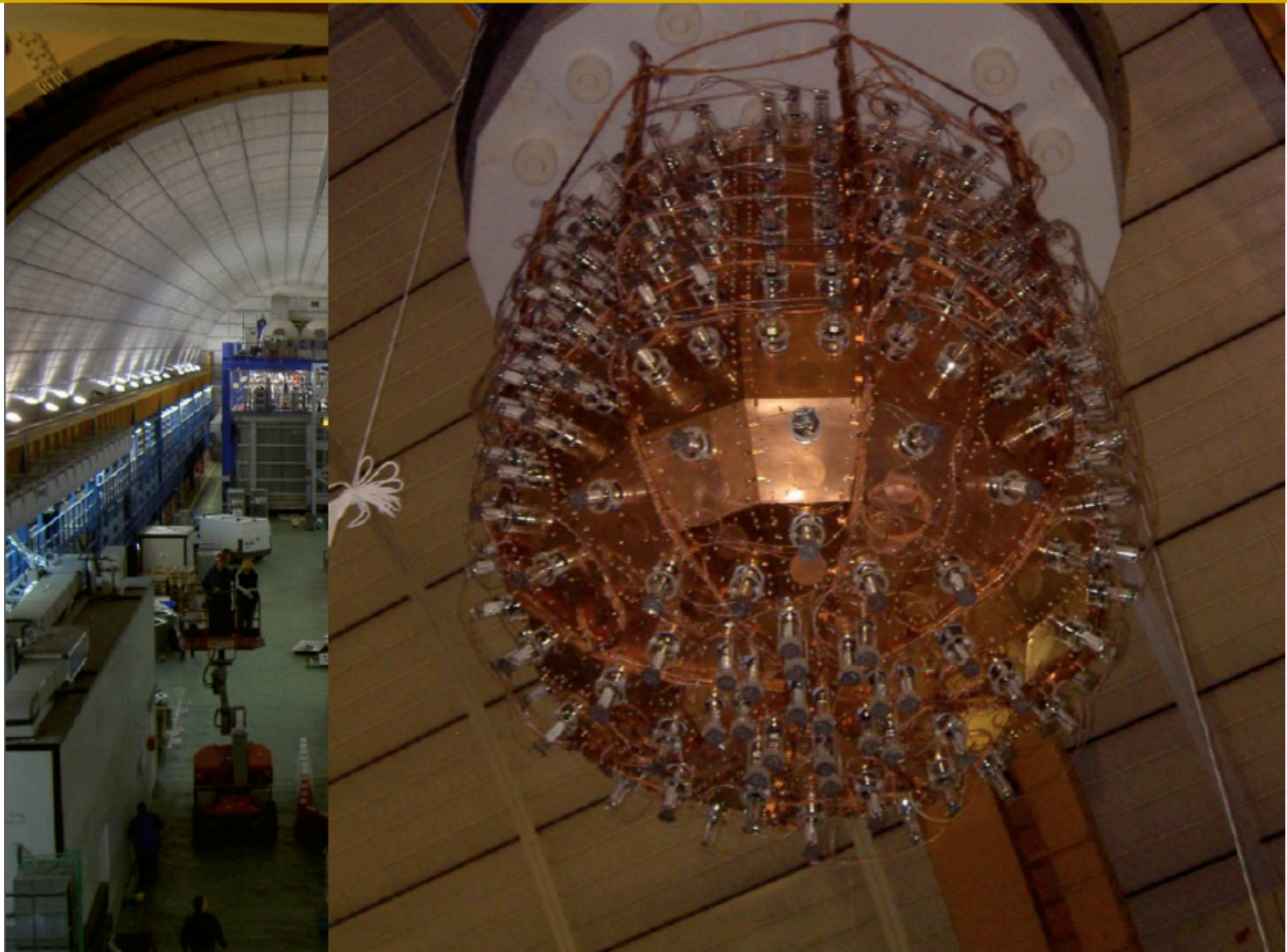
# Photomultipliers

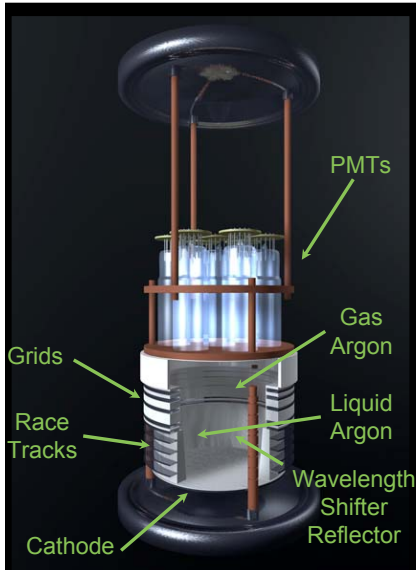
- 3" and 2" Bialkali Photomultipliers developed in co-operation with **Electron Tubes EMI** to work at **LAr temperature (ETL D750UKFLA, D757UKFLA)**
- **7% coverage** in the active veto
- **10% coverage** in the inner detector
- Low activity (0.2 Bq/PMT) and high QE (19% on average)



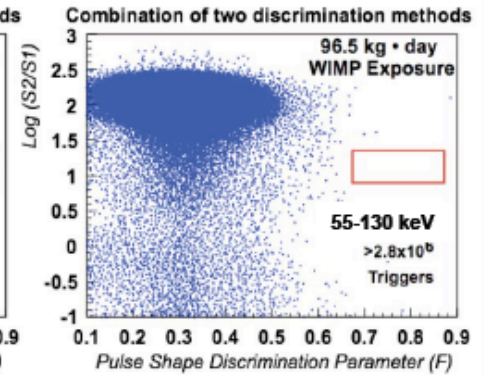
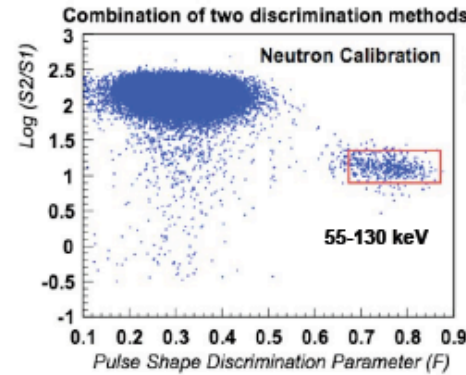
More than 400 PMTs verified to work at cryogenic temperature (77 K) in Napoli INFN laboratories and delivered to LNGS







Live-time = 52.8 days ( $\approx 96.5 \text{ kg} \cdot \text{day}$ )  
 Total triggers  $\approx 2.8 \times 10^7$   
 Trigger threshold  $\approx 3.5 \text{ keV}_{\text{rec}}$   
 Threshold for data selection =  $20 \text{ keV}_{\text{rec}}$   
 Threshold for WIMP analysis =  $40 \text{ keV}_{\text{rec}}$



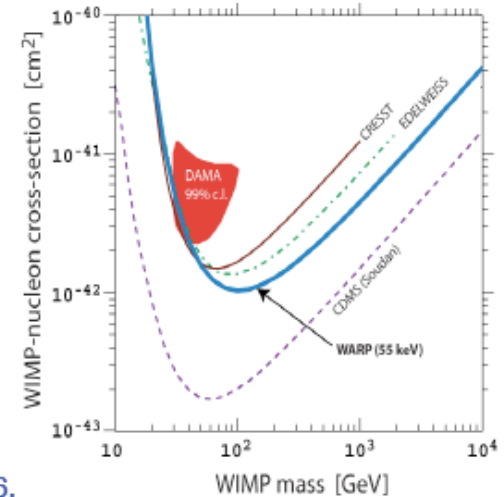
Selected events in the n-induced single-recoils window during the WIMP search run:

**5 events above 40 keV, 0 events above 55 keV  $\Rightarrow$**   
 $\approx 0.1 \text{ events / day}$  ( $0.05 \text{ events / (kg} \cdot \text{day)}$ )

The main beta background comes from the intrinsic  $^{39}\text{Ar}$ .

Excellent results from study of discrimination power between nuclear recoils and  $\gamma$ -betas:

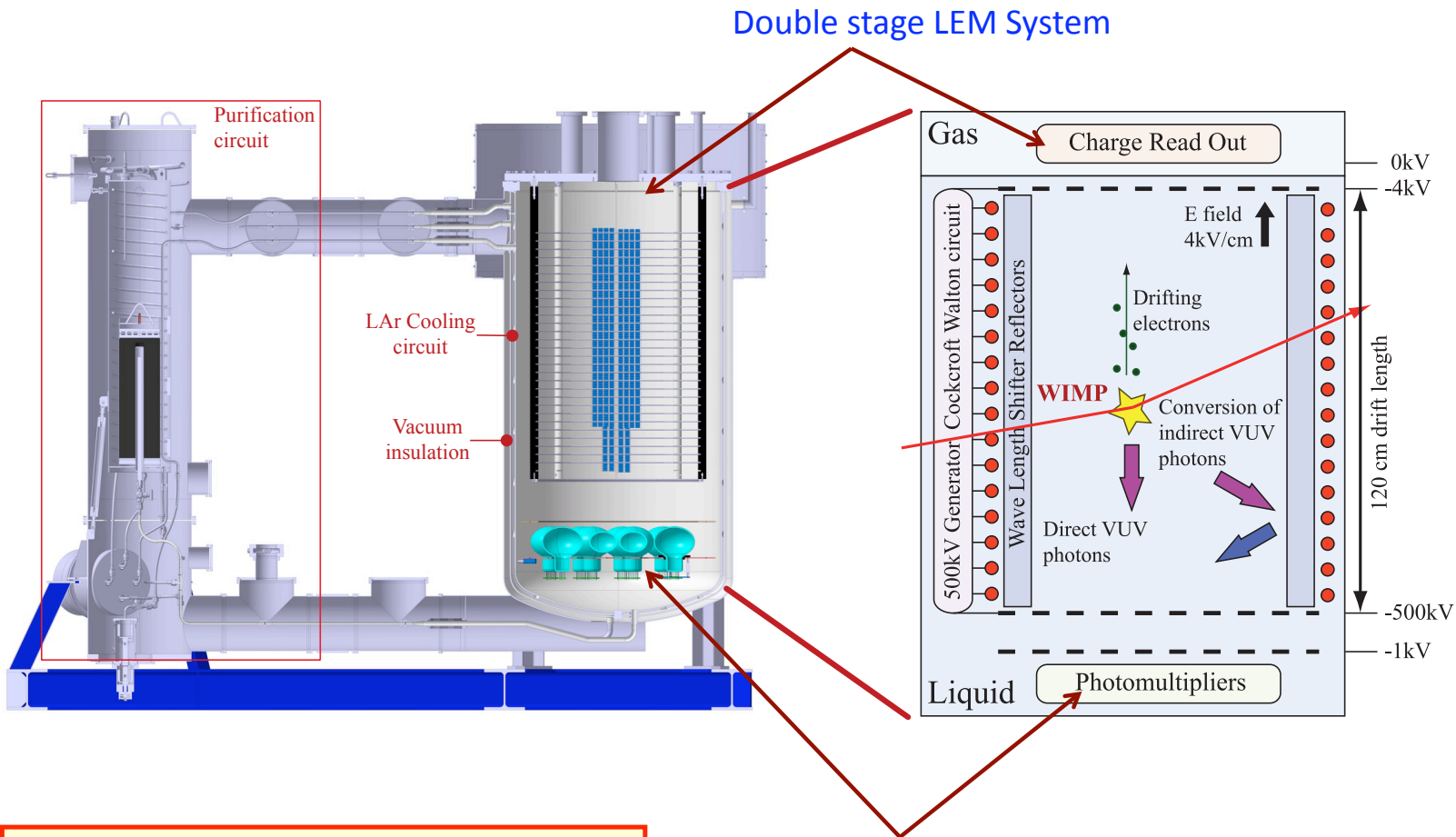
$10^{-8}$  pulse shape discrimination,  
 $5 \times 10^{-3}$  ionization/scintillation



P. Benetti et al., *Astrop. Phys.* 28 (2008) 6.

# The ArDM Experiment @ LSC

## General Layout of the Experiment



A. Rubbia, "ArDM: a Ton-scale liquid Argon experiment for direct detection of dark matter in the universe", *J. Phys. Conf. Ser.* 39 (2006)

129

14 Cryogenic PMTs to detect the scintillation light

P.Otyugova (UniZH)



## The ArDM main parameters

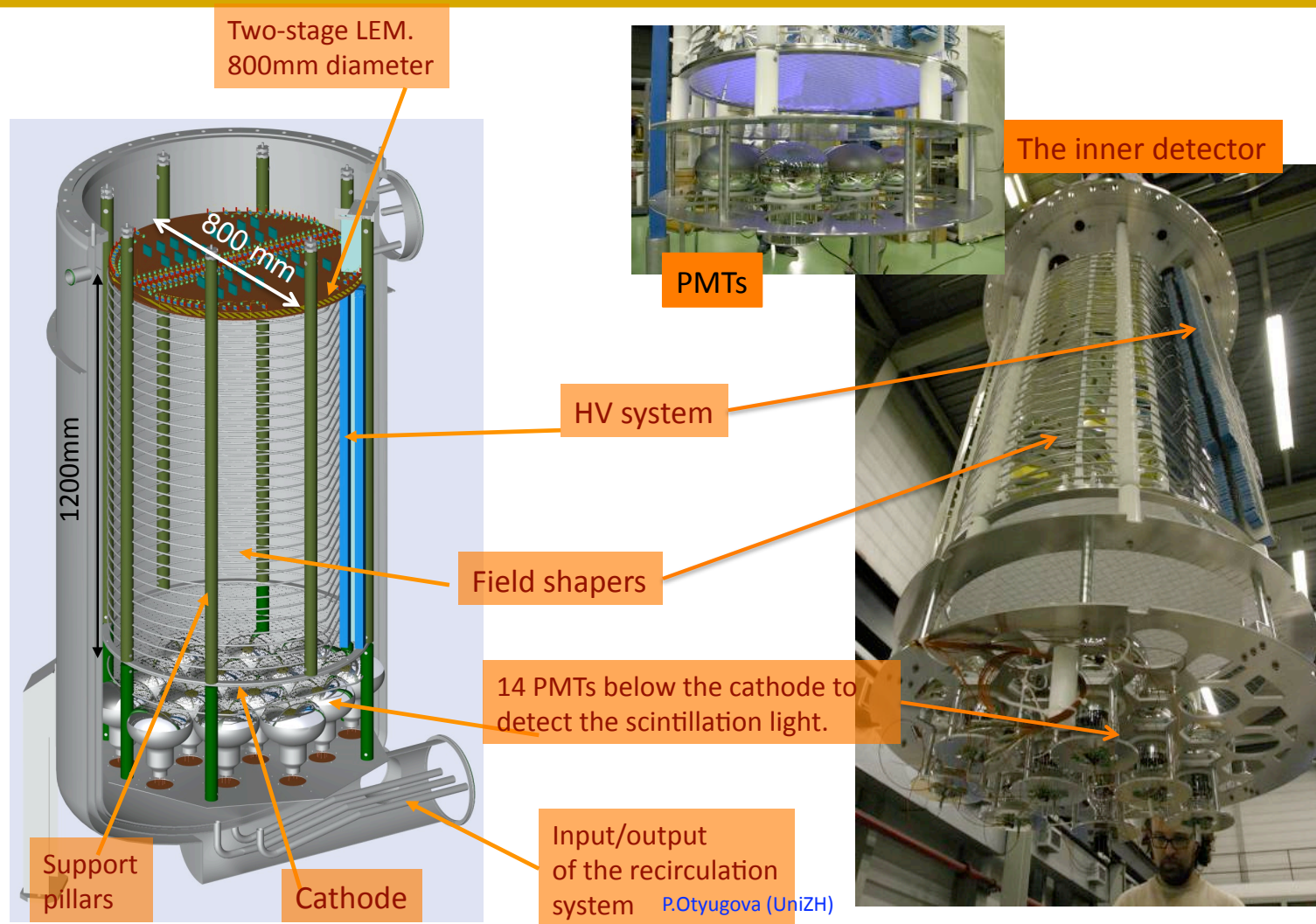
Detector	
Max. drift length	120 cm
Target mass	850 kg
Readout method	Independent readout of charge & light
Drift field	1÷4 kV/cm
Charge readout	
Charge gain	500÷10 <sup>3</sup> per e <sup>-</sup>
Light readout	
Global light collection efficiency	1÷2%

Background rejection is based on:

1. Light pulse shape discrimination. Different light structure for WIMP-like (nuclear recoil) events and e/ $\gamma$ -like events.
2. Different Light/Charge ratios for WIMP and e/ $\gamma$ -like events

P.Otyugova (UniZH)

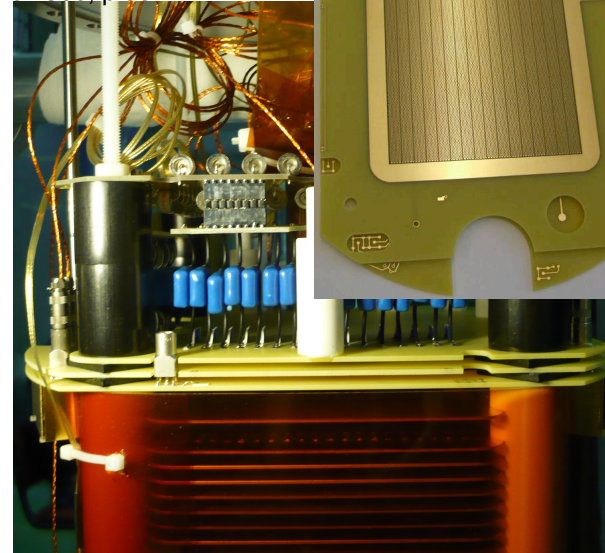
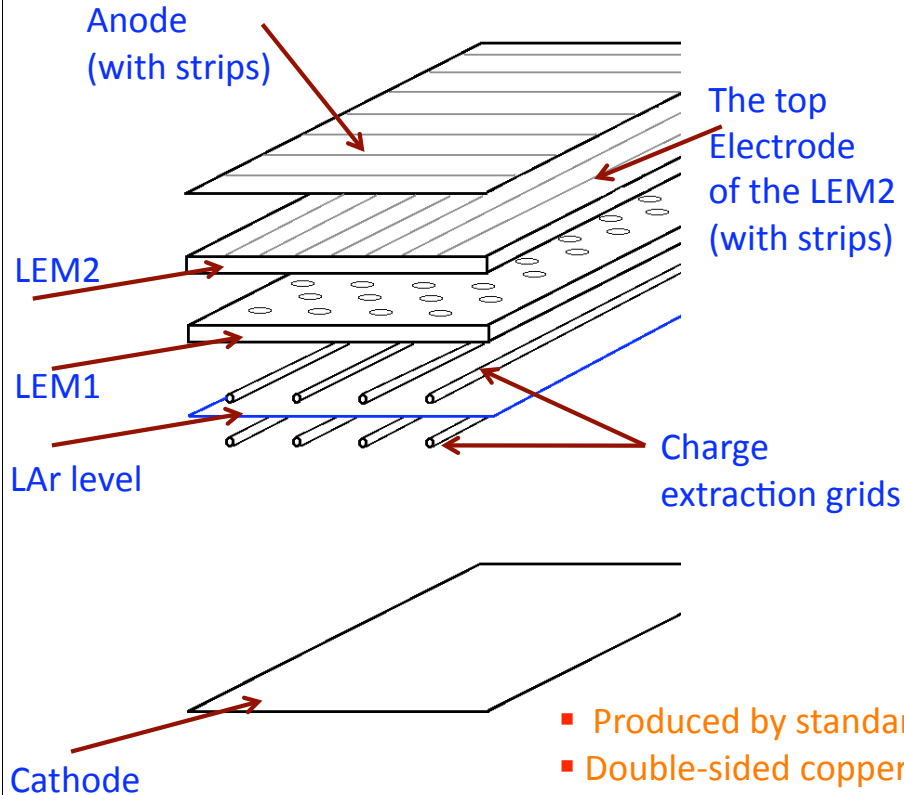
## The ArDM design and assembly



## Charge Read-Out System: Large Electron Multiplier (LEM)

LEM is a thick macroscopic GEM Ref. F.Sauli, NIM A, 1997, vol. 386, p.351

LAr LEM-TPC principal



**10 x 10 cm<sup>2</sup> prototype chamber**  
**Two planes 16 strips, 6 mm width**

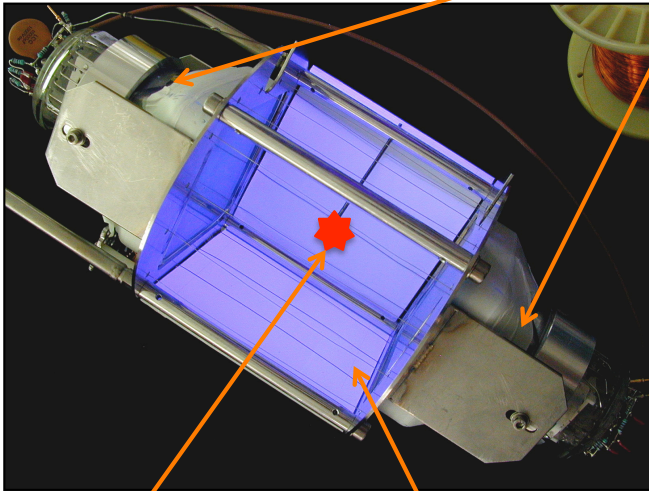
F.Resnati  
Proceeding for IEEE-NSS 2008

- Produced by standard PCB technique
- Double-sided copper-clad (18  $\mu\text{m}$  layer) FR4 plates
- Precision holes made by drilling
- Gold deposition on Cu ( $< \sim 1 \mu\text{m}$  layer) to avoid oxidization

P.Otyugova (UniZH)

## Dual- PMT cell for the pulse-shape events separation

2 Cryogenic PMTs.

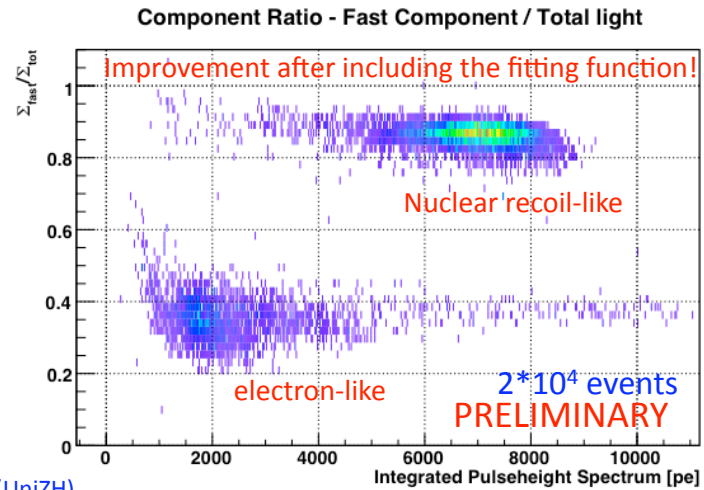
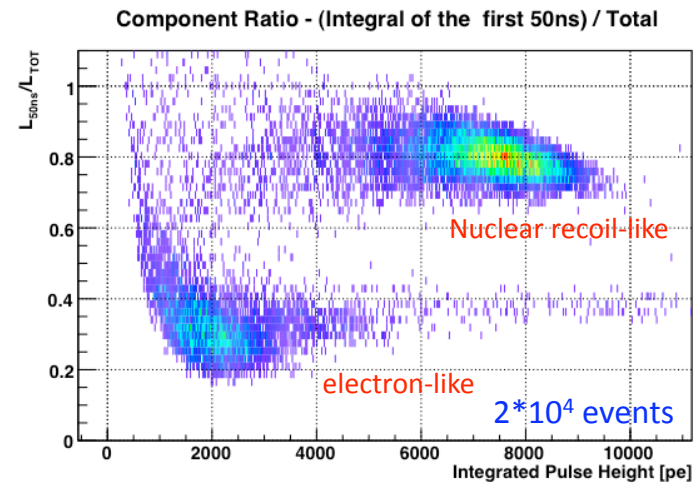


Pb<sup>210</sup> source  
50 Bq.

TPB coated reflectors.

Inner volume: 6 l of LAr  
Radioactive source: <sup>210</sup>Pb,  
 $\alpha$  5.3 MeV,  $\beta$  1.16 MeV  
 $\alpha$  and  $\beta$  events are clearly  
separated.

P.Otyugova (UniZH)



## Preliminary Light Yield Prediction

	Gaseous Argon	Liquid Argon
6lt.Test Cell (only 1 PMT)	~800phe/5.3MeV $\alpha$ ~0.150phe/keVne (measured)	~4500phe/5.3MeV $\alpha$ ~0.8phe/keVne (measured)
1ton ArDM	~550phe/4.4MeV $\alpha$ ~0.125phe/keVne (measured)	~3000phe/4.4MeV $\alpha$ ~0.7phe/keVne (expected)

A quenching factor ( $\sim 0.3$ ) has to be taken in account for hadrons, because of high ionization density

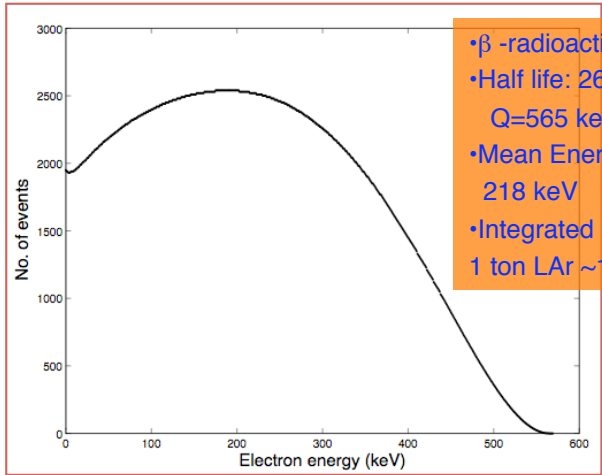
Electrons and gamma events have almost no quenching, so the light yield will be 3.3 times higher, i.e. we should reach a light yield of 2phe/keVee (electron equivalent)

The expected light yield fulfills the ArDM requirements!

P.Otyugova (UniZH)

# Ar<sup>39</sup> and neutrons backgrounds

Natural argon from liquefaction of air contains small fractions of <sup>39</sup>Ar radioactive isotope.



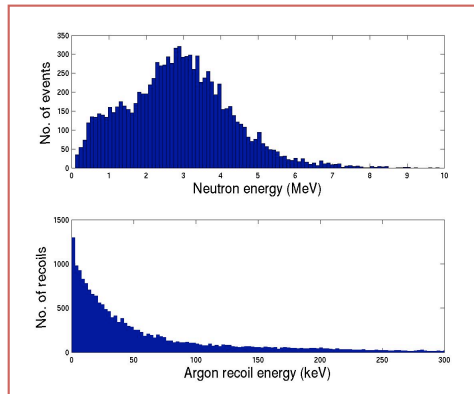
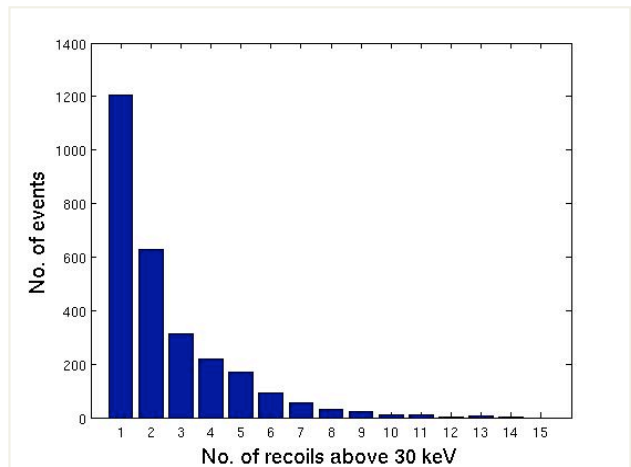
- β -radioactive isotope
- Half life: 269 years
- Q=565 keV
- Mean Energy: 218 keV
- Integrated rate in 1 ton LAr ~1kHz

Component	n per year	WIMP-like recoils per year
Container	~ 400	~ 30
LEM (std. mat.)	~ 10000	~ 900
LEM (low bg. mat.)	< 20	< 2
14 PMTs (std. mat.)	~ 12000	~ 1000
14 PMTs (low bg. mat.)	~ 600	~ 50

About 55% of the interacting neutrons scatter more than once at the threshold of 30keV.

Less than 10% of the emitted neutrons produce WIMP-like events single recoils, energy  $\in [30,100]$  keV).

The WIMP cross-section is very low, and it will scatter at most once.

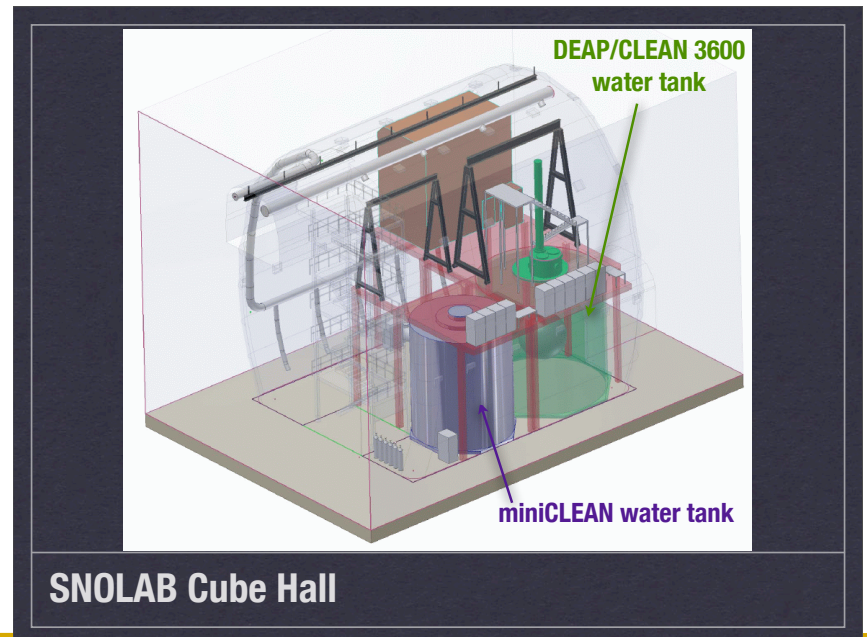
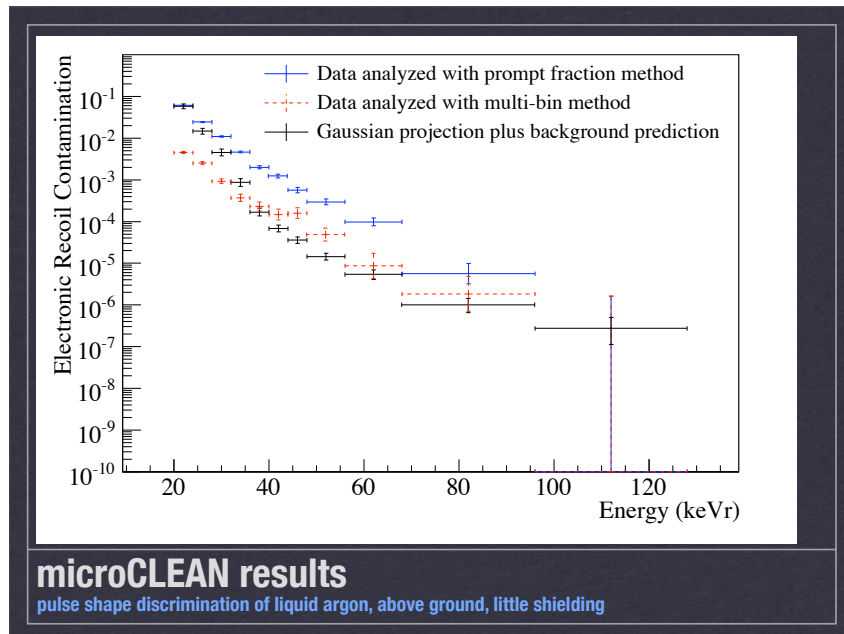
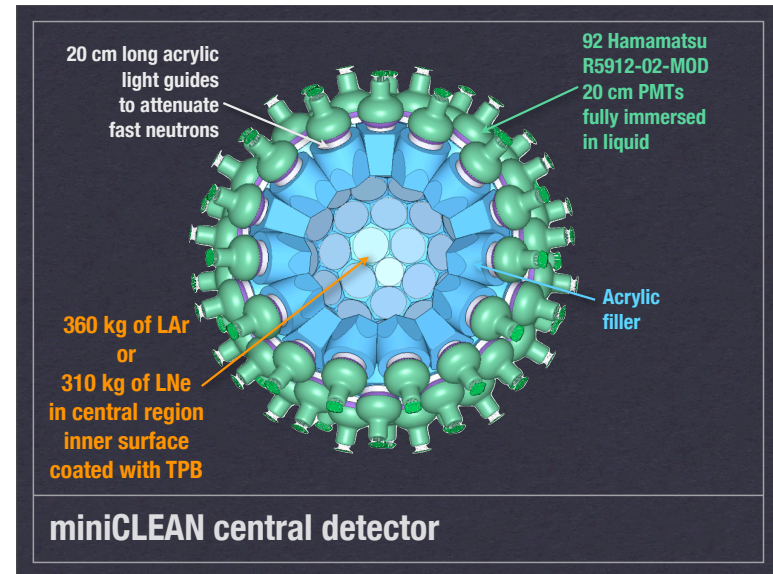


We need:  
Rejection power of  $10^8$   
OR  
use of <sup>39</sup>Ar-depleted argon

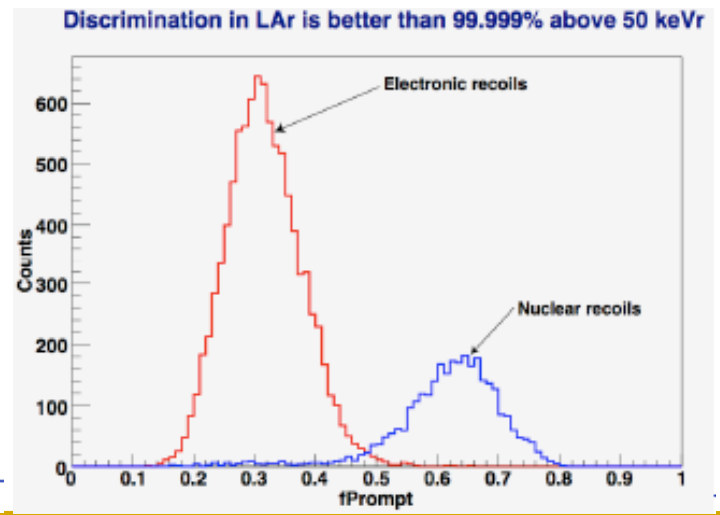
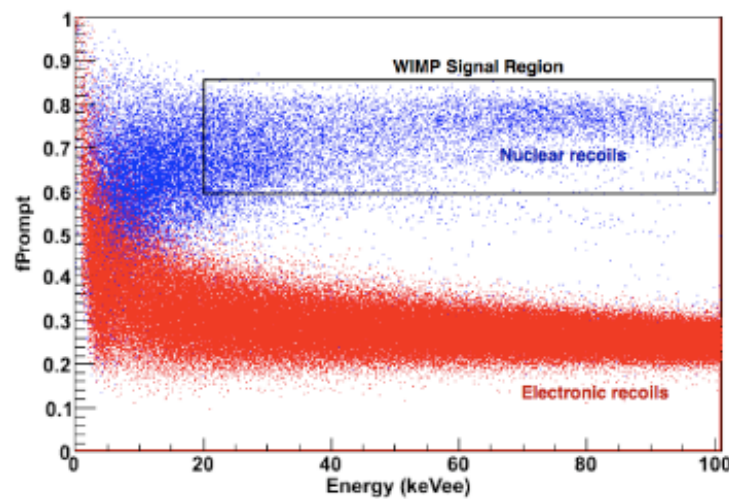
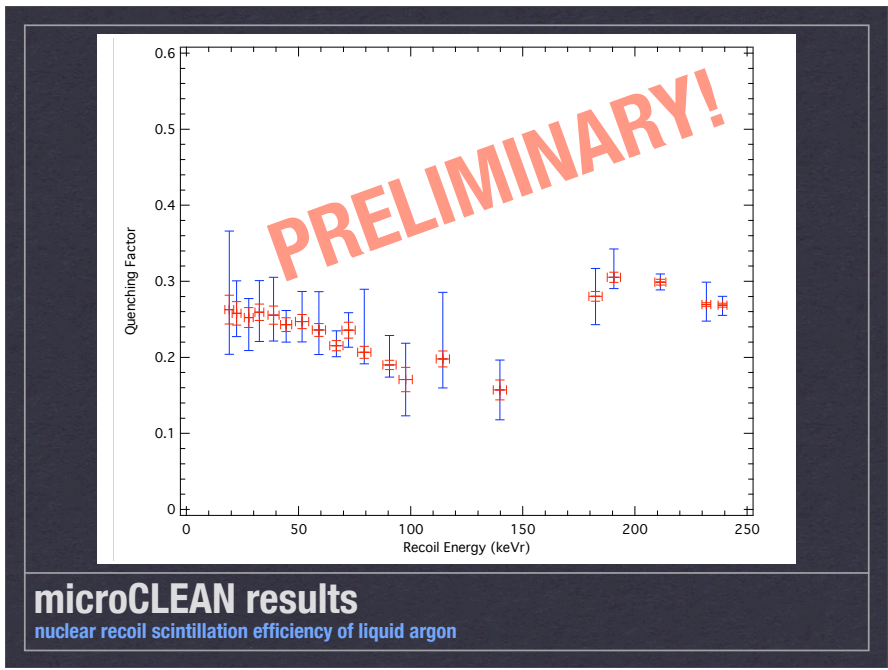
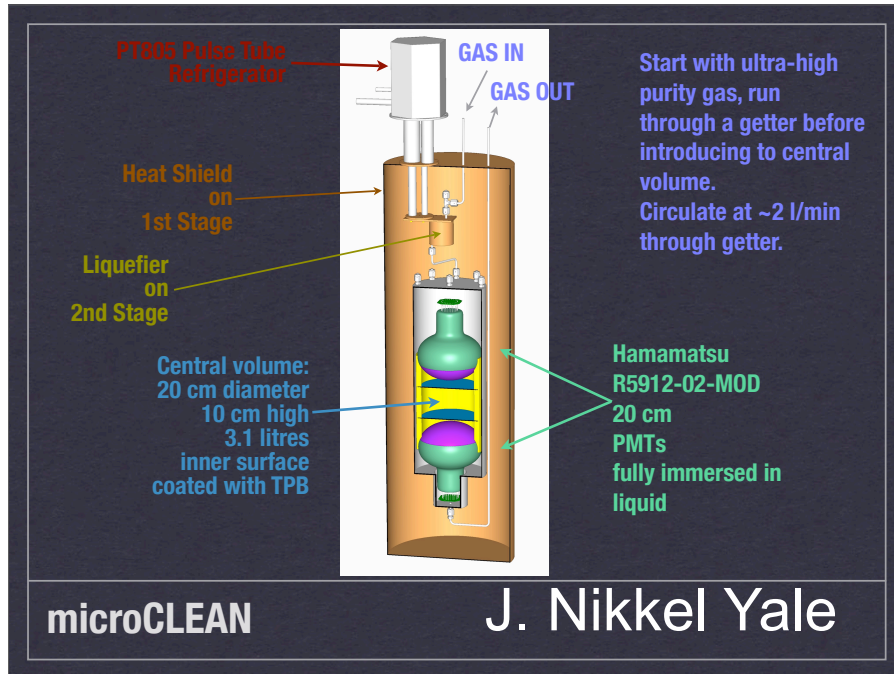
P.Otyugova (UniZH)

# DEAP & CLEAN: Single Phase LAr Detectors

- ton scale DEAP/CLEAN planned for SNOLAB
- proposed first phase: 100 kg mini-CLEAN with WIMP search goal of  $\sim 5 \times 10^{-45} \text{ cm}^2$  or  $\sim 10$  events/yr
- To reject gamma background from PMTs and Ar-39 a discrimination better than  $10^{-8}$  for  $ER > 50 \text{ keVr}$  is required
- Current data from small ( $\sim 7 \text{ kg}$ ) DEAP-1 and micro-CLEAN detectors above ground demonstrate a discrimination of  $10^{-5}$  limited by neutron back in lab



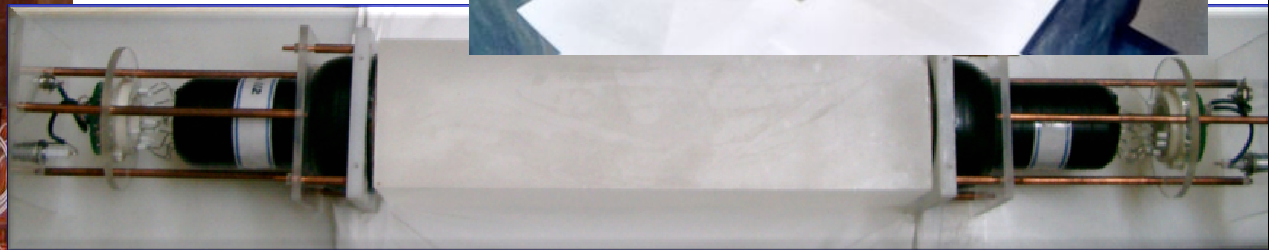
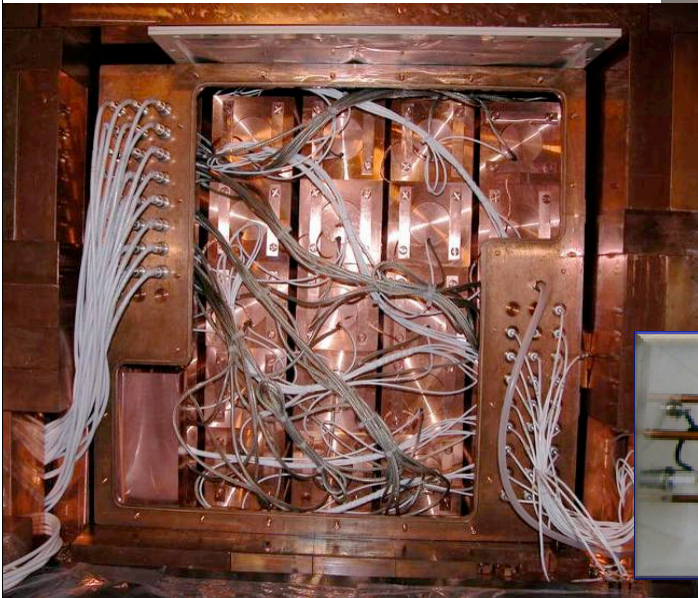
# Micro-CLEAN: PSD and QF





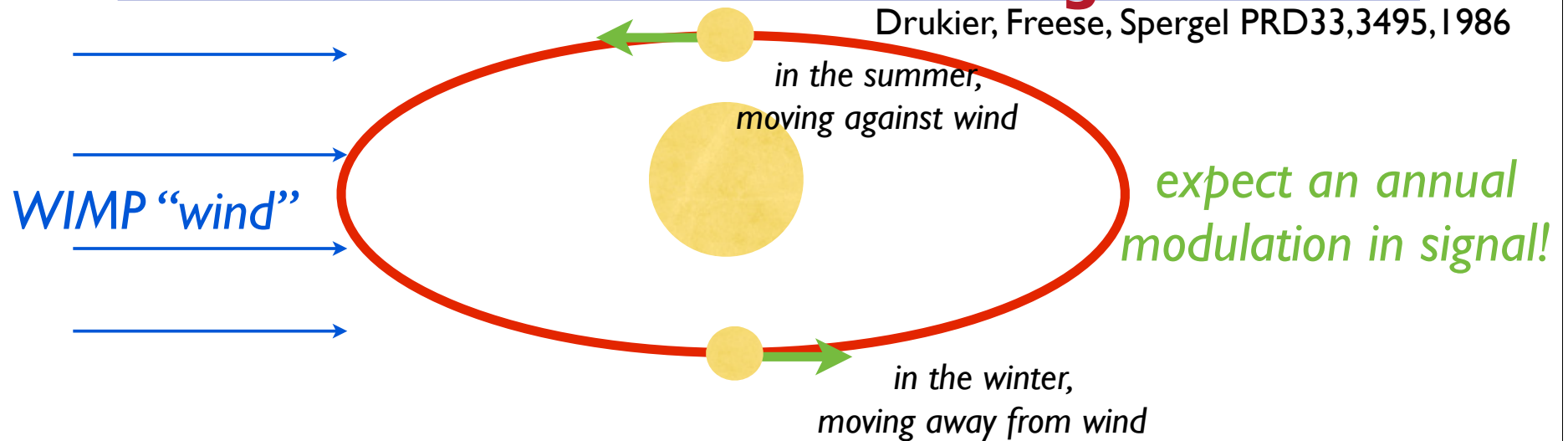
# Room Temperature Scintillation Experiments

- Inorganic alkali halide crystals (NaI (TI), CsI (TI)) : high density, high light output
- can be produced with high purity in large mass at affordable cost (annual modulation study)
- Sensitive to both SD and SI WIMP interactions
- PSD (better for CsI) but no discrimination between electron and nuclear recoils on an event-by-event basis
- Experiments: DAMA-LIBRA/Italy, KIMS/Korea, ANAIS/Spain (plan for 100kg NaI expt at Canfranc)

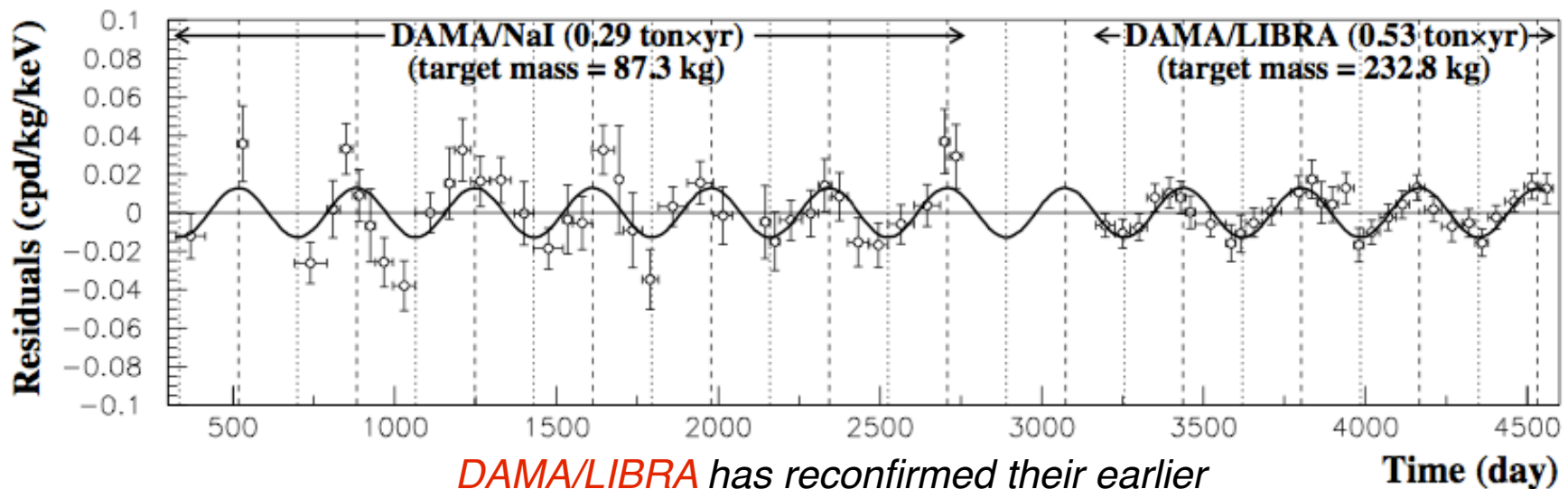


# Designed to search for Annual Modulation of WIMP Signal

Drukier, Freese, Spergel PRD33,3495,1986



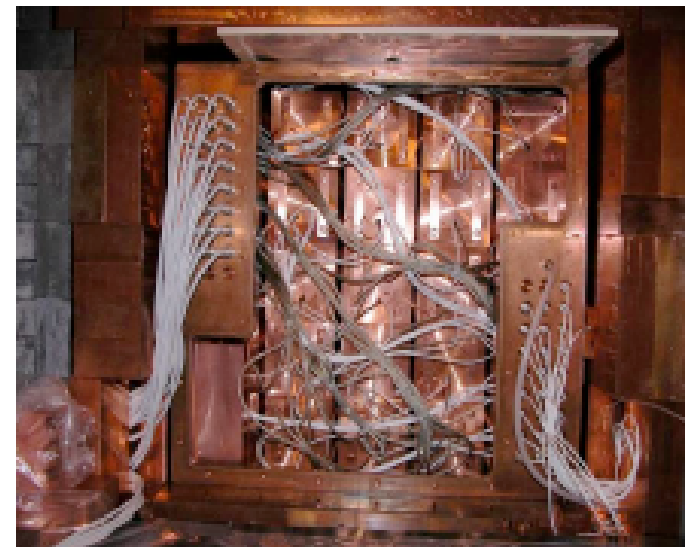
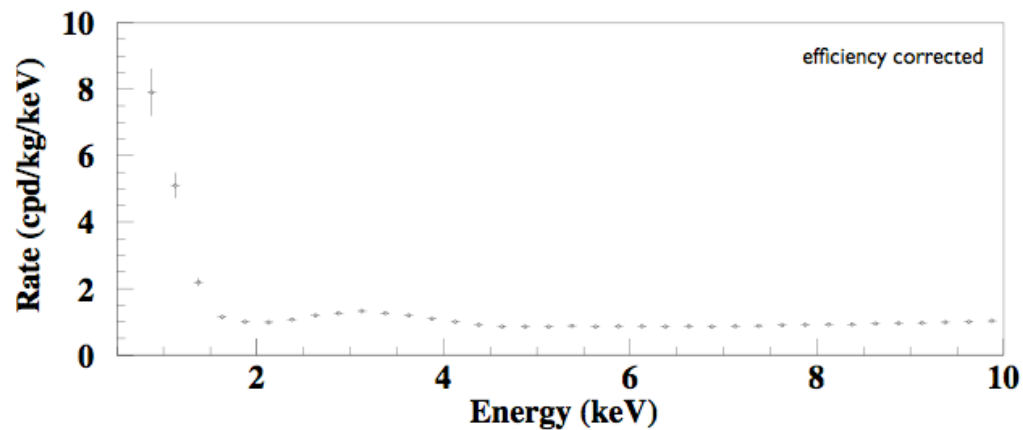
2-6 keV



*DAMA/LIBRA has reconfirmed their earlier observation of annual modulation signal [arXiv:0804.2741]*

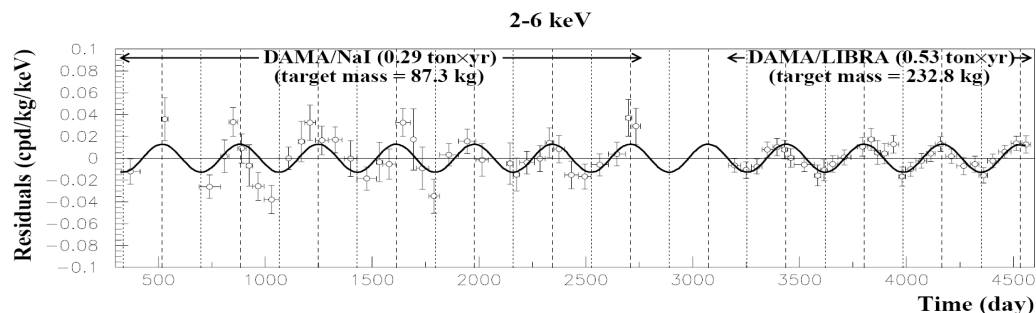
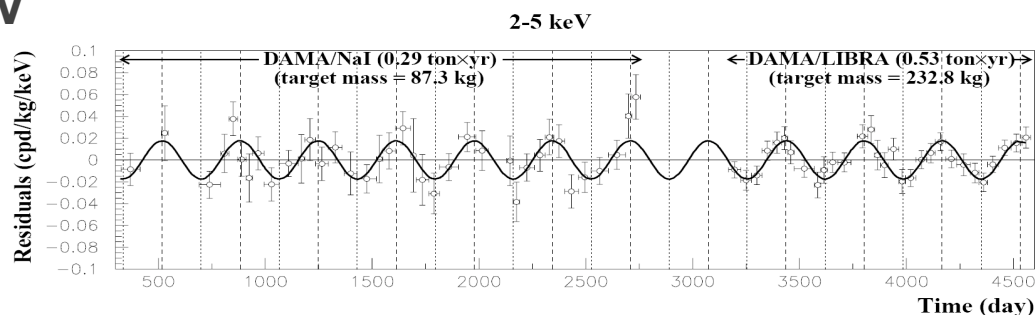
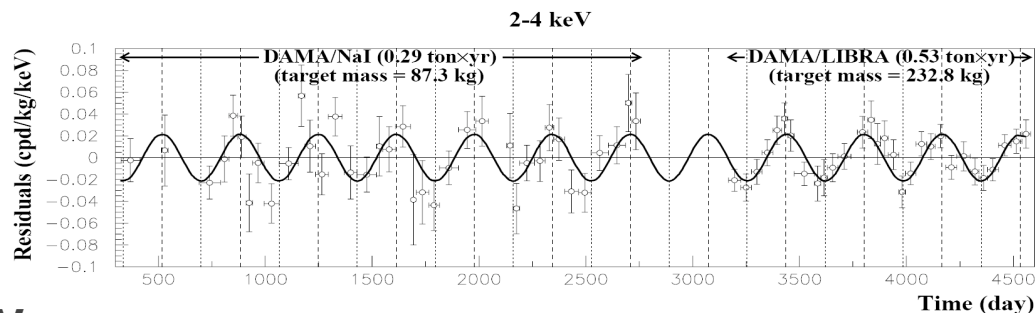
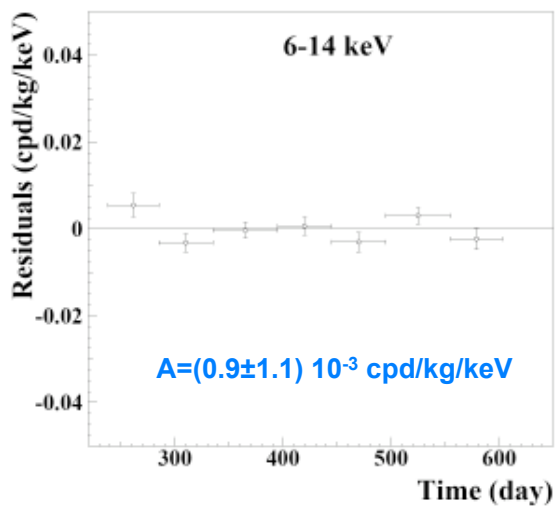
# DAMA/LIBRA @ LNGS

- **DAMA: 9 x 9.7 NaI (TI) crystals**
- BG level: 1-2 events/kg/d/keV
- $E_{\text{threshold}} \approx 2\text{keV}_{\text{ee}} \approx 25\text{keV}_r$
- **Data period:** 7 annual cycles, until July 2002; 0.29 ton x yr
- **LIBRA: 25 x 9.7 NaI (TI) crystals in 5 x 5 matrix**
- **Data period:** 4 annual cycles, 0.53 ton x year



# DAMA/LIBRA Results: A Strong Modulation Signal

- Total exposure: 0.82 ton x year
- **Modulation amplitude:**
- $A \cos [\omega(t-t_0)]$   
 $t_0 = 152.5 \text{ d}, T = 1 \text{ year}$
- **$A = (0.0215 \pm 0.0026) \text{ cpd/kg/keV}$**   
 (at  $8.3 \sigma \text{ CL}$ )
- No modulation above 6 keV



arXiv:0804.2738  
 arXiv:0804.2741

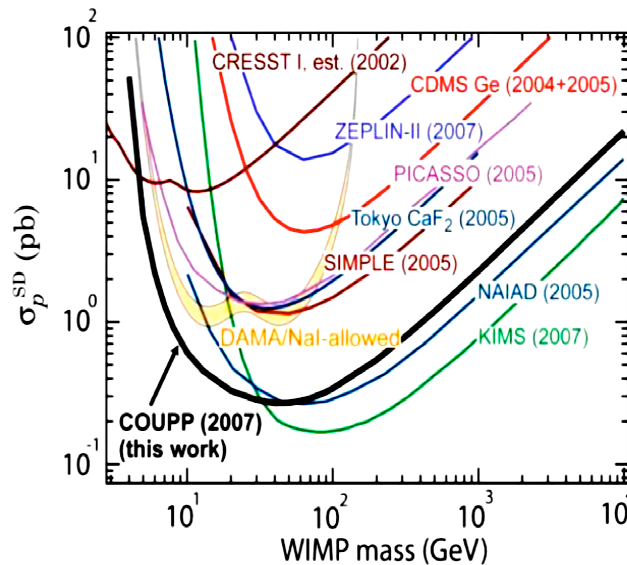
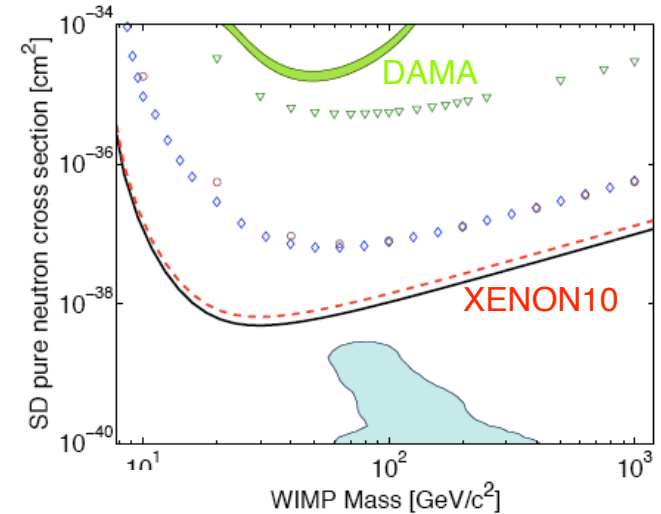
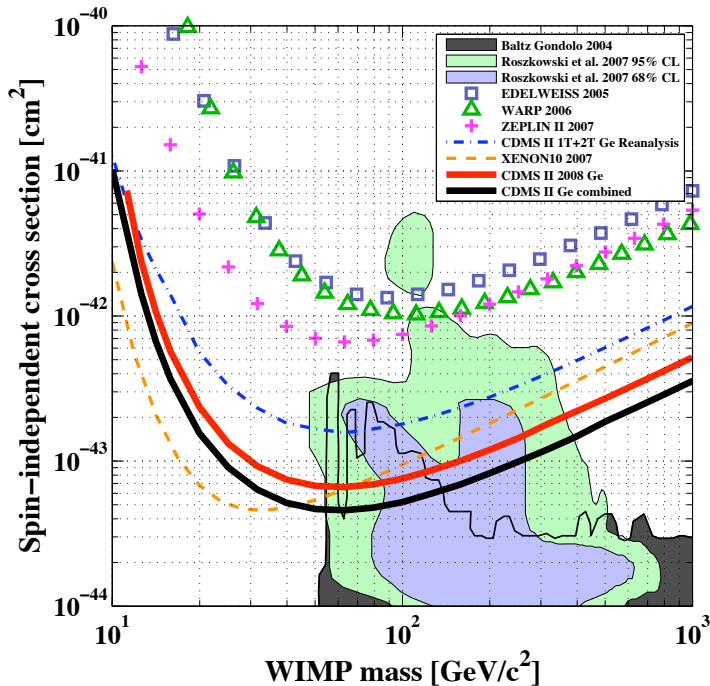
# Modulation result inconsistent with WIMP Recoils

WIMP recoils excluded by other experiments for both SD and SI

For a  $m_W = 50 \text{ GeV}$ ,  $\sigma_{SI} = 2 \times 10^{-6} \text{ pb}$ , the predicted rates are:

**XENON10:** 136 kg day, 4.5-27 keVr  $\Rightarrow$  162 events

**CDMS R123/124:** 397.8 kg day in Ge, 10-100 keVr  $\Rightarrow$  62 events



# Light Mass WIMPs also recently excluded!

- a 475 g Ge with a threshold of 0.33 keV ! WIMPs with mass  $<10$  GeV ruled out
- see Collar et al. <http://arxiv.org/abs/0807.0879>

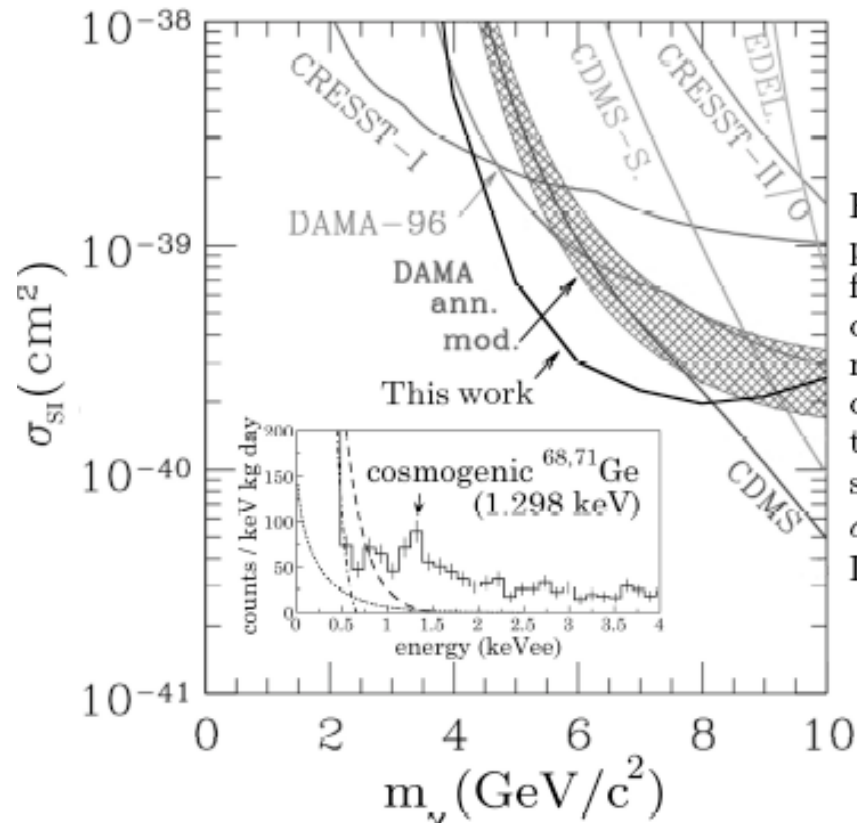


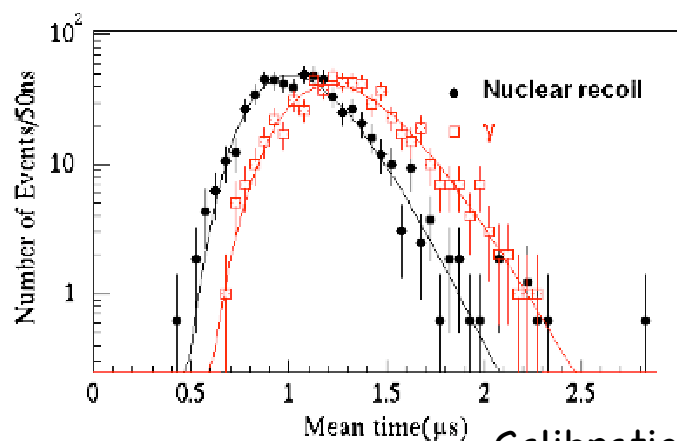
FIG. 2: Parameter space region (cross-hatched) able to explain the DAMA modulation via spin-independent couplings from an isothermal light-WIMP halo [5]. Lines delimit the coupling ( $\sigma_{SI}$ ) vs. WIMP mass ( $m_\chi$ ) regions excluded by relevant experiments [5]. All regions are defined at the 90% confidence level. Inset: PPC spectrum used for the extraction of present limits. Lines display the signals expected from some reference WIMP candidates (dotted:  $m_\chi = 8$  GeV/ $c^2$ ,  $\sigma_{SI} = 10^{-4}$  pb. Dashed:  $m_\chi = 6$  GeV/ $c^2$ ,  $\sigma_{SI} = 0.002$  pb. Dash-dotted:  $m_\chi = 4$  GeV/ $c^2$ ,  $\sigma_{SI} = 10^{-2}$  pb).

# KIMS @ Yang Yang Lab

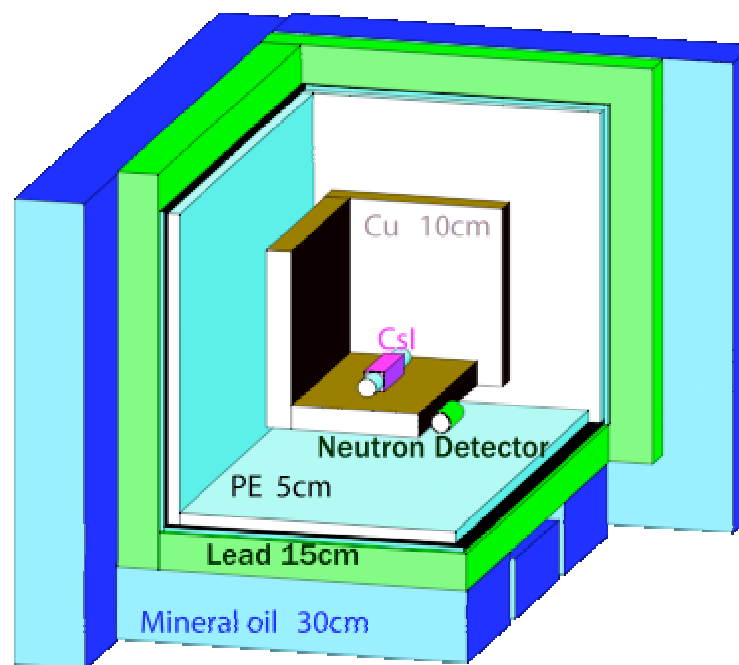


CsI(Tl) crystals at the YangYang Underground Lab in Korea (2000 m)  
 light yield:  $5 \times 10^4$  photons/MeV  
 QF: 8-15% between 10-100 keV<sub>e</sub>

	CsI(Tl)	NaI(Tl)
Density(g/cm <sup>3</sup> )	4.53	3.67
Decay Time(ns)	~1050	~230
Peak emission(nm)	550	415
Hygroscopicity	slight	strong



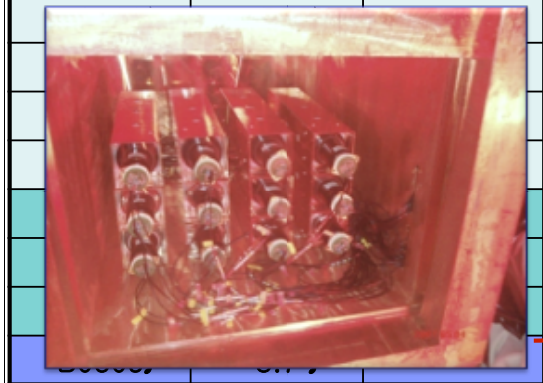
Calibration data



# KIMs Crystals

12 crystal array

Name	Weight (kg)	Data (kg days)
S0406	6.7	237
S0501A	8.7	1147
S0501B	8.7	1030
B0510A	8.7	616
B0510B	8.7	616



- CsI(Tl) Crystal 8x8x30 cm<sup>3</sup> (8.7 kg)

published  
PLB(2006)

Pilot run

published PRL(2007)

Engineering run

Total crystals  
8.7 kg x 12 = 104.4 kg  
started data taking

background level  
2~4 cpd/keV/kg (preliminary)

3" PMT (9269QA) quartz window, RbCs  
PC (green enhanced) ; 5 pe/keV

Sensitive to both SD and SI WIMP interactions  
Ge, Xe not sensitive to SD proton coupling  
complementary to CDMS, XENON10  
direct check of DAMA signal by I-127 recoil

Easy to get large mass at an affordable cost  
annual modulation study

High light yield ~60,000/MeV

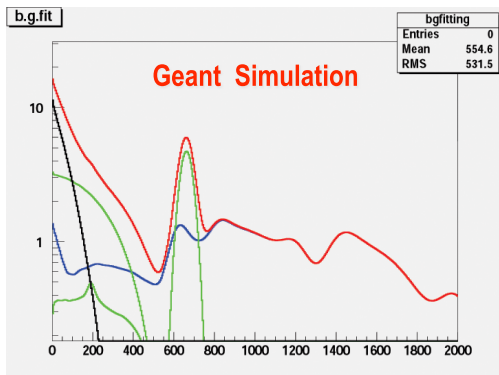
Pulse shape discrimination  
better gamma rejection than NaI(Tl)

Easy fabrication and handling



## Internal background

### Radioisotopes in the crystal



$^{137}\text{Cs}$  : 10 mBq/kg,  
0.35 cpd/mBq/kg @ 10 keV

$^{134}\text{Cs}$  : 20 mBq/kg  
0.07 cpd/mBq  
0.005 cpd/mBq

$^{87}\text{Rb}$  : 10 ppb  
1.07 cpd/ppb

**Cs-137 reduction – use ultra pure water in powder production; ~1.7 mBq/kg**

Rb reduction - recrystalization method ; < 1ppb

**Latest crystals are from ~2 cpd level powder**

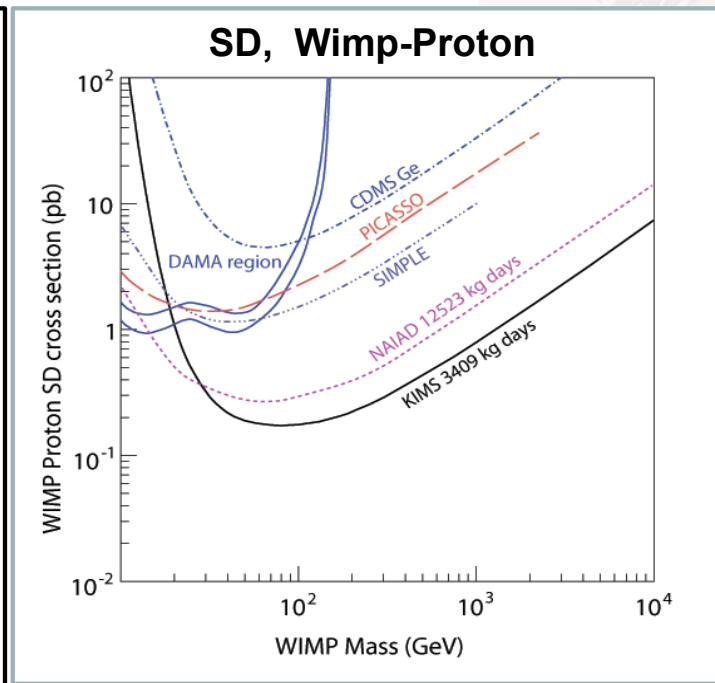
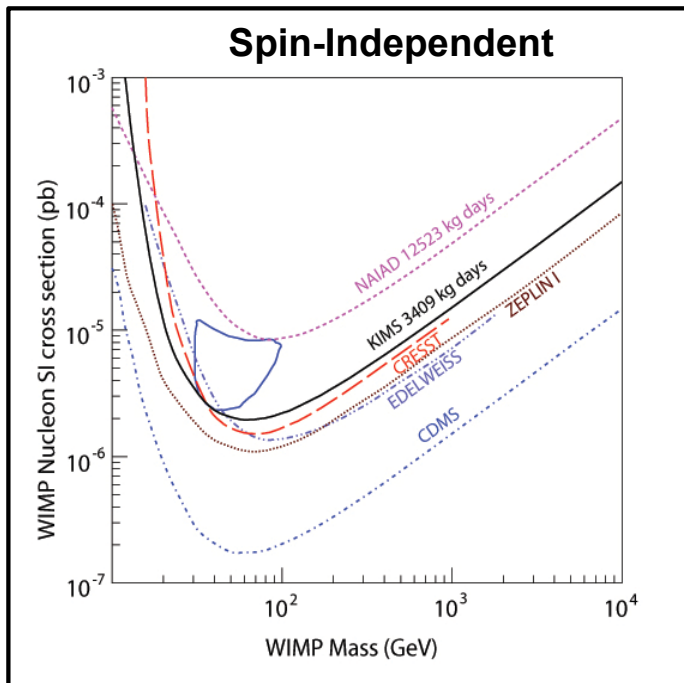
# KIMS Results

Direct comparison with DAMA (same nucleus) for SI coupling

Published data (3049 kg days) rule out DAMA signal region for both SD and SI interactions for WIMPs > 20GeV

Most stringent limit on SD interactions for pure proton coupling

## Cross-section upper limits



PRL 99, 091301 (2007)

$\rho_D = 0.3 \text{ GeV}/c^2/\text{cm}^3$   
 $v_0 = 220 \text{ km/s}$ ,  $v_{\text{esc}} = 650 \text{ km/s}$

Systematic uncertainty = 15%

**Nuclear recoil of  $^{127}\text{I}$   
of DAMA signal region**