

An analysis chain for CTA to test novel methods

Tino Michael

CEA Saclay, Irfu/SAp

SAP PostDoc Seminar

2016-12-06



2006 – 2011: Physics at TU Dresden

- specialisation: nuclear and particle physics
- diploma thesis on ATLAS detector

2012 – 2016: ANTARES at Nikhef, Amsterdam

- event reconstruction,
- neutrino point-source search
- first prototype optical module for KM3NeT
- **2016-05:** PhD thesis defended



**TECHNISCHE
UNIVERSITÄT
DRESDEN**

2016-04: CTA at CEA

- working with Thierry Stolarczyk
- implementing analysis chain for CTA
- testing impact of novel methods on:
 - event reconstruction, gamma-hadron separation, sensitivity
- Room Number: 120

2006 – 2011: Physics at TU Dresden

- specialisation: nuclear and particle physics
- diploma thesis on ATLAS detector

2012 – 2016: ANTARES at Nikhef, Amsterdam

- event reconstruction,
- neutrino point-source search
- first prototype optical module for KM3NeT
- 2016-05: PhD thesis defended

2016-04: CTA at CEA

- working with Thierry Stolarczyk
- implementing analysis chain for CTA
- testing impact of novel methods on:
 - event reconstruction, gamma-hadron separation, sensitivity
- Room Number: 120



2006 – 2011: Physics at TU Dresden

- specialisation: nuclear and particle physics
- diploma thesis on ATLAS detector

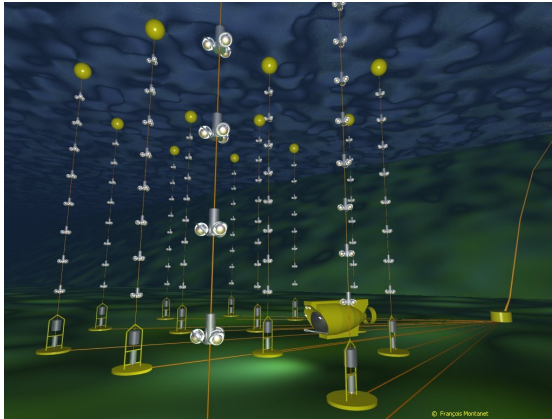
2012 – 2016: ANTARES at Nikhef, Amsterdam

- event reconstruction,
- neutrino point-source search
- first prototype optical module for KM3NeT
- **2016-05**: PhD thesis defended

2016-04: CTA at CEA

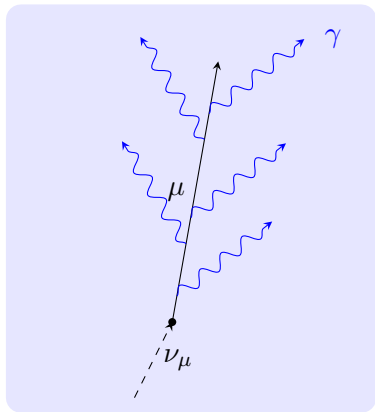
- working with Thierry Stolarczyk
- implementing analysis chain for CTA
- testing impact of novel methods on:
 - event reconstruction, gamma-hadron separation, sensitivity
- Room Number: 120



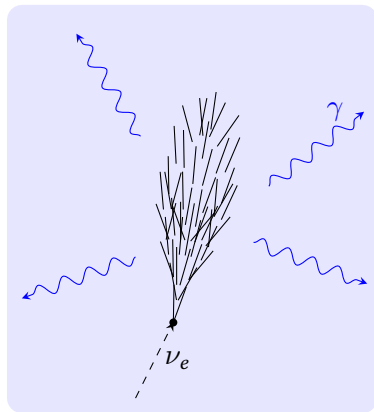


- 2500 m below sea surface
- 40 km off the coast of Toulon, France
- connected via electro-optical cable
- complete since 2008-05
- 12 Lines, 885 PMTs
- radius: 90 m
height: 400 m
- 0.5 ns time resolution
- < 10 cm acoustic positioning

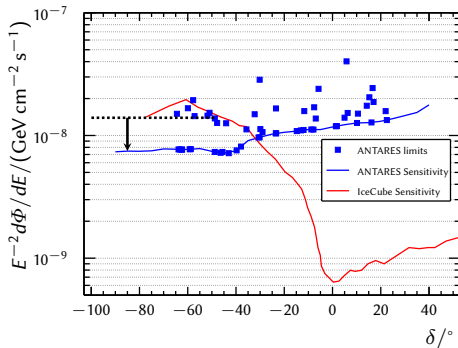
Muon:



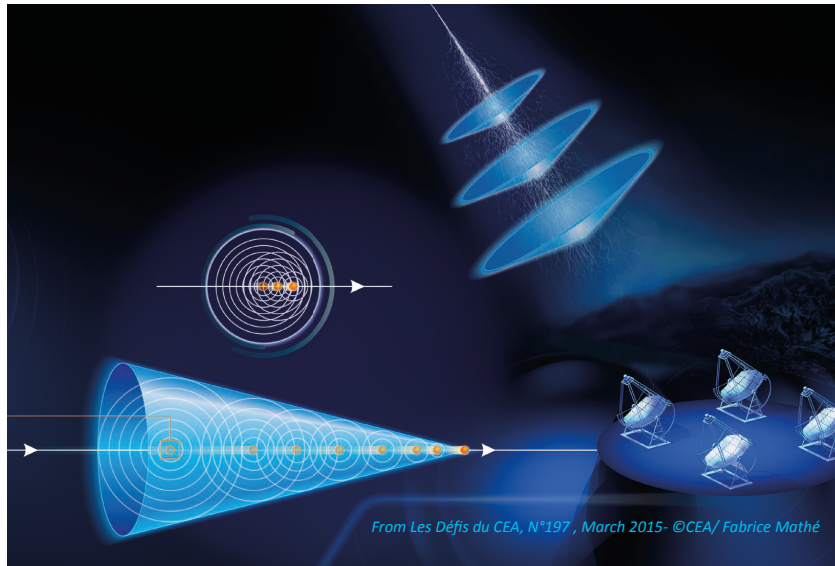
em-Shower:

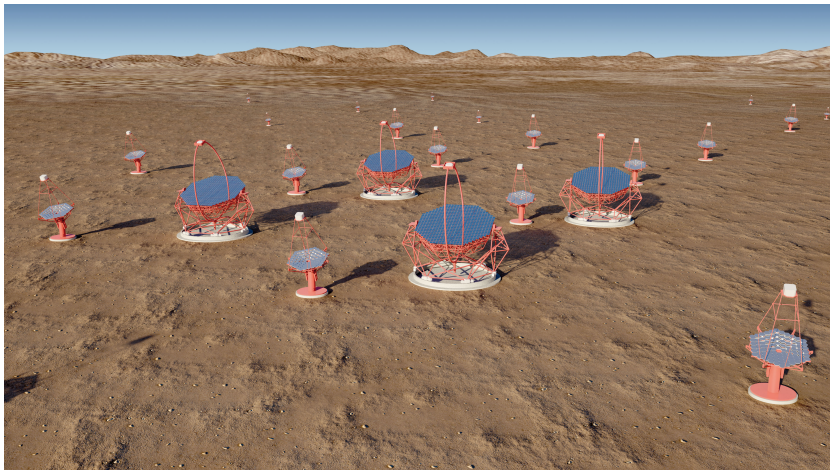


- no significant cluster in Full Sky Search
- Sensitivity improved from $1.4 \times 10^{-8} \text{ GeV cm}^{-2} \text{ s}^{-1}$ to $7.2 \times 10^{-9} \text{ GeV cm}^{-2} \text{ s}^{-1}$
- best limits for many candidates in galactic region from single experiment
- challenged only by combined ANTARES/IceCube analysis

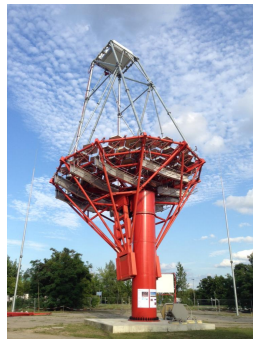








- sites chosen:
La Palma, Spain and
Paranal, Chile
- many prototypes already built
- construction of first foundation started



CTA – Telescope Sizes

LST

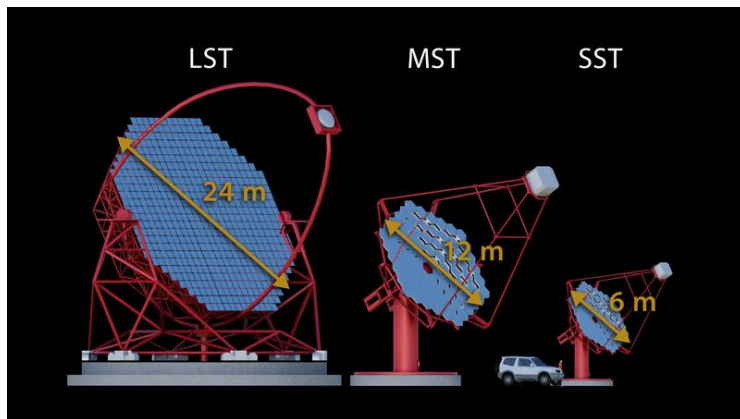
- 4 at both sites
- $E_\gamma < 0.1 \text{ TeV}$
- FoV: 4.5°

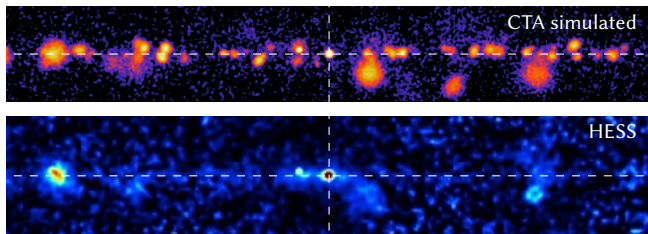
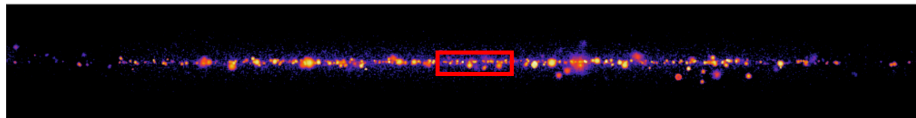
MST

- 25 at south; 15 at north site
- $0.1 < E_\gamma / \text{TeV} < 100$
- FoV: $> 7^\circ$

SST

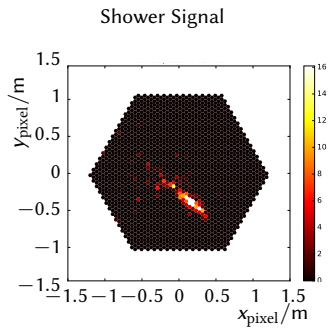
- 70 south site only
- $E_\gamma > 10 \text{ TeV}$
- FoV: 10°



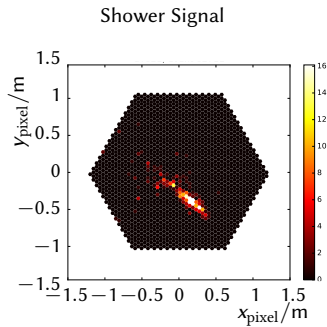


- better sensitivity and resolution w.r.t. HESS
- resolve overlapping extended sources
- viewing distance of 20 kpc
- expect population increase $\times 5$

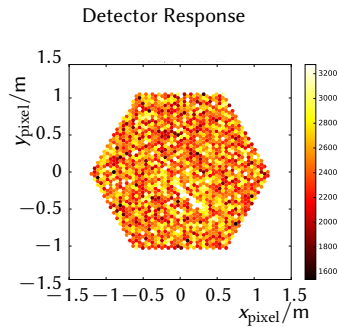
How the Signal looks like



How the Signal looks like

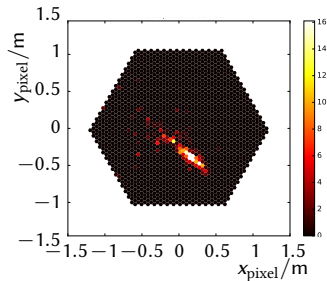


Stars, Clouds,
Electronics etc.



How the Signal looks like

Shower Signal

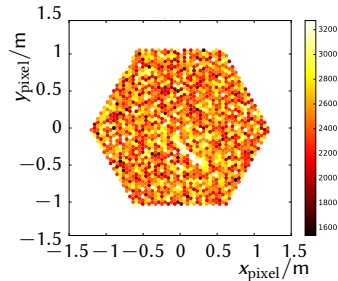


Stars, Clouds,
Electronics etc.

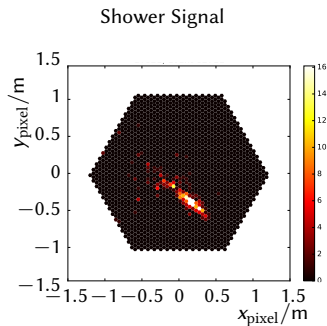


←
How to get
back?

Detector Response



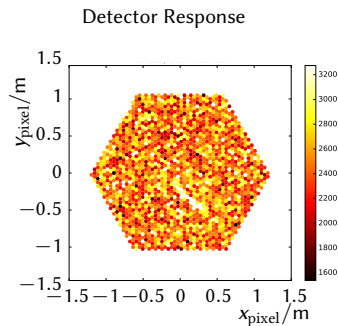
How the Signal looks like



Stars, Clouds,
Electronics etc.



←
How to get
back?



possible solutions:

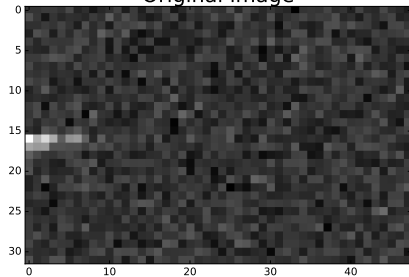
Tail-Cuts Used by H.E.S.S.: Cut away all pixels below a given threshold
– possibly recover neighbouring pixels through second, lower threshold.

Fourier Trans. Decompose image into Fourier coefficients and cut in Fourier space.

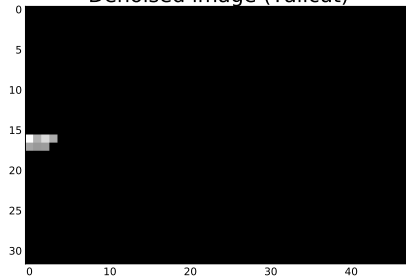
Wavelet Trans. Decompose image into waveletes (in contrast to waves) and cut there.

Comparing Methods

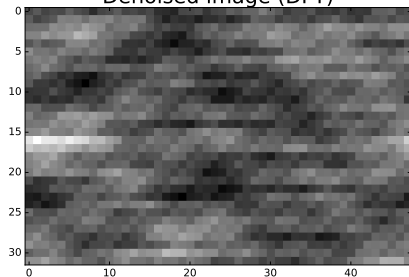
Original image



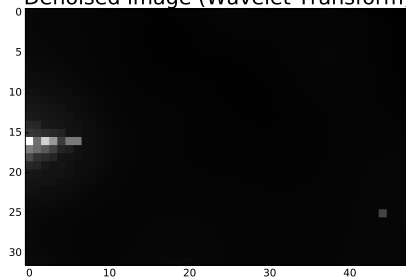
Denoised image (Tailcut)

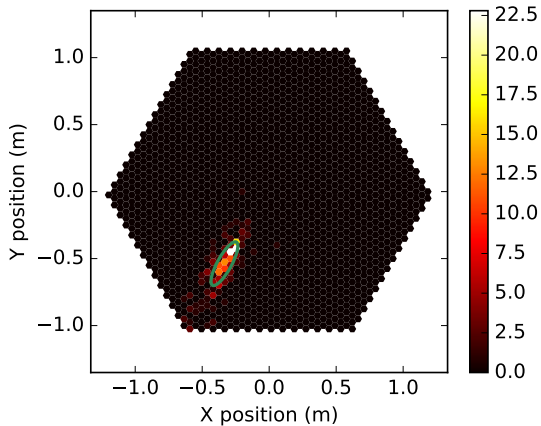


Denoised image (DFT)

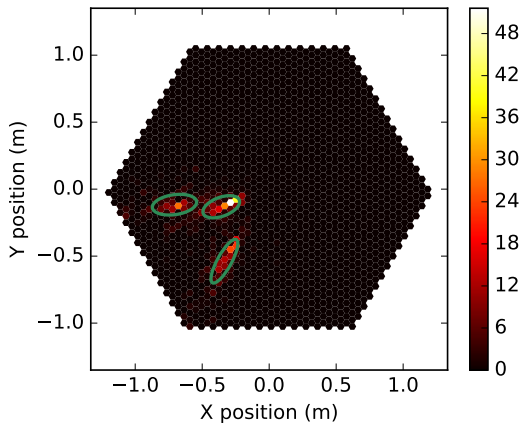


Denoised image (Wavelet Transform)

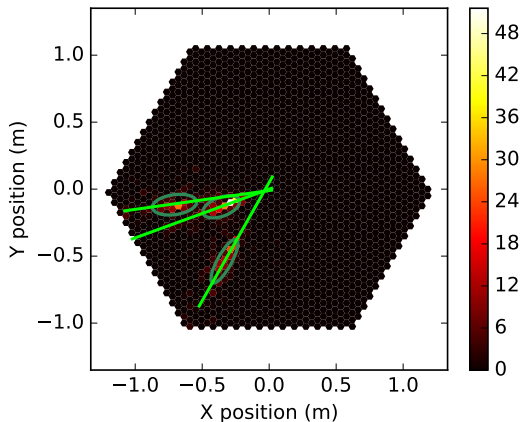




- construct an ellipsis with moments of the shower image:
Hillas Parametrisation



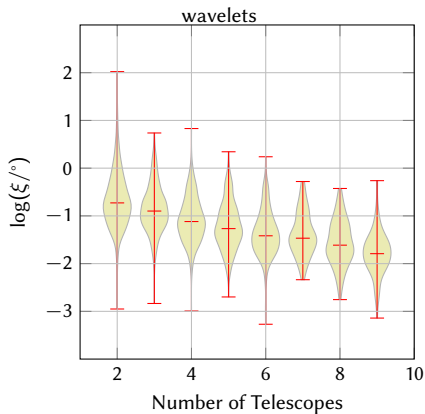
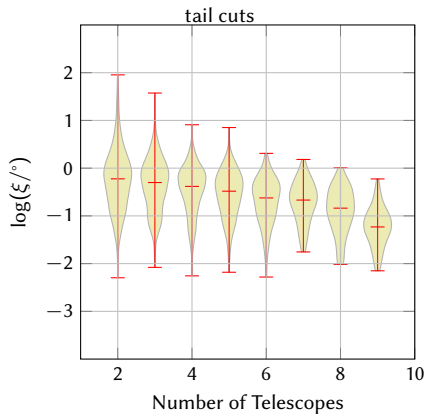
- construct an ellipsis with moments of the shower image:
Hillas Parametrisation
- combine images from different cameras



- construct an ellipsis with moments of the shower image:
Hillas Parametrisation
- combine images from different cameras
- intersection of their ellipsis axes is the shower origin

Shower Reconstruction – Performance

angle between reconstructed and simulated direction



Photon / Proton Discrimination

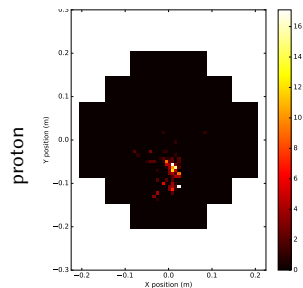
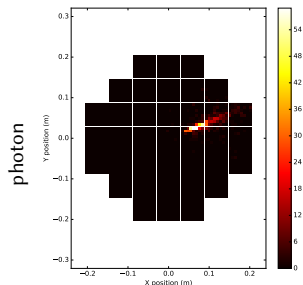
- Protons pose major background
- Event rate about 10^5 times above Photons

H.E.S.S. methode:

- reducing total signal on camera, length and width of ellipsis and their variances from all telescopes into one parameter to cut on

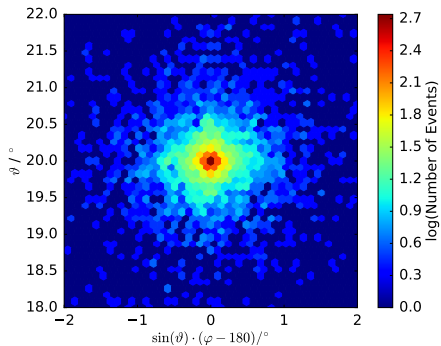
here instead:

- Discrimination with Random Forest Classifier fed with width, length, ... of Hillas ellipsis from each camera image

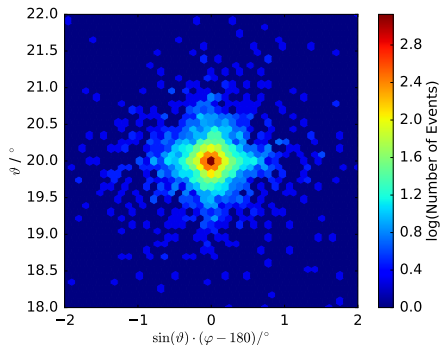


Skymap of a point-source

tail cuts



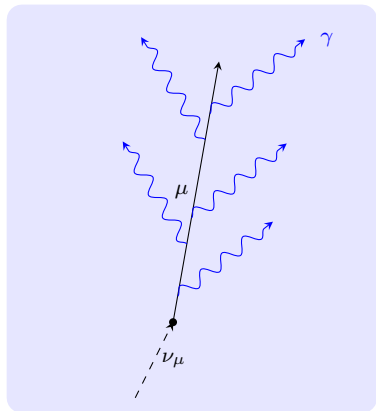
wavelets



- apply analysis chain to full array
- investigate
 - impact on overlapping sources
 - real-time analysis
 - → reduction of data flow
- for more information on wavelets:
see Jérémie's talk in February
- will be here for another 1,5 years

Backup

- muon track reconstruction already well established
 - muons can pass through detector
 - Cherenkov radiation along track
 - photons emitted at $\varphi_{\text{Ch}} \approx 42^\circ$
 - clean signature
-
- maximum likelihood fit based on hit time residuals
 - median angular resolution $\approx 0.4^\circ$
-
- limit us to $\nu_\mu \rightarrow \mu$ (and $\nu_\tau \rightarrow \tau \rightarrow \mu$) interactions



- shower events open window to

$$\nu_e \rightarrow e$$

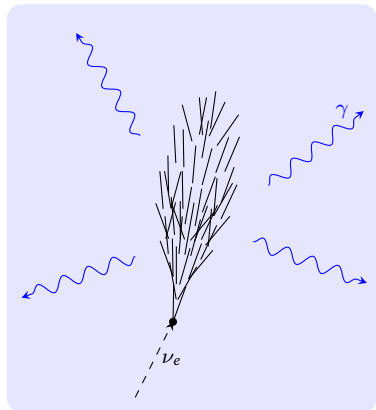
$$\nu_x \rightarrow \nu_x + \text{hadr.}$$

$$\nu_\tau \rightarrow \tau \rightarrow e/\text{hadr.}$$

- cascade of particles within few metres
- can be approximated as point source
- emits shell of light in all directions
- still, more light emitted under “Cherenkov angle”

in ice:

- effect almost completely gone due to scattering



- shower events open window to

$$\nu_e \rightarrow e$$

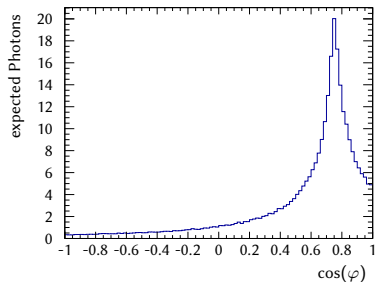
$$\nu_x \rightarrow \nu_x + \text{hadr.}$$

$$\nu_\tau \rightarrow \tau \rightarrow e/\text{hadr.}$$

- cascade of particles within few metres
- can be approximated as point source
- emits shell of light in all directions
- still, more light emitted under “Cherenkov angle”

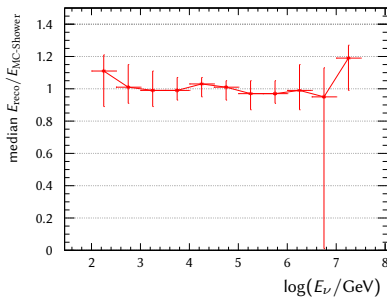
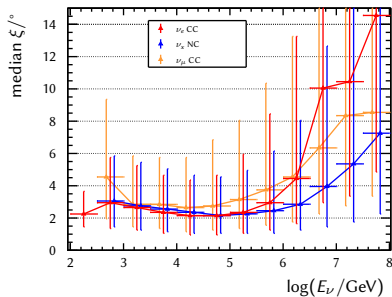
in ice:

- effect almost completely gone due to scattering



expected number of Photons from a 1 TeV
shower on a PMT in 100 m distance

Shower Reconstruction – Performance: Direction & Energy



- position of shower position reconstructed with accuracy of about 1 m
- median angular error $\xi \approx 2^\circ$ in relevant energy range (compared to 15° for IceCube)
- energy resolution of 5 % – 10 %



- 17'' glass sphere
- 31 3'' PMTs, 5 rings with 6 PMTs + 1
- 1 DOM equivalent to 3 ANTARES OMs (cheaper)
- but only one penetrator per floor (less risky)
- low power consumption (<10 W per DOM)
- covers almost whole solid angle
- directional information
- background reduction from PC segmentation
- prototype deployed on Instrumentation Line within ANTARES

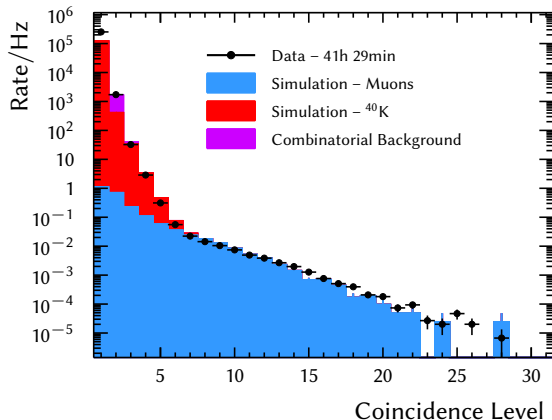


- 17'' glass sphere
- 31 3'' PMTs, 5 rings with 6 PMTs + 1
- 1 DOM equivalent to 3 ANTARES OMs (cheaper)
- but only one penetrator per floor (less risky)
- low power consumption (<10 W per DOM)
- covers almost whole solid angle
- directional information
- background reduction from PC segmentation
- prototype deployed on Instrumentation Line within ANTARES



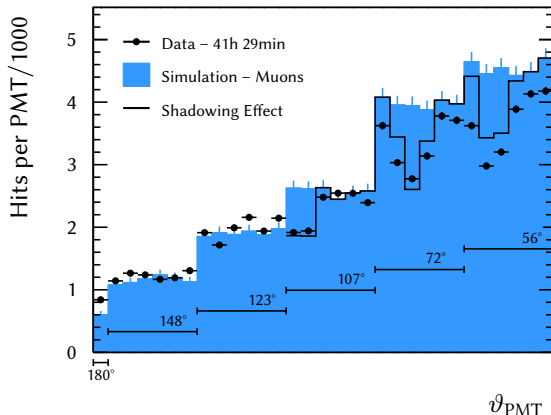
- 17'' glass sphere
- 31 3'' PMTs, 5 rings with 6 PMTs + 1
- 1 DOM equivalent to 3 ANTARES OMs (cheaper)
- but only one penetrator per floor (less risky)
- low power consumption (< 10 W per DOM)
- covers almost whole solid angle
- directional information
- background reduction from PC segmentation
- prototype deployed on Instrumentation Line within ANTARES

Rate of different Coincidence Levels



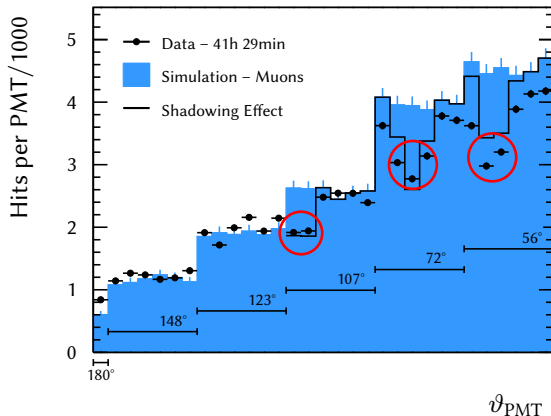
- Coincidence Level:
of Hits within 20 ns
 - Combinatorial Background:
estimate for coincidences caused by
multiple single hits
 - derived from average PMT rate in
data
- European Physical Journal C (2014)
74:3056

Hit Count according to PMT Zenith



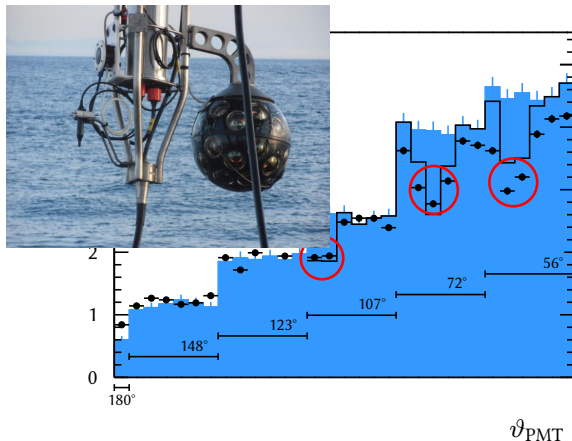
- Coincidence Level > 7
- clear asymmetry towards upwards looking PMTs
- drops due to LCM shadowing
- we can “see” the LCM
- \rightarrow directional information from single storey

Hit Count according to PMT Zenith



- Coincidence Level > 7
- clear asymmetry towards upwards looking PMTs
- drops due to LCM shadowing
- we can “see” the LCM
- \rightarrow directional information from single storey

Hit Count according to PMT Zenith



- Coincidence Level > 7
- clear asymmetry towards upwards looking PMTs
- drops due to LCM shadowing
- we can “see” the LCM
- \rightarrow directional information from single storey