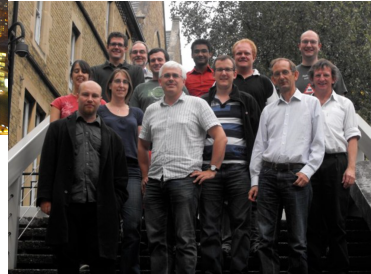


CMB component separation

Mel Irfan
Building 709, Room 274

14/02/17

About me...



MANCHESTER The University of Manchester
Jodrell Bank Centre for Astrophysics

altran



Overview

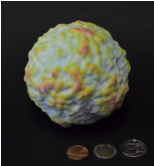
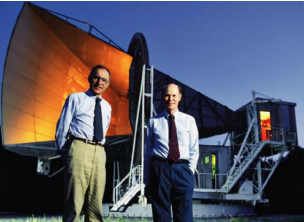
Introduction to the CMB

Diffuse foreground emissions

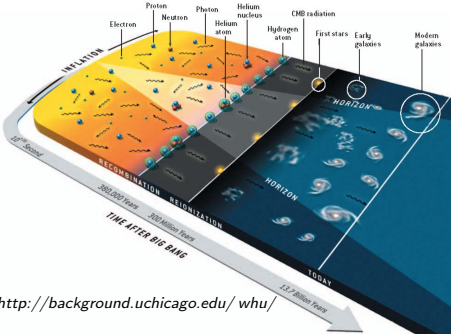
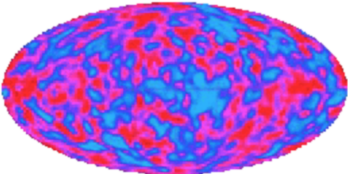
Data reduction for a typical radio survey

Early science results from C-BASS

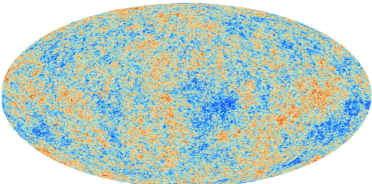
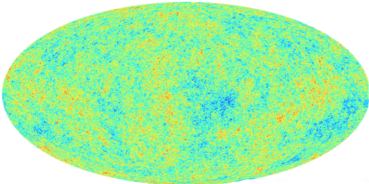
The CMB



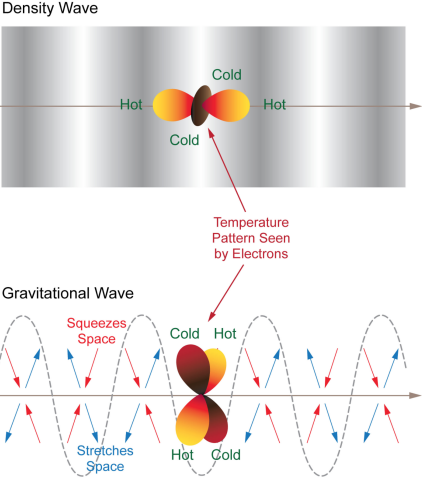
Clements et al. 2016



<http://background.uchicago.edu/whu/>

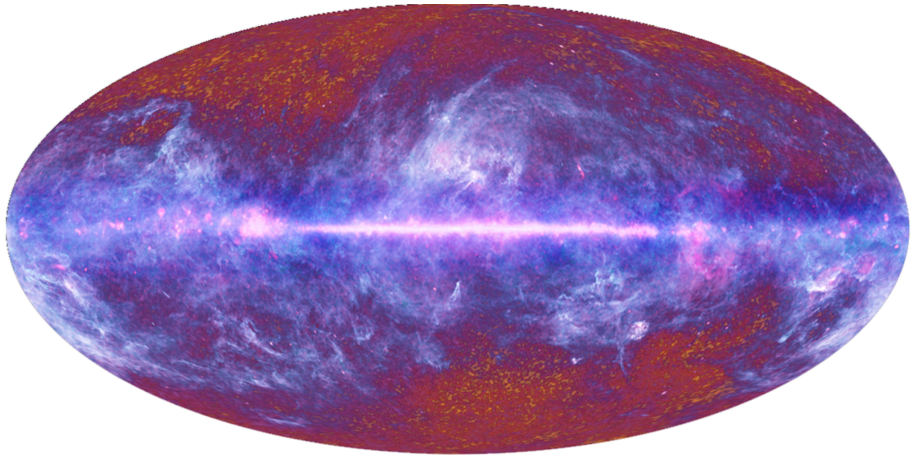


CMB anisotropies

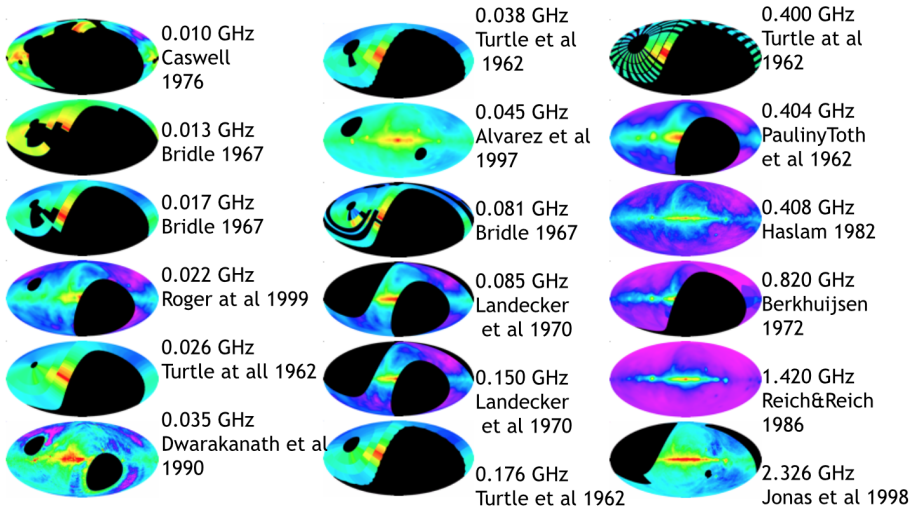


<http://bicepkeck.org/>

What you actually see...

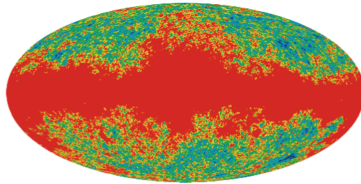


The solution...

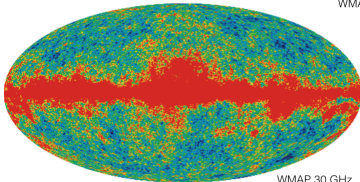


De Oliveira-Costa 2008

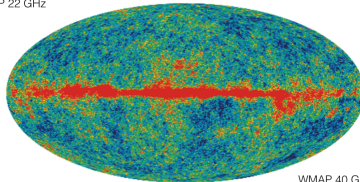
The solution...



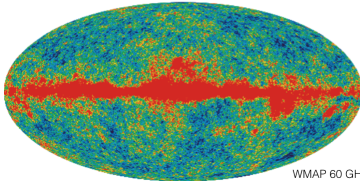
WMAP 22 GHz



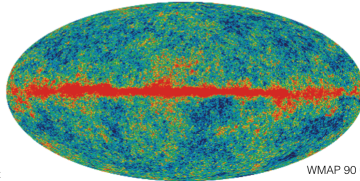
WMAP 30 GHz



WMAP 40 GHz

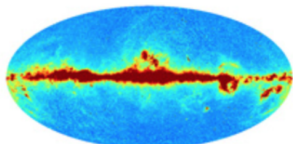


WMAP 60 GHz

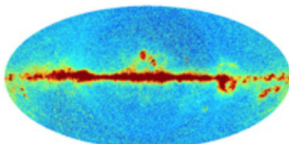


WMAP 90 GHz

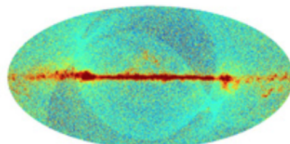
The solution...



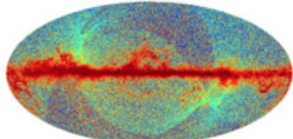
LFI 30 GHz



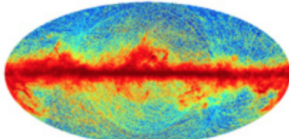
LFI 44 GHz



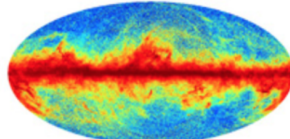
LFI 70 GHz



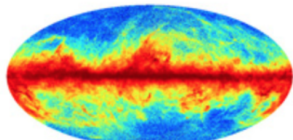
HFI 100 GHz



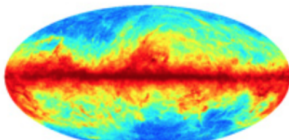
HFI 143 GHz



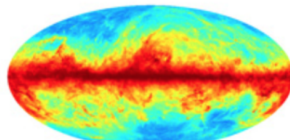
HFI 217 GHz



HFI 353 GHz

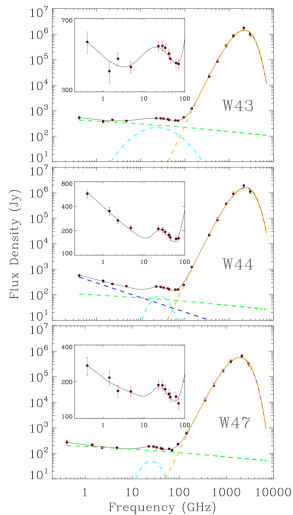


HFI 545 GHz



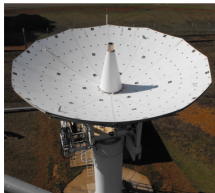
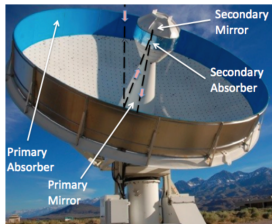
HFI 857 GHz

Physics of the ISM – AME

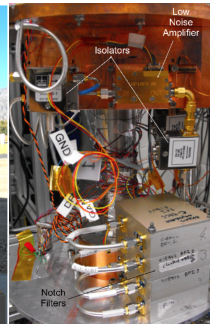
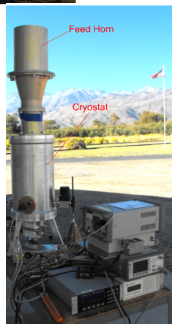


Irfan et al. 2015

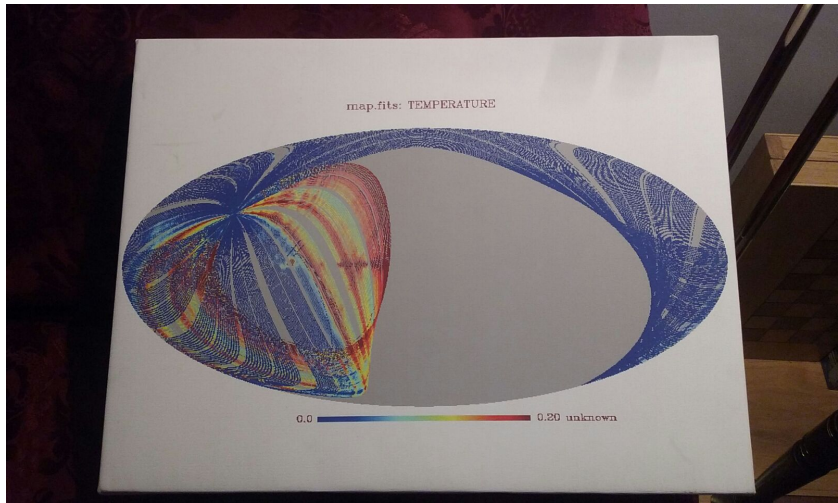
C-BASS



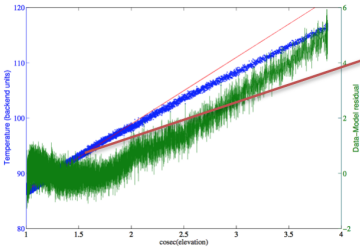
	Northern	Southern
Latitude	37°2	-30°7
Antenna type	Gregorian	Cassegrain
Diameter	6.1 m	7.6 m
FWHM	0°73	0°73
Main Beam Efficiency	80.0%	80.0%
Full Beam Efficiency	91.9%	91.3%
Intensity central frequency	4.76 GHz	-
Intensity Bandwidth	489 MHz	-
Total System Temperature	40 K	-



First ever map...



Calibration

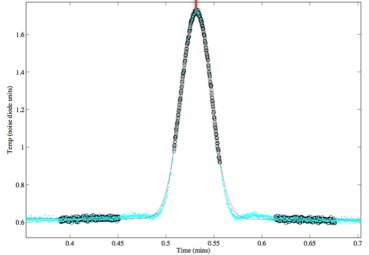


$$m_{ND} = \frac{T_{\text{atmos}} - T_{\text{Sky}}}{T_{ND}} \times \tau_0$$

$$y = \frac{T_{\text{sky},ND\text{on}} - T_{\text{sky},ND\text{off}}}{T_{\text{src}} e^{-\tau}} = \frac{T_{ND}}{T_{\text{src}}} e^{\tau} = \frac{T_{ND}}{T_{\text{src}}} \left(1 + \frac{\tau_0}{\sin(\theta_{\text{el}})} \right)$$

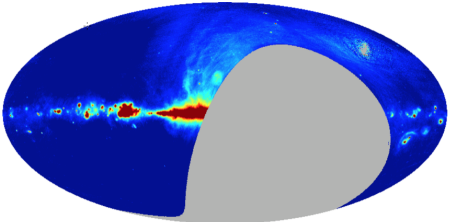
$$T_{ND} = \frac{-\beta + \sqrt{\beta^2 - 4\alpha\gamma}}{2\alpha}$$

where $\alpha = \frac{m_{ND}}{T_{\text{atmos}} - T_{\text{Sky}}}$, $\beta = \sin \theta_{\text{el}}$ and $\gamma = -y T_{\text{src}} \beta$

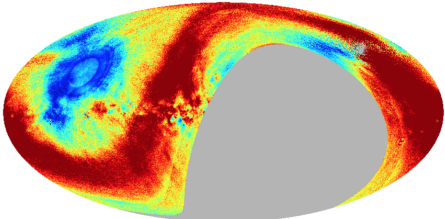


Baars 1977, Weiland et al. 2002

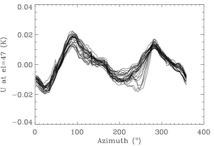
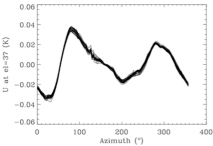
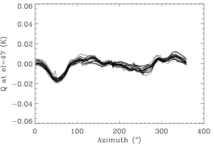
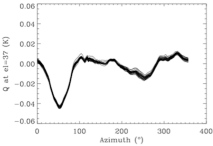
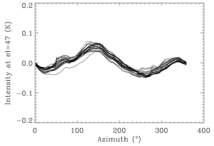
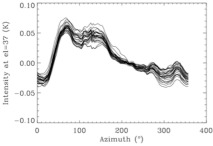
Ground Spillover



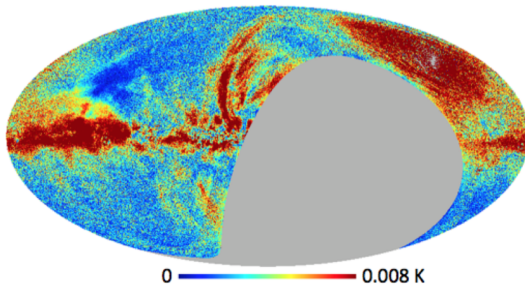
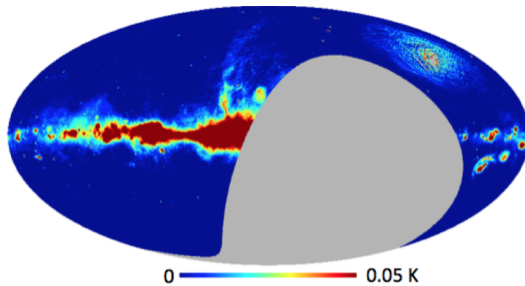
0.0 ——— 0.05 K



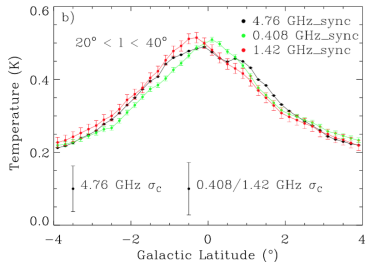
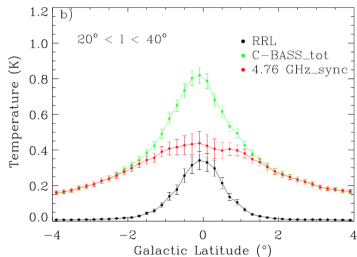
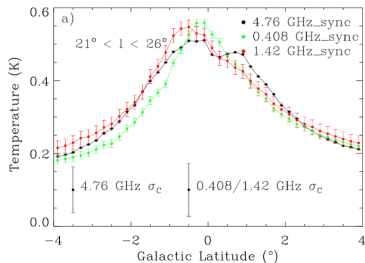
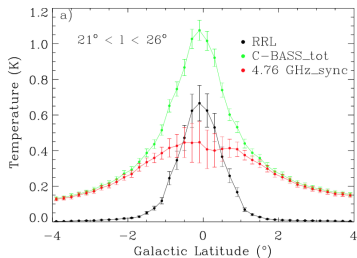
0.0 ——— 0.008 K



Just solar sidelobes left...

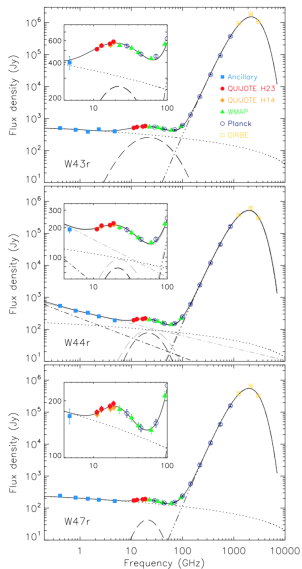


Synchrotron spectral index



Free-Free template from Alves et al. 2012

Constraining AME



Genova-Santos et al. 2017

Ongoing

