



Beams for European Neutrino Experiments (BENE) -- Midterm Scientific Report --

edited by the (extended) BENE Steering Group

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Abstract

The activities of BENE during 2004 and 2005 are reviewed. Neutrino oscillation experiments at accelerators offer the richest possibilities of precision studies of neutrino mixing and mass differences, with the potential of important discoveries including leptonic CP or T violation. Two main options for a major initiative have been studied.

- 1) a high energy Neutrino Factory coupled to a large dense magnetized detectors;
- 2) a lower energy Betabeam and conventional SuperBeam, coupled to a very large low- density, non-magnetic, detector.

Both offer significant scientific breakthroughs over other planned facilities. Much remains to be done to optimise and establish the cost, performance and feasibility of either solution so as to allow comparison and decision. The proposal of a FP7 Design Study of a Neutrino Facility to be completed by the end of the decade is being prepared. Its success will require strong support and engagement at CERN and other European laboratories and funding agencies. International contributors are already engaged in the framework of an international scoping study. The recommendations and milestones proposed by BENE towards a complete conceptual design are summarized.

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BENE

BENE (Beams for European Neutrino Experiments) is a FP6 Networking Activity (NA) approved by EC for the years 2004-2008 within the CARE Integrating Infrastructure Initiative. It is rooted in the ECFA study groups, chaired by Alain Blondel (Geneva) and active since 1998. Coordinated by Vittorio Palladino (Napoli), BENE assembles 220 members in Europe from 13 nodes, CERN, INFN, CEA, RAL, IN2P3, GSI, FZJ, PSI, three consortia of laboratories and universities coordinated by IC London in UK, Uni Geneva in Switzerland, TU Munich in Germany, CSIC in Spain, UC Louvain in Belgium. Groups from Poland, Greece, Finland and Bulgaria are in contact.

BENE integrates and coordinates the activities of the accelerator and particle physics communities working together, in a worldwide context, dedicated to the study of a forefront neutrino beam facility for Europe.

The final objectives are:

- to establish a road map for upgrade of our present facility and the design and construction of a new one;
- to assemble a community capable of sustaining the technical realisation and scientific exploitation of these facilities;
- to establish and propose the necessary R&D efforts to achieve these goals

This document summarizes progress made up to the end of the second of its five years of existence. A final report is due in December 2008.

BENE is organized in a Physics work package,

- the ECFA/BENE Neutrino Oscillations group coordinated by Mauro Mezzetto (Padova) and Pilar Hernandez (Valencia), recently joined also by Andrea Donini (Madrid) and the four accelerator packages of the European Neutrino Group coordinated by R. Edgecock (RAL) and H. Haseroth (CERN)
 - The proton driver work package, coordinated by Christian Cavata (CEA Saclay)
 - The high-power target work package coordinated by Roger Bennett (RAL)
 - The collector work package coordinated by Marcos Dracos (IRES Strasbourg)
 - A work package on novel neutrino beams, itself distributed in three activities:
 - muon front-end studies coordinated by Ken Long (Imperial College)
 - muon acceleration and storage ring coordinated by François Méot (CEA Grenoble)
 - beta-beam coordinated by Mats Lindroos (CERN)

At the occasion of the ISS a working group on detector design and R&D was also introduced.

BENE maintains close connection with the ongoing long baseline neutrino experiments in Europe (OPERA, ICARUS), or abroad (MINOS, T2K), with the proposed European projects (Double Chooz) and abroad (NOvA) and with the internationally conducted R&D experiments (HARP, MUSCAT, HIPPI, MUCOOL, PRISM, MICE and MERIT in particular).

BENE holds plenary meetings two or three times a year and participates actively in the preparation of the international workshops of the NUFACT series.

In preparation for the FP7 design study proposal, an International Scoping Study (ISS) was launched at the occasion of NUFACT05 in Frascati, in collaboration with the international partners of BENE. The ISS will seek to identify the R&D and design efforts that will be necessary in order to be ready to propose a major new neutrino initiative, that could be located in Europe, by about 2010.

Executive Summary

The aim of BENE is to investigate the feasibility and performance of an ambitious accelerator neutrino oscillation facility that could be located in Europe. Precision measurements of neutrino mixing angles and mass hierarchy, with the ultimate aim of discovery and study of leptonic CP violation, provide a very strong physics case, independent of the explorations at high energy colliders. This document summarizes the state of advancement of BENE after nearly two years.

The facilities that have been considered more specifically are:

- 1) A Neutrino Factory, based on a high brilliance muon beam would provide high-energy electron neutrinos (up to 20-50 GeV), aimed at magnetic detectors situated ~700 to 7000 km away. This is by far the best tool to perform very precise and unambiguous measurements of oscillation parameters, neutrino mass hierarchy, CP violation and tests of universality in the neutrino sector.
- 2) A neutrino beta-beam facility, that would provide from beta decay of specific ions a very clean beam of electron (anti) neutrinos of energies up to ~2 GeV. The very large detectors that are required may be the same as those needed to extend the search for proton decay and astrophysical neutrinos. In combination with a conventional muon-neutrino beam (“superbeam”) of the same energy, beta-beams could provide interesting sensitivity in search for leptonic CP and T violation.

Both possibilities would extend considerably our knowledge of neutrino masses and mixings over other proposals. Both offer appreciable synergies with other fields within or outside particle physics, and both options have proponents and experts in the international community. New promising ideas building on these primary concepts arise continuously. A number of open questions have been outlined concerning their feasibility, cost and performance, both in absolute terms and relative to each other. Answers to these questions require theoretical and conceptual work, a coordinated design study, and R&D experiments.

Considerable progress has been made over the last two years towards achieving these goals. BENE organized the MultiMegawatt Workshop in 2004 and two workshops in 2005, NNN05 and NuFact05. The HARP experiment to measure particle production on various targets and proton beam energies, has now produced its first results. The target experiment MERIT, to lay the foundations of a multi-megawatt neutrino target station, is now under construction at CERN; the MICE experiment, to test the practical feasibility of muon cooling, is progressing at RAL; all are carried out by international collaborations. The beta-beam design study is taking place in the context of EURISOL, and a preliminary feasibility study of a megaton detector in the Frejus laboratory is taking place.

The priority is now the preparation of the Design Study of a high performance Neutrino Facility, to be completed by the end of the decade. BENE is now engaged in the preparatory phase of this FP7 Proposal. To this effect, an International Scoping Study (ISS) was launched at the NuFact05 workshop, with the aim of establishing a strong international collaboration of experts. Physics studies, detector development and accelerator aspects are considered, and should deliver the program for the design study. The aim of the design study is to be able to propose a choice and a full conceptual design. The increased and coordinated contributions that will be required by European laboratories and funding agencies are outlined.

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Introduction: towards a precision neutrino facility

This is a great paradox of particle physics at the turn of the 21st century: the Standard Model (SM) met triumph with the precision measurements at LEP and SLD, the last missing quark was discovered at Fermilab, the quark mixing scheme was confirmed in a splendid manner at the B factories. At the same time, the observation of neutrino oscillations, demonstrating that neutrinos have mass and mix, gave the first direct signal of physics beyond the SM.

Neutrino masses could in principle be incorporated in a trivial extension of the SM, but this would require i) the addition of a new conservation law that is not now present in the SM, fermion number conservation, and ii) the introduction of an extraordinarily small Yukawa coupling for neutrinos, of the order of $m_\nu/m_{\text{top}} \cong 10^{-12}$. More natural theoretical interpretations, such as the see-saw mechanism, lead to the consequence that neutrinos are their own anti-particles, and that the smallness of the neutrino masses comes from their mixing with very heavy partners at the mass scale of Grand Unification (GUT). For the first time, solid experimental facts open a possible window of observation on physics at the GUT scale.

There are many experimental and fundamental implications of this discovery. Perhaps the most spectacular one is the possibility that the combination of fermion number violation and CP violations in the neutrino system could, via Leptogenesis, provide an explanation for the baryon asymmetry of the Universe.

The experimental implications are not less exciting. Fermion number violation, and the absolute mass scale of light neutrinos, should be testable in neutrinoless double beta decay. The direct measurement of the average mass of electron-neutrinos in beta decay could lead to an observable result. The precise values of mass differences, the ordering of masses and the determination of mixing angles must be accessed by neutrino oscillation experiments. Last but not least the discovery of CP or T violation in neutrino oscillations appears to be feasible, but it requires a new type of experimentation: precision appearance neutrino oscillation measurements involving electron-neutrinos. Precision neutrino oscillation experiments, and the CP asymmetry search in particular, require accelerator based neutrino beams, on which we concentrate in the following.

There is presently a vast experimental programme around the world that has begun this investigation, as described in the physics section of this report. The Fermilab-based NUMI beam turned on in 2005 with first results from the MINOS experiment expected in the course of 2006. The CERN based CNGS will start in 2006 with the goal of observing the tau neutrino appearance. The reactor experiment Double CHOOZ is expected to start in 2007/8. The first experiment dedicated to the search for the so far unobserved $\nu_e \leftrightarrow \nu_\mu$ oscillation, T2K, will begin in 2009, followed possibly by an off-axis upgrade of the NUMI experimentation, called NovA. The $\nu_e \leftrightarrow \nu_\mu$ (or $\nu_e \leftrightarrow \nu_\tau$) transition is remarkable in that it is driven both by the solar oscillation and by the atmospheric oscillation. The latter term is proportional to the yet unknown mixing angle θ_{13} . These experiments may have a chance to establish the neutrino mass hierarchy and should reach a sensitivity to θ_{13} of about 0.01 by 2010-2012. This information will be crucial in the final design of the wave of experimentation that will follow, and that will set its aim on the discovery of CP violation.

It is not too early to develop the techniques that will allow Europe to play a major role in the next generation of experiments, and be able to propose hosting a major neutrino facility. The combined requirements on flux intensity, purity and quality, combined with the requirements on detector mass and capabilities, are very challenging and require novel ideas.

There is growing confidence that novel neutrino beams can be produced from proton accelerators, overcoming the intrinsic limitations of conventional neutrino beams. Neutrino parents can be fully

selected, collimated, accelerated and stored in a decay ring. This can be done with muons in a Neutrino Factory and with radioactive ions in a Betabeam. These promise greatly superior performance with respect to conventional neutrino beams, based on a pion decay tunnel, even if the significant intensity upgrades presently envisaged will turn some of them into first generation "Superbeams".

Studies towards novel beams have taken place in Europe now for seven years and R&D projects (HARP, MUSCAT, HIPPI, MERIT, MICE and more) are being carried out by world wide collaborations. BENE has been approved in 2004 and a first Betabeam design started 2005 within the Eurisol Design Study. The fact that these studies are pursued despite the scarcity of resources available at the time of construction of LHC tells a great deal about the motivation and determination of the people involved.

The comparison, as of today, of the CP discovery potential of Superbeams, Betabeams and Neutrino Factories, in combination with appropriate far neutrino detectors (Fig 17 of next section), somewhat favours the Neutrino Factory for small values of θ_{13} , and points to the sensitivity to systematic errors for large values of θ_{13} . This conclusion cannot be considered final. Cost, timescale, performance and practicality of different accelerator systems, performance and optimisation of detectors, choice of optimal baselines, systematic errors and more must be again addressed systematically and coherently by a complete Design Study, leading to a choice by the beginning of next decade.

The challenge of very large neutrino detectors is being posed on a yet larger scale and the Design Study will have to be in close contact with existing R&D efforts and to promote others efforts if necessary. Water Cherenkov, liquid argon and possibly liquid scintillator tanks fit well a Betabeam/Superbeam. Magnetized segmented calorimeters and emulsions cloud chambers are for a Neutrino Factory, where magnetized liquid argon could also be used to great benefit.

The Design Study should cover also the opportunities offered by a new neutrino facility in other sectors of particle physics like search for forbidden transitions of slow muons or physics of neutrino interactions with near detector stations.

Turning to accelerator aspects, a driver with several MegaWatt of proton beam power is the backbone of any future high intensity facility, both for neutrino and other applications. Preliminary studies indicate that the best choice of the proton kinetic energy ranges from 3-5 GeV for a CERN-Frejus Superbeam, 5-30 GeV for a Neutrino Factory. HARP data on pion production are now becoming available and will finally settle these values.

R&D for the first stage (180 MeV or so) of a future high power driver is being carried out in the HIPPI project. For the higher energy sections, there is a second conceptual study of a 4 MW Superconducting Proton Linac in progress at CERN, extending the proton energy to 3.5 GeV from the 2.2 GeV of the first study. A 8 GeV Linac is being designed at Fermilab. Design work of multi MW Rapid Cycling Synchrotrons rings of 5, 15 GeV and more has been done at RAL. Upgrade of the power of the JPARC 50 GeV synchrotron is being vigorously pursued. Finally, FFAGs are now also being considered for high power proton beams.

Flexibility and multiple options will prove very important as the choice of the optimal power, energy and technology of the proton driver will be taken in a larger context, with special emphasis on LHC performance and attention to other applications.

The design of a pion production target and collector, operating in a multi megawatt environment with a long enough lifetime, and of the safety and radiation containment station around it, is probably the most difficult challenge in this sector. No established solution exists for the large energy density involved, 300-1000 J cm⁻³. The MERIT experiment will run at CERN in 2007 a proof of principle test of a liquid metal (mercury) jet at full design flow, 15T magnetic field and pulse power commensurate

to that of a neutrino factory or superbeam. Fundamental studies of a variety of problems for neutrino and EURISOL targets at CERN, solid target studies at BNL on low thermal expansion materials at BNL and tantalum target studies at 2000 K at RAL all deserve strong support. More work is necessary on the design of solenoid collector beyond that of US study II and on horn collectors beyond the CNGS design.

In the US and CERN schemes for a neutrino factory, a “front-end” prepares the muon beam for acceleration. Energy spread is reduced by phase rotation and emittance by ionization cooling, that results from the combination of energy loss in low Z material and reacceleration by RF. The MICE experiment at RAL is being built to prove the feasibility of ionisation cooling and study its properties extensively. The main priorities for the Design Study in this sector are to complete the MICE experimental program its second phase and to revive design work at CERN.

Muon acceleration studies have led to the cost effective FFAG option, rediscovered in Japan in the last few years. Mastering this technique in Europe seems essential. An alternative type of FFAG, the so-called non-scaling FFAG, has recently been invented and shown to have a number of advantages. A proof-of-principle machine should show that non-scaling optics work and plans are being made to build an electron model (EMMA) at the Daresbury Laboratory in the UK to do this. Optimisation studies of a muon storage ring are also necessary.

A design study has started with the objective of establishing feasibility of a $\gamma=100$ Betabeam at CERN, with a frozen parameter list, re-using a maximum of the existing CERN accelerator infrastructure. The technical challenges being addressed are sufficient production of ^{18}Ne and ^6He , ionization and bunching of the produced ions at unprecedented intensities, collimation and magnet protection, space charge and physical aperture limitations, the small duty cycle required for rejection of atmospheric background. A low energy accumulator ring has been added to the design so to increase design flexibility and boost ion intensities. The study of much higher intensity and γ betabeams, promising enhanced physics reach, is not included. It requires additional resources and should be done before next neutrino facility is chosen.

In summary, although the content of the Design Study will be defined by the on-going scoping study, it appears likely to include, in the accelerator sector, the conceptual feasibility study of the following components:

- a high power proton driver with energy up to about 4-5 GeV or more
- the engineering of the handling, containment and safety aspects of a high power target and collection station
- a performing and cost effective muon phase rotation and cooling channel
- non-scaling FFAGs for acceleration of muons (and possibly protons)
- an optimised storage ring for muons
- higher gamma and higher intensity beta beams

and a number of technical preparatory (R&D) projects, aiming at demonstrating:

- the existence of at least one adequate choice of target
- an extended lifetime of the horn prototypes at high rate and radiation
- muon ionization cooling, by completion of the MICE experimental program
- operation of a non-scaling FFAG
- RF cavities and kicker magnets for fast manipulation of muon beams

At the same time it will require specific detector R&D and design efforts on a number of topics:

- photo detector development for very large far detectors

- developments of the liquid argon technique including the presence of magnetic field
- study and tests of a magnetic calorimeter susceptible to be built with a mass of up to 100 kton.
- detectors dedicated for tau detection such as the emulsion cloud chamber
- last but not least the necessary near detector concepts and beam instrumentation that are crucial for the precise monitoring of neutrino and antineutrino event rates needed for CP violation measurements

The above structure reproduces that of existing working groups that are enthusiastic to realize this programme. A strong support from CERN and other funding agencies will be crucial for the success of this enterprise.

Physics of massive neutrinos and the role of accelerator experiments

1 Current status of neutrino masses and oscillations

1.1 Neutrino Masses

1.1.1 Laboratory limits

Direct laboratory limits on neutrino masses are obtained from kinematical studies. The most stringent current upper limit is that on the $\bar{\nu}_e$ mass, coming from the Mainz experiment measuring the end-point of the electron energy spectrum in Tritium beta decay [1]

$$m_{\bar{\nu}_e} \leq 2.2 \text{ eV (95\%CL)}$$

The Troitsk group has also published a similar limit [2]:

$$m_{\nu_e} \leq 2.1 \text{ eV (95\%CL)}$$

however they must include an ad-hoc step function near the endpoint to avoid the problem of negative mass squared.

The proposed KATRIN experiment aims to improve the sensitivity to $m_{\bar{\nu}_e} \sim 0.3 \text{ eV}$ [3]. Similar sensitivities are the goal of the longer term MARE experiment [4] based on an array of several thousand of microbolometers. These measurements are sensitive to:

$$m_{\bar{\nu}_e} = \left(\sum_i |U_{ei}^2| m_i^2 \right)^{1/2} = \left(\cos^2 \theta_{13} (m_1^2 \cos^2 \theta_{12} + m_2^2 \sin^2 \theta_{12}) + m_3^2 \sin^2 \theta_{13} \right)^{1/2} \quad (1)$$

An important constraint on Majorana neutrino masses arises from neutrinoless double- β decay, in which an (A, Z) nucleus decays to $(A, Z + 2) + 2 e^-$, without any neutrino emission. This process can be used to constrain the combination

$$|m_{\beta\beta}| = \left| \sum_i U_{ei}^{*2} m_i \right| = \left| \cos^2 \theta_{13} (m_1 \cos^2 \theta_{12} + m_2 e^{2i\alpha} \sin^2 \theta_{12}) + m_3 e^{2i\beta} \sin^2 \theta_{13} \right| \quad (2)$$

which involves a coherent sum over all the different Majorana neutrino masses m_i , weighted by their mixings with the electron flavour eigenstate, which may include CP-violating phases, as discussed below. This observable is therefore distinct from the quantity observed in Tritium β decay.

The interpretation of neutrinoless double- β decay data depends on calculations of the nuclear matrix elements entering in this process.

A claim for a neutrinoless double- β signal has been made by [5] analyzing the Heidelberg-Moscow data on ^{76}Ge :

$$T_{1/2}^{0\nu} = 1.19 \cdot 10^{25} \text{ years}$$

corresponding to

$$\langle m_{\beta\beta} \rangle = 0.05 - 0.85 \text{ eV (95\%CL)}$$

the uncertainty coming from the choice of the nuclear matrix element calculation.

This result is in contrast with the limit computed with a combined analysis of a subset of the Heidelberg-Moscow data and IGEX experiments [6] and to what reported by a separate group of the original collaboration [7], reporting no evidence for a signal.

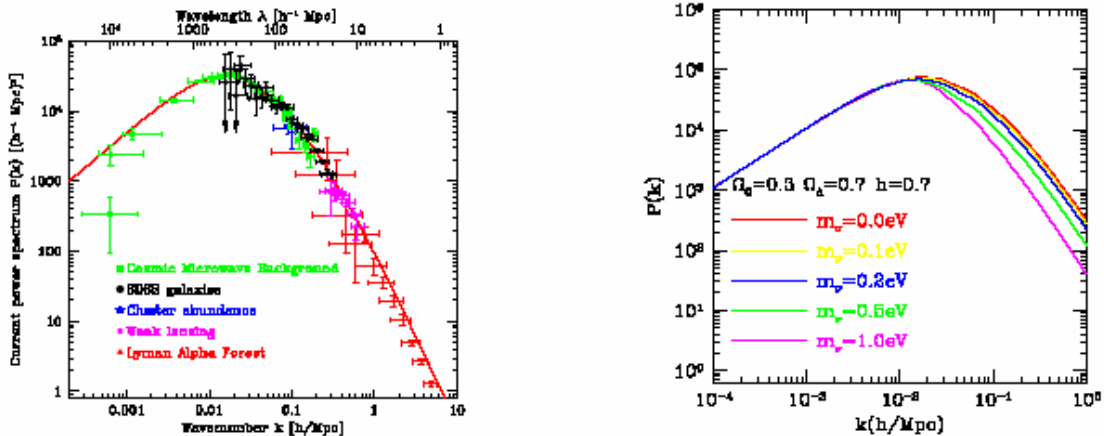


Figure 1: *Left: observational power spectrum from CMB fluctuations, galaxy clustering, gravitational lensing shear, cluster abundances and Lyman α clouds, compared with the prediction of Λ CDM model. The figure is taken from the SDSS paper, Tegmark et al. [22]. Right: Power spectrum predicted in the Λ CDM model including massive neutrinos with a degenerate mass specified in the legend. The top curve represents the case of massless neutrinos*

where $h = 0.71$ is the conventional notation for the Hubble constant.

A successful model is obtained for cosmic structure formation without having massive neutrinos. This means that massive neutrinos only disturb the agreement between theory and observations, hence leading to a limit on the neutrino mass.

The well-known effect of massive neutrinos is relativistic free streaming that damps fluctuations within the horizon scale. One electron volt neutrinos are still relativistic at matter-radiation equality, which takes place at $T \approx 1 \text{ eV}$, and then tend to smear fluctuations up to $\sim 100 \text{ Mpc}$ comoving scale, thus diminishing the power of $P(k)$ for these scales. This effect becomes stronger as the cosmological mass density of neutrinos, hence the neutrino mass, increases (see Fig. 1, right) (see [24, 25]). Therefore, the empirical knowledge of $P(k)$ across large to small scales gives a constraint on the summed mass of neutrinos:

(1) Limits derived from galaxy clustering varies from $\sum m_\nu < 0.6 \text{ eV}$ [19] to 2.1 eV [26] at a 95% confidence, both authors using 2dFGRS data (see also [27]). The limit using the SDSS data is $\sum m_\nu < 1.7 \text{ eV}$ [20]. (See also [28], where $\sum m_\nu < 0.75 \text{ eV}$ is concluded using SDSS and 2dFGRS.)

(2) With the Lyman α cloud absorption power derived from fluctuating optical depths, one can explore $P(k)$ at the smallest scale [29, 30], giving the strongest limit $\sum m_\nu < 0.42 \text{ eV}$ (95%) [31] (see [32] for the earlier work). This result, however, is more model dependent in the sense that one must invoke simulations to extract $P(k)$ from observed flux power spectrum, which suffer from significant uncertainties associated with modelling and simulations.

(3) With gravitational lensing one can directly explore mass fluctuations. For the moment the statistical error is not sufficiently small, but this provides us with a promising method. This may also be used to set the normalisation of the power spectrum derived from galaxy clustering.

Elgarøy and Lahav [33] give a summary of limits obtained in the literature.

1.3 Neutrino Oscillations

The observation of neutrino oscillations has now established beyond doubt that neutrinos have mass and mix. This existence of neutrino masses is in fact the first solid experimental fact requiring physics beyond the Standard Model.

The present status of the field is as follows. Since 1989, we know from LEP [34] that there are only three families of active light neutrinos coupling to the weak interaction.

Since the early 1970's we have hints from solar neutrino experiments that electron neutrinos produced in the sun undergo disappearance on their way to earth, and, from the different disappearance rates for experiments sensitive to different neutrino energy ranges (Chlorine [35], Gallium [36, 37], Water Čerenkov [38]) we have indication that matter effects in the sun play an important role. This "solar neutrino puzzle" was closed in 2002 with the results from the SNO experiment [39], which allowed simultaneously i) a direct measurement of the total flux of active neutrinos by the neutral current reaction in agreement with solar neutrino flux calculations, and ii) a measurement of the electron neutrino component, by the charged current reaction, determining that they represent less than half of the total flux, in agreement with previous observations. The KamLAND experiment [40] provided at the same time a measurement of disappearance of electron anti-neutrinos from nuclear fission reactors in Japan and Korea, providing, in combination with the solar neutrino results, a precise determination of the relevant neutrino mixing angle and of the corresponding mass difference, see Table 1.

Since the late 80's there has been indication from atmospheric neutrino experiments that the muon neutrinos undergo disappearance when going through the earth; this was finally unambiguously demonstrated by the Super-Kamiokande experiment in 1998 [41], a result well supported by Soudan2 [42] and Macro [43] experiments. This disappearance takes place at a much shorter quarter-wavelength than for solar neutrinos ($L/E \sim 500 \text{ km/GeV}$) [44]; it is not seen for electron neutrinos, a fact that has been best established by the CHOOZ reactor experiment [45]. This disappearance has been confirmed by the K2K experiment in Japan [46], the first accelerator neutrino long baseline experiment designed since neutrino masses have been established, and a prototype for future ones.

The above experimental observations are consistently described by three families ν_1, ν_2, ν_3 with mass values m_1, m_2 and m_3 that are connected to the flavor eigenstates ν_e, ν_μ and ν_τ by a mixing matrix U , usually parameterized as

$$U(\theta_{12}, \theta_{23}, \theta_{13}, \delta_{\text{CP}}) = \begin{pmatrix} c_{13}c_{12} & c_{13}s_{12} & s_{13}e^{-i\delta_{\text{CP}}} \\ -c_{23}s_{12} - s_{13}s_{23}c_{12}e^{i\delta_{\text{CP}}} & c_{23}c_{12} - s_{13}s_{23}s_{12}e^{i\delta_{\text{CP}}} & c_{13}s_{23} \\ s_{23}s_{12} - s_{13}c_{23}c_{12}e^{i\delta_{\text{CP}}} & -s_{23}c_{12} - s_{13}c_{23}s_{12}e^{i\delta_{\text{CP}}} & c_{13}c_{13} \end{pmatrix} \quad (6)$$

where the short-form notation $s_{ij} \equiv \sin \theta_{ij}, c_{ij} \equiv \cos \theta_{ij}$ is used. As a result, the neutrino oscillation probability depends on 3 mixing angles, $\theta_{12}, \theta_{23}, \theta_{13}$, see Fig. 2, 2 mass differences, $\Delta m_{12}^2 = m_2^2 - m_1^2$, $\Delta m_{23}^2 = m_3^2 - m_2^2$, and a CP phase δ_{CP} . Additional phases are present in case neutrinos are Majorana particles, but they do not influence at all neutrino flavor oscillations. Furthermore, the neutrino mass hierarchy, the ordering with which mass eigenstates are coupled to flavor eigenstates, can be fixed by measuring the sign of Δm_{23}^2 . In vacuum the oscillation probability between two neutrino flavors α, β is:

$$P(\nu_\alpha \rightarrow \nu_\beta) = -4 \sum_{k>j} \text{Re}[W_{\alpha\beta}^{jk}] \sin^2 \frac{\Delta m_{jk}^2 L}{4E_\nu} \pm 2 \sum_{k>j} \text{Im}[W_{\alpha\beta}^{jk}] \sin^2 \frac{\Delta m_{jk}^2 L}{2E_\nu}, \quad (7)$$

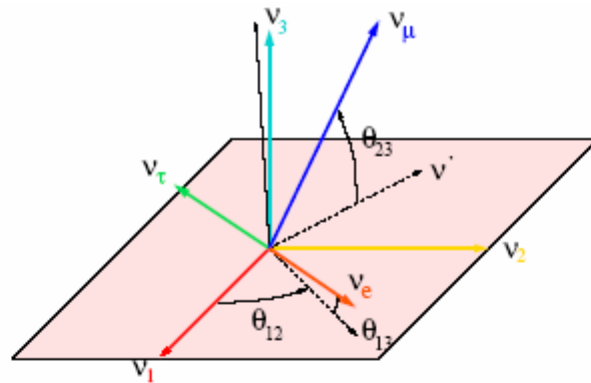


Figure 2: Representation of the 3-dimensional rotation between the flavor and mass neutrino eigenstates

Table 1: Neutrino oscillation parameters as of *NUFACT05* [49]

'solar parameters'	$\delta m^2 = \pm(7.92 \pm 0.72) \cdot 10^{-5} \text{ eV}^2$	$\sin^2 \theta_{12} = 0.314_{-0.025}^{+0.030}$
'atmospheric parameters'	$\Delta m^2 = \pm(2.4_{-0.6}^{+0.5}) \cdot 10^{-3} \text{ eV}^2$	$\sin^2 \theta_{23} = 0.44_{-0.10}^{+0.18}$
'absolute mass'	$m_\nu \leq 2.2 \text{ eV}$ (tritium decay) $\sum m_\nu \leq \mathcal{O}(1\text{eV})$ (cosmology)	

where $\alpha = e, \mu, \tau$, $j = 1, 2, 3$, $W_{\alpha\beta}^{jk} = U_{\alpha j} U_{\beta j}^* U_{\alpha k}^* U_{\beta k}$. In the case of only two neutrino flavor oscillation it can be written as:

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \cdot \sin^2 \frac{1.27 \Delta m^2 (\text{eV}^2) \cdot L (\text{km})}{E_\nu (\text{GeV})}. \quad (8)$$

When neutrinos pass through matter, the oscillation probability is perturbed [47]. Conventionally $\Delta m_{12}^2 \equiv m_2^2 - m_1^2 > 0$ and $\theta_{12} > 0$ [48]. Since atmospheric neutrino disappearance has only been observed for muon neutrinos, which couple weakly to electron neutrinos at the relevant wavelength, we cannot (yet) tell the sign of the mass difference Δm_{13}^2 , or Δm_{23}^2 .

The present values of oscillation parameters are summarized in Table 1.

Before leaving this section on the status of the field, it is worth remembering a further indication of $\nu_\mu \rightarrow \nu_e$ oscillations with a Δm^2 of $0.3 - 20 \text{ eV}^2$ which comes from the beam dump LSND experiment detecting a 4σ excess of $\bar{\nu}_e$ interactions in a neutrino beam produced by π^+ decays at rest where the $\bar{\nu}_e$ component is highly suppressed ($\sim 7.8 \cdot 10^{-4}$) [50]. The KARMEN experiment [51], with a very similar technique but with a lower sensitivity (a factor 10 less for the lower Δm^2), and the NOMAD experiment at WANF of CERN SPS [52] (for $\Delta m^2 > 10 \text{ eV}^2$) do not confirm the result, excluding a large part of the allowed region of the oscillation parameters. The LSND result doesn't fit the overall picture of neutrino oscillations and several non-standard explanations, as for instance sterile neutrinos, have been put forward to solve this experimental conflict. The MiniBooNE experiment at FNAL, presently taking data, is designed to settle this puzzle with a 5σ sensitivity [53]. If it will become true this will lead to even more exciting phenomenology.

1.3.1 Present generation of long-baseline experiments ²

Over the next five years the present generation of oscillation experiments at accelerators with long-baseline ν_μ beams (Table 2), K2K at KEK [46], MINOS [55] at the NuMI beam from FNAL [56] and ICARUS [57] and OPERA [58] at the CNGS beam from CERN [59] are expected to confirm the atmospheric evidence of oscillations and measure $\sin^2 2\theta_{23}$ and $|\Delta m_{23}^2|$ within 10 ÷ 15 % of accuracy if $|\Delta m_{23}^2| > 10^{-3} \text{ eV}^2$. K2K and MINOS are looking for neutrino disappearance, by measuring the ν_μ survival probability as a function of neutrino energy while ICARUS and OPERA will search for evidence of ν_τ interactions in a ν_μ beam, the final proof of $\nu_\mu \rightarrow \nu_\tau$ oscillations. K2K has already completed its data taking at the end of 2004, while MINOS has started data taking beginning 2005. CNGS is expected to start operations in the second half of 2006.

Table 2: *Main parameters for present long-baseline neutrino beams*

Neutrino facility	Proton momentum (GeV/c)	L (km)	E_ν (GeV)	pot/yr (10^{19})
KEK PS	12	250	1.5	2
FNAL NuMI	120	735	3	20 ÷ 34
CERN CNGS	400	732	17.4	4.5 ÷ 7.6

In all these facilities conventional muon neutrino beams are produced through the decay of π and K mesons generated by a high energy proton beam hitting needle-shaped light targets. Positive (negative) mesons are sign-selected and focused (defocused) by large acceptance magnetic lenses into a long evacuated decay tunnel where ν_μ 's ($\bar{\nu}_\mu$'s) are generated. In case of positive charge selection, the ν_μ beam has typically a contamination of $\bar{\nu}_\mu$ at few percent level (from the decay of the residual π^- , K^- and K^0) and $\sim 1\%$ of ν_e and $\bar{\nu}_e$ coming from three-body K^\pm , K_0 decays and μ decays. The precision on the evaluation of the intrinsic ν_e to ν_μ contamination is limited by the knowledge of the π and K production in the primary proton beam target. Hadroproduction measurements at 400 and 450 GeV/c performed with the NA20 [60] and SPY [61] experiments at the CERN SPS provided results with 5 ÷ 7% intrinsic systematic uncertainties.

The CNGS ν_μ beam has been optimized for the $\nu_\mu \rightarrow \nu_\tau$ appearance search. The beam-line design was accomplished on the basis of the previous experience with the WANF beam at CERN SPS [62]. The expected muon neutrino flux at the Gran Sasso site will have an average energy of 17.4 GeV and $\sim 0.6\%$ ν_e contamination for $E_\nu < 40$ GeV. Due to the long-baseline (L=732 Km) the contribution to neutrino beam from the K^0 and mesons produced in the reinteraction processes will be strongly reduced with respect to the WANF [63]: the ν_e/ν_μ ratio is expected to be known within $\sim 3\%$ systematic uncertainty [64].

Current long-baseline experiments with conventional neutrino beams can look for $\nu_\mu \rightarrow \nu_e$ even if they are not optimized for θ_{13} studies. MINOS at NuMI is expected to reach a sensitivity of $\sin^2 2\theta_{13} = 0.08$ [55] integrating $14 \cdot 10^{20}$ protons on target (pot) in 5 years according to the FNAL proton plan evolution [65]. MINOS main limitation is the poor electron identification efficiency of the detector. OPERA [58] can reach a 90% C.L. sensitivity $\sin^2 2\theta_{13} = 0.06$ ($\Delta m_{23}^2 = 2.5 \cdot 10^{-3} \text{ eV}^2$, convoluted to CP and matter effects) [66, 67], a factor ~ 2 better than Chooz for five years exposure to the CNGS beam at nominal intensity for shared operation $4.5 \cdot 10^{19}$ pot/yr. A plot of θ_{13} sensitivities is reported in Fig. 3. According to the CERN PS and SPS upgrade studies [68], the CNGS beam intensity could be improved by a factor 1.5, allowing for more sensitive neutrino oscillation searches for the OPERA experiment.

²Material for this Section is mainly taken from ref. [54]

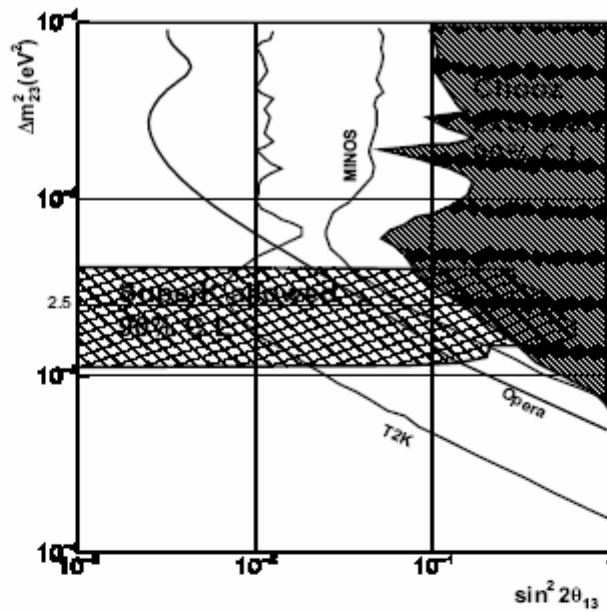


Figure 3: *Expected sensitivity on θ_{13} mixing angle (matter effects and CP violation effects not included) for MINOS, OPERA and for the next T2K experiment, compared to the Chooz exclusion plot.*

It is worth mentioning that the sensitivity on θ_{13} measurement of the current long-baseline experiments with conventional neutrino beams, like NuMI and CNGS, will be limited by the power of the proton source which determines the neutrino flux and the event statistics, by the not optimized L/E_ν and by the presence of the ν_e intrinsic beam contamination and its related systematics. This is particular true for CNGS where the neutrino energy, optimized to overcome the kinematic threshold for τ production and to detect the τ decay products, is about ten times higher the optimal value for θ_{13} searches at that baseline.

Another approach to search for non vanishing θ_{13} is to look at $\bar{\nu}_e$ disappearance using nuclear reactors as neutrino source. A follow-up of Chooz, Double Chooz [69], has been proposed to start in 2008 with a two detectors setup, aiming to push systematic errors down to 0.6 % and to reach a sensitivity on $\sin^2 2\theta_{13} \simeq 0.024$ (90% C.L., $\Delta m_{23}^2 = 2.5 \cdot 10^{-3}$) in a 3 years run.

A sketch of θ_{13} sensitivities as a function of the time, following the schedule reported in the experimental proposals, computed for the approved experiments, is reported in Fig. 4.

1.3.2 Three families oscillations and CP or T violation

It was soon realized that with three families and for a favorable set of parameters, it would be possible to observe violation of CP or T symmetries in neutrino oscillations [70]. This observation reinforced the considerable interest for precision measurements of neutrino oscillation parameters. We know since 2002 and the results from SNO [39] and KamLAND [40] that the neutrino parameters belong to the so-called LMA solution which suggests that leptonic CP violation could be large enough to be observed in high-energy neutrino oscillation appearance experiments.

This has led to extensive studies, such as those published recently in the CERN yellow report[71], or in a recent BENE [72] workshop on physics at a high intensity proton driver [73]. The phenomenon of CP (or T) violation in neutrino oscillations manifests itself by a difference in the oscillation probabilities of say, $P(\nu_\mu \rightarrow \nu_e)$ vs $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ (CP violation), or $P(\nu_\mu \rightarrow \nu_e)$ vs $P(\nu_e \rightarrow \nu_\mu)$ (T

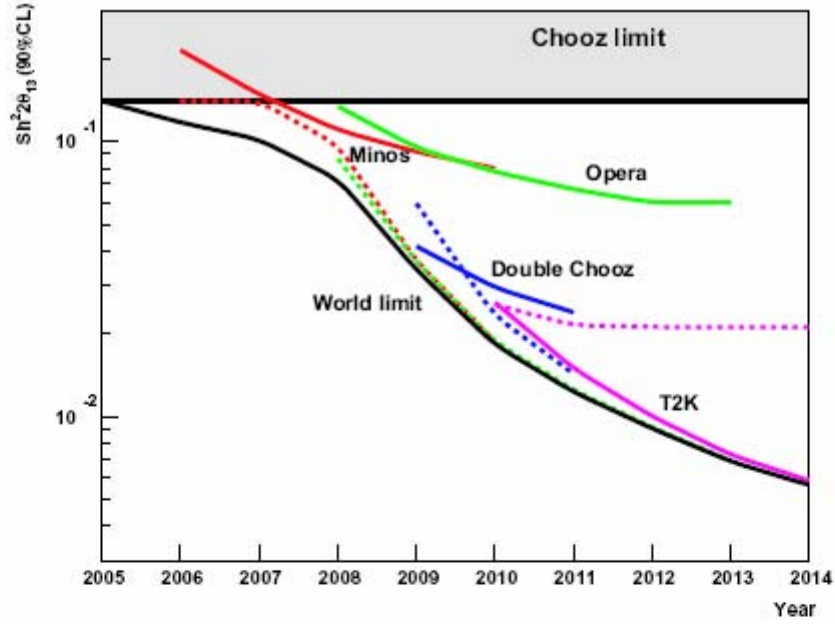


Figure 4: Evolution of sensitivities on $\sin^2 2\theta_{13}$ as function of time. For each experiment are displayed the sensitivity as function of time (solid line) and the world sensitivity computed without the experiment (dashed line). The comparison of the two curves shows the discovery potential of the experiment along its data taking. The world overall sensitivity along the time is also displayed. The comparison of the overall world sensitivity with the world sensitivity computed without a single experiment shows the impact of the results of the single experiment. Experiments are assumed to provide results after the first year of data taking.

violation).

It can be observed right away that observation of this important phenomenon requires appearance experiments; indeed a reactor or solar neutrino experiment, sensitive to the disappearance $P(\nu_e \rightarrow \nu_e)$ which is clearly time-reversal invariant, would be completely insensitive to it. This can be seen as an advantage in view of a precise and unambiguous measurement of the mixing angles; for the long term goal of observing and studying CP violation, we are confined to appearance experiments.

The $\nu_\mu \rightarrow \nu_e$ transition probability in case of small matter effects can be parameterized as [74]:

$$\begin{aligned}
P(\nu_\mu \rightarrow \nu_e) &= 4c_{13}^2 s_{13}^2 s_{23}^2 \sin^2 \frac{\Delta m_{13}^2 L}{4E_\nu} \times \left[1 \pm \frac{2a}{\Delta m_{13}^2} (1 - 2s_{13}^2) \right] \\
&+ 8c_{13}^2 s_{12} s_{13} s_{23} (c_{12} c_{23} \cos \delta_{\text{CP}} - s_{12} s_{13} s_{23}) \cos \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} \\
&\mp 8c_{13}^2 c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta_{\text{CP}} \sin \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} \\
&+ 4s_{12}^2 c_{13}^2 \{ c_{13}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{23} s_{13} \cos \delta_{\text{CP}} \} \sin \frac{\Delta m_{12}^2 L}{4E_\nu} \\
&\mp 8c_{12}^2 s_{13}^2 s_{23}^2 \cos \frac{\Delta m_{23}^2 L}{4E_\nu} \sin \frac{\Delta m_{13}^2 L}{4E_\nu} \frac{aL}{4E_\nu} (1 - 2s_{13}^2).
\end{aligned} \tag{9}$$

The first line of this parameterization contains the term driven by θ_{13} , the second and third contain CP even and odd terms respectively, and the fourth is driven by the solar parameters. The last

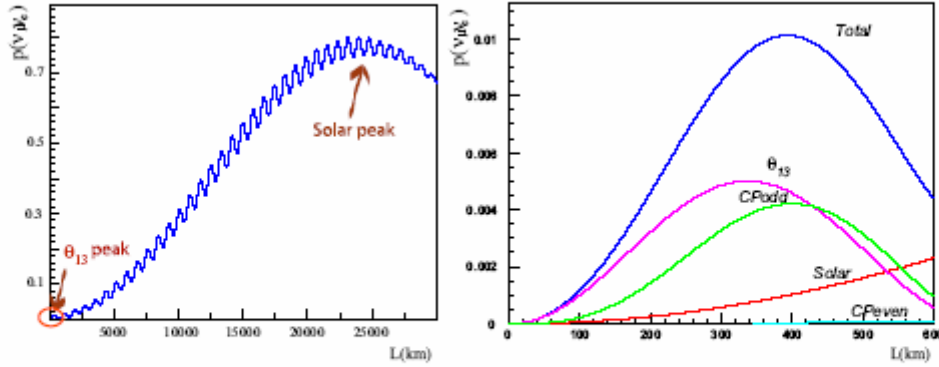


Figure 5: Sketch of $P(\nu_\mu \rightarrow \nu_e)$ as function of the baseline computed for monochromatic neutrinos of 1 GeV in the solar baseline regime for $\delta_{\text{CP}} = 0$ (left) and in the atmospheric baseline regime for $\delta_{\text{CP}} = -\pi/2$ (right), where the different terms of Eq. (9) are displayed. The following oscillation parameters were used in both cases: $\sin^2 2\theta_{13} = 0.01$, $\sin^2 2\theta_{12} = 0.8$, $\Delta m_{23}^2 = 2.5 \cdot 10^{-3} \text{ eV}^2$, $\Delta m_{12}^2 = 7 \cdot 10^{-5} \text{ eV}^2$. From ref. [54]

line parameterizes matter effects developed at the first order where $a[\text{eV}^2] = \pm 2\sqrt{2}G_F n_e E_\nu = 7.6 \cdot 10^{-5} \rho[\text{g/cm}^3] E_\nu[\text{GeV}]$. The \pm and \mp terms refer to neutrinos and antineutrinos. A sketch of $P(\nu_\mu \rightarrow \nu_e)$ as a function of L for 1 GeV neutrinos is shown in Fig. 5.

When matter effects are not negligible, following Eq. (1) of [75] the transition probability $\nu_e \rightarrow \nu_\mu$ ($\bar{\nu}_e \rightarrow \bar{\nu}_\mu$) at second order in perturbation theory in θ_{13} , $\Delta m_{12}^2/\Delta m_{23}^2$, $|\Delta m_{12}^2/a|$ and $\Delta m_{12}^2 L/E_\nu$ (see also [76]) is:

$$P^\pm(\nu_e \rightarrow \nu_\mu) = X_\pm \sin^2(2\theta_{13}) + Y_\pm \cos(\theta_{13}) \sin(2\theta_{13}) \cos\left(\pm\delta - \frac{\Delta_{23}L}{4E_\nu}\right) + Z, \quad (10)$$

where \pm refers to neutrinos and anti neutrinos, respectively. The coefficients of the two equations are:

$$\begin{cases} X_\pm &= \sin^2(\theta_{23}) \left(\frac{\Delta m_{23}^2}{|a - \Delta m_{23}^2|}\right)^2 \sin^2\left(\frac{|a - \Delta m_{23}^2|L}{4E_\nu}\right), \\ Y_\pm &= \sin(2\theta_{12}) \sin(2\theta_{23}) \left(\frac{\Delta m_{12}^2}{a}\right) \left(\frac{\Delta m_{23}^2}{|a - \Delta m_{23}^2|}\right) \sin\left(\frac{aL}{4E_\nu}\right) \sin\left(\frac{|a - \Delta m_{23}^2|L}{4E_\nu}\right), \\ Z &= \cos^2(\theta_{23}) \sin^2(2\theta_{12}) \left(\frac{\Delta m_{12}^2}{a}\right)^2 \sin^2\left(\frac{aL}{4E_\nu}\right) \end{cases} \quad (11)$$

(remember that a changes sign by changing neutrinos with antineutrinos and that $P(\nu_e \rightarrow \nu_\mu) = -P(\bar{\nu}_e \rightarrow \bar{\nu}_\mu)$).

θ_{13} searches look for experimental evidence of ν_e appearance in excess to what expected from the solar terms. These measurements will be experimentally hard because the Chooz limit on the $\bar{\nu}_e$ disappearance, $\theta_{13} < 11^\circ$ for $\Delta m_{23}^2 \simeq 2.5 \cdot 10^{-3} \text{ eV}^2$, translates into a $\nu_\mu \rightarrow \nu_e$ appearance probability less than 10% at the appearance maximum in a high energy muon neutrino beam.

One of the interesting aspects of this formula is the occurrence of matter effects which, unlike the straightforward θ_{13} term, depends on the sign of the mass difference $\text{sign}(\Delta m_{23}^2)$. These terms should

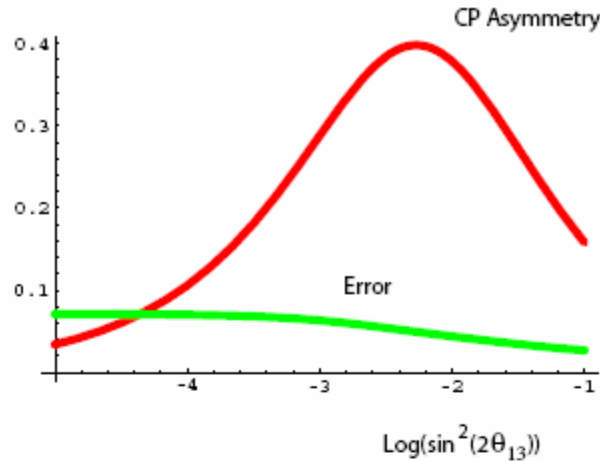


Figure 6: *Magnitude of the CP asymmetry at the first oscillation maximum, for $\delta = 1$ as a function of the mixing angle $\sin^2 2\theta_{13}$. The curve marked 'error' indicates the dependence of the statistical+systematic error on such a measurement. The curves have been computed for the baseline Beta Beam option at the fixed energy $E_\nu = 0.4$ GeV, $L=130$ km, statistical + 2% systematic errors.*

allow extraction of the mass hierarchy, but could also be seen as a background to the CP violating effect, from which they can be distinguished by the very different neutrino energy dependence, matter effects being larger for higher energies, with a 'matter resonance' at about 12 GeV. The CP violation can be seen as interference between the solar and atmospheric oscillation for the same transition. Of experimental interest is the CP-violating asymmetry A_{CP} :

$$A_{CP} = \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} \quad (12)$$

displayed in Fig. 6, or the equivalent time reversal asymmetry A_T .

The asymmetry can be large and its value increases for decreasing values of θ_{13} up to the value when the two oscillations (solar and atmospheric) are of the same magnitude. The following remarks can be made:

1. The ratio of the asymmetry to the statistical error is fairly independent on θ_{13} for large values of this parameter, which explains the relative flatness of the sensitivity curves.
2. This asymmetry is valid for the first maximum. At the second oscillation maximum the curve is shifted to higher values of θ_{13} so that it could be then an interesting possibility for measuring the CP asymmetry, although the reduction in flux is considerable (roughly factor 9).
3. The asymmetry has opposite sign for $\nu_e \rightarrow \nu_\mu$ and $\nu_e \rightarrow \nu_\tau$, and changes sign when going from one oscillation maximum to the next.
4. The asymmetry is small for large values of θ_{13} placing a challenging emphasis on systematics.

The richness of the $\nu_\mu \rightarrow \nu_e$ transition is also its weakness: it will be very difficult for pioneering experiments to extract all the genuine parameters unambiguously. Correlations are present between θ_{13} and δ_{CP} [75]. Moreover, in absence of information about the sign of Δm_{23}^2 [77, 78] and the approximate $[\theta_{23}, \pi/2 - \theta_{23}]$ symmetry for the atmospheric angle [79], additional clone solutions rise

up. In general, the measurement of $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ will result in eight allowed regions of the parameter space, the so-called eightfold-degeneracy [78].

As already pointed out, the $\nu_\mu \rightarrow \nu_e$ experimental sensitivity with conventional ν_μ beams is limited by an unavoidable ν_e beam contamination of about 1%. The ν_μ to ν_τ oscillations, with E_ν above the τ mass production threshold, generate background due to a significant number of ν_τ charged current interactions where a large fraction of τ 's decay into electrons. Finally, neutral pions in both neutral current or charged current interactions can fake an electron providing also a possible background for the ν_e 's.

Therefore the measurement of θ_{13} mixing angle and the investigation of the leptonic CP violation will require:

- neutrino beams with high performance in terms of intensity, purity and low associated systematics. Event statistics, background rates and systematic errors will play a decisive role in detecting ν_e appearance;
- the use of detectors of unprecedented mass, granularity and resolution. Again event statistics is the main concern, while high detector performances are necessary to keep the event backgrounds (as π^0 from ν_μ neutral current interactions, mis-identified as ν_e events) at low as possible rate;
- ancillary experiments to measure the meson production (for the neutrino beam knowledge), the neutrino cross-sections, the particle identification capability. The optimization of proton driver characteristics and the best possible estimation of the systematic errors will require this kind of dedicated experiments. The HARP hadroproduction experiment at CERN PS [80] took data for primary protons between 3 and 14.5 GeV in 2001 and 2002 with different target materials. These data contribute to the proton driver optimization, the determination of the K2K [81] and MiniBooNE neutrino beam fluxes and to the study of atmospheric neutrino interaction rates.

2 DESCRIPTION OF THE FACILITIES

According to the present experimental situation, conventional neutrino beams can be improved and optimized for the $\nu_\mu \rightarrow \nu_e$ searches. The design of a such new SuperBeam facility for a very high intensity and low energy ν_μ flux will demand:

- a new higher power proton driver, exceeding the megawatt, to deliver more intense proton beams on target;
- a tunable L/E_ν in order to explore the Δm_{23}^2 parameter region as indicated by the previous experiments with neutrino beams and atmospheric neutrinos;
- narrow band beams with $E_\nu \sim 1 \div 2$ GeV;
- a lower intrinsic ν_e beam contamination which can be obtained suppressing the K^+ and K^0 production by the primary proton beam in the target.

An interesting option for the SuperBeams is the possibility to tilt the beam axis a few degrees with respect to the position of the far detector (Off-Axis beams) [82, 83]. According to the two body π -decay kinematics, all the pions above a given momentum produce neutrinos of similar energy at a given angle $\theta \neq 0$ with respect to the direction of parent pion (contrary to the $\theta = 0$ case where the neutrino energy is proportional to the pion momentum). These neutrino beams have several advantages with respect to the corresponding on-axis ones: they are narrower, lower energy and with a smaller ν_e contamination (since ν_e mainly come from three body decays) although the neutrino flux can be significantly smaller.

The intrinsic limitations of conventional neutrino beams are overcome if the neutrino parents can be fully selected, collimated and accelerated to a given energy. This can be attempted within the muon or a beta decaying ion lifetimes. The neutrino beams from their decays would then be pure and perfectly predictable. The first approach brings to the Neutrino Factories [84], the second to the BetaBeams [85]. However, the technical difficulties associated with developing and building these novel conception neutrino beams suggest for the middle term option to improve the conventional beams by new high intensity proton machines, optimizing the beams for the $\nu_\mu \rightarrow \nu_e$ oscillation searches (SuperBeams).

2.1 Off axis SuperBeams: T2K, T2HK and NO ν A

The T2K (Tokai to Kamioka) experiment [82] will aim neutrinos from the Tokai site to the Super-Kamiokande detector 295 km away. The neutrino beam is produced by pion decay from a horn focused beam, with a system of three horns and reflectors. The decay tunnel length (130 m long) is optimized for the decay of 2-8 GeV pions and short enough to minimize the occurrence of muon decays. The neutrino beam is situated at an angle of 2-3 degrees from the direction of the Super-Kamiokande detector, assuring a pion decay peak energy of 0.6 GeV. The beam line is equipped with a set of dedicated on-axis and off-axis detectors at the distance of 280 meters.

The main goals of the experiment are as follows:

1. The highest priority is the search for ν_e appearance to detect sub-leading $\nu_\mu \rightarrow \nu_e$ oscillations. It is expected that the sensitivity of the experiment in a 5 years ν_μ run, will be of the order of $\sin^2 2\theta_{13} \leq 0.006$ [82].

2.2 SPL SuperBeam

In the CERN-SPL SuperBeam project [89, 90, 91] the planned 4MW SPL (Superconducting Proton Linac) would deliver a 2.2 GeV/c proton beam, on a Hg target to generate an intense π^+ (π^-) beam focused by a suitable magnetic horn in a short decay tunnel. As a result an intense ν_μ beam, will be produced mainly via the π -decay, $\pi^+ \rightarrow \nu_\mu \mu^+$ providing a flux $\phi \sim 3.6 \cdot 10^{11} \nu_\mu/\text{year}/\text{m}^2$ at 130 Km of distance, and an average energy of 0.27 GeV. The ν_e contamination from K will be suppressed by threshold effects and the resulting ν_e/ν_μ ratio ($\sim 0.4\%$) will be known within 2% error. The use of a near and far detector (the latter at $L = 130$ Km of distance in the Frejus area [92], see Sec. 2.2.1) will allow for both ν_μ -disappearance and $\nu_\mu \rightarrow \nu_e$ appearance studies. The physics potential of the 2.2 GeV SPL SuperBeam (SPL-SB) with a water Čerenkov far detector with a fiducial mass of 440 kton, has been extensively studied [90].

New developments show that the potential of the SPL-SB potential could be improved by rising the SPL energy to 3.5 GeV [93], to produce more copious secondary mesons and to focus them more efficiently. This seems feasible if status of the art RF cavities would be used in place of the previously foreseen LEP cavities [94].

The focusing system (magnetic horns) originally optimized in the context of a Neutrino Factory [95, 96] has been redesigned considering the specific requirements of a Super Beam. The most important points are the phase spaces that are covered by the two types of horns are different, and that for a Super Beam the pions to be focused should have an energy of the order of $p_\pi(\text{MeV})/3 \approx E_\nu \geq 2L(\text{km})$ to obtain a maximum oscillation probability. In practice, this means that one should collect 800 MeV/c pions to get a mean neutrino energy of 300 MeV, see Fig. 8. At higher beam energy, the kaon rates grow rapidly compared to the pion rates, an experimental confirmation [80] of such numbers would be strongly needed.

In this upgraded configuration neutrino flux could be increased by a factor ~ 3 with respect to the 2.2 GeV configuration, the number of expected ν_μ charged current is about 95 per kton \cdot yr

A sensitivity $\sin^2 2\theta_{13} < 0.8 \cdot 10^{-3}$ is obtained in a 5 years ν_μ plus 5 year $\bar{\nu}_\mu$ run ($\delta = 0$ intrinsic degeneracy accounted for, sign and octant degeneracies not accounted for), allowing to discovery CP violation (at 3σ level) if $\delta_{\text{CP}} \geq 25^\circ$ and $\theta_{13} \geq 1.4^\circ$ [97, 98]. The expected performances are shown in Fig. 10 and 17 along with those of other setups.

2.2.1 The MEMPHYS detector

The MEMPHYS (MEgaton Mass Physics) detector is a Megaton-class water Čerenkov in the straight extrapolation of the well known and robust technique used for the Super-Kamiokande detector. It is designed to be located at Frejus, 130 km from CERN and it is an alternative design of the UNO [99] and Hyper-Kamiokande [82] detectors and shares the same physics case both from non accelerator domain (nucleon decay, SuperNovae neutrino from burst event or from relic explosion, solar and atmospheric neutrinos) and from accelerator (SuperBeam, BetaBeam) domain. For the physics part not covered by this document, this kind of megaton water detector can push the nucleon decay search up to 10^{35} yrs in the $e^+\pi^0$ channel and up to few 10^{34} yrs in the $K^+\nu$ channel, just to cite these benchmark channels. MEMPHYS can register as many as 150,000 events from a SN at 10 kpc from our galaxy and 50 events or so from Andromeda. To detect relic neutrinos from past SuperNovae explosions one can use pure water and get a flux of 250 evts/10 y/500 kton or increase this number by a factor 10 by adding gadolinium salt.

A recent civil engineering pre-study to envisage the possibly of large cavity excavation located

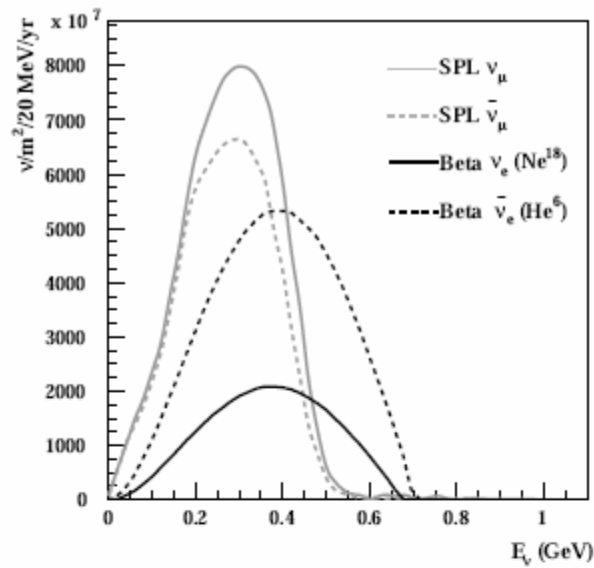


Figure 8: *Neutrino flux of β -Beam ($\gamma = 100$) and CERN-SPL SuperBeam, 3.5 GeV, at 130 Km of distance.*

under the Frejus mountain (4800 m.e.w.) near the present Modane Underground Laboratory has been undertaken. The main result of this pre-study is that MEMPHYS may be built with present techniques as 3 or 4 shafts modular detector, 250000 m³ each with 65 m in diameter, 65 m in height for the total water containment. Each of these shafts corresponds to about 5 times the present Super-Kamiokande cavity. For the present physical study, the fiducial volume of 440 kton which means 3 shafts and an Inner Detector (ID) of 57 m in diameter and 57 m in height. Each ID may be equipped with photodetectors (PMT, HPD,...) with a surface coverage at least 30%. The Frejus site, 4800 m.w.e, offers a natural protection against cosmic rays by a factor 10⁶.

The decision for cavity digging is fixed at 2010 after an intense Detector Design Study (eg. cavity excavation, photodetector R&D) performed in parallel of the digging of at least a Safety Gallery in the Frejus road tunnel. One may note that this key date may also be decisive for SPL construction as well as for the choice of the EURISOL site. After that, the excavation and PMT production are envisaged to take seven years or so, and the non accelerator program can start before the rise up of the accelerator program (SuperBeam and BetaBeam) which may start before 2020.

A first estimate of the costs of such a detector is reported in Table 3

Table 3: *Preliminary cost estimate of the MEMPHYS detector*

3 Shafts	240 ME
Total cost of 250k 12" PMTs	250 ME
Infrastructure	100 ME
Total	590 ME

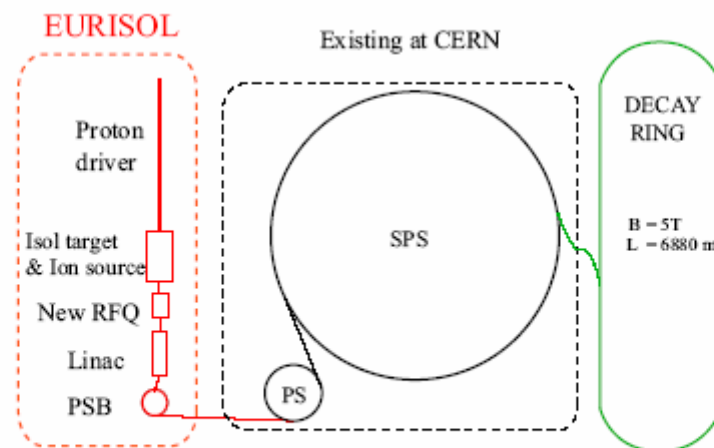


Figure 9: A schematic layout of the BetaBeam complex. At left, the low energy part is largely similar to the EURISOL project [100]. The central part (PS and SPS) uses existing facilities. At right, the decay ring has to be built.

2.3 BetaBeams

BetaBeams have been introduced by P. Zucchelli in 2001 [85]. The idea is to generate pure, well collimated and intense ν_e ($\bar{\nu}_e$) beams by producing, collecting, accelerating radioactive ions and storing them in a decay ring in 10 ns long bunches, to suppress the atmospheric neutrino backgrounds. The resulting BetaBeam fluxes could be easily computed by the properties of the beta decay of the parent ion and by its Lorentz boost factor γ and would be virtually background free. The best ion candidates so far are ^{18}Ne and ^6He for ν_e and $\bar{\nu}_e$ respectively. The schematic layout of a Beta Beam is the following (see also Fig. 9):

Ion production Protons are delivered by a high power Linac. Beta Beam targets need $100\ \mu\text{A}$ proton beam, at energies between 1 and 2 GeV.

In case the Super Proton Linac (SPL) [89] would be used, Beta Beams could be fired to the same detector together with a neutrino SuperBeam [90]. SPL is designed to deliver 2mA of 2.2 GeV (kinetic energy) protons, in such a configuration Beta Beams would use 10% of the total proton intensity, leaving room to a very intense conventional neutrino beam.

The ^6He target consists either of a water cooled tungsten core or of a liquid lead core which works as a proton to neutron converter surrounded by beryllium oxide [101], aiming for 10^{15} fissions per second. ^{18}Ne can be produced by spallation reactions, in this case protons will directly hit a magnesium oxide target. The collection and ionization of the ions is performed using the ECR technique [102].

This stage could be shared with nuclear physicists aiming to a source of radioactive ions of the same intensity to what needed by a Beta Beam. A design study has been recently approved by E.U.: Eurisol [100], where both nuclear and neutrino physics issues will be studied.

Ion acceleration The CERN PS and SPS can be used to accelerate the ions. There is a well established experience at CERN about ion accelerators. Ions are firstly accelerated to MeV/u by a Linac and to 300 MeV/u, in a single batch of 150 ns, by a rapid cycling synchrotron. 16 bunches (consisting of $2.5 \cdot 10^{12}$ ions each in the case of ^6He) are then accumulated into the PS, and reduced to 8 bunches during their acceleration to intermediate energies. The SPS will finally accelerate the 8 bunches to the desired energy using a new 40 MHz RF system and the existing 200 MHz RF system, before ejecting them in batches of eight 10 ns bunches into the decay ring. The SPS could accelerate

${}^6\text{He}$ ions at a maximum γ value of $\gamma_{{}^6\text{He}} = 150$.

Decay ring The decay ring has a total length of 6880 m and straight sections of 2500 m each (36% useful length for ion decays). These dimensions are fixed by the need to bend ${}^6\text{He}$ ions up to $\gamma = 150$ using 5 T superconducting magnets. Due to the relativistic time dilatation, the ion lifetimes reach several minutes, so that stacking the ions in the decay ring is mandatory to get enough decays and hence high neutrino fluxes. The challenge is then to inject ions in the decay ring and merge them with existing high density bunches. As conventional techniques with fast elements are excluded, a new scheme (asymmetric merging) was specifically conceived [103].

Summarizing, the main features of a neutrino beam based on the BetaBeams concept are:

- the beam energy depends on the γ factor. The ion accelerator can be tuned to optimize the sensitivity of the experiment;
- the neutrino beam contains a single flavor with an energy spectrum and intensity known a priori. Therefore, unlike conventional neutrino beams, close detectors are not necessary to normalize the fluxes;
- neutrino and anti-neutrino beams can be produced with a comparable flux;
- differently from SuperBeams, BetaBeams experiments search for $\nu_e \rightarrow \nu_\mu$ transitions, requiring a detector capable to identify muons from electrons. Moreover, since the beam does not contain ν_μ or $\bar{\nu}_\mu$ in the initial state, magnetized detectors are not needed. This is in contrast with the neutrino factories (see Sec.2.4) where the determination of the muon sign is mandatory.

A baseline study for a Beta Beam complex (Fig. 9) has been carried out at CERN [104]. The reference βB fluxes are $5.8 \cdot 10^{18}$ ${}^6\text{He}$ useful decays/year and $2.2 \cdot 10^{18}$ ${}^{18}\text{Ne}$ decays/year if a single ion specie circulates in the decay ring.

The water Čerenkov could be a suitable technology for a large detector. The physics potential has been initially computed in [105, 106] for $\gamma_{{}^6\text{He}} = 60$, $\gamma_{{}^{18}\text{Ne}} = 100$ and with a 440 kton detector at 130 km.

The most updated sensitivities for the baseline Beta Beam are computed in a scheme where both ions are accelerated at $\gamma = 100$, the optimal setup for the CERN-Frejus baseline of 130 km, [107], see Fig. 11. The θ_{13} sensitivity curve, computed with a 6 parameters fit minimized over the solar and the atmospheric parameters and projected over θ_{13} , is shown in Fig. 10 [107]. Degeneracies induced by the unknown values of $\text{sign}(\Delta m_{23}^2)$ and θ_{23} are not accounted for in these first plots.

The leptonic CP violation discovery potential (LCPV) has been computed with the following procedure. For any choice of a true value of θ_{13} , $\bar{\theta}_{13}$, a loop on test values of δ_{CP} , $\bar{\delta}_{CP}$, is initiated, until the fit around $(\bar{\theta}_{13}, \bar{\delta}_{CP})$ is 3σ away from any solution at $\delta_{CP} = 0$ and $\delta_{CP} = \pi$. While in ref.[108] this procedure is performed in the full $(\theta_{13}, \delta_{CP})$ space (3σ corresponding to $\Delta\chi^2 = 11.8$), here the solution is searched having marginalized out θ_{13} (GLOBES function `GblChiDelta` [109]). The LCPV at 3σ ($\Delta\chi^2 = 9.0$) is shown in Fig. 11. It takes into account all the parameter errors and all the possible degeneracies [107]. As it is common practice in literature $\theta_{23} = 40^\circ$ has been used, to leave room for the octant $(\pi/2 - \theta_{23})$ degeneracy. Each of the 4 true values of $\text{sign}(\Delta m_{23}^2)$ and θ_{23} : normal: $\text{sign}(\Delta m_{23}^2) = 1$, $\theta_{23} < \pi/4$; sign: $\text{sign}(\Delta m_{23}^2) = -1$, $\theta_{23} < \pi/4$; octant: $\text{sign}(\Delta m_{23}^2) = 1$, $\theta_{23} > \pi/4$; mixed: $\text{sign}(\Delta m_{23}^2) = -1$, $\theta_{23} > \pi/4$. has been fitted with the 4 possible fit combinations, the worst case is then taken. Also shown are the leptonic CP violation discovery potentials neglecting the degenerate solutions (that is choosing the right combination of $\text{sign}(\Delta m_{23}^2)$ and θ_{23} for the fit). Effect of degeneracies are sometimes visible for high values of θ_{13} , precisely the region where they

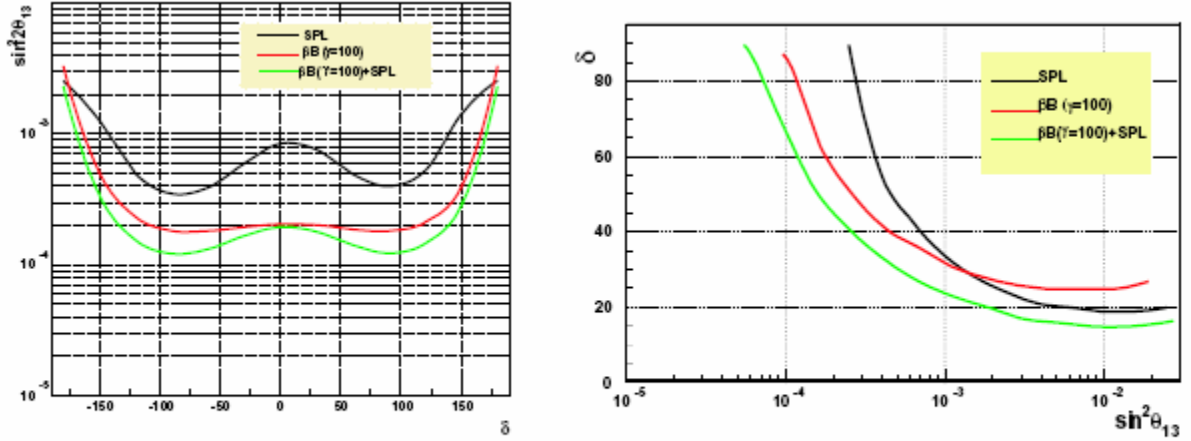


Figure 10: *LEFT: θ_{13} 90% C.L. sensitivity as function of δ_{CP} for $\Delta m_{23}^2 = 2.5 \cdot 10^{-3} eV^2$, $\text{sign}(\Delta m_{23}^2) = 1$, 2% systematic errors. SPL-SB sensitivities have been computed for a 10 years ν_μ run, βB and $\beta B_{100,100}$ for a 10 years $\nu_e + \bar{\nu}_e$ run. RIGHT: δ_{CP} discovery potential at 3σ (see text) computed for 10 years running time (5 years ν + 5 years $\bar{\nu}$ for both the facilities). The SPL-SB 3.5 GeV, BetaBeam with $\gamma = 100, 100$ and their combination are shown.*

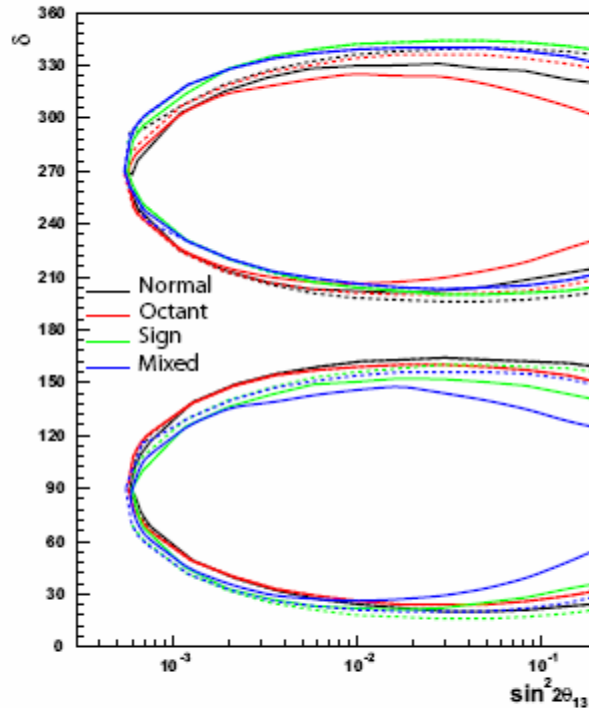


Figure 11: *Leptonic CP violation discovery potential at $3\sigma(\Delta\chi^2 > 9.0)$ of the BetaBeam with $\gamma = 100$ computed for the 4 different options about the true values of $\text{sign}(\Delta m_{23}^2)$ and θ_{23} and including all the possible degeneracies (see text). Dotted curves are computed neglecting the effects of the clone solutions. Taken from ref. [107]*

can be reduced by a combined analysis with atmospheric neutrinos [110].

BetaBeams require a proton driver in the energy range of 1-2 GeV, 0.5 MWatt power. The SPL can be used as injector, at most 10% of its protons would be consumed. This allows a simultaneous β B and SPL-SB run, the two neutrino beams having similar neutrino energies (cfr. Fig. 8). The same detector could then be exposed to 2×2 beams (ν_μ and $\bar{\nu}_\mu \times \nu_e$ and $\bar{\nu}_e$) having access to CP, T and CPT violation searches in the same run. This is particularly important because CP and T channels would have different systematics and different backgrounds, allowing for independent checks of the same signal. Furthermore, the SPL ν_μ and $\bar{\nu}_\mu$ beams would be the ideal tool to measure signal cross sections in the close detector.

With this combination of neutrino beams a sensitivity to $\sin^2 2\theta_{13} \geq 2 \cdot 10^{-4}$ (90%CL) exploiting a CP violation discovery potential at 3σ if $\delta_{CP} \geq 18^\circ$ and $\theta_{13} \geq 0.55^\circ$ [97] (Figs. 10).

BetaBeam capabilities for the maximum values of γ available with the SPS, $\gamma^{6\text{He}} = 150$ have been computed in [111].

BetaBeam capabilities for ions accelerated at higher energies than those allowed by SPS have been firstly computed in [112] and subsequently in [111, 113, 114]. These studies assume that the same ion fluxes of the baseline scenario can be maintained. However, this is not the case if the number of stored bunches is kept constant in the storage ring. On the other hand, by increasing γ (i.e. the neutrino energy) the atmospheric neutrinos background constraint on the total bunch length [85] tends to vanish. Studies are in progress at CERN in order to define realistic neutrino fluxes as a function of γ [115].

Performances of these higher energy Beta Beams would be anyway extremely interesting, as shown in the plots of Fig. 12, [114], in comparison with several other setups.

It is worth noting that if a high intensity Beta Beam with $\gamma \sim 300 \div 500$ (requiring a higher energy accelerator than SPS, like the Super-SPS[116]) can be built, a 40 kton iron calorimeter located at the Gran Sasso Laboratory will have the possibility to discover a non vanishing δ_{CP} if $\delta_{CP} > 20^\circ$ for $\theta_{13} \geq 2^\circ$ (99% C.L.) and measure the sign of Δm_{23}^2 [117].

A very recent development of the Beta Beam concept is the conceptual possibility to have monochromatic, single flavor neutrino beams thanks to the electron capture process [118, 119]. A suitable ion candidate exists: ^{150}Dy , whose performances have been already delineated [118].

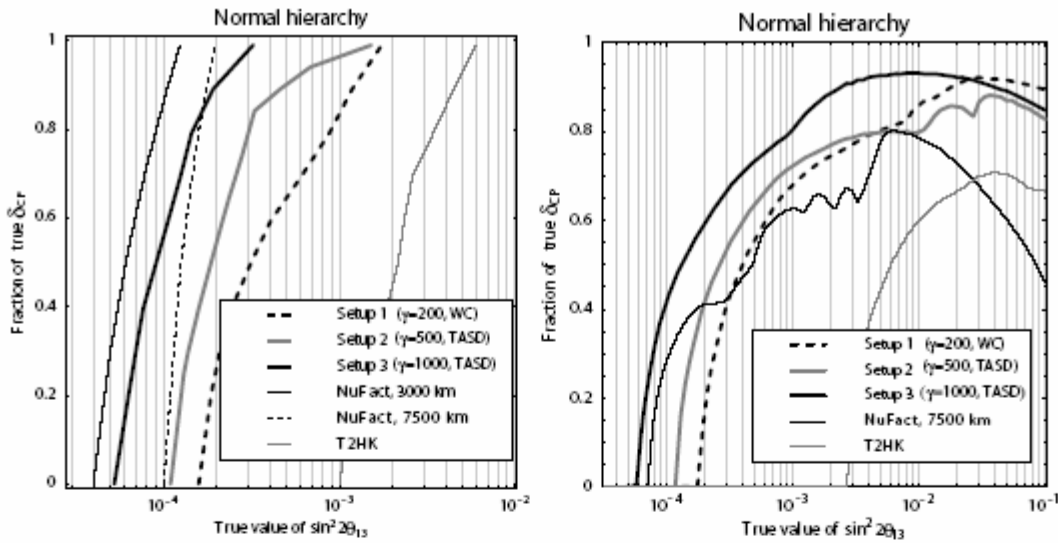


Figure 12: *LEFT: The $\sin^2 2\theta_{13}$ discovery reach (including systematics and correlations) for different setups as function of the true values of $\sin^2 2\theta_{13}$ and δ_{CP} (3σ confidence level). The values of δ_{CP} are “stacked” to the fraction of δ_{CP} , i.e., $\sin^2 2\theta_{13}$ will be discovered for a certain fraction of all possible values of δ_{CP} . For a uniform probability contribution in δ_{CP} , the fraction of δ_{CP} directly corresponds to the probability to discover $\sin^2 2\theta_{13}$. Taken from [114]. RIGHT: The sensitivity to CP violation for the normal (left) and inverted (right) mass hierarchy for different experiments as function of the true values of $\sin^2 2\theta_{13}$ and δ_{CP} at the 3σ confidence level. Taken from [114]*

2.4 The Neutrino Factory ³

In a Neutrino Factory [121] muons are accelerated from an intense source to energies of several GeV, and injected in a storage ring with long straight sections. The muon decays:

$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu \quad \text{and} \quad \mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$$

provide a very well known flux with energies up to the muon energy itself. The overall layout is shown in Fig. 13.

Neutrino Factory designs have been proposed in Europe [122], [123], the US [124] [125][126], and Japan [127]. Of these designs, the American one is the most developed, and we will use it as an example in general with a few exceptions. The conclusions of these studies is that, provided sufficient resources, an accelerator complex capable of providing about 10^{21} muons per year can be built. The Neutrino Factory consists of the following subsystems:

Proton Driver. Provides 1-4 MW of protons on a pion production target. For the Neutrino Factory application the energy of the beam within 4-30 GeV is not critical, since it has been shown that the production of pions is roughly proportional to beam power. The time structure of the proton beam

³Material for this Section is mainly taken from ref. [120]

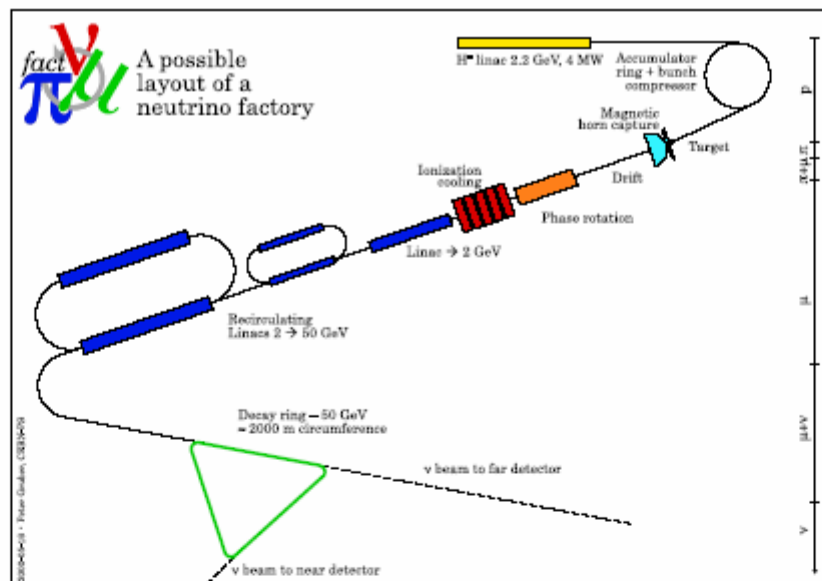


Figure 13: Schematic layout of a Neutrino Factory.

has to be matched with the time spread induced by pion decay (1-2 ns); for a Linac driver such as the SPL, this requires an additional accumulator and compressor ring.

Target, Capture and Decay. A high-power target sits within a 20T superconducting solenoid, which captures the pions. The high magnetic field smoothly decreases to 1.75T downstream of the target, matching into a long solenoid decay channel. A design with horn collection has been proposed at CERN for the Neutrino Factory, with the benefit that it can be also used for a SuperBeam design.

Bunching and Phase Rotation. The muons from the decaying pions are bunched using a system of RF cavities with frequencies that vary along the channel. A second series of RF cavities with higher gradients is used to rotate the beam in longitudinal phase-space, reducing the energy spread of the muons.

Cooling. A solenoid focusing channel with high-gradient 201 MHz RF cavities and either liquid-hydrogen or LiH absorbers is used to reduce the transverse phase-space occupied by the beam. The muons lose, by dE/dx losses, both longitudinal- and transverse-momentum as they pass through the absorbers. The longitudinal momentum is replaced by re-acceleration in the RF cavities.

Acceleration. The central momentum of the muons exiting the cooling channel is 220 MeV/c. A superconducting Linac with solenoid focusing is used to raise the energy to 1.5 GeV. Thereafter, a Recirculating Linear Accelerator raises the energy to 5 GeV, and a pair of Fixed-Field Alternating Gradient rings accelerate the beam to at least 20 GeV.

Storage Ring. A compact racetrack geometry ring is used, in which 35% of the muons decay in the neutrino beam-forming straight section. If both signs are accelerated, one can inject in two superimposed rings or in two parallel straight sections. This scheme produces over $6 \cdot 10^{20}$ useful muon decays per operational year and per straight section in a triangular geometry.

Table 4: Comparison of unloaded Neutrino Factory costs estimates in M\$ for the US Study II design and improvement estimated for the latest updated US design (20 GeV/c muons). Costs are shown including A: the whole complex; B no Proton Driver and C: no proton driver and no Target station in the estimates. Table from Ref. [128].

Costs in M\$	A	B	C
Old estimate from Study II	1832	1641	1538
Multiplicative factor for new estimate	0.67	0.63	0.60

The European Neutrino Factory design is similar to the US design, but differs in the technologies chosen to implement the subsystems.

The Japanese design is quite different, and uses very large acceptance accelerators. Cooling, although it would improve performance, is not considered mandatory in this scheme.

An important Neutrino Factory R&D effort is ongoing in Europe, Japan, and the U.S. since a few years. Significant progress has been made towards optimizing the design, developing and testing the required components, and reducing the cost.

To illustrate this progress, the cost estimate for a recent update of the US design [128] is compared in Table 4 with the corresponding cost for the previous "Study II" US design [126]. In this design the Neutrino Factory would accelerate protons up to 20 GeV/c, with a flux of $1.2 \cdot 10^{20}$ muon decays per straight section per year for a proton driver power of 1 MW ($4.8 \cdot 10^{20}$ μ decays year at 4 MW).

It should be noted that the Study II design cost was based on a significant amount of engineering input to ensure design feasibility and establish a good cost basis. Neutrino Factory R&D has reached a critical stage in which support is required for two key international experiments (MICE [129] and Targetry [130]) and a third-generation international design study. If this support is forthcoming, a Neutrino Factory could be added to the Neutrino Physics roadmap by the end of the decade.

2.4.1 Oscillations physics at the Neutrino Factory

Considering a Neutrino Factory with simultaneous beams of positive and negative muons, 12 oscillation processes can in principle be studied, Table 5.

Table 5: Oscillation processes in a Neutrino Factory

$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$	$\mu^- \rightarrow e^- \bar{\nu}_e$	
$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$	$\nu_\mu \rightarrow \nu_\mu$	disappearance
$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$\nu_\mu \rightarrow \nu_e$	appearance (challenging)
$\bar{\nu}_\mu \rightarrow \bar{\nu}_\tau$	$\nu_\mu \rightarrow \nu_\tau$	appearance (atm. oscillation)
$\nu_e \rightarrow \nu_e$	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	disappearance
$\nu_e \rightarrow \nu_\mu$	$\bar{\nu}_e \rightarrow \bar{\nu}_\mu$	appearance: "golden" channel
$\nu_e \rightarrow \nu_\tau$	$\bar{\nu}_e \rightarrow \bar{\nu}_\tau$	appearance: "silver" channel

Of course the neutrinos coming from decays of muons of different charge must not be confused with each other, this can be done by timing provided the storage ring is adequately designed.

One of the most striking features of the Neutrino Factory is the precision with which the characteristics of all components of the beam could be known. This was studied extensively in the CERN

Yellow Report [131], where the following were considered

- beam polarization effects, and its measurement with a polarimeter, allowing extraction of the beam energy, energy spread and verification that the polarization effects on the neutrino fluxes average out to zero with high precision;
- beam divergence effects, with the preliminary, conceptual study of a Čerenkov device to monitor the angular distribution of muons in the beam; [132]
- radiative effects in muon decay;
- absolute normalization to be obtained both from a beam monitor, with the added possibility of an absolute cross-section normalization using the inverse muon decay reaction, $\nu_\mu e^- \rightarrow \mu^- \nu_e$, in the near detector;

with the conclusion that, in principle, a normalization of fluxes and cross-sections with a precision of 10^{-3} can be contemplated. Some of these features should also be present for a BetaBeam, and for any facility in which a stored beam of well defined optical properties is used to produce neutrinos. This will be an essential difference with respect to the SuperBeams, where the knowledge of relative neutrino-vs-antineutrino cross-sections and fluxes will rely on the understanding of the initial particle production.

The Neutrino Factory lends itself naturally to the exploration of neutrino oscillations between ν flavors with high sensitivity. The detector should be able to perform both appearance and disappearance experiments, providing lepton identification and charge discrimination which is a tag for the initial flavor and of the oscillation. In particular the search for $\nu_e \rightarrow \nu_\mu$ transitions (“golden channel”) [133] appears to be very attractive at the Neutrino Factory, because this transition can be studied in appearance mode looking for μ^- (appearance of wrong-sign μ) in neutrino beams where the neutrino type that is searched for is totally absent (μ^+ beam in νF).

The emphasis has been placed so far on small mixing angles and small mass differences. With two 40 kton magnetic detectors (MINOS like) at 700 (or 7000) and 3000 km, with a conservative high energy muon detection threshold of 5 GeV, exposed to both polarity beams and 10^{21} muon decays, it will be possible to explore the θ_{13} angle down to 0.1° opening the possibility to measure the δ_{CP} phase [133, 75, 134], as it is shown by the plots of Fig.17.

On the other hand, the relative high energies of neutrinos selected by placing such a high threshold on muon energies require very long baselines (several thousands kilometers) for Neutrino Factories experiments, and at these baselines CP asymmetries are dominated by matter effects [135]. Taking advantage of the matter effects, such an experiment will determine unambiguously $\text{sign}(\Delta m_{23}^2)$ for large enough θ_{13} ($\theta_{13} \geq 2^\circ$). However, we remind that, such as for other facilities, the determination of (θ_{13}, δ) at the Neutrino Factory is not free of ambiguities: up to eight different regions of the parameter space can fit the same experimental data. In order to solve these ambiguities, a single experimental point on a single neutrino beam is not enough.

One possibility at the Neutrino Factory is to make use of the rich flavour content of the beam. This imply an optimized network of detectors with different characteristics. Indeed, a specific disadvantage of the considered magnetized iron detector when dealing with degeneracies is the following: the lower part of the Neutrino Factory spectrum (say, $E_\nu \in [0, 10]$ GeV) cannot be used due to the extremely low efficiency in this region of the detector. This part of the spectrum, on the other hand, is extremely useful to solve degeneracies, as it has been shown in several papers [136].

One possibility is to envisage that all of BetaBeams, SuperBeams and Neutrino Factories will be available. Several investigations on how to solve this problem have been carried out, as reported

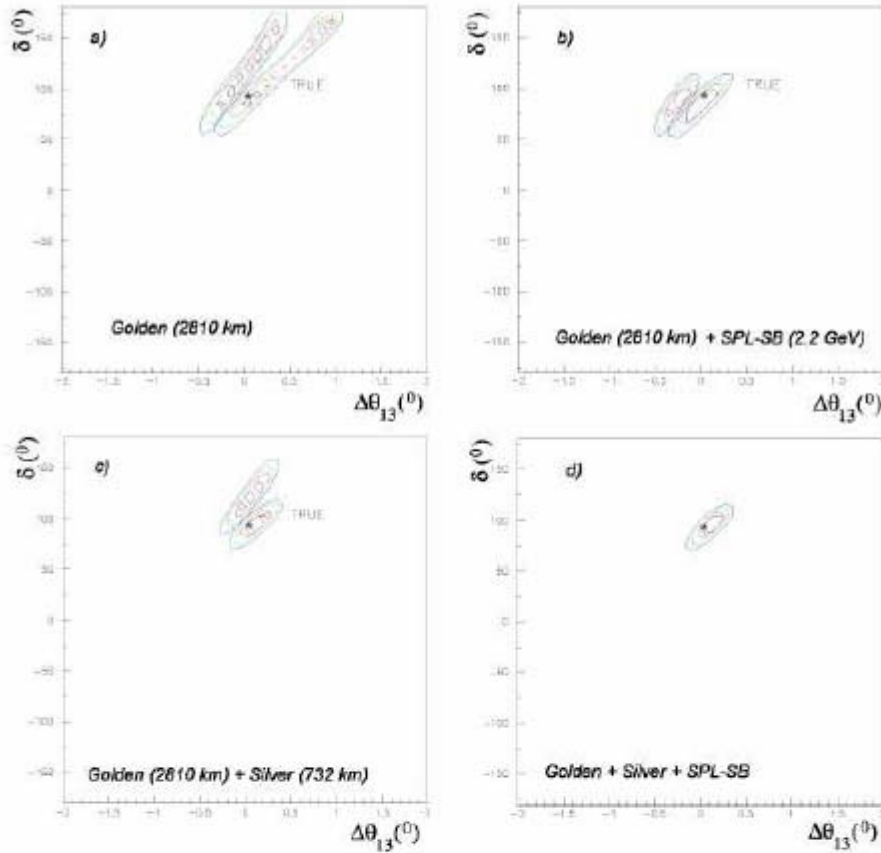


Figure 14: The results of a χ^2 fit for $\bar{\theta}_{13} = 2^\circ$; $\delta_{CP}^- = 90^\circ$. Four different combinations of experimental data are presented: a) magnetized iron detector (MID) at a νN ; b) MID plus SPL-SB; c) MID plus hybrid emulsion (HE) at νF ; d) the three detectors together. Notice how in case (d) the eightfold-degeneracy is solved and a good reconstruction of the physical θ_{13}, δ_{CP} values is achieved. From ref. [138]

in [137] and references therein. As an example the result of such an analysis combining the golden and the silver ($\nu_e \rightarrow \nu_\tau$) νF channels with the SPL-SB, taken from reference [138], is shown in Fig. 14. It should be stressed, however, that the analysis presented in this figure does not include systematic errors and it is based on a non-updated simulation of the magnetized iron detector, [139].

A more interesting but challenging task will be to assume that only one of these facilities will become available (a more economical assumption!) and to investigate its ability to solve these ambiguities.

There are several handles to this problem at a Neutrino Factory. Clearly one should use more than just the wrong sign muons. Such a study was performed assuming the feasibility of a liquid argon detector [140]. By separating the events into several classes, right sign muon, wrong sign muon, electron and neutral current; and by performing a fine energy binning down to low energies; it was shown that the matter resonance could be neatly measured as shown in Fig. 15. The simultaneous observation of the four aforementioned channels was shown to allow resolution of ambiguities to a large extent.

The tau appearance channel *silver channel* [138] has been advocated as a powerful means of solving ambiguities. This can be readily understood since this channel has the opposite-sign dependence

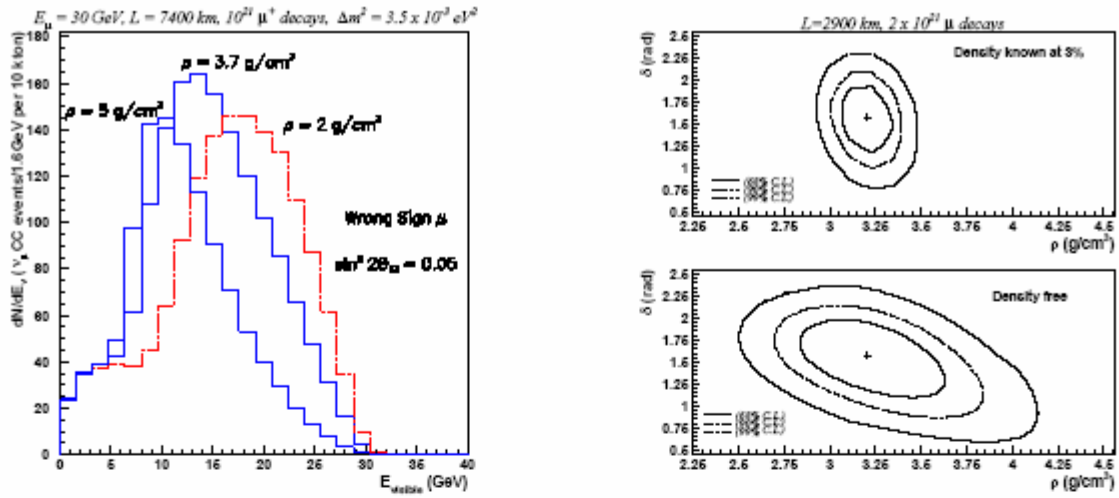


Figure 15: Left: variation of the MSW resonance peak for wrong sign muons as a function of Earth density. The plot is normalized to $10^{21} \mu^+$ decays. Right: result of a simultaneous fit to the matter density and to the CP phase for a 10kton detector situated 7000km away from a 30 GeV muon storage ring Neutrino Factory. Both plots from [140].

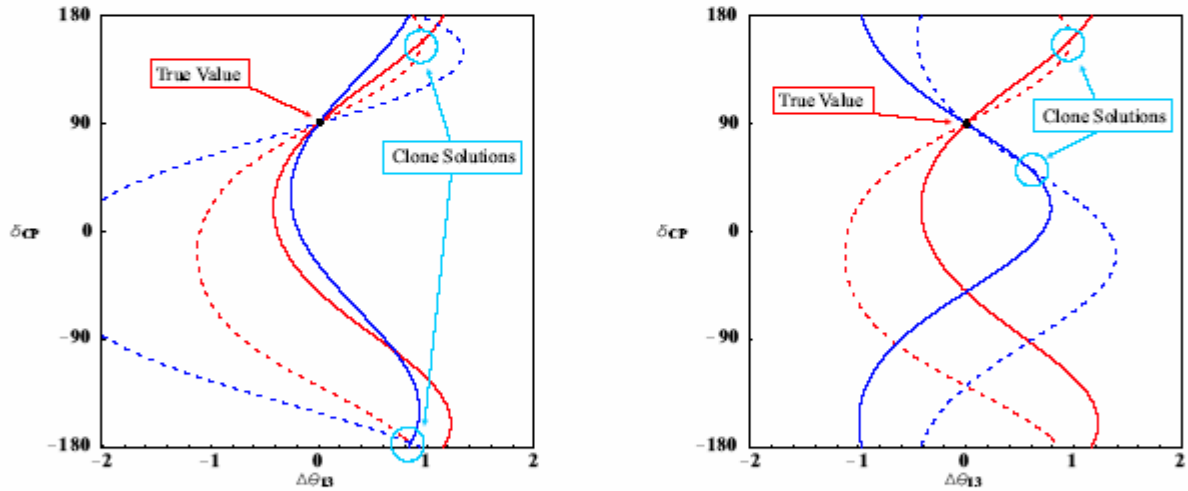


Figure 16: Solving the intrinsic degeneracy: two baseline $L=730$ and 3500 km, same channel example on the left, vs two channels $\nu_e \rightarrow \nu_\mu$ vs $\nu_e \rightarrow \nu_\tau$ same baseline example on the right.

on δ_{CP} than the golden one, while having similar dependence on matter effects and θ_{13} . Another channel to be used is the ν_μ disappearance channel, rather effective for large values of θ_{13} [87]. The principle of degeneracies-solving using several baselines, binning in energies and both silver and golden channels is explained on Fig. 16. The full demonstration that a Neutrino Factory alone with a complete set of appropriate detectors and two baselines could unambiguously do the job remains however to be worked out.

According to Table 5, Neutrino Factory potential could be further improved with a detector capa-

ble of measuring the charge of the electrons. R&D efforts for a liquid argon detector embedded in a magnetic fields are ongoing [141]; the first curved tracks were recently observed in a 10 liters Liquid Argon TPC embedded in magnetic field [142].

3 Comparison

The comparison of performances of different facilities cannot be considered concluded. Several different aspects still need to be clarified before a final comparison can be performed.

- Costs, timescales, fluxes of the different accelerators systems are not yet fully worked out.
- Performances and optimization of the detectors are not known at the same level: for water Čerenkov detectors full simulation and full reconstructions of the events are available, based on the experience of Super-Kamiokande, but the optimization in terms of photo-detector coverage is still to be performed. The optimization of the magnetic detector (and of the assumed cuts) for the Neutrino Factory was performed for very small values of θ_{13} , and the performances are for the moment based on parametrization of the MINOS performances, and very low values of backgrounds from hadron decays (pions, kaons, and charm) are claimed possible. The performances of the emulsion detector for the silver channel of Nufact or of the liquid argon detector are based on full simulations, but will need to be bench-marked using the performances of OPERA and ICARUS respectively.
- Several different measurements can be defined as significant for the facility, and they cannot be optimized all together (see also reference [114]). For instance the following measurements bring to different optimizations: sensitivity to θ_{13} , discovery of subleading $\nu_{\mu} \rightarrow \nu_e$ oscillations, unambiguous measurement of θ_{13} , measure of $\text{sign}(\Delta m_{23}^2)$, discovery of leptonic CP violation, unambiguous measurement of δ_{CP} .
- The final extraction of the oscillation parameters can significantly change based on technical aspects of the fitting programs (like choice of the input parameters, treatment of the errors on the neutrino oscillation parameters, treatment of degeneracies etc.).

The GLoBES program [109] represents a major improvement: it allows to compare different facilities keeping the same the fitting program, and it makes explicit the description of the performances of the detectors. We strongly recommend that new developments in this field will make use of GLoBES, in view of a more transparent comparison of the different proposals.

- Systematic errors that strongly influence performances, for instance sensitivity to leptonic CP violation for large values of θ_{13} , are not substantially discussed in the literature. We are confident that facilities where neutrino fluxes can be known a-priori, as the case of Beta Beams and Neutrino Factories, will have smaller systematic errors (and smaller backgrounds) than e.g. neutrino SuperBeams. This difference is not known quantitatively today.

The concept of the near detector station(s) and flux monitoring systems has to be proposed together with the facility, in particular for low energy (few 100 MeV) BetaBeam and SuperBeam where the issues of muon mass effect, Fermi motion and binding energy uncertainty are highly non-trivial.

Finally, for the Neutrino Factory, the question of systematics on the prediction of matter effects is essential for the performance at large values of θ_{13} .

- Overall performances will depend on the combination of several different inputs. For instance low energy Super Beams and Beta Beams can profit of atmospheric neutrino oscillations, detected with a large statistics in the gigantic water Čerenkov detector, to solve degeneracies and measure $\text{sign}(\Delta m_{23}^2)$, as shown in the pioneering work of [110]. Neutrino Factory can profit of the combination of different channels as the golden and the silver one (see Section 2.4), as well as detectors at different baselines. A full exploration of these possibilities is an ongoing process and the results available at today cannot be considered final.

Having said that, a comparison of the facilities that at present are described in the GLoBES library [143], as far as concerns and leptonic CP violation discovery potential, is shown in Fig. 17.

The plots show the sensitivity to CP violation at 3σ CL ($\Delta\chi^2 = 9$). Sensitivity to CP violation is defined, for a given point in the θ_{13} - δ_{CP} plane, by being able to exclude $\delta_{\text{CP}} = 0$ and $\delta_{\text{CP}} = \pi$ at the given confidence level. All plots have been prepared with GLoBES [109].

Degeneracies and correlations are fully taken into account. For all setups the appropriate disappearance channels have been included. The Beta Beam is lacking muon neutrino disappearance, but the result does not change if T2K disappearance information is included in the analysis. In all cases systematics between neutrinos, anti-neutrinos, appearance and disappearance is uncorrelated. For all setups with a water ν Čerenkov detector the systematics applies both to background and signal, uncorrelated.

The Neutrino Factory assumes $3.1 \cdot 10^{20} \mu^+$ decays per year for 10 years and $3.1 \cdot 10^{20} \mu^-$ decays for 10 years. It has one detector with $m=100$ kton at 3000 km and another 30 kton detector at 7000 km. The density errors between the two baselines are uncorrelated, sensitivities are computed for 2% and 5% systematic error on matter density. The systematics are 0.1% on the signal and 20% on the background, uncorrelated. The detector threshold and the other parameters are taken from [136] and approximate the results of [139].

The Beta Beam assumes $5.8 \cdot 10^{18} {}^6\text{He}$ 6 decays per year for five years and $2.2 \cdot 10^{18} {}^{18}\text{Ne}$ 18 decays per year for five years. The detector mass is 500 kton. The detector description and the glb-file is from [107].

The SPL setup is taken from [98], and the detector mass is 500 kton.

The T2HK setup is taken from [136] and closely follows the LoI [82]. The detector mass is 1000 kton and it runs with 4MW beam power, 6 years with anti-neutrinos and 2 years with neutrinos. The systematic error on both background and signal is 5%.

The oscillation parameters were [144, 145]: $\delta m_{23}^2 = 0.0024 \text{ eV}^2$, $\delta m_{12}^2 = 0.00079 \text{ eV}^2$, $\theta_{23} = \pi/4$, $\theta_{12} = 0.578$. The input errors are (at 1σ): 10% on δm_{23}^2 , 10% on θ_{23} , 10% on θ_{12} , 4% on δm_{12}^2 , 5% on ρ (unless otherwise stated).

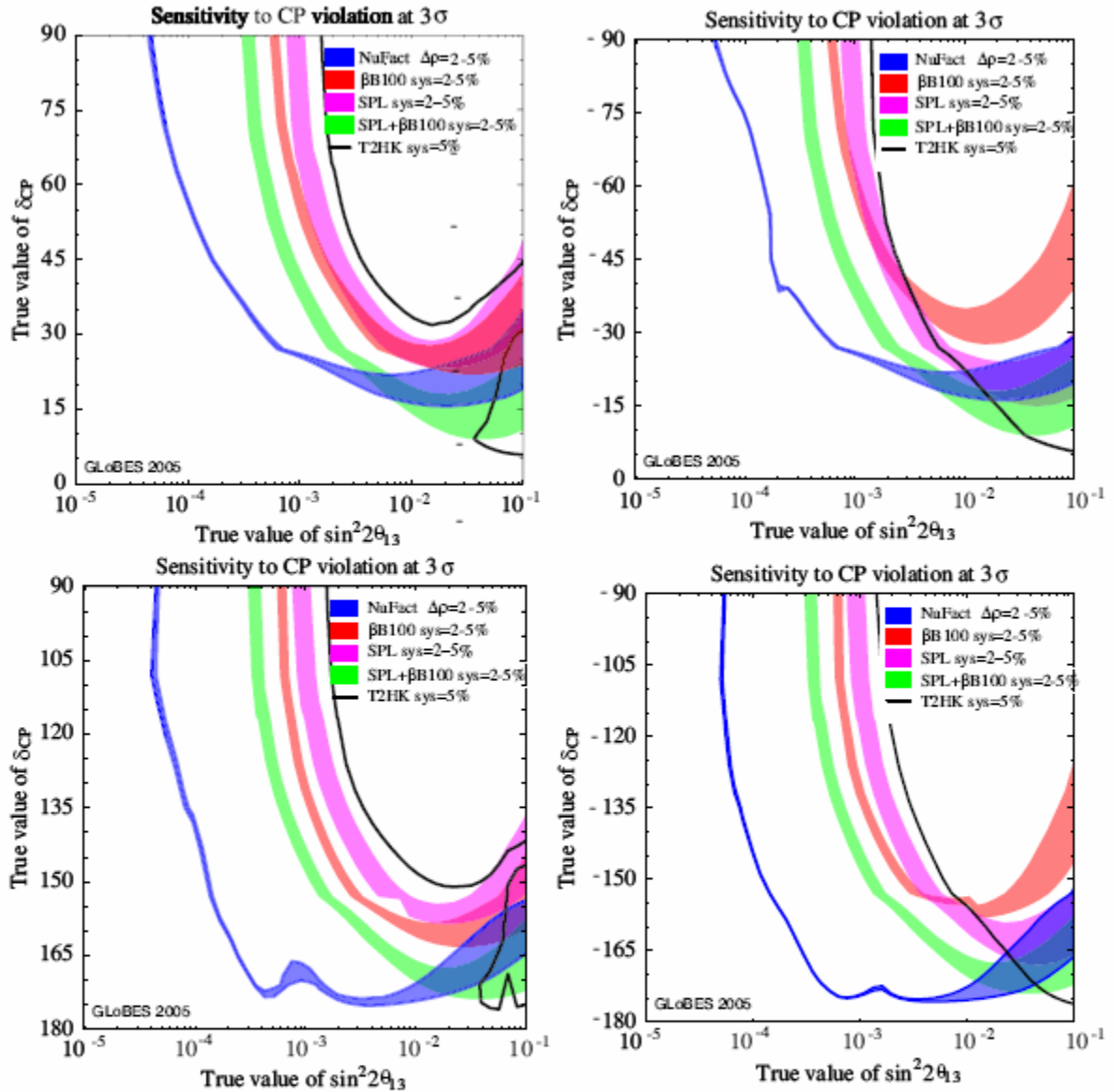


Figure 17: δ_{CP} discovery potential at 3σ (see text) computed for 10 years running time. For explanation of the facilities see the text. The four plots represent the four possible quadrants of δ_{CP} values, performances of the different facilities are not at all the same in the different quadrants. The width of the curves reflects the range of systematic errors: 2% and 5% on signal and background errors for SPL-SB and Beta Beam, 2% and 5% for the matter density. Other systematic errors are 5% on signal and background of T2HK, 0.1% for nufact signal, 20% for nufact backgrounds. Description of the facilities can be found in [143]

References

- [1] C. Weinheimer, Nucl. Phys. Proc. Suppl. **118** (2003) 279.
- [2] V. M. Lobashev *et al.*, Nucl. Phys. Proc. Suppl. **91** (2001) 280. V. M. Lobashev, proceeding of "Neutrino Telescopes 2005", Venice, pg. 507-517
- [3] KATRIN Coll., hep-ex/0109033.
- [4] A. Monfardini *et al.*, arXiv:hep-ex/0509038.
- [5] H. V. Klapdor-Kleingrothaus, A. Dietz and I. V. Krivosheina, Found. Phys. **32** (2002) 1181 [Erratum-ibid. **33** (2003) 679] [arXiv:hep-ph/0302248].
- [6] Y. G. Zdesenko, F. A. Danevich and V. I. Tretyak, Phys. Lett. B **546** (2002) 206.
- [7] A. M. Bakalyarov, A. Y. Balysh, S. T. Belyaev, V. I. Lebedev and S. V. Zhukov [C03-06-23.1 Collaboration], Phys. Part. Nucl. Lett. **2** (2005) 77 [Pisma Fiz. Elem. Chast. Atom. Yadra **2** (2005) 21] [arXiv:hep-ex/0309016].
- [8] C. Arnaboldi *et al.*, Phys. Rev. Lett. **95** (2005) 142501 [arXiv:hep-ex/0501034].
- [9] R. Arnold *et al.* [NEMO Collaboration], arXiv:hep-ex/0507083.
- [10] R. Ardito *et al.*, arXiv:hep-ex/0501010.
- [11] S. Schonert *et al.* [GERDA Collaboration], Nucl. Phys. Proc. Suppl. **145** (2005) 242.
- [12] M. Fukugita, arXiv:hep-ph/0511068.
- [13] M. Fukugita and T. Yanagida, Phys. Lett. B **174**, 45 (1986).
- [14] W. Buchmüller, P. Di Bari and M. Plümacher, Nucl. Phys. B **665**, 445 (2003).
- [15] G. F. Giudice, A. Notari, A. Riotto and A. Strumia, Nucl. Phys. B **685**, 89 (2004).
- [16] W. Buchmüller, P. Di Bari and M. Plümacher, Ann. Phys. **315**, 303 (2005).
- [17] C. L. Bennett *et al.*, Astrophys. J. Suppl. **148**, 1 (2003).
- [18] D. G. York *et al.*, Astron. J. **120**, 159 (2000).
- [19] D.N. Spergel *et al.*, Astrophys. J. Suppl. **148**, 175 (2003).
- [20] M. Tegmark *et al.*, Phys. Rev. D **69**, 103501 (2004).
- [21] D. Eisenstein *et al.*, arXiv:astro-ph/0501171.
- [22] M. Tegmark *et al.*, Astrophys. J. **606**, 702 (2004).
- [23] S. Cole *et al.*, Mon. Not. Roy. astr. Soc. **362**, 505 (2005).
- [24] W. Hu, D. J. Eisenstein and M. Tegmark, Phys. Rev. Lett. **80**, 5255 (1998).
- [25] R. Valdarnini, T. Kahniashvili and B. Novosyadlyj, Astron. Astrophys. **336**, 11 (1998).

- [26] S. Hannestad, JCAP 5. 004 (2003).
- [27] S.W. Allen, R.W. Schmidt and S.L. Bridle, Mon. Not. Roy. astr. Soc. 346, 593 (2003).
- [28] V. Barger, D. Marfatia and A. Tregre, Phys. Lett. B595, 55 (2004).
- [29] R.A.C. Croft, D.H. Weinberg, N. Katz and L. Hernquist, Astrophys. J. 495, 44 (1998).
- [30] P. McDonald et al., arXiv:astro-ph/0405013.
- [31] U. Seljak et al., arXiv:astro-ph/0407372.
- [32] R.A.C. Croft, W. Hu and R. Davé, Phys. Rev. Lett. 83, 1092 (1999).
- [33] Ø. Elgarøy and O. Lahav, New J. Phys. 7, 61 (2005).
- [34] Precision Electroweak Measurements on the Z Resonance The LEP collaborations, hep-ex/0509008, to appear as physics report.
- [35] B. T. Cleveland *et al.*, Astrophys. J. **496** (1998) 505.
- [36] M. Altmann *et al.* [GNO Collaboration], Phys. Lett. B **490** (2000) 16 [arXiv:hep-ex/0006034].
- [37] J. N. Abdurashitov *et al.* [SAGE Collaboration], J. Exp. Theor. Phys. **95** (2002) 181 [Zh. Eksp. Teor. Fiz. **122** (2002) 211] [arXiv:astro-ph/0204245].
- [38] S. Fukuda *et al.* [Super-Kamiokande Collaboration], Phys. Lett. B **539** (2002) 179 [arXiv:hep-ex/0205075].
- [39] S. R. Ahmad et al., [SNO Collaboration], nucl-ex/0204008. S. N. Ahmed *et al.* [SNO Collaboration], Phys. Rev. Lett. **92** (2004) 181301 [arXiv:nucl-ex/0309004].
- [40] K. Eguchi *et al.* [KamLAND Collaboration], Phys. Rev. Lett. **90** (2003) 021802 [arXiv:hep-ex/0212021]. T. Araki *et al.* [KamLAND Collaboration], Phys. Rev. Lett. **94** (2005) 081801 [arXiv:hep-ex/0406035].
- [41] Y. Fukuda *et al.* [Super-Kamiokande Collaboration], Phys. Rev. Lett. **81** (1998) 1562 [arXiv:hep-ex/9807003].
- [42] M. C. Sanchez *et al.* [Soudan 2 Collaboration], Phys. Rev. D **68** (2003) 113004 [arXiv:hep-ex/0307069].
- [43] M. Ambrosio *et al.* [MACRO Collaboration], Phys. Lett. B **566** (2003) 35 [arXiv:hep-ex/0304037].
- [44] Y. Ashie *et al.* [Super-Kamiokande Collaboration], Phys. Rev. Lett. **93** (2004) 101801 [arXiv:hep-ex/0404034].
- [45] M. Apollonio *et al.* [CHOOZ Collaboration], Eur. Phys. J. C **27** (2003) 331, [arXiv:hep-ex/0301017].
- [46] E. Aliu *et al.* [K2K Collaboration], Phys. Rev. Lett. **94** (2005) 081802 [arXiv:hep-ex/0411038].
- [47] L. Wolfenstein, Phys. Rev. D17 (1978) 2369. S.P. Mikheev and A.Y. Smirnov, Nuovo Cim. C9 (1986) 17.

- [48] see for instance P. Lipari, Phys. Rev. D **64** (2001) 033002 [arXiv:hep-ph/0102046].
- [49] G. L. Fogli, E. Lisi, A. Marrone and A. Palazzo, arXiv:hep-ph/0506083.
- [50] A. Aguilar *et al.* [LSND Collaboration], Phys. Rev. D **64** (2001) 112007 [arXiv:hep-ex/0104049].
- [51] B. Armbruster *et al.* [KARMEN Collaboration], Phys. Rev. D **65** (2002) 112001 [arXiv:hep-ex/0203021].
- [52] P. Astier *et al.* [NOMAD Collaboration], Phys. Lett. B **570** (2003) 19 [arXiv:hep-ex/0306037].
- [53] E. Church *et al.* [BooNe Collaboration], arXiv:nucl-ex/9706011.
- [54] A. Guglielmi, M. Mezzetto, P. Migliozi and F. Terranova, arXiv:hep-ph/0508034.
- [55] E. Ables *et al.* [MINOS Collaboration], Fermilab-proposal-0875 G. S. Tzanakos [MINOS Collaboration], AIP Conf. Proc. **721**, 179 (2004).
- [56] The Fermilab NuMI Group, "NumI Facility Technical Design Report", Fermilab Report NuMI-346, 1998.
- [57] F. Arneodo *et al.* [ICARUS Collaboration], Nucl. Instrum. and Meth. A **461** (2001) 324; P. Aprili *et al.*, "The ICARUS experiment", CERN-SPSC/2002-27, SPSC-P-323.
- [58] OPERA Collaboration, CERN-SPSC-P-318, LNGS-P25-00; H. Pessard [OPERA Collaboration], arXiv:hep-ex/0504033. M. Guler *et al.* [OPERA Collaboration], "OPERA: An appearance experiment to search for $\nu_\mu \rightarrow \nu_\tau$ oscillations in the CNGS beam. Experimental proposal," CERN-SPSC-2000-028.
- [59] G. Acquistapace *et al.*, "The CERN neutrino beam to Gran Sasso", CERN 98-02, INFN/AE-98/05 (1998); CERN-SL/99-034(DI), INFN/AE-99/05 Addendum.
- [60] H. W. Atherton *et al.*, "Precise measurements of particle production by 400 GeV/c protons on Beryllium targets," CERN-80-07.
- [61] G. Ambrosini *et al.* [NA56/SPY Collaboration], Eur. Phys. J. C **10** (1999) 605.
- [62] L. Casagrande *et al.*, "The alignment of the CERN West Area neutrino facility," CERN-96-06.
- [63] A. Guglielmi and G. Collazuol, "Monte Carlo Simulation of the SPS WANF Neutrino Flux", INFN/AE-03/05 (2003).
P. Astier *et al.* [NOMAD Collaboration], Nucl. Instrum. Meth. A **515** (2003) 800 [arXiv:hep-ex/0306022].
- [64] A. Ferrari, A. Guglielmi and P. R. Sala, Nucl. Phys. Proc. Suppl. **145** (2005) 93 [arXiv:hep-ph/0501283].
- [65] Report to the Fermilab Director by the Proton Committee, November 9, 2004, http://www.fnal.gov/directorate/programplanning/Nov2004PACPublic/Draft_Proton_Plan.v2.pdf
- [66] M. Komatsu, P. Migliozi and F. Terranova, J. Phys. G **29** (2003) 443 [arXiv:hep-ph/0210043].

- [67] P. Migliozi and F. Terranova, *Phys. Lett. B* **563** (2003) 73 [arXiv:hep-ph/0302274].
- [68] M. Benedikt, K. Cornelis, R. Garoby, E. Metral, F. Ruggiero and M. Vretenar, "Report of the High Intensity Protons Working Group," CERN-AB-2004-022-OP-RF.
- [69] F. Ardellier *et al.* [Double-CHOOZ Collaboration], arXiv:hep-ex/0405032.
- [70] A. De Rujula, M.B. Gavela, P. Hernandez, *Nucl. Phys. B* **547** (1999) 21, arXiv: hep-ph/9811390.
- [71] 'ECFA/CERN studies of a European Neutrino Factory complex', A. Blondel et al, ed CERN 2004-002 (2004) <http://preprints.cern.ch/cernrep/2004/2004-002/2004-002.html>
- [72] Beams for European Neutrino Experiments is a Networking Activity (<http://bene.na.infn.it/>) supported by the EC and most major European Agencies in the framework of the FP6 Integrated Activity CARE (Coordinated Accelerator R&D in Europe) <http://esgard.lal.in2p3.fr/Project/Activities/Current/>. The scope of BENE is also described in more detail in the following document <http://esgard.lal.in2p3.fr/Project/Activities/Current/Networking/N3/BENE-downsized-11.doc>
- [73] Workshop on Physics at a Multi Megawatt Proton driver, CERN-SPSC-2004-SPSC-M-722,
- [74] B. Richter, SLAC-PUB-8587 [arXiv:hep-ph/0008222], and references therein.
- [75] J. Burguet-Castell, M. B. Gavela, J. J. Gomez-Cadenas, P. Hernandez and O. Mena, *Nucl. Phys. B* **608** (2001) 301 [arXiv:hep-ph/0103258].
- [76] K. Kimura, A. Takamura and H. Yokomakura, *Phys. Rev. D* **66** (2002) 073005 [arXiv:hep-ph/0205295]. E. K. Akhmedov, R. Johansson, M. Lindner, T. Ohlsson and T. Schwetz, *JHEP* **0404** (2004) 078 [arXiv:hep-ph/0402175]. M. Freund, *Phys. Rev. D* **64** (2001) 053003 [arXiv:hep-ph/0103300].
- [77] H. Minakata and H. Nunokawa, *JHEP* **0110** (2001) 001 [arXiv:hep-ph/0108085].
- [78] V. Barger, D. Marfatia and K. Whisnant, *Phys. Rev. D* **65** (2002) 073023 [arXiv:hep-ph/0112119].
- [79] G. L. Fogli and E. Lisi, *Phys. Rev. D* **54** (1996) 3667 [arXiv:hep-ph/9604415].
- [80] M.G. Catanesi *et al.* [HARP Collaboration], CERN-SPSC/2001-017, SPSC/P322, May 2001.
- [81] M. G. Catanesi *et al.* [HARP Collaboration], arXiv:hep-ex/0510039.
- [82] Y. Itow *et al.*, "The JHF-Kamiokande neutrino project", arXiv:hep-ex/0106019.
- [83] The E889 Collaboration, "Long Baseline Neutrino Oscillation Experiment at the AGS", Brookhaven National Laboratory Report BNL No. 52459, April 1995. A. Para and M. Szeleper, arXiv:hep-ex/0110032.
- [84] S. Geer, *Phys. Rev. D* **57** (1998) 6989 [Erratum-ibid. *D* **59** (1999) 039903], [hep-ph/9712290].
- [85] P. Zucchelli, *Phys. Lett. B* **532** (2002) 166.

- [86] H. Minakata, M. Sonoyama and H. Sugiyama, Phys.Rev.D70:113012,2004; hep-ph/0406073. A. Donini, D. Meloni and S. Rigolin, hep-ph/0506100.
- [87] A. Donini, E. Fernandez-Martinez, D. Meloni and S. Rigolin, arXiv:hep-ph/0512038.
- [88] D. S. Ayres *et al.* [NOvA Collaboration], arXiv:hep-ex/0503053.
- [89] B. Autin *et al.*, CERN-2000-012
- [90] J. J. Gomez-Cadenas *et al.*, Proceedings of "Venice 2001, Neutrino telescopes", vol. 2*, 463-481, arXiv:hep-ph/0105297. A. Blondel *et al.*, Nucl. Instrum. Meth. A **503** (2001) 173. M. Mezzetto, J. Phys. G **29** (2003) 1771 [arXiv:hep-ex/0302005].
- [91] M. Apollonio *et al.*, arXiv:hep-ph/0210192.
- [92] L. Mosca, Nucl. Phys. Proc. Suppl. **138** (2005) 203.
- [93] J. E. Campagne and A. Cazes, arXiv:hep-ex/0411062.
- [94] R. Garoby, "The SPL at CERN," CERN-AB-2005-007.
- [95] S. Gilardoni et al, AIP Conf. Proc. **721** (2004) 334.
- [96] A. Blondel et al, CERN-NUFACT-Note-78 (2001)
- [97] M. Mezzetto, Nucl. Phys. Proc. Suppl. **149** (2005) 179.
- [98] J.E. Campagne, hep-ex/0511005.
- [99] C. K. Jung [UNO Collaboration] arXiv:hep-ex/0005046.
- [100] <http://www.ganil.fr/eurisol/>
- [101] J. Nolen, NPA 701 (2002) 312c
- [102] P. Sortais, presentations at the Moriond workshop on radioactive beams, Les Arcs (France) 2003 "ECR technology", <http://moriond.in2p3.fr/radio>
- [103] M. Benedikt, S. Hancock and J-L. Vallet, CERN note AB-Note-2003-080 MD
- [104] B. Autin *et al.*, arXiv:physics/0306106. M. Benedikt, S. Hancock and M. Lindroos, Proceedings of EPAC 2004, <http://accelconf.web.cern.ch/AccelConf/e04>.
- [105] M. Mezzetto, J.Phys. G **29** (2003) 1781 [arXiv:hep-ex/0302007]. J. Bouchez, M. Lindroos and M. Mezzetto, AIP conference proceedings, **721** (2003) 37 [arXiv:hep-ex/0310059]. M. Mezzetto, Nucl. Phys. Proc. Suppl. **143** (2005) 309 [arXiv:hep-ex/0410083].
- [106] A. Donini, E. Fernandez-Martinez, P. Migliozzi, S. Rigolin and L. Scotto Lavina, Nucl. Phys. B **710** (2005) 402 [arXiv:hep-ph/0406132].
- [107] M. Mezzetto, arXiv:hep-ex/0511005.
- [108] A. Donini, E. Fernandez-Martinez and S. Rigolin, Phys. Lett. B **621** (2005) 276 [arXiv:hep-ph/0411402].

- [128] "Neutrino Factory and Beta Beam Experiments and Developments", (Eds. S. Geer and M. Zisman), Report of the Neutrino Factory and Beta Beam Working Group, APS Multi-Divisional Study of the Physics of Neutrinos, July 2004.
- [129] The International Muon Ionization Experiment MICE, <http://hep04.phys.iit.edu/cooldemo>
- [130] R.J. Bennett et al, 'Studies of a target system for a 4MW 24 GeV proton beam', CERN-INTC proposal 2003-033, April 2004.
- [131] A. Blondel et al eds., *ECFA/CERN studies of a European Neutrino Factory complex*, CERN 2004-002 (2004) <http://preprints.cern.ch/cernrep/2004/2004-002/2004-002.html>
- [132] A. Broncano and O. Mena, *Eur.Phys.J. C* **29** (2003) 197-206; hep-ph/0203052
- [133] A. Cervera et al, *Nucl.Phys.B* **579**:17-55,2000, Erratum-*ibid.* **593**:731-732,2001; hep-ph/0002108
- [134] P. Huber and W. Winter; *Phys. Rev. D* **68**, 037301(2003); hep-ph-0301257.
- [135] S. Geer, *Comments Nucl. Part. Phys. A* **2** (2002) 284 [arXiv:hep-ph/0008155].
- [136] P. Huber, M. Lindner and W. Winter, *Nucl.Phys.B* **645**:3-48,2002; hep-ph/0204352
- [137] A. Donini, *AIP Conf. Proc.* **721** (2004) 219 [arXiv:hep-ph/0310014].
- [138] A. Donini, D. Meloni, P. Migliozzi, *Nucl. Phys. B* **646** (2002) 321, hep-ph/0206034; D. Autiero *et al.*, *Eur. Phys. J. C* **33** (2004) 243 [arXiv:hep-ph/0305185]. S. Rigolin, *Rencontres de Moriond 2004*, hep-ph/0407009.
- [139] A. Cervera, F. Dydak and J. Gomez Cadenas, *Nucl. Instrum. Meth. A* **451** (2000) 123.
- [140] A. Bueno, M. Campanelli, A. Rubbia, *Nucl. Phys. B* **589** (2000) 577 [arXiv:hep-ph/0005007].
- [141] A. Ereditato and A. Rubbia, arXiv:hep-ph/0510131.
- [142] A. Badertscher et al, physics/0505151.
- [143] <http://www.hep.ph.ic.ac.uk/~7E1ongkr/UKNF/Scoping-study/ISS-www-site/WG1-PhysPhen/Workshops/2005-11/AEDL/AEDL-Catalogue.htm>
- [144] M. Maltoni, T. Schwetz, M.A. Tortola, J.W.F. Valle, *New J.Phys.* **6**:122,2004, [hep-ph/0405172].
- [145] G.L. Fogli, E. Lisi, A. Marrone, A. Palazzo, A.M. Rotunno, hep-ph/0506307.

Superior Neutrino Beams: overview & recommendations

As described in the previous sections, the discovery of neutrino oscillations forms one of the most important results in particle physics in the last decade [1]. It has led to the creation of many new projects to investigate oscillations in more detail. For example, a number of so-called long baseline neutrino oscillation facilities, in which the neutrinos are produced by an accelerator and detected several hundred kilometres away, have recently started taking data or are under construction, e.g. MINOS [2], CNGS [3] and T2K [4]. However, these projects are unlikely to have much impact on the so far unmeasured parameters of neutrino oscillations, θ_{13} , the CP-violation angle δ and the sign of Δm_{23}^2 . For these, new, more advanced neutrino beams will be required and these are the subjects of this section.

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There are three candidate facilities to provide these advanced beams: superbeams, a Neutrino Factory (see figure 1) and a Beta-beam (see figure 2). In a superbeam facility, e.g. an upgraded T2K or CERN to Frejus, the neutrinos are produced in the “standard” manner, by bombarding a target with a proton beam and focussing the pions created. However, the superbeam projects are different in two respects from the long baseline facilities currently under construction: (1) the proton beam power is higher, around 4MW, and (2) the neutrino detector is much bigger, typically around 1 MT. Some will also employ the so-called off-axis technique, in which the pion beam is pointed a few degrees away from the detector. This has the effect of kinematically suppressing the high energy neutrino tails and producing more neutrinos of a useful energy. In a Neutrino Factory, the neutrinos come from the decay of muons in a storage ring at an energy of between 20 and 50 GeV. This has many advantages over a conventional neutrino beam, in particular the composition, intensity and energy spectra of the neutrino beams are precisely known and high intensities are possible. A beta-beam has a number of benefits in common with a Neutrino Factory. However, the stored particles in this case are beta-emitters, in particular ${}^6\text{He}$ and ${}^{18}\text{Ne}$ which produce electron anti-neutrinos and neutrinos, respectively. As for the Neutrino Factory, the result is a neutrino beam of precisely known composition, intensity and energy spectra.

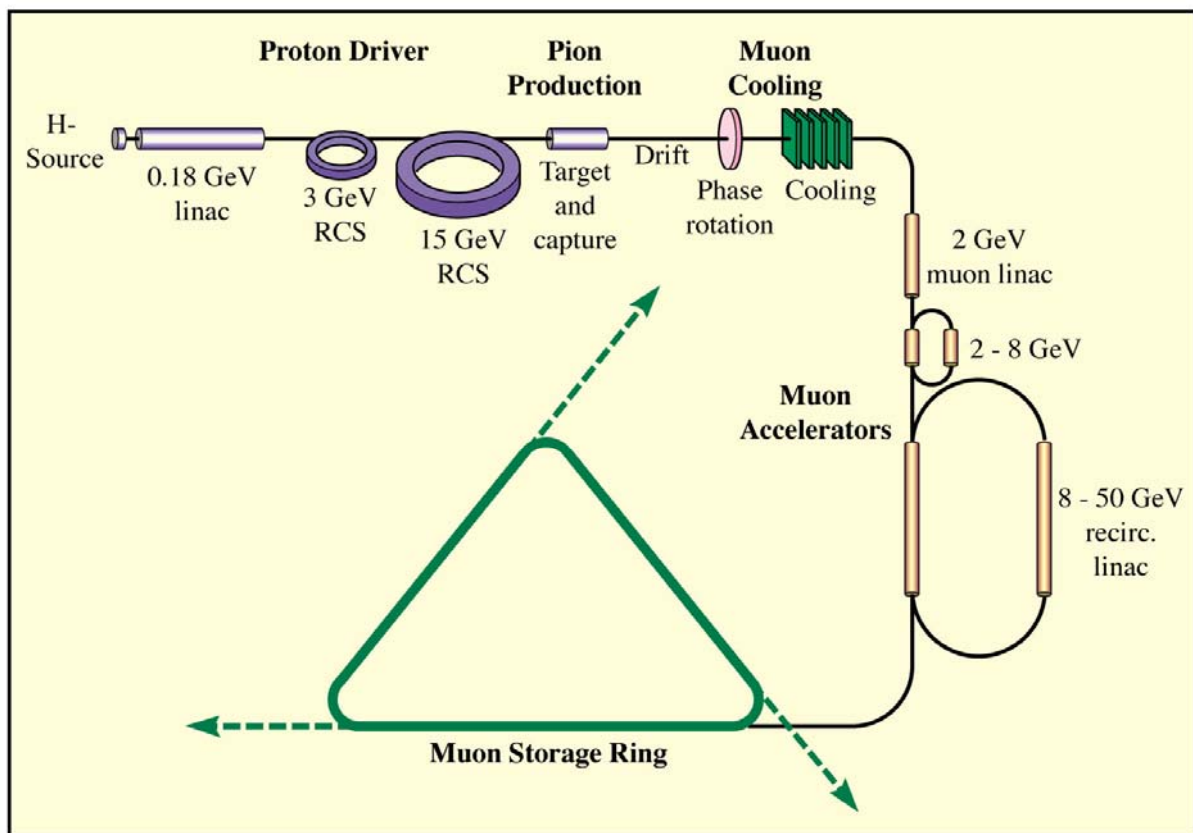
As current preliminary simulations indicate that the Neutrino Factory has the best sensitivity to the three unmeasured parameters of neutrino oscillations [5], this section will concentrate on this machine. It will give an overview of the facility, introduce the main technical challenges and describe the theoretical and experimental work that is or should be undertaken to solve these. It should be noted that a superbeam could form the first stage in a Neutrino Factory project. A possible betabeam facility is described later in this report.

1 Neutrino Factory Layout

A possible layout for a Neutrino Factory is shown in figure 1. As outlined above, the primary aim is the production of intense neutrino beams for precise long baseline neutrino oscillation measurements from the decay of muons in a storage ring. The muons are made by firing an intense proton beam into a target to make pions. As many of these pions as possible are focussed magnetically into a decay

channel in which they decay to give muons. The muons produced have a large spread in energy and this must be reduced to increase the fraction captured in the subsequent accelerators. This compression takes place in two stages: phase rotation and cooling. The muons are then accelerated in a number of steps before injection into the storage ring.

Theoretical studies suggest that the muons must have an energy of at least 20 GeV and that two different baselines are desirable, preferably one around 3000 km and another of the order of either 1000 or 7000 km [6]. Typically these studies assume a total of 10^{21} muon decays per year in the storage ring, with up to 40% being in a straight-section and hence useful for physics. These parameters determine the performance of rest of the elements of the accelerator complex and lead to many challenges in their design and construction. These, and the R&D being undertaken to solve the problems, are described in the following sections. Recommendations for additional work that must be carried out over the next few years to enable the production of a conceptual design are also given



Neutrino Factory at RAL

Figure 1: The layout of a Neutrino Factory, consisting of a proton driver, target, pion collection channel, muon frontend, muon accelerators and a storage ring.

1.1 Proton Driver

To achieve a sufficient number of muons, the proton driver should have a beam power of 4 MW, bigger than any equivalent existing machine. At this power level it is essential to minimise beam losses at higher energies where they would cause significant activation, resulting in problems with access. This imposes stringent requirements on the preparation and handling of the beam. In addition, to aid in the compression of the energy spread of the muons via phase rotation, the proton bunch should be about 1 ns long. Thus, the large proton current required to achieve 4 MW must be compressed into a very short bunch.

In Europe, two different approaches have been taken for the proton driver. In the CERN layout, a linac is used for acceleration, while buncher and compressor rings are used to produce the correct bunch structure and bunch length. At RAL, synchrotron rings are used for the bulk of the acceleration, while a linac is used only for injection. A number of designs have been studied for a variety of proton energies.

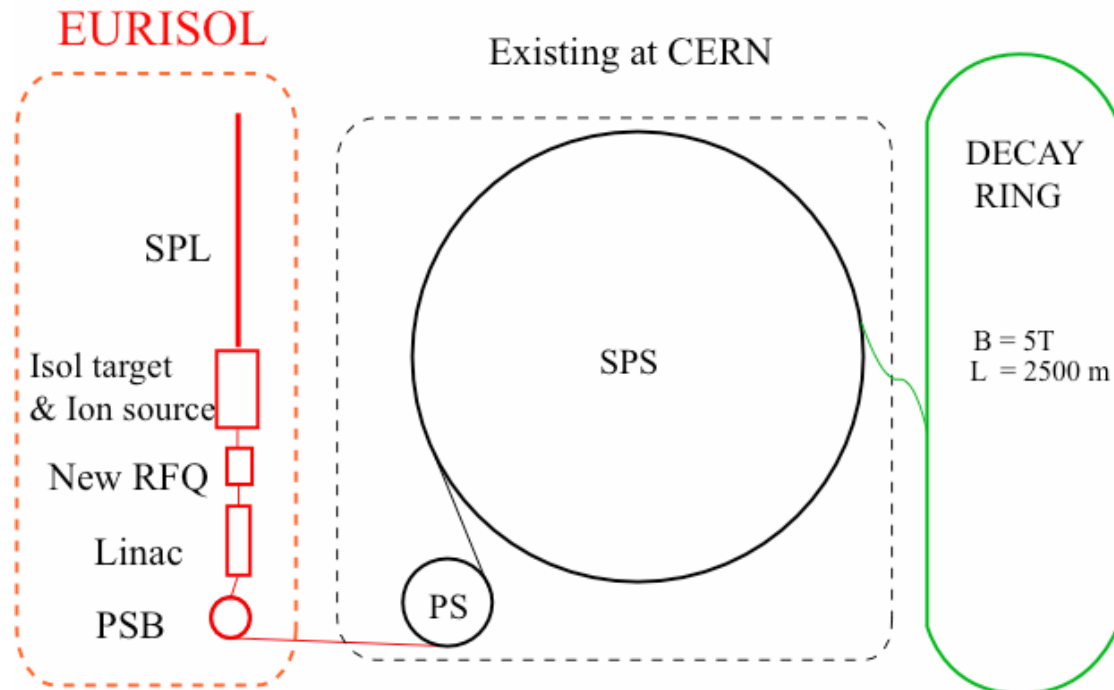


Figure 2: Layout of a Beta-beam facility at CERN, consisting of a proton driver (the SPL), an ion source, an ion acceleration system and the storage ring.

The R&D activity on the proton driver is currently focussed on the low energy components of the linac injector (see figure 3). This is because these are crucial for the preparation of the beam to avoid losses at higher energy and it is essential to check their performance at the high beam currents required. Two test stands are being constructed as part of the HIPPI project [7]: one primarily intended for HEP uses at CERN and a more generic version at RAL. Both will consist of a H⁻ ion source, Low Energy Beam Transport, RFQ and chopper and will have the necessary diagnostics to check their operation. Both should be operational within the next couple of years.

Along with this R&D, work continues on the design of the higher energy sections of the driver. In particular, a second conceptual design of the linac proposed at CERN, the SPL, is being undertaken and a new design using a non-scaling FFAG has recently been produced at RAL.

Recent studies of pion production in the target with MARS and Geant 4 indicate that the optimum proton energy might be somewhat above the current 3.5 GeV of the SPL, in the range 6-10 GeV, and give a substantial improvement in the captured pion flux, a factor of 2 or more. Thus the practicality of a higher energy proton driver at CERN needs to be investigated. In addition, studies of shock in the target show that this can be significantly reduced if the proton beam is split into a number of bunches, with a bunch spacing of about 10 μ s or more. As a result of this and a potentially higher energy proton beam, the accumulator and compressor rings in the CERN layout also need further study.

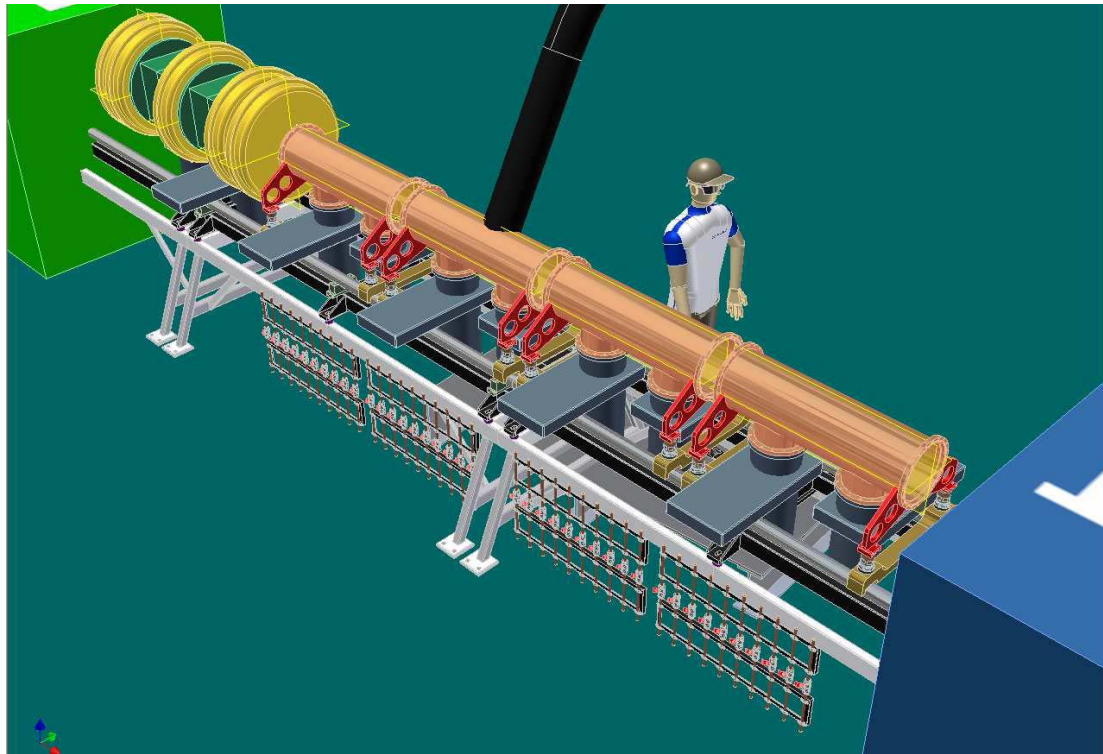


Figure 3: Drawing of the Radio-Frequency Quadrupole and chopper for the RAL proton Front-End Test Stand.

1.2 Target and Pion Collection

The target is probably the most difficult problem for a Neutrino Factory and the area that should receive the highest priority for R&D. As described above, the large proton power has to be compressed into a short bunch. In addition, to maximise the production of pions, the target cannot be too large transversely to prevent large losses from re-interactions. Thus, the energy density from the beam in the target is very large. This causes sudden heating, leading to severe stress in a solid target, and huge activation. Although targets of similar energy density already exist, they are run at a much lower frequency (~ 1 Hz compared to up to 50 Hz) and there are indications that they are damaged by the beam [8]. In addition, the activation of the target and surroundings will be huge and it is likely that the target area will require the same safety precautions as for a nuclear reactor. Safety is probably the single most difficult problem in the design of the target station and a potential show-stopper for the whole project.

As a result, R&D on the target is essential. Two types of target have been proposed and are currently under study: a liquid metal (mercury) jet and a solid, replaced frequently. For the former, a liquid mercury jet has been tested both, but separately, by impinging a high intensity beam on to it and passing it through a strong magnetic field to simulate that used to capture the pions [9]. However, the effect of a beam and the magnetic field have not been tested together and, in general, the parameters of the tests have been less stringent than required for the Neutrino Factory. An experiment, MERIT, is now under development to perform the beam and magnet tests simultaneously and will use parameters closer to those of the real situation. This will be housed in the TT2A beamline at CERN and use a beam from the PS with approximately the right energy density. It **should demonstrate** the feasibility of a liquid mercury jet as a high power target.

For solid targets, the most important problems are the lifetime due to shock induced by the sudden heating from the proton beam and the resulting temperature. Both of these are eased by changing the target between proton pulses in some way, so that different targets are exposed to different proton

bunches. Nevertheless, it is important to determine what the lifetime of such a target would be. Theoretical simulations suggest this could only be one proton pulse, though the existence of higher energy density targets implies this is not the whole story. A programme of tests on a number of materials, including tantalum, graphite and tungsten, has already started. This will use proton beams, lasers and large electric currents to induce shock and use a variety of techniques to measure the effect and determine the target lifetime.

These programmes will hopefully demonstrate the feasibility of at least one target technology. What will still be missing is the practical incorporation of the selected target technology in the target station, including a method of rapidly changing the solid targets. In addition, the design of a target station that will satisfy all the relevant safety requirements will be a major issue.

There are two possible methods for pion collection. The default is to use a 20 T solenoidal magnetic field around the target which is adiabatically ramped down to match the pion decay channel. The radius of the channel is increased to ensure that no captured pions are lost. The magnet has a superconducting outer core which is shielded from radiation by a normally conducting inner core. The alternative method is to use a magnetic horn.

The main problems for the solenoid are the heating and long term damage from radiation and its incorporation in the target station, particularly in relation to the beam dump. For the horn, there are additional problems arising from a huge current requirement, $\sim 300\text{kA}$, and the need to pulse it at up to 50 Hz.

1.3 Muon Frontend

The “muon frontend” is the section of the machine from the start of the pion decay channel to the first muon accelerator. A solenoidal decay channel is envisaged to transport the pion beam captured from the target and the large emittance muon beam resulting from the pion decay. The remainder of the frontend is used to prepare the muon beam for acceleration. In particular, the muon energy spread, and possibly emittance, must be reduced. The energy compression is performed using phase rotation, in which RF accelerating cavities are used to speed up the slower muons and to slow down the faster muons.

The emittance reduction would be performed by ionisation cooling. In this, the muons are passed through some material, call an absorber, in which they will lose both longitudinal and transverse momentum via standard ionisation energy loss. The lost longitudinal momentum is restored with RF-cavities after the absorber, resulting in a net reduction in transverse momentum and a net transverse cooling. Of course, things are never that simple and as well as cooling coming from the energy loss there is also heating coming from multiple scattering. The net cooling is a balance between the cooling and heating effects. This balance is tilted in favour of cooling by using a low Z absorber and creating a highly convergent or divergent beam using a strong magnetic field. As a result, a cooling cell would be a technically complicated device.

Due to this complication, the Muon Ionisation Cooling Experiment (MICE) is being built at RAL to show that ionisation cooling really works and to learn about it in detail.

Over recent years, the amount of cooling included in Neutrino Factory designs has decreased from more than a factor of 10 to 1.7 in the most recent study in the US, primarily due to the use of larger acceptance FFAGs for muon acceleration. As a result, one of the most important questions concerning the frontend is whether cooling is actually necessary. As well as the simplification and cost reduction coming from not having this channel, the rest of the frontend will also become easier and cheaper as it will be unnecessary to transmit such a large emittance muon beam. In addition, the existing frontend designs in Europe have fallen behind the US and, in particular, have not developed in line with the changes in the muon accelerators. It is important this problem is addressed in the near future.

1.4 Muon Acceleration

The acceleration of the muons needs to be fast, due to the muon lifetime, but the cost has to be kept under control. The requirement for rapid acceleration essentially eliminates conventional synchrotrons because of the time required to cycle the magnets (the current fastest cycling synchrotron is ISIS at RAL [10], which has a frequency of 50 Hz and hence cycle time of 20 ms). On the other hand, although acceleration via linear accelerators is fast, it is expensive for the muon energies required. As a result, two other techniques are under study: re-circulating linear accelerators [11] and Fixed Field Alternating Gradient synchrotrons (FFAGs) [12]. The former comprises a linac injector into a ring consisting of two linacs connected together with a number of separate magnetic arcs, each set for a different momentum. This re-uses each linac a number of times (4) and but avoids the need to ramp the magnets in the arcs. Nevertheless, the second US Feasibility Study showed this to be an expensive option. On the other hand, an FFAG, as originally envisaged, employs large aperture magnets with a strong gradient in the magnetic field as a function of radius (see figure 4) so that as they are accelerated, the muons move to a larger radius and see a stronger magnetic field. It again has the advantage that the magnets do not have to be ramped and that the RF-cavities are passed through many times (typically 10). In addition, a FFAG naturally has a large transverse acceptance and, if low enough frequency RF is used, it can also have a large longitudinal acceptance. As a result, it may be unnecessary to cool the muons if FFAGs are used for the acceleration. In fact, the Japanese Neutrino Factory layout utilises four FFAG rings, accelerating in steps from 300 MeV/c to 20 GeV/c, without cooling.



Figure 4: The 150 MeV scaling FFAG built at the KEK laboratory in Japan.

The FFAGs used in the Japanese layout are called scaling FFAGs. Recently, an alternative type, called non-scaling, has been invented. These have a number of advantages over their scaling cousins for the acceleration of muons. In particular:

- the non-scaling optics allows a much smaller magnet aperture
- the magnetic field variation is a linear function of the magnet radius rather than being a high power of it
- higher frequency RF can be used, ~ 200 MHz, rather than ~ 6 MHz
- it is possible to run the RF at a fixed frequency, rather than having to modulate it during acceleration, resulting in very fast acceleration

- there is a large transverse acceptance
- the accelerators are more compact

In short, non-scaling FFAGs are ideal for the muon acceleration and are rapidly becoming the default for a Neutrino Factory.

Non-scaling FFAGs are **not**, however, without problems. None have ever been built and they have a number of unique features which have never been tested before. Therefore, it is now essential that a proof-of-principle machine is built to show that non-scaling optics work, to study these features in detail and learn how to build the optimum accelerators for a Neutrino Factory. It is planned to construct an electron model of the muon accelerators at the Daresbury Laboratory in the UK to do this.

2 Recommendations

As noted above, the following are the areas that most urgently need addressing by the European Neutrino Factory community. All currently require additional resources. They are split into accelerator design work and R&D activities.

Design studies:

- 1) FFAGs: non-scaling FFAGs have become the baseline for the acceleration of muons for a Neutrino Factory. Currently, no work has been done to incorporate them into a European layout. It is essential for such work to start soon, as Europe is already far behind the US and Japan in this and they are likely to bring a significant reduction in the cost of the facility.
- 2) Muon frontend: this is another area in which little work has been done in Europe for a number of years and we are again far behind the US and Japan. The existing designs need urgent development to allow for the large acceptance FFAGs for acceleration.
- 3) Proton driver: due to the emerging information about pion production and shock in the targets, the higher energy parts of the proton driver in the CERN layout require further study. In addition, the potential of non-scaling FFAGs for proton acceleration needs understanding.
- 4) Storage ring: there has also been little work on this for a number of years. It is clearly the most crucial part of the facility as far as neutrino oscillations are concerned. The existing studies need to be updated to take into account recent requirements coming from the physics.
- 5) Once the investigations of the individual sections of the accelerator are more advanced, detailed end-to-end studies must be performed. These should lead to an overall cost optimisation of the entire complex, including the neutrino detectors.

R&D:

- 1) Target and the target station: this is the most critical item for R&D for a Neutrino Factory and other high power proton projects. It is essential that the feasibility of at least one target technology is demonstrated over the next few years and the work currently being undertaken should do this, though additional resources are required to do a thorough job. It is also critical that a serious engineering study of a target station is undertaken to show that this target can be incorporated within the station and to investigate all aspects of safety. At the moment, no work is being done on this.

- 2) FFAGs: it is essential to build and successfully operate a non-scaling FFAG to show that it works and to demonstrate the applicability to a Neutrino Factory.
- 3) Proton frontend test stands: these are essential for preparing the intense proton beam for acceleration and their successful operation is very important. The resources are in place to do this.
- 4) RF cavities and kicker magnets: these have potential application in several parts of the machine and require common R&D. In particular, due to the muon lifetime and beam emittance, large aperture, 200 MHz cavities with large gradients need to be developed for rapid muon acceleration. If the muons are to be injected into a ring, e.g. for cooling or acceleration, substantial kickers will be required due to the large emittance and relatively small ring circumference. Similar R&D for other components of the machine is likely to be identified as work advances.
- 5) MICE: if ionisation cooling is used in a Neutrino Factory, it will important to have a good understanding of how it works and how to build the cooling channel. Currently only the first phase of MICE, which will not demonstrate cooling, is funded.
- 6) Horns: if magnetic horns are used for the pion focussing, it needs to be demonstrated that they have a high enough lifetime resulting from the current and high pulsing rate and from radiation damage. There are insufficient resources do this at full current and at 50 Hz.

References

- [1] M.Mezzetto, talk at European Neutrino Group meeting, see <http://cern.ch/psdata/www/nfwg/minutes/ENG-1/mezzetto.pdf>
- [2] MINOS Collaboration, NuMI-L-375 (1998)
- [3] D.Duchesneau, hep-ex/0209082 (2002)
- [4] Y.Itow et al, hep-ex/0106019 (2001)
- [5] P.Huber et al, Nucl. Phys. B645, (2002) 3-48
- [6] A.Cervera et al, Nucl. Phys. B579, (2000) 17-55
- [7] See <http://cern.ch/mgt-hippi/>
- [8] http://www.cap.bnl.gov/mumu/conf/target-030908/tuesday_pm/Morgan.pdf
- [9] A.Fabich and J.Letry, Nucl. Instrum. Methods Phys. Res. A503, (2003) 336-339
- [10] D.J.Adams et al, see <http://cern.ch/accelconf/e00/PAPERS/THP4A08.pdf>
- [11] S.A.Bogacz, J. Phys. G29, (2003) 1723-1726
- [12] S.Machida et al, see <http://cern.ch/accelconf/e00/PAPERS/MOP1B20.pdf>

MultiMegaWatt Proton Drivers

A high power proton driver is the backbone of any future neutrino facility. Two main technical options are being studied in Europe, USA and Japan:

- Superconducting Proton Linacs (SPL)
- Rapid Cycling Cyclotrons (RCS)

In addition, recently, Fixed Field Alternating Gradient (FFAG) accelerators have been proposed as an attractive alternative¹.

The selection of the optimum technology for the proton driver producing intense neutrino beams is strongly correlated with the method of neutrino production that will be selected. Following the recent SPSC recommendations² (Villars04), Europe is pursuing its R&D on two neutrino production techniques:

- beta-beam, or neutrinos from a beta emitter beam, where beta emitters (⁶He and ¹⁸Ne) produced by a third generation radio-active ion beam (RIB) facility (driven by a 1-2 GeV SPL) are accelerated in the CERN SPS to $\gamma=100$ and sent to a decay ring to produce electron neutrinos.
- Neutrino factory, or neutrinos from a muon beam³, where muons originating from the decay of pions produced from the interaction of a high power proton driver on a Mercury target, are accelerated up to an energy greater than 20 GeV, and sent to a decay ring to produce muon and electron neutrinos.

The availability of a superbeam (neutrinos originating from decays of parents (pions and kaon) without reacceleration, produced from the interaction of a high power proton driver on a Mercury target) would benefit both options.

In terms of proton energy, neutrinos from beta-beam would only require a modest energy (1-2 GeV), a moderate energy will be optimal for a super-beam proton driver (3-4 GeV), whereas a neutrino factory would require an energy greater or equal to 5 GeV. In addition, the proton driver parameters will have to be defined through a global optimization including RIB or pion and/or muon capture and muon front-end.

We review in this paper the R&D activities regarding high power proton drivers for intense neutrino beam in Europe, USA and Japan, as they were presented at the most recent relevant conference, NUFAC05⁴.

1 Proton driver activities in Europe

1.1 Accelerator R&D for a proton driver of a neutrino factory in the UK⁵

¹ G. Rees, Nufact05,

http://www.lnf.infn.it/conference/nufact05/talks2/WG3/Rees_WG3.ppt

² J.B. Dainton, Villars report,

<http://committees.web.cern.ch/Committees/SPSC/Villars-Report.pdf>

³ muon decay weakly into an electron, a muon neutrino and an electron antineutrino

⁴ <http://www.lnf.infn.it/conference/2005/nufact05/>

⁵ C. Prior, Nufact05,

http://www.lnf.infn.it/conference/nufact05/talks2/WG3/Prior_WG3.pdf.

The UK is actively supporting accelerator studies and developments aiming both at the upgrade of the ISIS facility and at a UK-based neutrino factory (Fig 1).

A classical means of accelerating particles to high energy is through the use of synchrotrons. To reach high beam power, they must be “Rapid Cycling Synchrotrons” (RCS), equipped with fast ramping magnets and power supplies, ceramic vacuum chambers to avoid Eddy current and unusually high RF power for acceleration.

Such proton drivers for a neutrino factory are under study at the Rutherford Appleton Laboratory, complementing a similar investigation using a high energy linac plus accumulator and compressor rings at CERN (see below). Two models are being considered:

- Driver I: a 5 GeV, 4 MW driver with 4 bunches of $2.5 \cdot 10^{13}$ protons per pulse at 50 Hz (Fig 2)
- Driver II: a 15 GeV, 4 MW driver (upgradeable to 5 MW) with 6 bunches of $1.11 \cdot 10^{13}$ protons per pulse at 25 Hz (Fig 3)

The linac is common to both designs. In both cases, the 180 MeV linac beam of negative hydrogen ions is accumulated into the first rapid cycling synchrotron (RCS) using charge exchange injection.

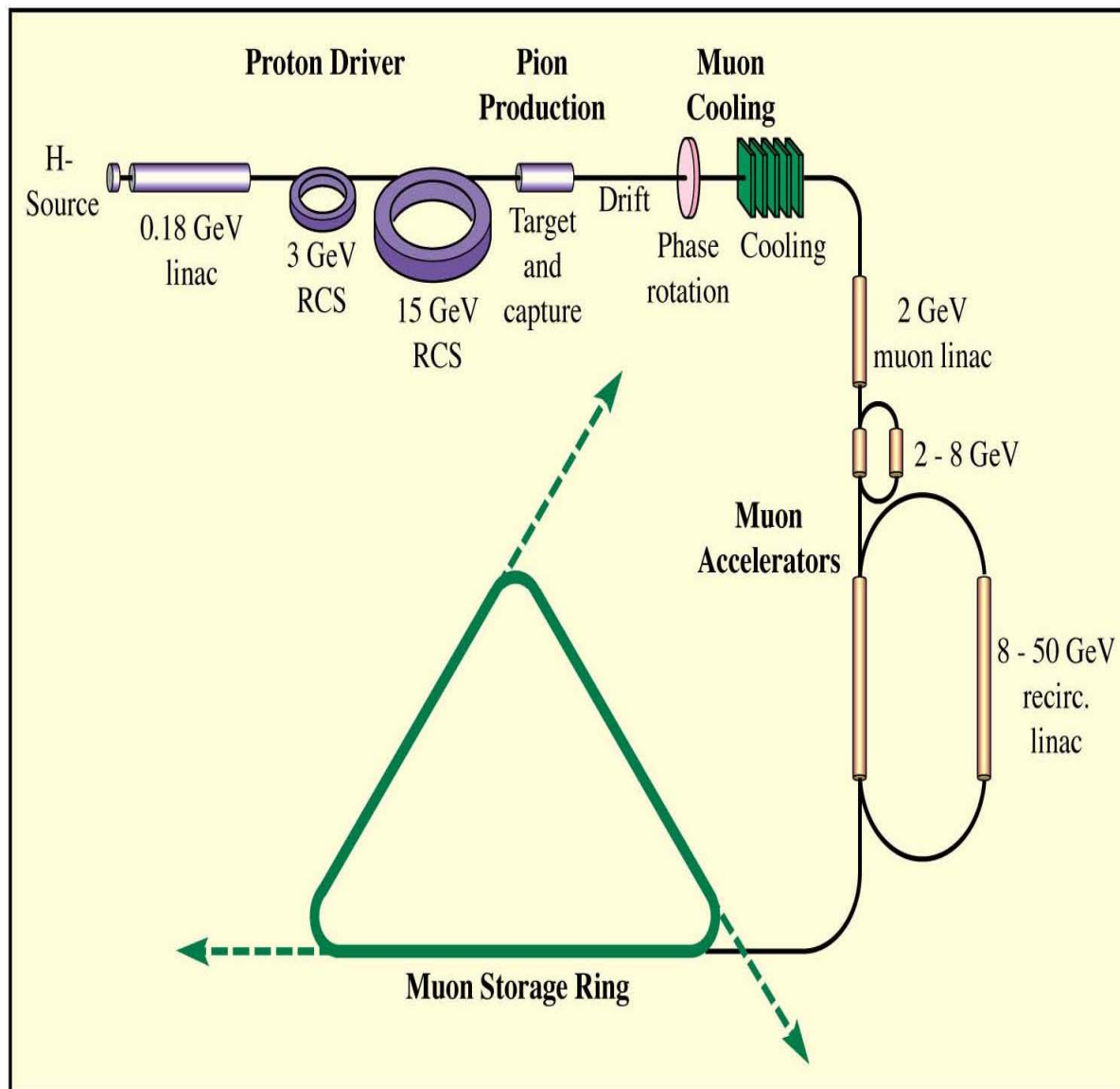
The main requirement of the driver is to produce bunches with a longitudinal emittance of $\epsilon_L \sim 1 \text{ eV}\cdot\text{s}$ and of 1 ns rms duration at ejection. The low emittance is best obtained with a linac energy in the range 150-200 MeV. Injection presents different problems than compression and is carried out in separate booster rings. The design of the high-energy synchrotrons is determined by the method of final bunch compression. 5 GeV is considered the lowest practical energy at which compression can be achieved. However, a crude figure of merit for the pion target is given by the product of output energy \times repetition rate for a fixed driver power, and this suggests some bias towards the 15 GeV option.

The main design issues to be addressed in an RCS proton driver are:

- getting 4MW of beam power at a repetition rate appropriate for the pion target
- minimization of beam loss, keeping uncontrolled loss below 1 W/m:
 - beam instabilities (electron cloud)
 - space charge effects
- beam accumulation, namely:
 - balance between linac current and number of injection turns
 - design of the charge exchange injection (phase space painting, foil heating & life-time, H₀ dump...)
- bunch compression at the nanosecond level.

The initial design of the UK Neutrino Factory linac was based on the low energy part of European Spallation Source linac, with an additional DTL section to 180MeV. A subsequent design made in 2004 was based on a DTL to 90 MeV at a frequency of 234.8 MHz, followed by an Side-Coupled Linac (SCL) section operating on the third harmonic (704.4 MHz). This SCL is common to the CERN scheme for its linac4 (see below). In the most recent design, the Japanese JPARC accelerator frequency of 324 MHz is preferred at low energy, and its third harmonic (972 MHz) is used at high energy, with commercial TOSHIBA klystrons as RF sources.

The injector linac has a length of 129 m and accelerates a 57 mA H- beam to 180 MeV. The transfer line to the rings includes a four period 180 degrees achromatic arc with reverse bends to give a high peak dispersion for momentum collimation in a limited space. Two buncher cavities provide momentum ramping for injection painting.



Neutrino Factory at RAL

Fig 1 – Neutrino Factory at RAL

Some details of the two RCS designs are given below.

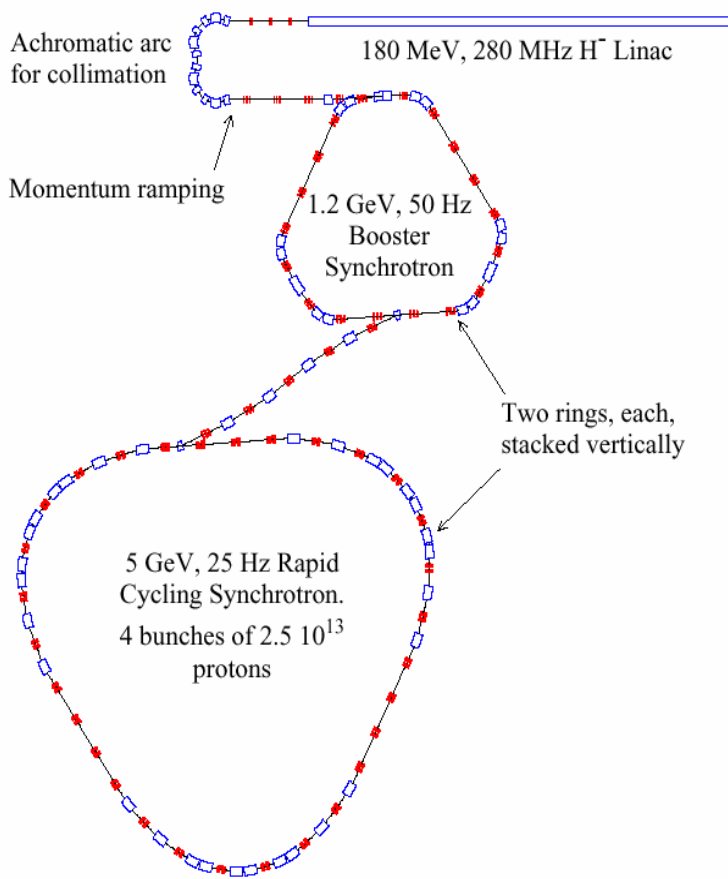


Fig 2 Schematic of the 5 GeV, 50 Hz RCS Design.

1. 5 GeV, 50 Hz 4MW RCS design

Beam from the linear accelerator is injected into two 50 Hz/1.2 GeV/200 m circumference synchrotrons, operating almost in phase and stacked one over the other. The synchrotrons accelerate the beam with a peak RF voltage of 0.25 MV on $h=2$. This generates two bunches per ring which are extracted sequentially and injected in alternate cycles into a pair of 5 GeV synchrotrons operating at 25 Hz and $h=8$. The 5 GeV rings are also vertically stacked, but operate in phase opposition. The beam is accelerated to 5 GeV with maximum voltages of 0.575 MV. In the final stages of the acceleration process, the beam is compressed in time by the addition of 0.5 MV on $h=24$. The combined output of the latter two rings is (on target) 5 GeV at 50 Hz.

2. 15 GeV, 25 Hz, 4MW RCS Driver

The scheme is similar to the 5 GeV design but with half the repetition rate. The linac beam is injected into the first pair of rings which simultaneously accelerate to 3 GeV on $h=3$. The 3 GeV beam is alternatively send to the 15 GeV rings which operate in phase opposition at 12.5 Hz. They accelerate on $h=36$ and their combined outputs are sent onto the target at 25 Hz. The required short bunch duration results from adiabatic compression of the bunch. This design can be upgraded to 5MW.

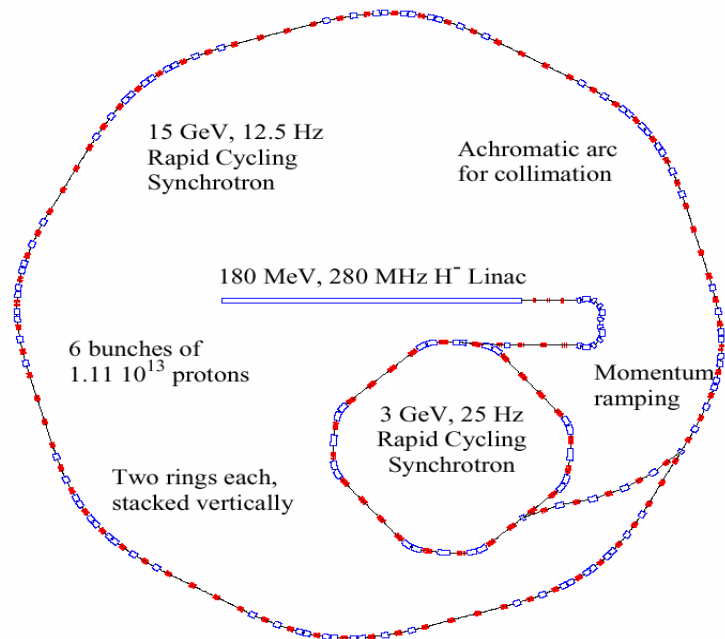


Fig. 3 Layout of the 15 GeV, 25 Hz RCS Design.

In both designs, the H- linac beam is chopped with a 70% duty factor at the ring revolution frequency. It is stripped passing through an Al_2O_3 foil and painting is done in all phase planes of the booster synchrotron. The linac beam pulse to fill one ring is 200 μs (Driver I) or 134 μs (Driver II) long, and the second ring is filled immediately after the first one. A normalised dispersion $\Delta x/\beta\sqrt{x}=1.8$ couples the painting into transverse horizontal phase space, while vertical orbit bump magnets help reduce foil traversals by the circulating beam.

In both designs, the bunches from the two vertically-stacked booster synchrotrons are extracted and transferred together to either the upper or lower of the two main synchrotrons. They are then accelerated to top energy (5 GeV or 15 GeV) and, as soon as compressed, ejected onto the pion target which is then supplied with beam at twice the repetition rate of a single ring.

Final bunch compression to 1 ns (rms) is carried out over the last 500 revolutions of the accelerating cycle. At top energy the rings are close to, but just below, transition energy, and the lattices are designed so as to minimise the depression in γ_t under transverse space charge. Longitudinal inductive impedances of 5-10 Ω for Z/n are envisaged for the metallic and ceramic vacuum chambers to reduce longitudinal space charge effects. In Driver I, a higher order harmonic voltage ($h=24$) is raised progressively to assist the compression, which is achieved with a peak voltage of $\sim 0.6\text{MV}$. Driver II relies on adiabatic bunch compression at a single harmonic ($h=36$) to achieve similar results. The final longitudinal emittance is $\epsilon_L \sim 1.0$ eV.s, with $\Delta p/p \sim 1.6\%$ or 1.0 % in the 5 and 15 GeV rings respectively. Chromaticity is corrected with an arrangement of sextupoles and these can also be used to reduce second-order momentum effects.

1.2 Accelerator R&D for the proton driver of a neutrino complex at CERN⁶

During the CERN council meeting in December 2004, the Director General of CERN outlined the “seven point CERN strategy” :

1. Completion of the LHC project on schedule.
2. Consolidation of existing infrastructure at CERN to guarantee reliable operation of the LHC.
3. An examination of a possible future experimental programme apart from the LHC.
4. A role for CERN in the growing coordination of research in Europe.
5. The construction of a linear accelerator injector at CERN to provide more intense beams for the LHC.
6. An accelerated R&D effort towards CLIC, CERN’s novel new accelerator technology, which could open the way to much higher energies than are available today.
7. A comprehensive review of CERN’s long-term activity to be available by 2010, when results from the LHC will have given a first description of the particle physics landscape for years to come.

The following events that happened in 2005 are also relevant for the proton drivers activity devoted to a neutrino complex at CERN :

- Publication of SPSC Villars’s report⁷

⁶ R. Garoby, Nufact05,

http://www.lnf.infn.it/conference/nufact05/talks2/WG3/Garoby_WG3.ppt

⁷ <http://committees.web.cern.ch/Committees/SPSC/Villars-Report.pdf>

- Continuation of the negotiation for additional resources from non-member states (India & China) in favour of Linac4, the low energy front end of the SPL, which should replace the existing CERN Linac2.
- Announcement of a special INTC meeting (NuPAC⁸ – Geneva, Oct. 2005) about the needs of ISOLDE and nToF for the next decade, which will provide a new set of recommendations, including on proton drivers for such facilities.
- Creation of two CERN working groups⁹ (POFPA: “Physics Opportunities with Future Proton Accelerators” and PAF: “Proton Accelerators of the Future”) to analyze scenarios and contribute to the definition of a baseline scheme for 2010. Relevant excerpts of the PAF mandate are given in Annex.

1.2.1 R&D activities for the SPL at CERN

R&D activities related to the Super Proton Linac (SPL, Fig 4)) as a proton driver for a neutrino complex at CERN are organized around 4 main projects:

- The SPL Study, with contributions from KFZ Juelich, CERN, RAL, GSI, Tokyo Univ., LANL.
- The IPHI-SPL collaboration : CEA (DSM/DAPNIA @ Saclay) + CNRS (IN2P3 @ Orsay & Grenoble)
- The HIPPI JRA (inside CARE, supported by the European Union) : CEA (F), CERN (CH), Frankfurt University (D), GSI (D), INFN-Milano (I), IN2P3 (F), RAL (GB), KFZ Juelich (D)
- ISTC projects #2875, 2888 and 2889 : BINP (Novosibirsk), IHEP (Protvino), IHEP (Moscow), VNIIEF (Sarov), VNIITF (Snezinsk)

The original characteristics of the SPL, as published in the Conceptual Design Report¹⁰, were “optimized” for a neutrino factory and assumed the use of LEP cavities & klystrons up to the highest energy. They have been revised (Table 1) taking into account the latest physics requests, and assuming the use of state-of-the-art bulk Niobium cavities above 180 MeV. A second conceptual design report is therefore in preparation (CDR-2) in collaboration with scientist from the CEA-Saclay & the INFN-Milano. It should be published at the end of 2005. Up-to-date information is available at <http://project-spl.web.cern.ch/project-spl/>.

The RF system will be based on 704 MHz bulk Niobium cavities. Three families of cavities will be used, centered at $\beta=0.5, 0.85, 1.0$, with accelerating gradients of 15, 18, 30 MV/m and using 5, 6 and 7 cells per cavity.

To minimize cold/warm transitions and maximize real estate gradient, quadrupoles will be cold (2 K) because imbedded in the cryomodules, and independently aligned from the cavities. The design assumes cryomodules of maximum length (between 10 and 15 m), containing n cavities and (n+1) quadrupoles. Beam Diagnostics, steering etc. will be located between cryomodules. The length of the cavities is limited by fabrication and handling considerations. The proposed number of cells per cavity is therefore 5, 6 and 7 for the three sections. The SPL will be built assuming 2 MW maximum power per coupler.

⁸ <http://indico.cern.ch/conferenceDisplay.py?confId=a052>

⁹ <http://paf.web.cern.ch/paf/>

¹⁰ Conceptual Design Report of the SPL, <http://doc.cern.ch/yellowrep/2000/2000-012/p1.pdf>

Table 1 The CDR2 SPL characteristics

Ion species	H-
Kinetic energy	3.5 GeV
Mean current during the pulse	40 mA
Mean beam power	4 MW
Pulse repetition rate	50 Hz
Pulse duration	0.57 ms (0.76 ms?)
Bunch frequency	352.2 MHz
Duty cycle during the pulse	62 (5/8)%
rms transverse emittances	0.4 π mm mrad
Longitudinal rms emittance	0.3 π deg MeV

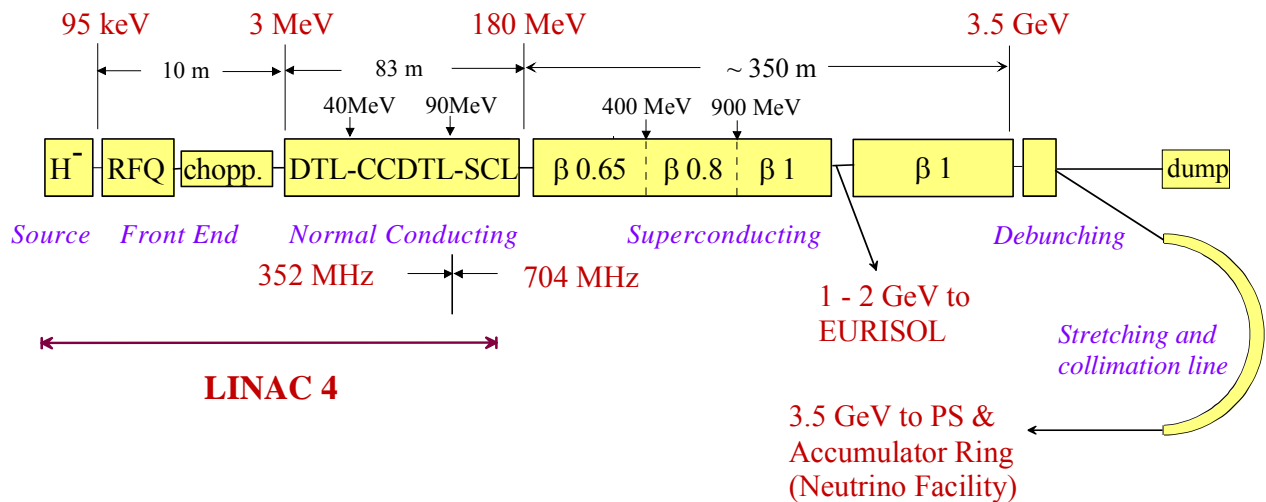


Fig 4 – The SPL

The main goals of the SPL at CERN are to:

- increase the performance of the CERN high energy accelerators (PS, SPS & LHC)
- address the needs of future experiments with neutrinos and radio-active ion beams.

The present R&D program concentrates on the low-energy items (Linac4), wherever possible in collaboration with other laboratories.

1.2.3 Linac4 related activities

In this context, Linac4, a new injector for the CERN booster synchrotron, is proposed to improve the beam delivered to the LHC, ease operation, help reach the ultimate luminosity, and increase the flux to ISOLDE.

With a length of ~90m, it will be located in an existing experimental hall and will extensively re-use LEP RF equipment (klystrons etc.)

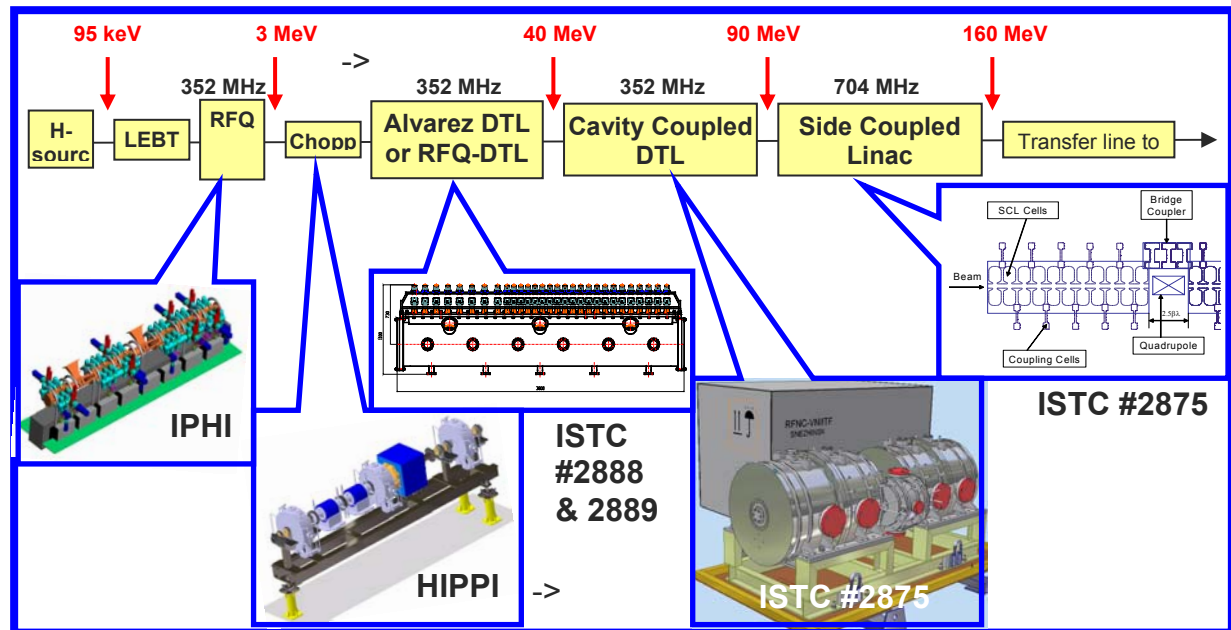


Fig 5 – LINAC4

The main collaborations related to Linac4 are

- **IPHI collaboration**: between CEA, CERN and IN2P3 whose main goal is the construction of the 3 MeV – high duty factor RFQ to be delivered at the end of 2007.

- **ISTC project #2875¹¹**: collaboration between BINP (Novosibirsk), VNIITF (Snezinsk) and CERN whose aims are:

- The development of the technological basis for serial production of CCDTL structures in the energy range of 40-100 MeV for the SPL project.
- The feasibility study of effective application of normal conducting SCL structures up to the energy of 150-180 MeV.

This ISTC project, which costs k\$ 550, has started in July 2004. The Prototype of CCDTL structure should be delivered at CERN for high power testing by mid-2006, whereas the technological model of SCL structure is to be tested in 2006.

- **ISTC project #2888¹²**: collaboration between ITEP (Moscow) + VNIIEF (Sarov) and CERN for the development of the technological basis for construction of Alvarez-type structures for the room-temperature part of the CERN SPL. After approval in February 2005 (for a cost of k\$ 460), the prototype of the first DTL tank (3-10 MeV) should be delivered at CERN for high power testing by the end of 2006. In the frame of this project, only one drift tube will be equipped with a Permanent Magnet Quadrupole (PMQ) and the prototype will not be capable to accelerate beam.

- **ISTC project #2889¹³**: collaboration between IHEP (Protvino) + VNIIEF (Sarov) + CERN to design and manufacture a DTL-RFQ accelerating structure prototype for a 3-40 MeV H- linac of the SPL project. At a cost of k\$ 477, this project was also approved in February 2005. A prototype of RFQ-DTL structure is expected at CERN for high power testing by the end of 2006.

¹¹ <http://science.istc.ru/ISTC/sc.nsf/html/projects.htm?open&id=2875>

¹² <http://science.istc.ru/ISTC/sc.nsf/html/projects.htm?open&id=2888>

¹³ <http://science.istc.ru/ISTC/sc.nsf/html/projects.htm?open&id=2889>

Other collaborations are currently in negotiation:

- with China (IHEP – Beijing), concerning a 352 MHz buncher, Linac4 quadrupoles, and Linac4-PSB transfer line magnets,
- with India (BARC-Bombay & CAT-Indore) concerning pulsed power supplies for the LEP klystrons.

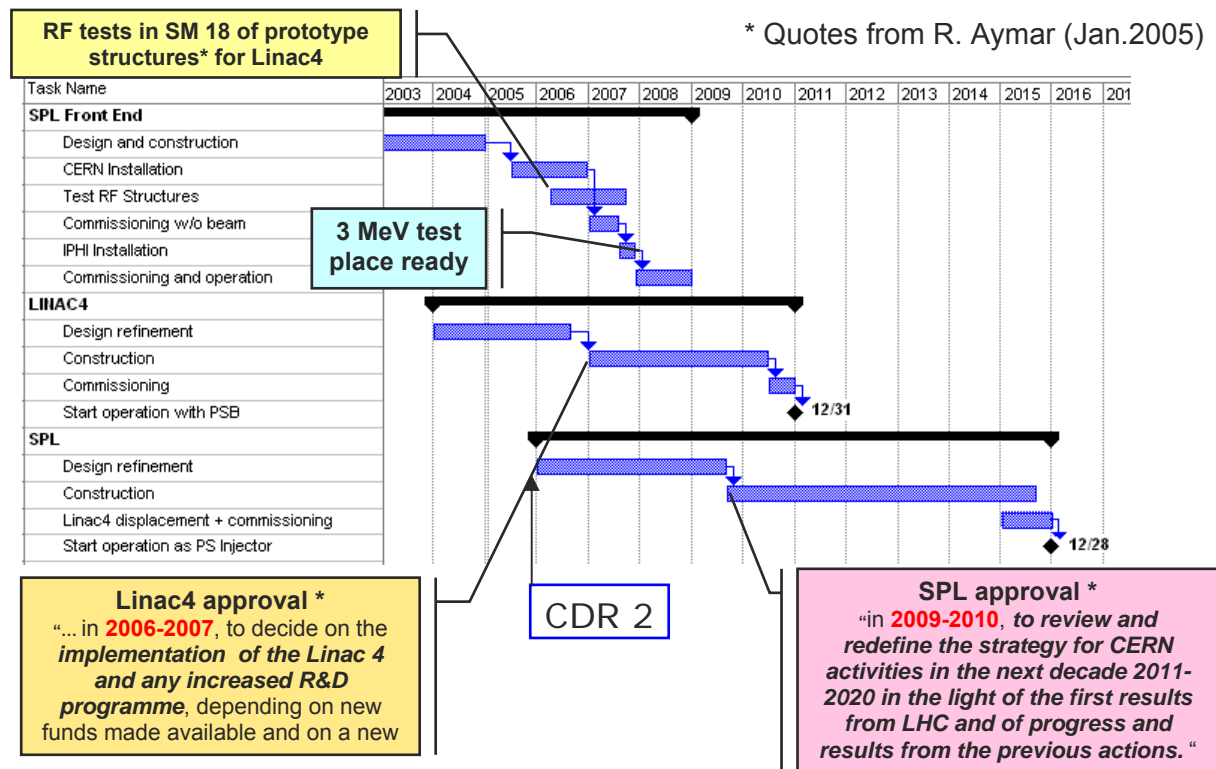


Fig 5. The LINAC4 and SPL planning

1.2.4 Other potential applications of the SPL R&D

Nuclear Physics application: there exists a strong demand to establish in Europe Radioactive Ion Beam (RIB) facilities, with intensities 3 orders of magnitude higher than presently available, for nuclear and solid state physics, biophysics, nuclear astrophysics, etc. Two parallel programs are under study: in-flight at GSI and Ion Separation On-Line (ISOL) in a new laboratory. In the ISOL-type facility, radioactive ions will be produced by a high-intensity 1-2 GeV proton beam hitting either a direct production target or a converter target irradiating the production material with spallation neutrons. For direct production, 100 kW of proton beam power are sufficient, while up to 5 MW are required when using a converter. The driver would operate in a CW or pulsed mode, at a frequency higher than 50 Hz.

In the frame of the EU funded R&D for a European third generation RIB machine (EURISOL Design Study) the preferred option is to have a dedicated driver, which would then operate continuously, with a Normal Conducting 5 MeV Front-end at 352 MHz followed by an Intermediate energy SuperConducting part at 704 MHz. The alternative option of using a pulsed linac like the SPL is also analyzed.

NEUTRON production application

Superconducting linacs are the preferred sources of protons to produce intense fluxes of neutrons for:

- basic science and condensed matter studies (as in the US SNS project (Fig 6), see www.sns.gov, that is hosting the highest power (1.4 MW), proton linac)
- driving subcritical reactors and “burn” nuclear waste and possibly generate electricity (Accelerator Driven Systems)...



Fig 6 The US SNS proton linac

Regarding ADS, the accelerator group working for the EU program, XADS, has investigated the characteristics of a transmutation driver for 600 MeV, 6 mA, and less than 5 trips/year. The conclusion is that CW SC linacs are preferred to cyclotrons because of their stronger reliability and upgradeability. An Integrated Project, EUROTRANS now submitted to EU, to study reliability and build a demonstrator will seek funding for construction in next FP (2008)

3 Proton driver activities in USA¹⁴

In matter of neutrino physics, the recent US events are on one side the publication of the results of the American Physical Society on the Physics of Neutrinos. The report made public at the end of 2004, “the neutrino matrix”¹⁵, underlines clear recommendations for the future:

“We recommend, as a high priority, a phased program of sensitive searches for neutrinoless nuclear double beta decay”. [...] We recommend, as a high priority, a comprehensive U.S. program to complete our understanding of neutrino mixing, to determine the character of the neutrino mass spectrum and to search for CP violation among neutrinos.

This comprehensive program would have several components: an experiment built a few kilometers from a nuclear reactor, a beam of accelerator-generated neutrinos aimed towards a detector hundreds of kilometers away, and, in the future, a neutrino ‘superbeam’ program utilizing a megawatt-class proton accelerator. [...]

And on the other side, the launching in march 2005 by the DOE and the NSF of a subcommittee of HEPAP and NSAC on neutrino physics, NUSAG for a duration of two years, whose charge is to “make recommendations on the specific experiments that should form part of the broad US neutrino science program”.

In the USA, two neutrino superbeam proposals have been made: at Brookhaven (Fig. 7), based on the AGS and a 1.2 GeV superconducting proton linac, and at Fermilab (Fig. 8), based on a new 8 GeV superconducting proton linac.

¹⁴ G. Apollinari, Nufact05,

http://www.lnf.infn.it/conference/nufact05/talks/Plenary/Apollinari_Plenary.ppt and <http://protondriver.fnal.gov/>

¹⁵ <http://www.aps.org/neutrino/>

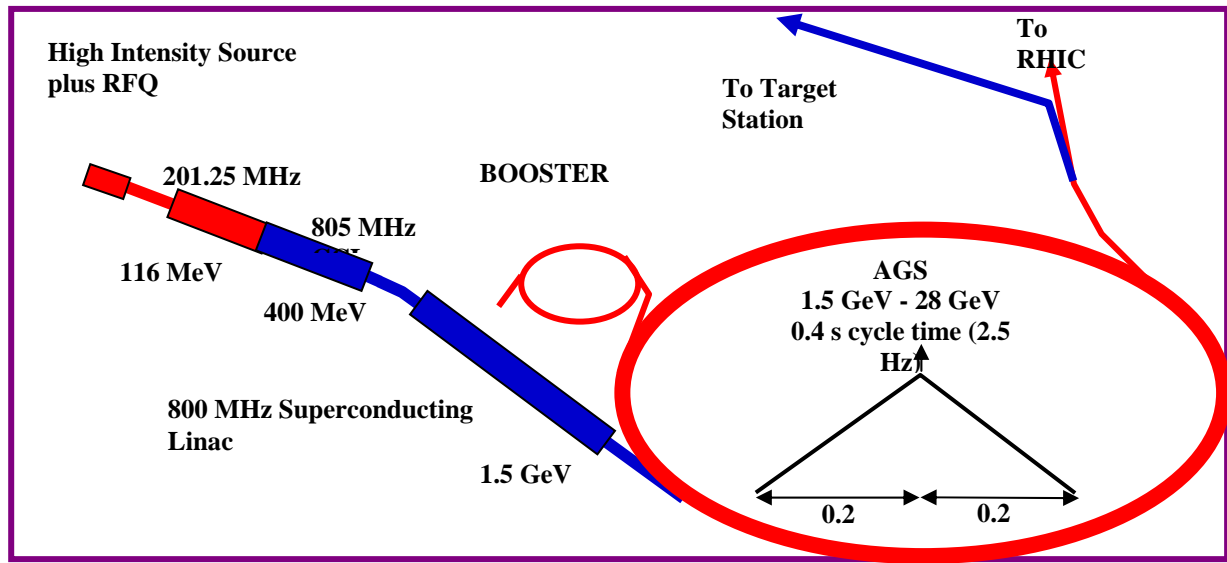


Fig 7. Upgrade plans at BNL

It is based on upgrading the existing (in red) facilities at Brookhaven by adding a Coupled Cavity Linac (CCL) to the existing 116 MeV Linac, and bring the energy up to 400 MeV, and continue with a Superconducting linac to reach 1.5 GeV. Only one type of cavity, cryomodule and klystron, similar to the one in operation for SNS, will be used. The AGS Repetition rate has to be increased from 1/3 Hz to 2.5 Hz, which implies to triple the existing main magnet power supply and current feeds and double the RF power and accelerating gradient. More details are given in Table 2 and 3.

Table 2 Comparison between the existing AGS parameters, with the upgraded AGS and the Japanese JPARC project.

	AGS (now)	AGS(1MW)	JPARC
Total beam power [MW]	0.14	1.00	0.75
Injector Energy [GeV]	1.5	1.2	3
Beam energy [GeV]	24	28	50
Average current [μ A]	6	36	15
Cycle time [s]	2	0.4	3.4
No. of protons per fill	0.7×10^{14}	0.9×10^{14}	3.3×10^{14}
Ave. circulating current [A]	4.2	5.0	12
No. of bunches at extraction	6	23	8
No. of protons per bunch	1×10^{13}	0.4×10^{13}	4×10^{13}
No. of protons per 10^7 sec.	3.5×10^{20}	23×10^{20}	10×10^{20}

Table 3 Characteristics of the 1.2 GeV BNL Superconducting linac, based on the SNS experience, are in table 3

Beam energy	0.2 → 0.4 GeV	0.4 → 0.8 GeV	0.8 → 1.2 GeV
RF frequency	805 MHz	1610 MHz	1610 MHz
Acc. Gradient	10.8 MeV/m	23.5 MeV/m	23.5 MeV/m
Length	37.8 m	41.4 m	38.3 m
Beam power (exit)	17 kW	34 kW	50 kW

According to Fermilab management¹⁶, FNAL will be the center of US HEP in ~2010. The strategy will depend strongly on the output of the on going Global Design Effort for the International Linear Collider (ILC). If the **ILC Cost** published in the CDR in 2006 is considered affordable, construction could start at FNAL as soon as 2010, together with a neutrino program based on the main injector (120 GeV/1 MW). If not, the construction of a superconducting 8 GeV Proton Driver could start as early as 2008. It would be followed by implementation of 30-120 GeV and 8 GeV beams of 2-4 MW after 2012. This would constitute a stepping-stone to a delayed construction of the ILC, starting in ~2012.

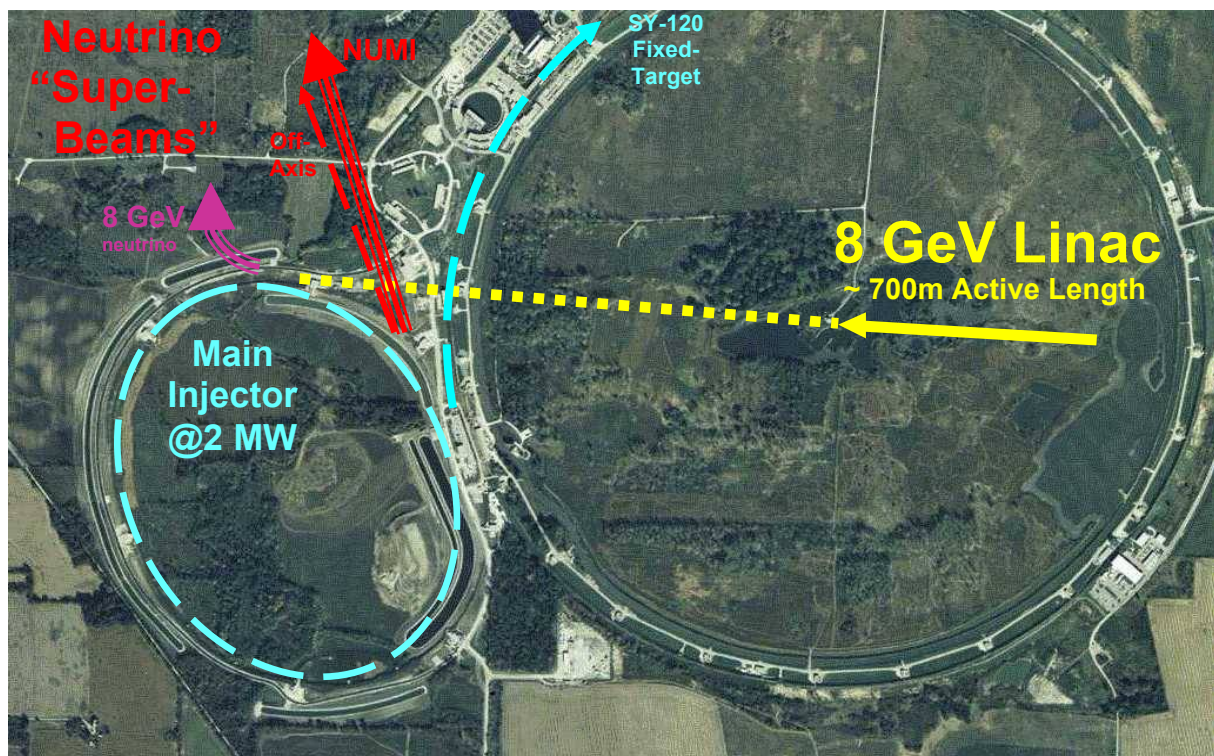


Fig 8. The Fermilab p complex with a 8 GeV SPL

The Fermilab proton driver design will benefit from new ideas and concepts emerging from the ILC study, the Spallation Neutron Source construction, and the RIA and APT projects. It is inspired by the SNS, RIA, and JPARC Linac designs up to 1.3 GeV, and uses the ILC Cryomodules in the energy range 1.3 - 8 GeV. Charge exchange injection is used to accumulate the 8 GeV H- beam in the main injector. The small emittances delivered by the SC linac proton driver will assure small losses in the following machine.

“Super Beams” at Fermilab will be based on 2 MW Beam power at 8 GeV with the linac alone, and up to 120 GeV when using the main injector.

The Fermilab linac design presents two phases :

- Phase A: 0.5 MW beam power ($8.3 \text{ mA} \times 3 \text{ msec} \times 2.5 \text{ Hz} \times 8 \text{ GeV} = 0.5 \text{ MW}$) requiring 12 Klystrons.

¹⁶ http://www.fnal.gov/directorate/Longrange/Long_range_planning.html

- Phase B: 2 MW beam power (25 mA x 1 msec x 10 Hz x 8 GeV = 2.0 MW) requiring 33 Klystrons.

Either phase is able to get 2 MW of beam power from the main injector ($1.5 \cdot 10^{14} \times 0.7 \text{ Hz} \times 120 \text{ GeV} = 2 \text{ MW}$).

Moreover, there is a very valuable synergy (Fig 9) between the superconducting proton linac and the International Linear Collider program at Fermilab. The technology is claimed to be the same in the energy range 1.5 - 8 GeV. Therefore, for a small percentage of the total cost of the ILC, this machine would be an excellent demonstration of the ILC technology and would represent a very valuable basis for a reliable cost estimate. In addition, it would be a seed project for U.S. industrialization of SCRF. Indeed, the building block of the 8 GeV Linac is the TESLA RF Station, made-up of 36 SCRF cavities, 4 Cryomodules, 1 Klystron and 1 Modulator, to reach ~1 GeV of Beam Energy.

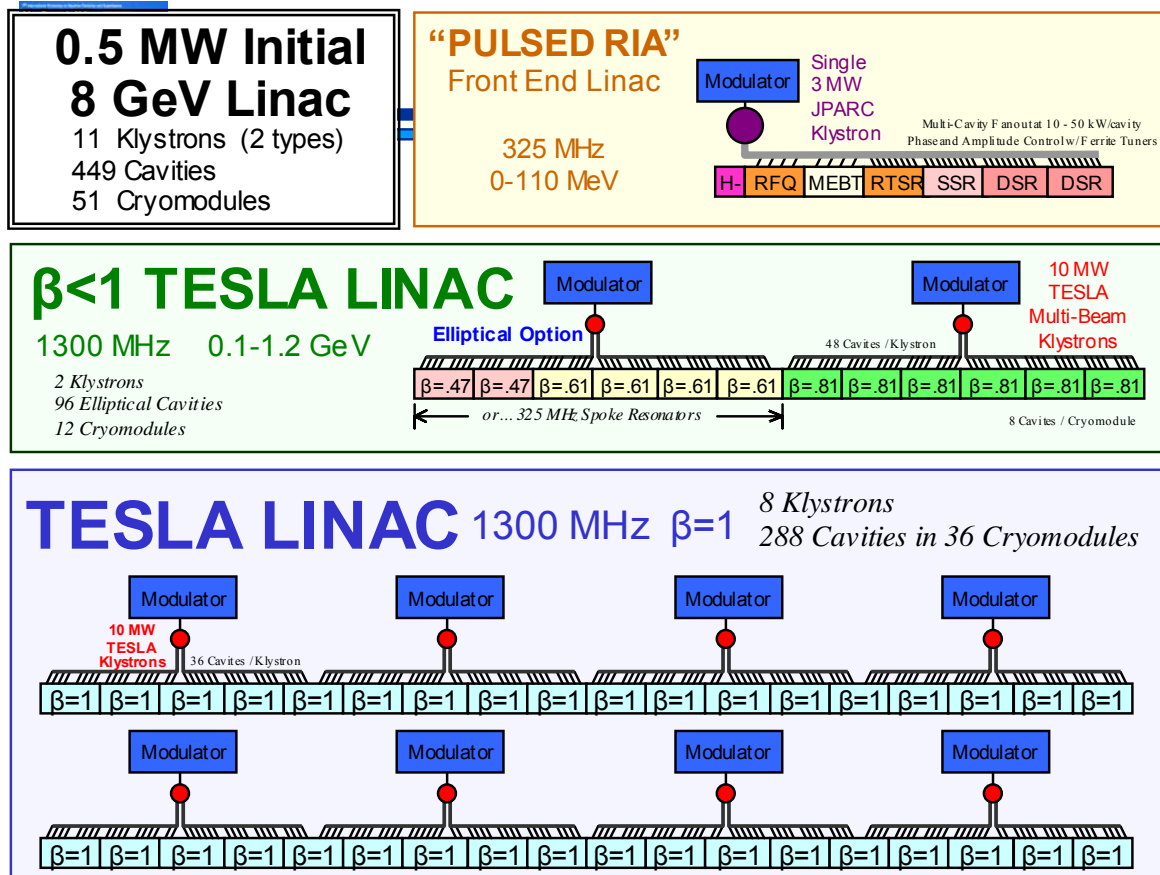


Fig 9. Possible synergy at Fermilab between SPL and ILC

Following the selection of the cold SCRF Option for the ILC, FNAL has chosen TESLA/XFEL compatible frequencies:

Main Injector Beam	($1.5 \cdot 10^{14}$, 1.5 sec, 2 MW)
Pulse Parameters	(8 mA x 3 msec x 2.5 Hz)
Ultimate Upgrade	(25 mA x 1 msec x 10 Hz)
Operating Frequency	(1300 MHz / 325 MHz)
Copper to SCRF transition	(15 MeV)
Spokes-to-Elliptical transition	(110 – 400 MeV)

- 1300 MHz Main Linac (= ILC TESLA / XFEL)

- 325 MHz (=1300MHz/4) Front-End Linac (= JPARC) (a gift !)
to take benefits from the multi-year R&D on SRF cavities, RF couplers, cryomodules, klystrons etc ...

2 Proton driver activities in Japan¹⁷

In Japan, proton driver activities are related to the construction of the JPARC multi purpose facility (Fig 10, 11, 12). The accelerator complex consists of a linac, a 3 GeV synchrotron, and a 50 GeV synchrotron. Four experimental areas are foreseen, focused respectively on nuclear transmutation, materials and life science, hadron beam physics, and neutrino generation.

They correspond to the three main goals of the JPARC complex. One is nuclear and particle physics using 50 GeV beams, that includes neutrino physics. Another is materials and life science using 3 GeV beams, and the other is R&D towards transmutation using 0.6 GeV linac beam. Phase I of the project, that excludes the nuclear transmutation facility and the superconducting part of the linac, is expected to be completed at the end of JFY2007.

Commissioning of the NC Linac (Fig 10) has already started at KEK Tsukuba where it has been designed and built. The measured transmission through the first DTL was 100%. Duty cycle is still lower than nominal, but there is no obstacle preventing from reaching the design goal.



Fig 10. The JPARC normal conducting linac

Emittance from the linac almost meet the design goal. Other crucial components such as RF chopper and S-DTL are fabricated and tested. Linac installation in the JAERI site has started.

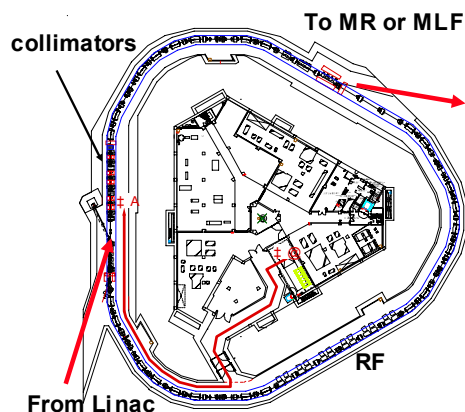
¹⁷ G. Apollinari, Nufact05,

http://www.lnf.infn.it/conference/nufact05/talks/Plenary/Apollinari_Plenary.ppt

3 GeV Synchrotron (RCS)

- Rapid Cycle (25Hz)
- Ceramics vacuum chamber
- Stranded conductor coil for B and Q magnets
- High field MA loaded cavity
- Long lived carbon foil for charge exchange injection

•Circumference	348.3m
•Repetition	25Hz(40ms)
•Injection Energy	180/400 MeV
•Output Energy	3GeV
•Beam Power	0.6/1MW
•Harmonic	2
•Bunch Number	2
•Nominal Tune	(6.72, 6.35)
•Transition γ_t	9.14
•S.C. Tune Shift	-0.2



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The next machine is the 3 GeV synchrotron. To reduce Eddy current effect in the magnets, aluminium stranded coil has been developed with radiation-hard insulation. Ceramic vacuum chambers with RF shield outside and TiN coating inside have been developed and tested. High power test of RF cavity using magnetic alloy was successful. For the RF cavity, direct water cooling method is adopted. Tunnel and buildings are ready for RCS installation.

More than half of the magnets of the 50 GeV synchrotron have been measured. Bipolar kickers and septum for fast extraction are now under fabrication.

In the initial phase, for cost saving reasons, the linac energy will be lower than nominal (200 instead of 400 MeV).

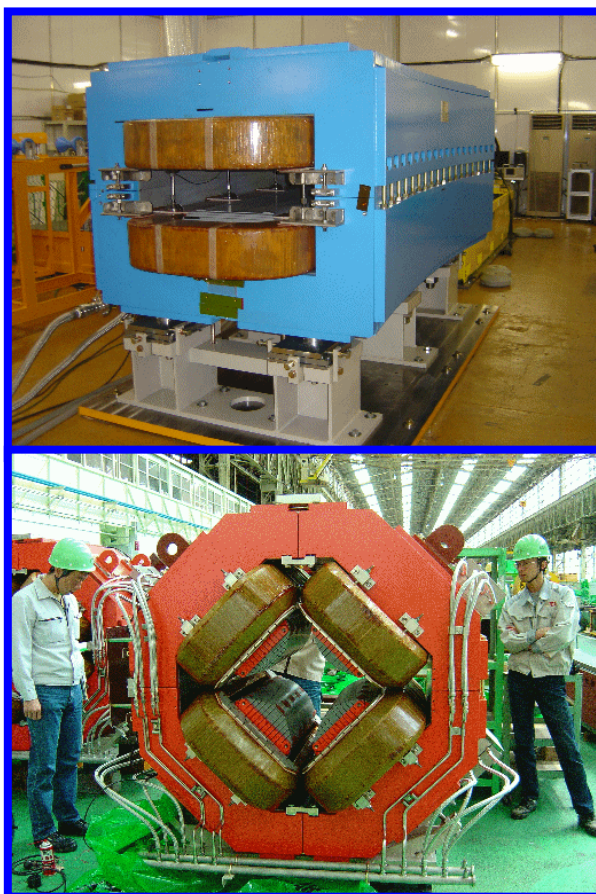


Fig 11 and 12. Layout of the JPARC 3 GeV machine and magnet production for the 50 GeV

With the nominal injection energy in the 3 GeV synchrotron, the 50 GeV synchrotron was expected to deliver 0.75 MW of beam power. With the reduced injection energy, the current in the 3 GeV synchrotron will be limited by space-charge effects and the expected beam power from the 50 GeV synchrotron will not exceed 0.4 MW. As long as the linac energy will not be nominal, the only possibility is to arbitrate between the beam power available at 3 GeV and at 50 GeV. Plans based on increasing the number of pulses accumulated in the 50 GeV give the hope to approach the foreseen power from this machine, at the detriment of the power at 3 GeV.

A further increase of the beam power by a factor of two is planned to be achieved by increasing the repetition rate. Four major upgrades will be needed. First, an energy storage system has to be added. Secondly, an additional magnet power supply system has to be built. Thirdly, more RF cavities have to be installed. Fourthly, the water cooling system has to be upgraded.

To go beyond 1.5 MW and reach a few MW, more ambitious improvements are necessary and many studies remains to be made, although the basis of some upgrades are already under investigation. Barrier buckets are envisaged, for example, to reduce the space charge tune shift at injection and possibly increase the number of pulses accumulated.

3 Final remarks

Around the end of the year 2006, important study teams (ISS for neutrino factories with the support of BENE, PAF+POFPA for CERN and NUSAG for the USA) will have published their conclusions/recommendations/guidelines concerning high intensity proton drivers and future neutrino facilities. Future decisions will undoubtedly be strongly influenced.

Regarding technology choice selection, a good summary was given by B. Weng during his concluding talk¹⁸ on accelerators at the Physics with a Multi-MW Proton source workshop held at CERN in may 2004:

“The choice of an SPL for the CERN proton driver is a realistic and competitive option which has no show-stoppers in its current design. [...]

Current conceived applications are too varied for an effective project. Intensive discussion among accelerator experts and physicists has to take place to identify realistic phase-I experiments to select proper accelerator configuration (SCL vs RCS), intensity, energy, pulse length, and target/horn design for optimal secondary beam spectrum. Then, other applications and further upgrades can be contemplated.”

Of course, the specificities of the host site of a future EU neutrino facility will strongly influence the choice between the various options.

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http://physicsatmwatt.web.cern.ch/physicsatmwatt/Presentations/Weng_summary.ppt

ANNEX: Excerpts from the PAF mandate

“In preparation for the strategic decisions foreseen to be taken in 2006 and 2010 concerning future facilities at CERN, [...] an inter-departmental working group aiming at the definition of a baseline scenario of the possible development and upgrade of the present Proton Accelerator Complex is mandated below. [...] The group reports to the DG; its findings will be discussed in the Executive Board.

The study is a natural extension of the analysis already done by the High Intensity Proton (HIP) Working Group which focused on intensity upgrade (CERN-AB-2004-022 OP/RF). Its scope is widened to cover other parameters such as beam energy and the needs of all possible users of CERN facilities. It is expected to make use of the EU supported initiatives, namely the Networks HHH and BENE, the Joint Research Activity HIPPI and the Design Studies EURISOL and DIRAC (FAIR project).

The working group will:

- Collect performance requests of the future users, taking into account the foreseen LHC upgrade, the possible Fixed Target Physics programme (including future options for neutrino physics) as recently discussed by the SPSC (Villars workshop) in the report CERN SPSC -2005-010 and the Nuclear Physics programme which will be discussed by the INTC (outcome of the future workshop in September 2005).
- Analyse the various development and upgrade options of the overall CERN proton complex including possible replacement of some of the present accelerators with Rapid Cycling Synchrotrons (RCS) and/or Fixed Field Alternating Gradient (FFAG) accelerators.
- Identify technical bottlenecks and identify R&D that would be required to validate the various options if necessary.
- Identify synergies of R&D with non-CERN studies and projects.
- Report to the DG results from the above studies before the end of 2005. [...]
- Define a preferred scenario together with a suggested implementation schedule, staged in time, and provide a preliminary estimate of the necessary resources (budget, manpower and expertise). A first presentation is expected by mid 2006 as an input for the critical decisions by the management in 2006 on a possible LINAC4. [...]

MultiMegaWatt Pion Production Target and Pion Collector

1. Introduction

High intensity neutrino beams [PJK00, studyII] (superbeams, neutrino factories and muon colliders) require a high power target as a pion source and an efficient pion collection system. Developments in high power targets [hpt03,] are also needed in several different scientific and technological areas - linear collider positron sources [ilc_web, ilcps05], radioactive beam facilities [ria_web, eurisol_web], pulsed spallation neutron sources [sns_web, ess_web], accelerator production of tritium [apt_web] and accelerator transmutation of radioactive waste [hpt05nea_web, tm_web]. Synergies already exist between some of these groups. However, pion collection is specific to neutrino beams.

Because the pions emerge from the target at a wide range of angles and momenta it is necessary to collect and focus as many as possible into a forward direction to optimize the neutrino intensity. The collector could be either a strong magnetic solenoid or horn surrounding the target. The intimate relation of target and collector mean that that they must be designed as a unit.

2. Target Design

The main problems of the target design are: heat dissipation, radiation damage and, in addition, with pulsed beams, thermal stress, shock and fatigue. The high radiation can bring problems of heating, activation and radiation damage to surrounding components.

The key issue is not total power dissipation but power density and, for pulsed beams, energy density per pulse. A high power dissipated in a large volume produces a small temperature rise and the heat is easily dissipated; similarly the radiation damage is low, although it can build up with time. The transmutation and tritium production facilities have the highest levels of beam powers - in the tens to hundreds of MW – but with the very large targets the energy density is comparatively low and hence the heating and radiation damage problems are eased.

The neutrino target will be bombarded by a 4 MW proton beam of 3-30 GeV, pulsed at 50 Hz with a pulse structure which will consist of one or more ~1 ns long micro-pulses within a ~1 μ s macro-pulse. The heavy metal targets are 1-2 cm in diameter and 20-30 cm long. The length is dictated by the desire for an acceptable pion beam longitudinal emittance and the diameter is restricted by self absorption of the pions. The target will dissipate ~1 MW of the beam power the remainder being dissipated in a beam dump. The energy density is 300-1000 J cm^{-3} per pulse, which is one of the highest of currently proposed targets for any type of machine.

To reduce the effective energy density, the target can be moved through the beam. Rotating solid targets are in use [psi_web] and a contained flowing liquid metal targets is being made at PSI [megapie_web]. The SNS [sns_web] is also committed to a contained flowing mercury target. However, the SNS has a pulsed proton beam, unlike the cw PSI beam, and the liquid metal suffers from shock effects. In particular, cavitation bubbles are formed and those near

the surface erode the container wall. Currently, the expected lifetime is only a few weeks at 1 MW [Haines05].

Another solution is to form a free metal jet [merit] which will disintegrate in the beam pulse, and then reform between pulses. The jet has many advantages since heating, radiation damage and shock are not a problem. However safety, containment and thin beam windows could still be issues for concern.

To reduce the thermal shock in solids it is advantageous to have a small target so that the transit time for the shock is small compared to the pulse length. For pulses of 1 ns length this requires a size of a few mm. Using this principle, a target consisting of small beads has been suggested by Sievers [Sievers03].

There is some freedom to choose the pulse configuration of the proton driver. If the macro-pulse is split up into several micro-pulses separated by at least the shock transit time then the energy density per pulse is effectively reduced by the number of micro-pulses in the macro-pulse. It appears feasible to have at least 10-20 micro-pulses thus reducing the energy density per pulse by an order of magnitude and reducing the effect of shock damage in the target accordingly [Bennett05].

2.1. The Granular Target

Peter Sievers [Sievers03] suggested a target consisting of small solid tantalum spheres (~2 mm diameter), cooled by flowing gas or liquid (see Figure 1). This reduces the effect of thermal shock that is experienced by a target constructed from a single piece of solid material. To further reduce the energy density in the target it was proposed to sweep the beam across 4 targets on successive pulses and to collect the pions and recombine them into a single beam in a sophisticated arrangement of optics. Currently, work on this target has ceased at CERN.

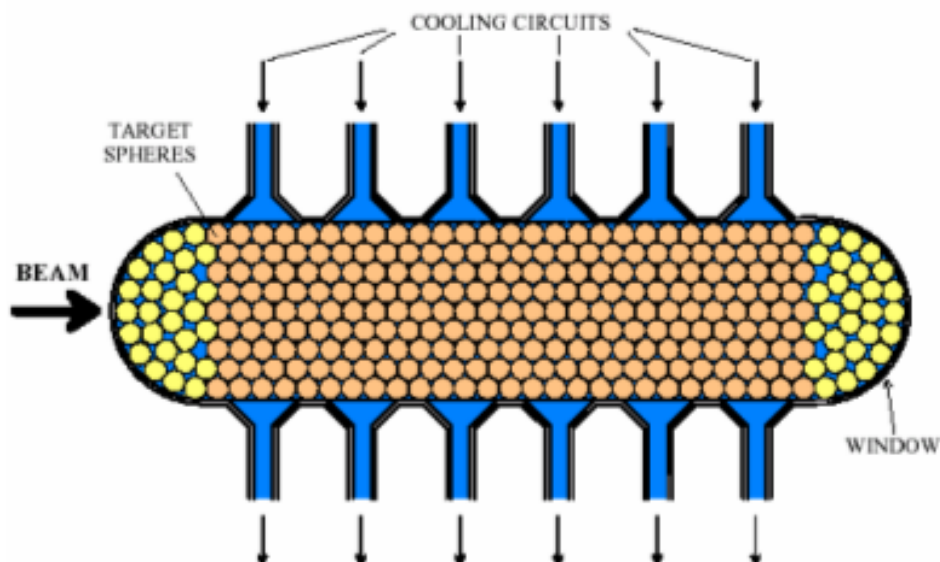


Figure 1. Schematic section through the granular target.

2.2. Solid target Studies at CERN

Jacques Lettry [hpt05] and his colleagues are pursuing a detailed programme of tests of solid targets for radioactive beams and neutrino sources. Some of this extensive work is expected to be published soon. An investigation of shock in various solid materials including the CNGS [Elsner] graphite target has been reported by Roman Wilfinger [Wilfinger05].

2.3. Solid target Studies in the USA

BNL [Simos05] are currently investigating materials (various nickel iron alloys) that have low thermal expansion to circumvent the problem of thermal shock. It has been found that under irradiation the materials lose the low expansion properties, although they can be recovered by modest heating. The search for better materials is continuing.

2.4. Solid Target Studies in the UK

The proposed target [Bennett05] for the neutrino factory is a rotating tantalum toroid (see Figure 2) operating at a temperature of ~ 2000 K to dissipate 1 MW of power produced by the impact of a 4 MW proton beam. Alternatives to the toroid are being considered - such as firing individual target bars through the beam or passing a “chain” of bars through the beam.

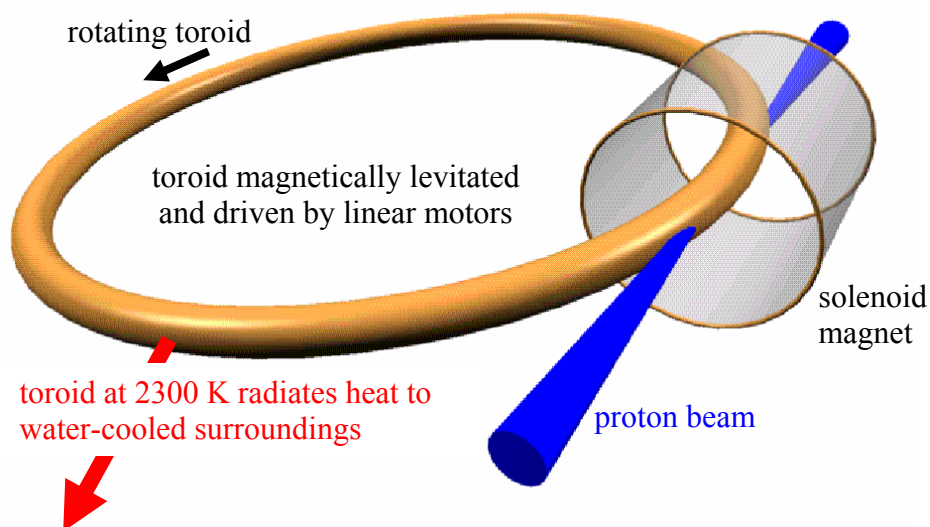


Figure 2. Schematic diagram of the rotating solid tantalum toroid

The UK programme of target developments is centred on the study of thermal shock and the long term fatigue effects. A high current electric pulse will be passed through a thin tantalum wire to provide a thermal shock by ohmic heating. The radial acceleration of the surface of the wire will be measured by a VISAR [Barker72] and the motion of the surface will be modelled using a commercial computer code, LS-DYNA [lstc_web]. In this way the constitutive equations of the tantalum at high temperatures under shock conditions will be realised and the full scale target can be modelled.

Finally, some in-beam testing is planned to confirm the VISAR measurements with the wire test and the resultant modelling. Tests at FNAL on the pbar target have already shown no disastrous effects at an energy density of 38000 J cm^{-3} (two orders of magnitude higher than the neutrino factory) for 1100 pulses.

Tantalum has been chosen for the toroid because it is refractory and resistant to radiation damage [Chen01]. However, the possibilities of carrying out studies of radiation damage at high temperature are currently being investigated.

2.5. Liquid Metal Jet Target Studies – The MERIT Experiment

Tests have been carried out at BNL [Kirk01, Lettry04, merit] to test the effect of an intense beam pulse on the jet and have confirmed that the jet does disintegrate with droplet velocities of $10\text{-}20\text{ m s}^{-1}$ (see Figures 3 and 4). The velocities increase with the energy deposited in the jet, but are not expected to cause damage to the surrounding walls. However, care will be needed to ensure that the fast droplets do not reach the beam windows, particularly the thin downstream pion window. Reforming the jet between pulses was not a problem. Tests at Grenoble [Lettry04, merit] showed that the jet was made more stable in an intense solenoidal magnetic field of 20 T.

Currently an experiment, MERIT [merit], is being constructed (Fig. 5 and 6) to test the jet in both a magnetic field and with beam of the correct intensity (up to 28×10^{12} protons per pulse at 24 GeV) to fully simulate the neutrino factory conditions. This experiment is a collaboration between laboratories in the USA, Japan and Europe. It will be installed in the nToF11 beam line at CERN for operation in 2007.

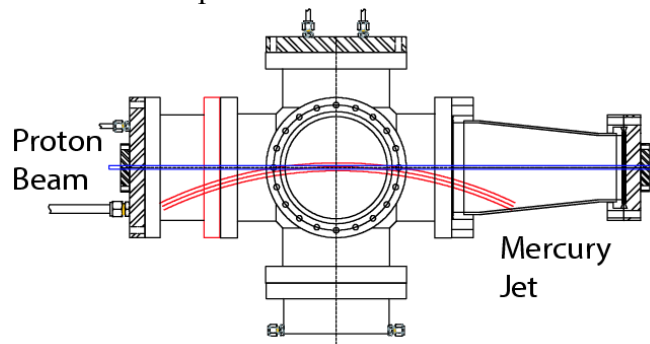


Figure 3. BNL experiment. Schematic diagram of the mercury jet hit by a proton beam; no magnetic field.

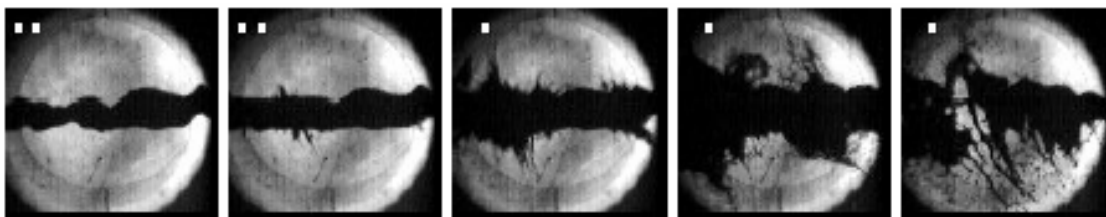


Figure 4. Pictures of the mercury jet disintegrating when hit by the proton beam. No magnetic field. Exposures of $25\text{ }\mu\text{s}$ at $t = 0, 0.75, 2, 7, 18\text{ ms}$ after a 1 cm diameter mercury jet was struck by a pulse of 2×10^{12} protons at 24 GeV.

Mercury has been chosen as the liquid for the jet although low melting point lead alloys are possible equally good alternatives. The jet is formed from a nozzle located at the entrance to the collector – in this case a pulsed magnetic solenoid.

The solenoid magnet is being fabricated from copper and cooled with liquid nitrogen to ~ 80 K to reduce the resistance and ohmic losses. The magnet is pulsed with a slow rise over ~ 10 s to reduce inductance effects and has a flat top of 1-2 s before slowly falling to zero. The cycle time between pulses is ~ 30 minutes to allow cooling to take place. The maximum field is calculated to be 15 T.

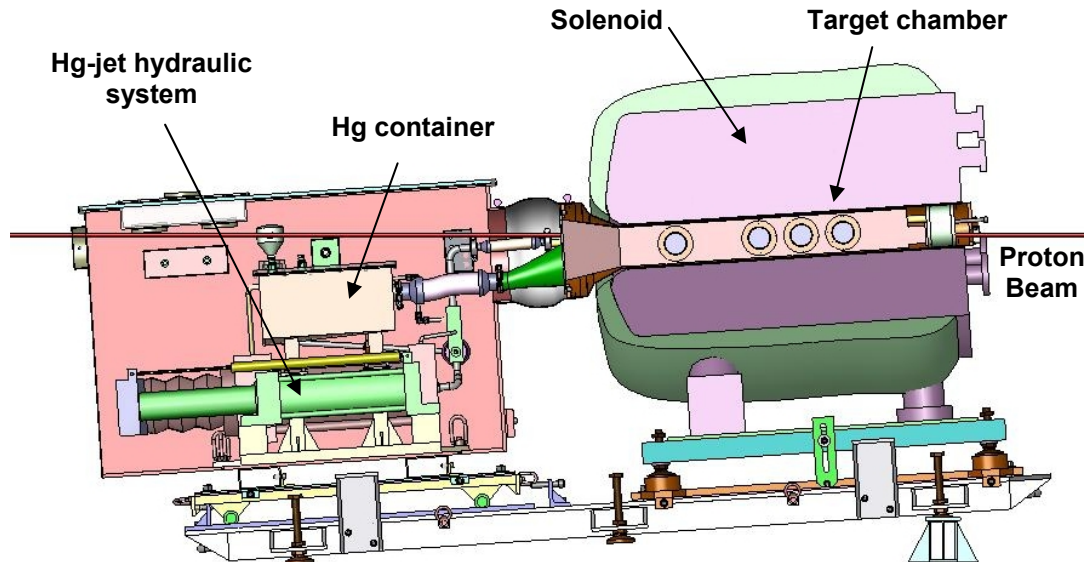


Figure 5: Schematic view of the MERIT experiment.

The jet will have a velocity of >10 m s $^{-1}$ and about 1 cm in diameter. Currently studies of the nozzle design are being undertaken at Princeton. The jet and proton beam are off the solenoid axis; the jet is at an angle of 100 mrad and the proton beam at 67 mrad. The proton beam size will be ≤ 1.5 mm rms. The diagnostic of the beam-target interaction will be optical through four windows located along the length of the mercury jet recorder by high speed cameras, while the secondary pion flux will be monitored by scintillator counters.

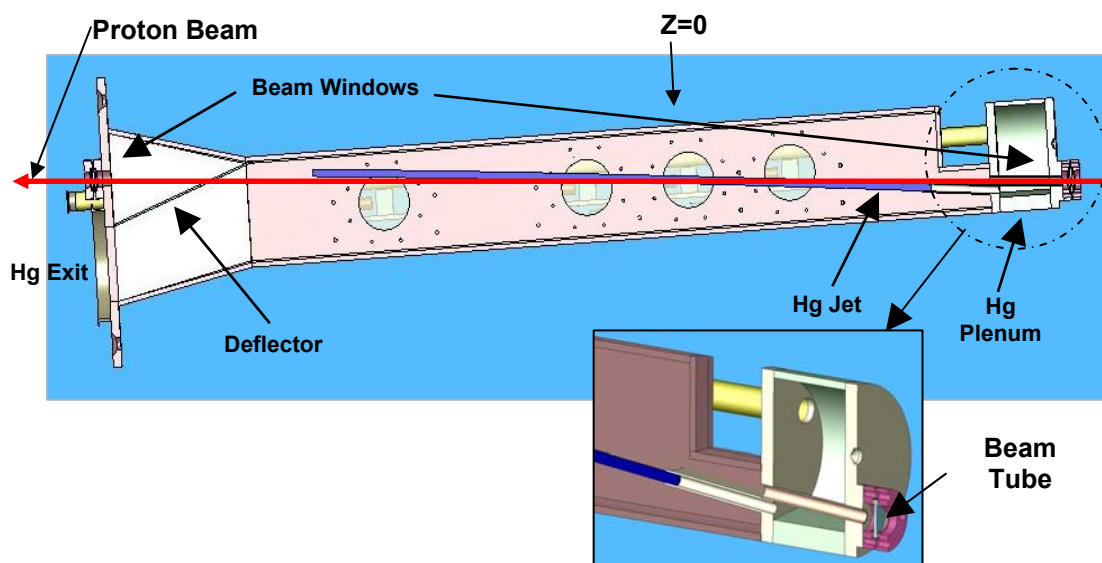


Figure 6 : Detail of the target volume showing the path of the Hg jet and of the beam. The four observation windows are also visible.

Only a few hundred pulses are required since, in principle, a single pulse will be able to show the success of the jet target concept. The goals of the experiment include studies of the jet entering the magnetic field, the jet dispersal by the beam in the magnetic field the influence of the jet entry angle on jet performance at fields from 0 to 15 T and the influence of the proton beam energy. Also, the CERN PS beam profile allows one to study the effect of varying the charge intensity from 5×10^{12} to 28×10^{12} protons per spill and to study the effects of the jet dispersal on the beam in long pulses. If the mercury has not dispersed sufficiently it could absorb pions. It is not possible directly to test the jet at 50 Hz operation, but it is possible at reduced beam energy (14 GeV) to extract 2 spills separated by 20 ms. This will be an important part of the test to confirm the correct re-establishment of the jet on successive beam pulses.

The MERIT experiment is an important proof of principle experiment which will hopefully show that reliable long-life targets of almost unlimited power dissipation can be built. There is still an issue of safety with radioactive mercury; the use of a low melting point metal alloy could improve this situation.

3. Pion Collection

Two options are discussed which relate to the neutrino facility: a superbeam would require a magnetic horn (see Figure 7), sign-selecting the pions to be focused in the decay channel, whereas a neutrino factory would favour a solenoid magnet focusing both sign pions at the same time.

From an initial design made at CERN on a horn prototype system (horn + reflector) [Ball] foreseen for a neutron factory, an optimization and a redesign has been made in a superbeam context [Campagne], driven by the physics case of a long baseline experiment (130 km) between CERN and Fréjus (MEMPHYS detector location).

The proton energy appears to be best suited between 3.5 and 4.5 GeV proton kinetic energy and the current is then required to be 300 kA in the horn and 600 kA in the reflector. The horn must have a conductor thickness of only 3 mm to minimise the energy deposition by the secondary protons in the neck (<30 kW); a real challenge. Integration and testing with a target would need a dedicated R&D effort in an experimental area designed to face the various safety aspects (chemistry of heavy metals, high radiation levels, high voltage, high current ...), which would also include the design of a complete remote handling installation for the horn and target maintenance and exchange. The power supply design is an extrapolation of existing facilities and a first prototype version at only a fraction of the nominal current and repetition rate has been built. Hence, it requires special attention and cost estimates due to the high frequency envisaged, 50 Hz. A dedicated facility with prototypes of the different electrical components and a load (e.g. horn prototype) is needed to test existing components provided by manufacturers, and also to validate the mechanical design that has been chosen so far. This work should incorporate the latest knowledge acquired with the Miniboone and CNGS horns.

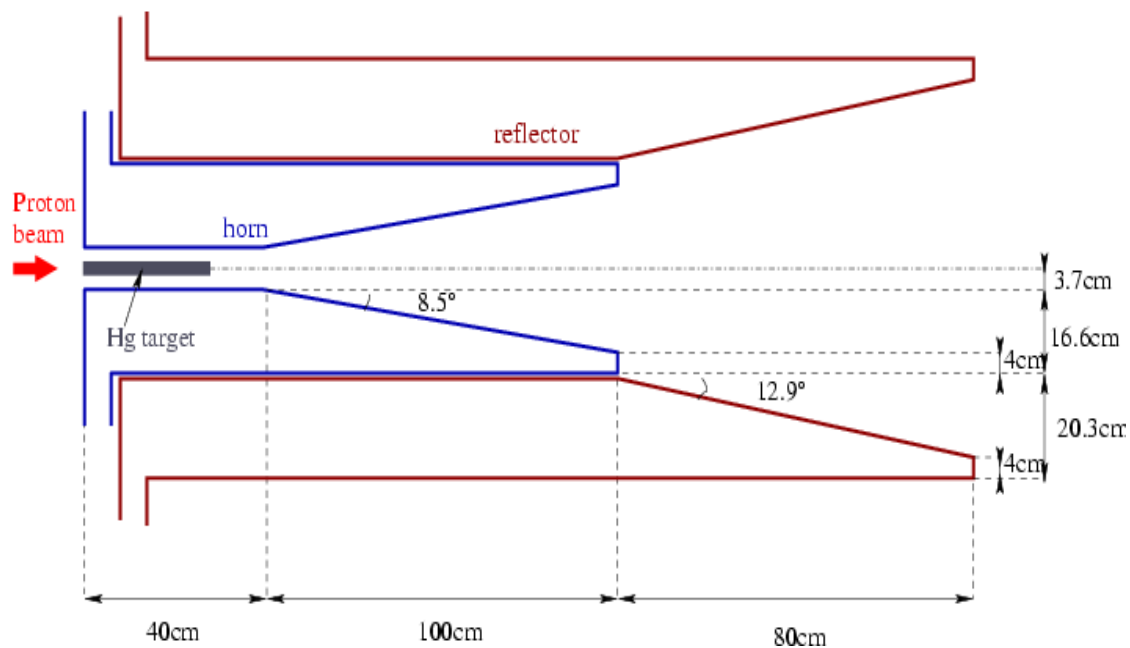


Figure 7. Schematic section of the horn for pion collection.

Generally the solenoid collection is divided into two stages. The first one, where the field is constant, capture pions with a maximum transverse momentum, whereas a second part where the field decreases to lower values, does the real focusing. Usually more pions are captured for a stronger magnetic field for a given aperture. A compromise between field and aperture has to be made, since the cost of the superconducting device increases with the magnetic volume. The US Design Study II [studyII], indicates that a 20 T solenoidal magnet is feasible, with a reasonable life time. The solenoid has water-cooled copper coils in the centre, which also act as radiation shields to the superconducting coils on the outside. A detailed study of the behaviour of the superconducting part under a quench would be mandatory.

The pion yield by 3-20 GeV protons is however still poorly known: Geant4 and Mars simulations give significantly different results (up to 50%) in the best phase space regions both for superbeams and neutrino factories. The HARP experiment is presently publishing its first results [harp] and should help clarify this point soon.

4. Integration Issues

There is an awareness that many of the target problems are common to several communities. These issues are being discussed at a number of gatherings, such as the High Power Target Workshops series [hpt03, hpt05]; this is helping to bring the various communities together.

Groups working on neutrino factories around the world are already collaborating closely. A good example of this is the MERIT mercury jet experiment [merit]; a collaboration between laboratories in the USA, Japan and Europe which reflects the worldwide effort to master the technological challenges towards a neutrino factory and a muon collider. Although MERIT

will explicitly test the system in a strong solenoidal field, the results in zero magnetic field will be very important for jet applications in a horn collector.

The absence of suitable in-beam test facilities is a severe problem. There has been a recent proposal [hpt05] from ORNL to try and set up an in-beam facility to allow target to be tested at high power. Radiation damage to materials is another area requiring proton beams. Pulsed facilities are scarce but PSI [psi_web] is providing a good, although limited, service using cw proton beams. To have an accurate estimate on the lifetime of the system: target and collector components, extensive mechanical tests with the pulsed power supply combined with the radiation field are mandatory.

Due to the absence of testing facilities and the dependence on other ongoing developments, in this R&D, the role of simulation is very important. It can enable the testing of different design concepts, reduce the experimental working load - due to the easier access to specific data as such as deformation, temperature, field, etc., - and provide a more cost-effective engineering of the system. It is intend to develop the use of an already existing simulation tool with the modelling of cross-discipline studies: thermo mechanical and electromagnetic computations in the dynamic and transient domains. Experiences from other academic and industrial fields such as RF cavity (modelling, design, tuning) and induction heating will be integrated in order to minimize the time to reach solution. A tuning of these simulations (fatigue, deformations, modal analyses, transient thermo mechanical excitation of the structure, skin effect and Joule heating, power dissipation, heat exchange and cooling, radiation resistance, etc.) could be done using input provided by the previously described tests and facilities. This software development has to be done in parallel and with a strong interaction with all related hardware tasks.

These studies may also indicate the road for a more effective use of the existing resources in term of hardware in the different laboratories.

5. Recommendations

- Currently progress is seriously limited by resources, so more resources would be of great benefit.
- Nevertheless some progress is being made with targets, particularly mercury jets and studies at CERN, BNL and in the UK on solid materials. This should continue with more resources for both targets and collectors.
- The opportunity for more interactions with other communities needing high power targets must be continued and expanded. This also applies to collectors.
- High power target R&D should be continued.
- Collector R&D should be continued, including a dedicated facility for the horn power supply for high frequency tests.
- A facility to study corrosion damage to the horn conductor material should be envisaged, since it seems to be the main cause of failure in the high frequency, water cooled horn.
- An in-beam test facility available to all should be pursued.

- The design of the target station, beam dump, remote handling, maintenance and replacement of components has hardly been addressed. This is an important area and will need considerable engineering and scientific resources.
- Related simulations have to be done in parallel.
- Address the safety issues and assimilate safety and maintenance considerations already in the early stages of the design. It takes a considerable time to obtain the necessary authorisations to operate these high power target stations.

References

- [apt_web] <http://apt.lanl.gov/system.html>
- [Ball] A. Ball et al. CERN-NUFACT-Note-42.
- [Barker72] L. M. Barker and R. E. Hollenbach, J. Appl. Phys. 43 (1972) 4669.
- [Bennett05] J. R. J. Bennett et al, NuFact05, LNF, Frascati (Rome), Italy, 21-26 June 2005, see web site: <http://www.lnf.infn.it/conference/2005/nufact05/>
- [Campagne] J.E. Campagne CERN-NUFACT-Note-138.
- [Chen01] J. Chen, H. Ulmaier, et al. J. Nucl. Mater. 298 (2001) 248.
- [Elsner] K. Elsener, CNGS - CERN neutrinos to Gran Sasso, CERN-SL-2000-018 EA, http://proj-cngs.web.cern.ch/proj-cngs/PDF_files/sl-2000-018.pdf
- [ess_web] The ESS web site, <http://www.kfa-juelich.de/ess/#>
- [eurisol_web] The EURISOL web site, <http://www.eurisol-ds.inl.infn.it/>
- [Haines05] John Haines; Second High Power Targetry Workshop, Oak Ridge, TN, USA, 10-14 October 2005, <http://www.cap.bnl.gov/mumu/conf/target-051010/>
- [harp] HARP Coll. to appear in Nuclear Phys. B.
- [hpt03] Workshop on High Power Targetry for Future Accelerators, Ronconcoma, Long Island, New York, USA, 8-12 September 2003, <http://www.cap.bnl.gov/mumu/conf/target-030908/>
- [hpt05] Second High Power Targetry Workshop, Oak Ridge, TN, USA, 10-14 October 2005, <http://www.cap.bnl.gov/mumu/conf/target-051010/>
- [ilc_web] The International Linear Collider, see; <http://www-project.slac.stanford.edu/ilc/default.htm>
- [ilcps05] 'Workshop on Positron Sources for the International Linear Collider', CCLRC Daresbury Laboratory, 11th-13th of April 2005; http://www.astec.ac.uk/id_mag/ID-Mag_Helical_ILC_Positron_Production_Workshop.htm
- [Kirk01] H.G.Kirk *et al.* Target studies with BNL E951 at the AGS, proc. 2001 Part. Accel. Conf. (Chicago, IL, June 2001), p. 1535.
- [Lettry03] J. Lettry *et al.*, Thermal shocks and magnetohydrodynamics in high power mercury targets, J. Phys. G: Nucl. and Part. Phys. 29 (2003) 1621.
- [lstc_web] Livermore Software Technology Corp. <http://www.lstc.com/>
- [megapie_web] The MEGAPIE web site, <http://megapie.web.psi.ch/>
- [merit] Studies of a Target System for a 4-MW, 24-GeV Proton Beam, CERN-INTC-2004-016 INTC-P-186, 26 April 2004, http://www.hep.princeton.edu/~mcdonald/mumu/target/cern_proposal.pdf
- [nea_web] <http://www.nea.fr/html/trw/>
- [PJK00] R.B. Palmer, C. Johnson, E. Keil, Nucl. Instrum. and Meth. A 451 (2000) 265.
- [psi_web] The PSI web site, http://www.psi.ch/index_e.shtml
- [ria_web] The RIA web site, <http://www.phy.anl.gov/ria/>
- [Sievers03] P. Sievers, Nucl. Instrum. and Meth. A 503 (2003) 344
- [Simos05] Nick Simos in Second High Power Targetry Workshop, Oak Ridge, TN, USA, 10-14 October 2005, <http://www.cap.bnl.gov/mumu/conf/target-051010/>
- [sns_web] The SNS web site, <http://www.sns.gov/>
- [studyII] S. Ozaki et al, Feasibility study II of a muon-based neutrino source, BNL report No. BNL-52623, June 2001.
- [tm_web] <http://apt.lanl.gov/index.html>
- [Wilfinger05] R. Wilfinger, ENG/BENE, Target Section Meeting, March 2005, <http://bene.web.cern.ch/bene/>

The front–end of a Neutrino Factory

1 Overview

The particles produced by the bombardment of the target form the input to the muon front-end of the Neutrino Factory. The front end is required to capture the pions produced in the target, to contain the muons produced by pion decay, to produce an appropriate energy and bunch structure in the beam, and to reduce the phase-space volume (cool) the muon beam so that it is well matched to the subsequent acceleration systems. Recent progress in the development of the various subsystems that make up the muon front-end is described in detail in [1,2] and is summarised below.

2 Pion capture and decay channel

Two schemes have been proposed by which the particles produced in the target may be captured. The first uses high-field solenoid magnets to capture both positive and negative particles at the same time [3–5]. The second calls for a magnetic horn to focus either positive or negative particles into the subsequent transport and decay sections [6]. The horn scheme has the advantage that the focussing element closest to the target itself is relatively simple. The advantage of the solenoid scheme is that an efficiency gain of a factor of two can be achieved if the downstream accelerator complex is designed to manipulate and store μ^+ and μ^- simultaneously [5]. In each case, significant engineering work needs to be carried out to ensure that the target station can be operated safely.

A large aperture (16 – 30 cm) focussing channel consisting of high-field (1.2 – 5 T) superconducting solenoids is used to transport the pions and to capture and transport the decay muons. The pion lifetime ($\tau \approx 2.6 \times 10^{-8}$ s) yields $c\tau \approx 10$ m, therefore decay channels with lengths of up to 100 m have been considered.

3 Phase rotation and bunching

By the end of the decay channel a strong energy-time correlation has developed in the longitudinal phase space of the muon beam; energetic muons arrive at the end of the channel early, low momentum muons arrive late. The system that receives the beam from the decay channel must accommodate this large longitudinal phase space. The Japanese scheme for a Neutrino Factory [7] calls for the beam emerging from the decay channel to be injected directly into a large acceptance fixed field alternating gradient (FFAG) accelerator of the PRISM type [8]. The FFAG effects both the rotation of the longitudinal phase space and the acceleration of the muon beam from 0.3 GeV/c to 1 GeV/c.

While the CERN Neutrino Factory scheme maintains the SPL bunch structure and so requires no bunching section, each of the US schemes employs an RF system to bunch the beam. The most recent study, Study IIa, proposes the use of RF cavities with frequencies decreasing from 333 MHz to 234 MHz to bunch the beam.

The various European and US Neutrino Factory schemes all call for phase rotation to reduce the energy spread of the muon beam in preparation for the ionisation cooling channel. The CERN scheme uses a series of 44 MHz or 88 MHz RF cavities. No buncher is required as the bunch structure of the proton driver (the SPL) is preserved through the capture channel. In

Study IIa, a sequence of RF cavities the frequencies of which decrease from 234 MHz to 201 MHz is used to perform phase rotation. The phase space produced is shown in figure 1. A relatively long muon pulse is produced in 201 MHz bunches that are well matched to the subsequent cooling channel.

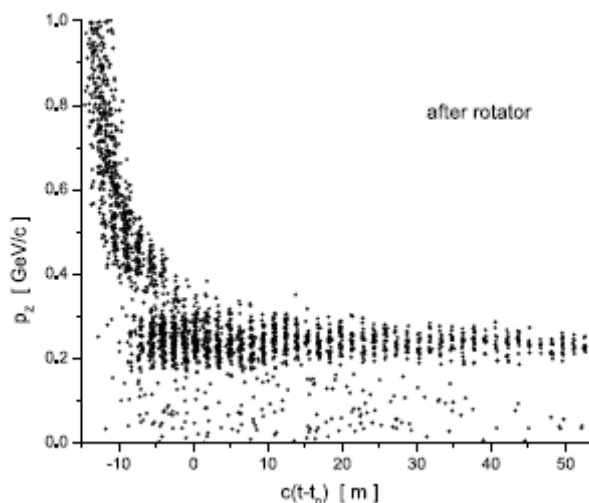


Figure 1: The longitudinal phase space of the muon beam at the end of the phase-rotation section of Study IIa.

4 Ionisation cooling

The muon beam that emerges from the phase-rotation and bunching sections fills a large volume in transverse phase space. For example, in US Study II the transverse emittance at the exit of the decay channel is 12 mm [9]. The spread of the muons in the longitudinal phase space is also very large (~ 60 mm in Study II). Efficient, cost effective, acceleration of the muon beam requires that the phase space be modified. The phase-rotation and bunching systems that follow the decay channel are required to produce a beam with an energy spread of ~ 60 MeV which is appropriately bunched to match the subsequent cooling sections.

Each of the five Neutrino Factory conceptual design studies have considered the benefit of reducing the emittance of the muon beam (cooling) before injecting it into the acceleration and storage systems. There are two principal motivations for this: to increase the number of muons inside the acceptance of the downstream accelerators; and to keep the cost of the muon acceleration system to a minimum.

At the end of the decay channel, the muons have a momentum of roughly 200 MeV/c. The time-dilated lifetime of the muon is short (~ 4.7 μ s) making it essential that cooling and acceleration take place as rapidly as possible. Ionisation cooling, a process in which the muon beam is caused to pass through an alternating series of liquid-hydrogen absorbers and accelerating RF-cavities, is the technique by which it is proposed to cool the muon beam prior to acceleration. Various 'gain factors' have been defined to quantify the gain in performance due to the cooling channel (see table 1). Systems that give gain factors of between 2 and 10 have been devised. Since a factor of Γ gain in stored muon-beam intensity implies a reduction, by a factor Γ , in the running time required to achieve a particular total neutrino flux, and a decrease in emittance of the muon beam entering the acceleration section is likely to lead to significantly lower costs for muon acceleration, it will be important to make a

careful optimisation, for performance and cost, of the cooling and acceleration systems. The engineering demonstration of the ionisation-cooling technique will be carried out by the international Muon Ionisation Cooling Experiment (MICE) collaboration [9]. The MICE experiment, which has been approved, will take place at the Rutherford Appleton Laboratory (RAL), using muons produced by the ISIS 800 MeV proton synchrotron. The status of the experiment is reviewed in the paragraphs that follow.

The US MuCool collaboration [2] is carrying out an R&D programme by which each of the components of the cooling channel will be built and operated. The principal components of the cooling channel that have been built and tested over the past year are the liquid hydrogen absorber and the 201 MHz cavity. During the past year, the Muon Test Area (MTA) at FNAL has become available. Component tests are presently underway in the MTA which will eventually be served by a high intensity proton beam so that the operation of the cooling channel components can be proved using realistic particle fluxes.

Table 1: Survey of the gain afforded using ionisation cooling in a number of conceptual design studies of the Neutrino Factory.

Design	Number of cooling cells	Gain factor per cell (%)	Cooling factor	Comment
Study II [4]	26	6	7	Increase in phase-space density in acceptance of downstream accelerator.
Study IIa [5]	26	2	2	Increase in number of muons in acceptance of downstream accelerator.
CERN [6]	36	10	7	Increase in muon yield at 2 GeV over optimised Neutrino Factory without cooling.
NuFact-J [7]	–	–	1.5 – 2	Acceleration based on FFAGs. Performance improvement when absorber is included in FFAG ring giving 6D cooling effect.

4.1 The international Muon Ionisation Cooling Experiment

The principal components of the MICE experiment are shown in figure 6.2 [9]. Two, functionally equivalent, spectrometers are placed upstream and downstream of a single lattice cell of the Study II cooling channel. In the Study II design approximately 10^{14} μ /s pass through the channel. The lateral dimensions of the beam are such that space-charge forces can be ignored making it possible to run MICE as a single particle experiment in which the Neutrino Factory bunch is reconstructed offline using an ensemble of particles recorded in the experiment. At the nominal input emittance of $\varepsilon_{in} = 6 \pi$ mm a cooling effect ($\varepsilon_{in}/\varepsilon_{out} - 1$, where ε_{out} is the output emittance) of $\sim 10\%$ is expected. The cooling effect will be measured with a precision of 1% (i.e. $\varepsilon_{in}/\varepsilon_{out} - 1$ will be measured with an absolute precision of 0.1%).

The MICE cooling channel consists of three absorber/focus-coil (AFC) modules and two accelerating-cavity/coupling-coil (RFCC) modules. The AFC modules each contain a 21 l liquid-hydrogen absorber inside a pair of superconducting coils that bring the beam to a focus in the centre of the absorber. Liquid hydrogen is the most efficient ionisation-cooling material because it has a large specific ionisation and a comparatively large radiation length. Safe operation of the system in the presence of liquid hydrogen leads to significant engineering constraints. The AFC modules and the hydrogen system each have both active and passive safety systems. The hydrogen will be stored in the form of metal hydride when the absorber is emptied. A vigorous R&D programme is underway to demonstrate the safe operation of the hydrogen system. The super-conducting coils and the liquid-hydrogen vessel itself are refrigerated using closed-cycle 'cryo-coolers' [10].



Figure 6.2: Drawing of the MICE experiment [9]. The beam enters the experiment from the bottom left-hand corner. The beam first passes through one of the scintillator hodoscopes that form the time-of-flight system. After passing through the upstream spectrometer, the beam passes through three absorber/focus-coil modules and two cavity/coupling-coil modules before it passes through the downstream spectrometer and, a second time-of-flight hodoscope, the downstream Cherenkov counter and is stopped in the electromagnetic calorimeter. The beamline and upstream instrumentation is not shown.

The RFCC module must restore the energy lost by the muons as they pass through the absorber. The coupling coil, a short, large diameter solenoid, provides the magnetic field that transports the muons through the module. The acceleration is produced by four 201 MHz copper cavities which produce a gradient of 8 MV/m. To produce the required field gradient, the cavities must be electrically closed, yet, to preserve the cooling effect, the amount of material through which the beam passes must be minimised. Thin beryllium windows have been developed for this purpose through the MuCool programme [2]. The degree of emission from the cavity surfaces is significantly enhanced by the Lorentz force produced by an intense magnetic field [11]. While reducing the field emission in a Neutrino Factory cooling channel, in which the cavities must operate at 16 MV/m, is a challenging problem, it has been estimated that for operation in MICE, the emission can be kept within acceptable bounds.

The muon beam that enters the experiment may contain a small pion contamination. The instrumentation upstream of the cooling channel is therefore required to distinguish pions from muons and to measure the phase space coordinates of the muons entering the channel.

Downstream of the cooling channel, the instrumentation is required to identify electrons produced in the channel by muon decays and to measure the muon phase-space coordinates. The upstream particle identification will be performed using a scintillator-based time-of-flight (TOF) system and a threshold Cherenkov counter. The TOF system will also be used to trigger the experiment and to determine the phase of the RF fields in the cavities as the muon traverses the experiment. The upstream and downstream spectrometers are each composed of a 4 T superconducting solenoid instrumented with a scintillating-fibre tracking device. Downstream of the cooling channel a final TOF station, a Cherenkov counter and a calorimeter are used to distinguish muons and electrons.

The MICE collaboration will take enough data to make the uncertainty on the measured cooling effect systematics limited. It is therefore crucial that the systematic errors are understood in detail. To do this, the experiment will be built up in stages. A first measurement of cooling, using the two spectrometers and one AFC module, is scheduled for 2008. The first RFCC module and a second AFC module will then be installed and the full MICE cooling channel will be assembled in 2009.

The MICE experiment will be mounted on ISIS at the CCLRC Rutherford Appleton Laboratory. The preparation of the MICE Muon Beam on ISIS, the MICE Hall and the first phase of the MICE experiment is proceeding to schedule. The first data-taking period, in which the muon beam will be characterised, the instrumentation calibrated and the relative systematics of the two spectrometers will be measured, will begin in April 2007.

Bibliography

1. R.C. Fernow, "Muon front-end for the Neutrino Factory", contributed paper to the 7th International Workshop on Neutrino Factories and Superbeams, NuFact05, Laboratori Nazionali di Frascati, Frascati (Rome), June 21 - 26, 2005.
2. Y. Torun, "Muon cooling: MuCool and MICE", contributed paper to the 7th International Workshop on Neutrino Factories and Superbeams, NuFact05, Laboratori Nazionali di Frascati, Frascati (Rome), June 21 - 26, 2005.
3. D.Finley, N.Holtkamp, eds, "Feasibility Study on a Neutrino Source Based on a Muon Storage Ring", (2000). See http://www.fnal.gov/projects/muon_collider/reports.html.
4. S. Ozaki, R. Palmer, M. Zisman, and J. Gallardo, eds, "Feasibility Study-II of a Muon-Based Neutrino Source", BNL-52623, June 2001.
5. C. Albright et al, "Neutrino Factory and Beta Beam Experiments and Development", BNL-72369-2004, FNAL-TM-2259, LBNL-55478. See <http://www.cap.bnl.gov/mumu/study2a/> and <http://www.cap.bnl.gov/mumu/studyii/FS2-report.html>.
6. A. Blondel, ed., "ECFA/CERN studies of a European neutrino factory complex", CERN-2004-022, <http://preprints.cern.ch/cernrep/2004/2004-002/2004-002.html>.
7. NuFactJ working group, "A Feasibility Study of A Neutrino Factory in Japan". See <http://www-prism.kek.jp/nufactj/nufactj.pdf>

8. The PRISM Working Group, "The PRISM Project - A Muon Source of the World-Highest Brightness by Phase Rotation", submitted to J-PARC Letter of Intent L24, unpublished, <http://www-ps.kek.jp/jhf-np/LOIlist/LOIlist.html>.
9. G. Gregoire et al., the MICE collaboration, "MICE, an international Muon Ionisation Cooling Experiment: proposal to the Rutherford Appleton Laboratory", submitted to CCLRC and PPARC on the 10th January 2003: <http://mice.iit.edu/cooldemo/micenotes/public/pdf/MICE0021/MICE0021.pdf>.
10. <http://www.janis.com/p-a4k10.html>.
11. J.Norem, "Breakdown in RF cavities", TPPT025, paper contributed to PAC05, the Particle Accelerator Conference, Knoxville, Tennessee, USA, 16 -- 20 May 2005.

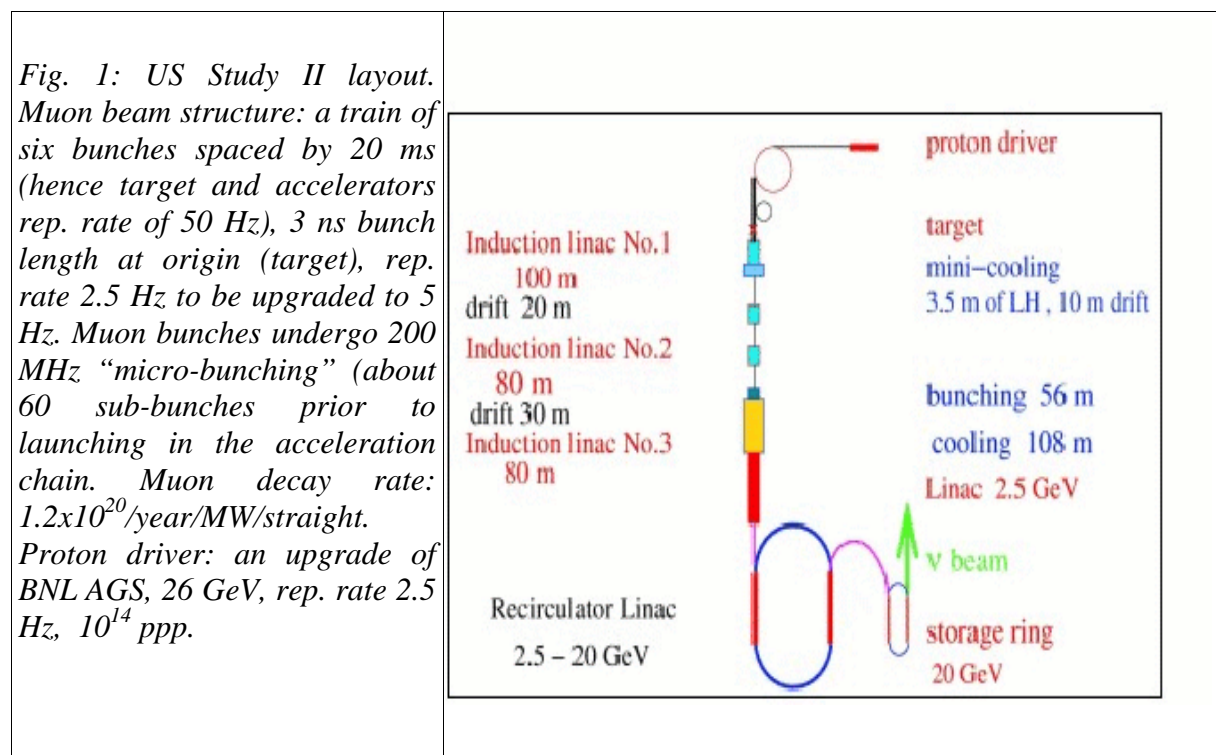
Muon Acceleration and Storage in a Neutrino Factory

Due to the manifest superiority of FFAGs over other techniques for muon acceleration in a Neutrino Factory and following decisions taken in December 2003, work in this area has focussed on the FFAG method. This year has again seen an increase in the participation of the EU in FFAG research and design. The active fields within CARE/BENE are: (i) design of an electron model of a non-scaling muon FFAG, (ii) design of an isochronous FFAG, (iii) development of simulation and tracking tools, (iv) participation in the design of the US linear muon FFAGs, and (v) collaboration in Japanese scaling FFAG studies. In addition, the EU is proposing to construct the first non-scaling FFAGs ever to be built. This and the NuFact International Scoping Study will be the next challenging activities. All of this work is the subject of internal reports and of contributions in Conferences and Workshops.

1 The context

Muon acceleration in the Neutrino Factory is profiting from modern accelerator technologies, which should permit the construction of a facility capable of bringing intense beams - in the 10^{21} muons/year range, to multi-GeV energies, 20-50 GeV – in the near future. Various muon acceleration methods have been investigated so far, and have yielded a number of different layouts, as described below. However, the accelerator community is now particularly concentrating on the method of FFAGs and is bringing new, modern ideas and concepts to the field, that are very promising both for muon acceleration and for proton driver applications.

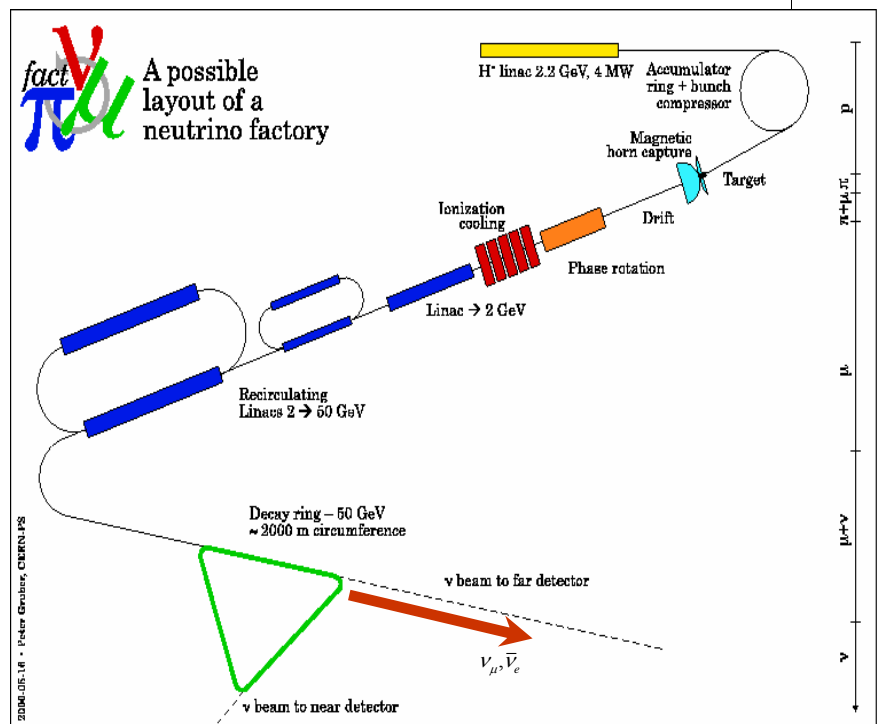
The “why” and “how” of this strong increase in activity in the FFAG domain will be addressed in the following sections, along with the corresponding engagement within BENE/MuEnd, as planned from the very beginning, and the consequences for the future plans.



2 Muon acceleration concepts

RLA based designs. Linac acceleration presents the essential property of allowing fast acceleration in a short distance, thanks to the use of high gradient / high frequency RF systems, compatible with the short lifetime of muons, thus ensuring a high transmission - typically, more than 80% at an average gradient of 5 MV/m. With regard to the large muon beam phase-space, superconducting Linacs also offer a good geometrical acceptance, owing in particular to the large iris of the ~ 200 MHz RF cavities, and the possibility of accelerating trains of bunches, all of which are properties that translate into 6-D acceptance performance.

Fig. 2: The CERN NuFact layout. Muon beam structure: 3.2 μ s long trains of 140, 23 ns spaced bunches (44 MHz structure), at a rep. rate of 50 Hz. 1 ns bunch length at origin (target). Bunch-to-bucket operation: next to capture, phase rotation and cooling, each muon bunch occupies a 220 MHz Linac RF bucket. 50 GeV top energy reached in three stages: Linac up to 3 GeV, followed by 3-11 and 11-50 RLAs. Muon rate: 10^{21} /year in the storage ring.

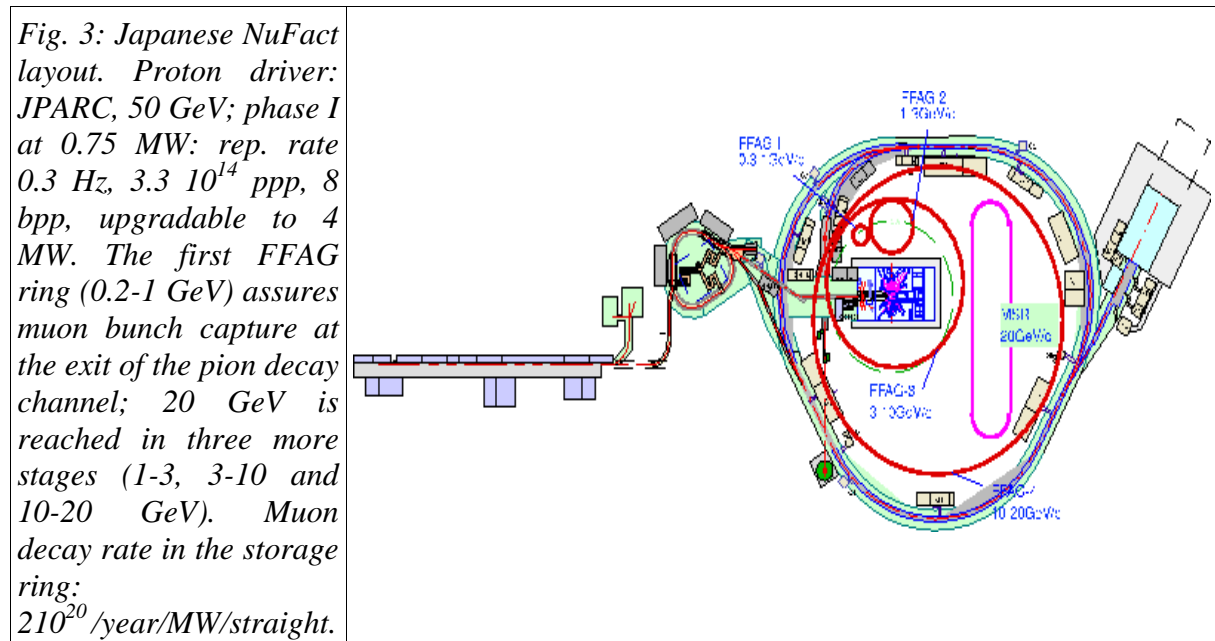


Two recirculating Linac accelerator (RLA) schemes have been produced, namely:

- The US Study I, in 2000 [Ref7-USSI], based on Linac pre-acceleration followed by two RLAs that bring the muon beam up to 50 GeV; a follow-on, Study II, in 2001 [Ref7-USSII] (Fig. 1), differed mostly, from the point of view of the acceleration, in having a lower muon energy (20 GeV), using a single RLA; and, in 2004, Study IIa [Ref7-USSIIa], which benefits in particular from cost-effective approaches, allows both signs of muons, and sees the first introduction of FFAGs in a combined Linac/RLA/FFAG acceleration scheme.
- The EU study, a design produced in 2004, with 50 GeV top energy, based on the acceleration of muons by a two-stage RLA system [Ref7-EUNuFact], and operating in a bunch-to-bucket mode (Fig. 2).

Amongst other conclusions, these design studies are said to have “demonstrated technical feasibility (provided the challenging component specifications are met), established a cost baseline and achieved the desired range of physics performance”, and to have shown that “progress is still needed [...] towards optimising the design, developing and testing the required accelerator components, and significantly reducing the cost” [Ref7-USSIIa]. The cost of the NuFact installation is very high with the muon accelerators representing about a third

of it, and the possibility that “FFAG rings could be also considered” [page 6.2, in Ref7-USSII], as a cheaper solution, has already arisen in the US Study IIa.



Scaling FFAG and Japan R&D. Independently, in 2001 a NuFact based on the JPARC 50 GeV proton installation and on the capture and acceleration of the muons to 20 GeV by scaling FFAGs was proposed (Fig. 3) [Ref7-JNuFact].

Acceleration in the high energy regime uses high gradient (up to 10 MV/m), low frequency (5-25 MHz) RF, and is based either on a huge-bucket method, with injection at the bottom and rotation upward in the manner of the RF manipulation in the PRISM experiment (see below) or, alternatively, on the use of frequency modulation. The FFAG method yields large transverse acceptance (on the order of 3π cm), large longitudinal acceptance (1.5 eV.s) and reasonably fast acceleration (1 MV/m on average), resulting in about 50 % muon survival.

Extensive R&D programmes have been pursued in Japan. They comprise, amongst other RF and component studies, the building of a proof of principle 500 keV proton machine [App. C, in Ref7-JnuFact], followed by a larger 150 MeV accelerator [Ref7-FFAG150], both based on a radial sector DFD triplet. These two machines proved in particular the feasibility of ultra-fast cycling, rep. rate in the kHz range, by means of high gradient RF systems based on special magnetic alloy cores [Ref7-HGRF]. This is now followed by the prototype ADS/Reactor experiment facility under construction at the KURRI institute, Kyoto (Fig. 7.5 [Ref7-KURRI]).

The first muon FFAG will be the PRISM project, a 2003-2007 programme [Ref7-PRISM]. The PRISM lattice is based on a DFD triplet similar to the KEK FFAGs design, and comprises 10 cells. This FFAG is intended for muon bunch phase rotation aiming at momentum spread compression, from $68 \text{ MeV}/c \pm 20\%$ down to $\pm 5\%$ in 6 turns, for use in a muon beam physics facility. The remarkable features of PRISM are its 2 MV/turn RF system, and its challenging injection and extraction systems, a benchmark towards the NuFact accelerators.

Non-scaling FFAGs. This represents the main field of activity within BENE/MuEnd. The concept was introduced in the US in the late 90's [Ref7-CarolMills] and uses a synchrotron type of cell, i.e. using linear optical elements, with fixed fields. Hence, the orbit position moves in the course of acceleration, and the tunes change (in contrast to scaling FFAGs) due to the change in beam rigidity. Compared to RLAs, these “linear, non-scaling FFAGs” allow more turns, and hence less RF, and in addition there are a smaller number of FFAG rings (2-3) than there are arcs in RLAs ($2 \times 4-5$ passes)

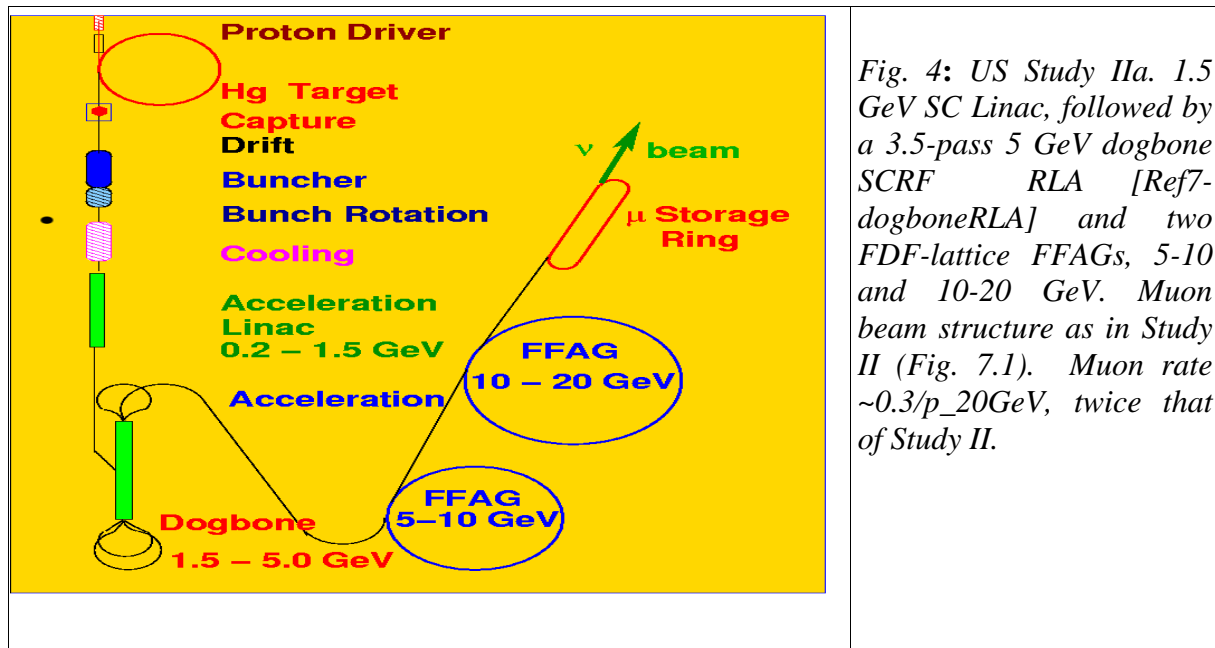


Fig. 4: US Study IIA. 1.5 GeV SC Linac, followed by a 3.5-pass 5 GeV dogbone SCRF RLA [Ref7-dogboneRLA] and two FDF-lattice FFAGs, 5-10 and 10-20 GeV. Muon beam structure as in Study II (Fig. 7.1). Muon rate $\sim 0.3/p_{20\text{GeV}}$, twice that of Study II.

Linear, non-scaling optics have a number of advantages: (i) large transverse acceptance due to the linear fields and large momentum acceptance due to the small dispersion function, up to the point where the necessity of cooling is now questionable, (ii) rapid acceleration (energy gain of 2 to 3 over about 10 turns), due to high frequency/high gradient RF and near-crest acceleration, (iii) reduced circumference (and muon decay loss) compared to scaling FFAGs, (iv) reasonable size magnets, due to the limited horizontal beam excursion. There are drawbacks to the method, however, such as strongly non-linear longitudinal motion, and resonance crossing during acceleration. This is addressed in the last Section.

Some preliminary conclusions have been drawn from the work undertaken so far: (i) linear FFAGs yield lower cost/GeV than RLAs above 5 GeV, and possibly also below, though this needs further investigation, (ii) a new muon acceleration scheme has emerged, US Study IIA [Ref7-USSIIa] (Fig. 4), combining Linac, dogbone RLA [Ref7-Bogacz] and non-scaling FDF-lattice FFAGs. The acceleration in these machines is based on 201 MHz SCRF and $3 \pi \text{ cm} / 0.05 \text{ eV.s}$ acceptance is reached (twice the transverse acceptance in Study II). Higher energy could be attained using additional FFAG rings.

Isochronous FFAG: Isochronism to high order allows on-crest acceleration, a cyclotron-like regime. This type of FFAG has been developed in the EU recently [Ref7-IsochroFFAG]. The current design is based on a 5-magnet cell (Fig. 5) using 3 different types of combined multipole magnets, see the figure below. The cell acts as a DFD triplet at low energy and as an FDF triplet at high energy. This use of two families of cells allows the design of a ring with insertions. The horizontal tune varies (less than 0.15 per cell), whereas the vertical tune

is about constant. A 8-20 GeV muon ring has been designed this way, accelerating in 16 turns using 201 MHz RF, with a circumference close to 1000 m. Alternative designs have been proposed as well [Ref7-HSNuFact07].

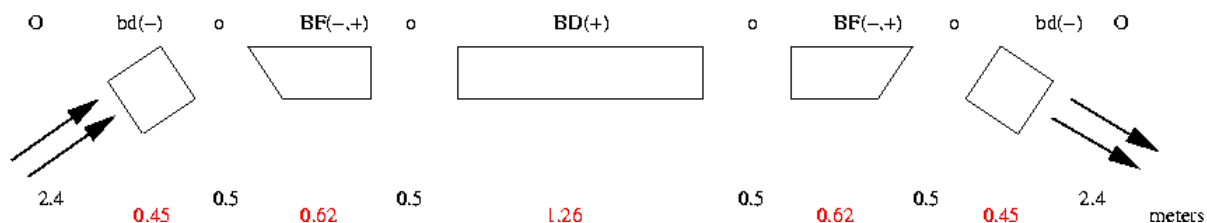


Fig. 5 Five-magnet cell of an isochronous FFAG

The isochronous FFAG with insertions has a number of advantages, in particular optimum, on-crest acceleration, resistive beam loading, the flexibility of insertions (for injection, extraction and collimation), fewer RF systems by the use of 4-cell cavities and no crossing of integer or half-integer vertical betatron resonances during acceleration (in contrast with linear FFAGs). Dedicated tracking tool developments [Ref7-MeotNIMA] have allowed numerical simulations of 6-D transmission in this FFAG [Ref7-NuFact05Lemuet].

3 Electron model

A proof-of-principle of non-scaling FFAGs is needed, with the goal of demonstrating in particular the viability of the concepts of (i) rapid acceleration, in 5 to 10 turns, (ii) resonance crossing and the related constraints on field and alignment defects, (iii) large acceptance. These considerations have motivated the emergence of the EMMA electron model project [Ref7-EMMA], a 10 to 20 MeV ring, with a 42 cell FODO lattice, about 15 meters in circumference, using 1.3 GHz RF. It is hoped to build EMMA in the Daresbury Laboratory in the UK, once the necessary funding has been obtained. It is also planned to construct a proton non-scaling FFAG as a prototype for hadron therapy in Oxford. Experiments on fixed field acceleration are also taking place at existing pulsed synchrotrons [Ref7-HIMAC].

4 Muon storage rings

Work on μ^+ and μ^- storage rings, addressed in the aforementioned NuFact study reports [Ref7-EUNuFact, Ref7-USSI, Ref7-JNuFact], has recommenced in the newly formed ISS (International Scoping Study), with C. Johnstone of FNAL and G H Rees of RAL, respectively, acting as coordinators for racetrack and triangular ring designs. A single tunnel is to house μ^+ and μ^- rings, which are designed for 20 GeV initially, but with an upgrade potential of 50 GeV.

Designs assume normalised ring acceptances of 67,500 mm mr, to allow loss collimation of the 30,000 mm mr, normalised emittance, MW, cooled, muon beams. The ratio for the rms divergences of the 20 GeV muon and the neutrino beam is set at 0.1 for an assumed (possibly incorrectly) muon, normalised, rms emittance of 4800 mm mr.

Weak bending magnets are placed at the ends of the production region(s) to remove the neutrinos emanating from the muons which decay in the ring matching sections, of larger divergence angles. Modified dispersion suppression is then required in the ring matching sections, with details depending on the 20 and 50 GeV lattice designs.

Injection kicker systems for the new designs have higher stored energy due to the larger ring acceptances and the shorter rise times, necessitated by multiple bunch train filling. A reduced circumference, triangular ring is obtained by use of an isosceles design with two, long production straights and one, short region for RF cavities and beam collimators.

5 Perspectives

It is evident from what precedes, that a large amount of work has already been undertaken in the quest for better performing and less expensive muon acceleration techniques, but that much is still required.

At the present stage of the research, it is clear that more collaborators are needed on the EU side, to match those in Japan and the US. Areas requiring particular attention are:

1. Lattice design.

- . A task force needs now be devoted to muon storage ring studies
- . Optimisation work is needed to improve the muon transmission in the accelerators
- . EU participation in both EMMA and the proton FFAG is currently very limited
- . There is limited activity in the design of the FFAGs for muon acceleration, with no contributions at all to the crucial question of linear FFAGs below 5GeV
- . Error studies are needed (field, alignment, tolerances).

2. Magnet.

. It is time now to launch magnet studies, 3D design, etc. Some work has already started, mainly at Fermilab and mostly focussed on EMMA. Designs relating to the muon machines would allow progress with the lattice design studies as well.

3. Model.

. There are 2 proposals, e (10->20MeV) (EMMA) or p (-> about 100MeV). EMMA would be a proof of linear FFAG acceleration of muons. The pModel can be considered relevant to proton driver case and would be the non-scaling equivalent of the machine our Japanese colleagues built in the late 1990s (POP) with a scaling lattice. This was a success and significant progress in the field.

Bibliography

Ref7-PlaneryNuFact05 F. Méot, The acceleration of muons in the neutrino factory, Procs. NuFact05 Workshop, Frascati (2005).

Ref7-HHHNuFact05 H. Haseroth, The technical challenges of Super-Beams and Neutrino Factories, Procs. NuFact05.

Ref7-OhmoriNuFact05 C. Ohmori, Achievements at NuFact05, Procs. NuFact05.

Ref7-USSI N. Holtkamp and D. Finley, eds., Tech. Rep. Fermilab-Pub-00/108-E, Fermilab (2000).

Ref7-USSII S. Ozaki, R. Palmer, M. Zisman, and J. Gallardo, eds., Tech. Rep., BNL-52623 (2001).

Ref7-USSIIa <http://www.cap.bnl.gov/mumu/study2a/REPORT/NF-BB-WG.pdf>

- Ref7-Bogacz A. Bogacz, Beam dynamics of low energy muon acceleration, Procs. NuFact05.
- Ref7-EUNuFact The study of a European neutrino factory complex, CERN/PS/2002-080 (PP).
- Ref7-JNuFact A Feasibility Study of a Neutrino Factory in Japan, NufactJ Working Group (2000).
- Ref7-FFAG150 M. Aiba et als., 150 MeV FFAG accelerator [...], EPAC-2002, Paris.
- Ref7-HGRF C. Ohmori, High field-gradient cavities loaded with magnetic alloys, PAC99 Conf.
- Ref7-KURRI M. Tanigaki et als., Construction of FFAG Accelerators in KURRI [...], PAC05 Conf.
- Ref7-PRISM Y. Kuriyama, Development status of PRISM-FFAG ring [...], Procs. NuFact05.
- Ref7-CarolMills F.E. Mills and C. Johnstone, 4th Int. Conf. on Physics at $\mu^+\mu^-$ colliders, San Francisco, UCLA (1999).
- Ref7-dogboneRLA A. Bogacz, Beam Dynamics of Low Energy Muon Acceleration, Procs. NuFact05.
- Ref7-IsochroFFAG G.H. Rees, FFAG Studies for Neutrino Factory Accelerators, Procs. NuFact05 ; F. Lemuet, F. Méot, G. Rees, Numerical 6-D acceleration in an isochronous FFAG ring, PAC05 Conf.
- Ref7-MeotNIMA F. Lemuet, F. Méot, developments in the ray-tracing code Zgoubi for 6-D multiturn tracking in FFAG rings, NIMA 547 (2005) 638-651.
- Ref7-NuFact05 F. Lemuet, 6D beam dynamics in an isochronous FFAG ring, Procs. NuFact05.
- Ref7-HSNuFact05 Horst Schonauer, An Isochronous 10-20 GeV Muon Ring with Constant Tunes, Operating above Transition, Procs. NuFact05.
- Ref7-EMMA E. Keil, for the EMMA collaboration, Electron model: lattice and performance, Procs. NuFact05. ;
R. Edgecock, for the EMMA collaboration, Introduction to the non-scaling electron model FFAG EMMA, ibid. ;
Grahame Rees Isochronous, FFAG Rings with Insertions for the Rapid Acceleration of Muons or Electrons, ibid.
- Ref7-HIMAC S. Machida, Experimental study of resonance crossing at HIMAC, Procs. NuFact05.

A Slow Muon Facility at a Neutrino Facility

High intensity slow muon beams can be obtained at the very front end of a Neutrino facility. They could be very useful in a wide variety of experiments which are usually divided in three categories: (1) precision measurements, (2) searches for rare processes and (3) applications. A non comprehensive list of some of these experiments subdivided in the three categories is shown in Table 1.

Table 1. Overview of experiments with slow muons

Table 1

Overview list of slow muon physics programs.

Categories	Topics	Comments	Beam
Precision measurements	muon lifetime	G_F determination	pulsed
	muon capture rates	nuclear physics	pulsed or DC
	muonic X-ray	nuclear physics	pulsed or DC
	muon $g - 2$	SM allowed, new physics	pulsed
Rare processes	muon EDM	SM suppressed, new physics	pulsed
	charged lepton mixing	SM forbidden, new physics	pulsed or DC
Application	catalyzed fusion	aim break-even	pulsed
	materials science	μ SR	DC

A more comprehensive list with details of the beam characteristics needed can be found in [Aysto01].

From the point of view of particle physics very important measurements are those of the muon anomalous magnetic dipole moment (MDM) $a_\mu = (g-2)/2$, of the possible muon electric dipole moment (EDM) and of rare decays in which charged lepton flavour is violated (CLFV). These measurements are often called “the muon trio” because in SUSY models their respective amplitudes can be derived from the slepton mass matrix as in the following Figure 1:

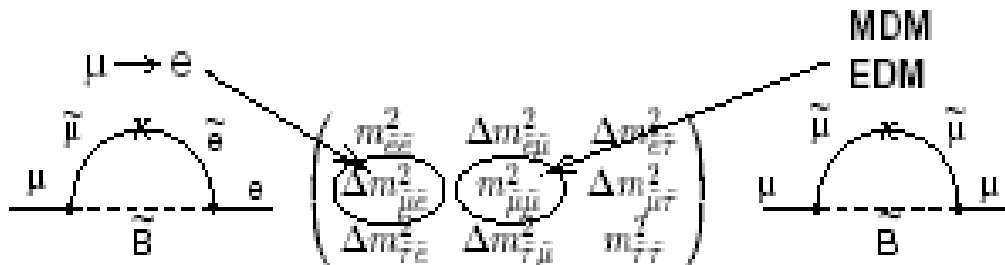


Fig.1: Slepton mass matrix and its contributions to the three processes: MDM, EDM, CLFV

where MDM and EDM are related to the real and imaginary part of the smuon diagonal element while CLFV is related to the off diagonal elements.

As it is well known the present best experimental MDM measurement [E821]:

$$a_{\mu} = \frac{g-2}{2} = 1.1659208(6) \times 10^{-10} \quad (0.5 \text{ ppm})$$

deviates 2.6 or 0.9 σ from the SM value depending if this is computed by using e^+e^- or τ data.. This possible deviation can be explained by SUSY contributions with parameters that will be explored by LHC experiments. It is possible to improve the experimental result down to 0.05 ppm [JPARC] with higher muon beam intensities, though to fully realize the potential of the improved experimental measurement the hadronic contribution would need to be known with an equivalent uncertainty.

The current limit [Bailey79] on the muon EDM ($\sim 10^{-18}$ e-cm) could also be improved by several order of magnitudes, down to 10^{-24} e-cm, by using high intensity muon beams in a storage ring with a suitable electric field [Farley04]. EDMs of this order of magnitude are again predicted by several SUSY models.

Relevant non diagonal terms of the slepton mass matrix (Fig. 1) are predicted in ‘‘SUSY-GUT’’ models, where these terms arise from radiative corrections from the Planck scale to the weak scale and in ‘‘SUSY-seesaw’’ models, where a suitable scheme of neutrino masses and chiralities is introduced consistently with the existing experimental data of neutrino oscillations experiments. These models predict CLFV with branching ratios just below to a few orders of magnitude below the current experimental upper limits. CLFV due to neutrino mixing included in the SM frame would be completely unobservable (for instance a $BR \approx 10^{-54}$ is predicted for $\mu \rightarrow e\gamma$). The detection of $\mu \rightarrow e\gamma$ events would thus be a clear, unambiguous sign of physics beyond the standard model, even including neutrino masses.

In $\mu \rightarrow e\gamma$ searches, a beam of positive muons is stopped in a thin target and a search is made for a back to back positron-photon couple with the right momenta and timing. The main background in present experiments comes from the accidental coincidence of independent positrons and photons within the resolutions of the used detectors. The best available detectors for low energy positrons and photons must therefore be employed. In the MEG experiment at PSI (see Fig. 2) a surface muon beam with an intensity greater than $10^7 \mu/s$ will be stopped in a thin target. A magnetic spectrometer, composed of a superconducting magnet and drift chambers, will be used for the measurement of the positrons trajectories. Positron timing will be measured by an array of scintillators. Photons will be detected by an innovative electromagnetic calorimeter in which a total of about 800 photomultipliers will detect the light produced by photons initiated showers in about 800 liters of liquid Xenon. In a recent test at PSI the design energy resolution of 4.5% FWHM was obtained in a 100 l liquid Xenon prototype for 55 MeV photons. The aim of this experiment is to reach a sensitivity down to BR of the order of 10^{-13} , with an improvement of two orders of magnitude with respect to the present experimental limit. The start of the data taking is foreseen in 2006.

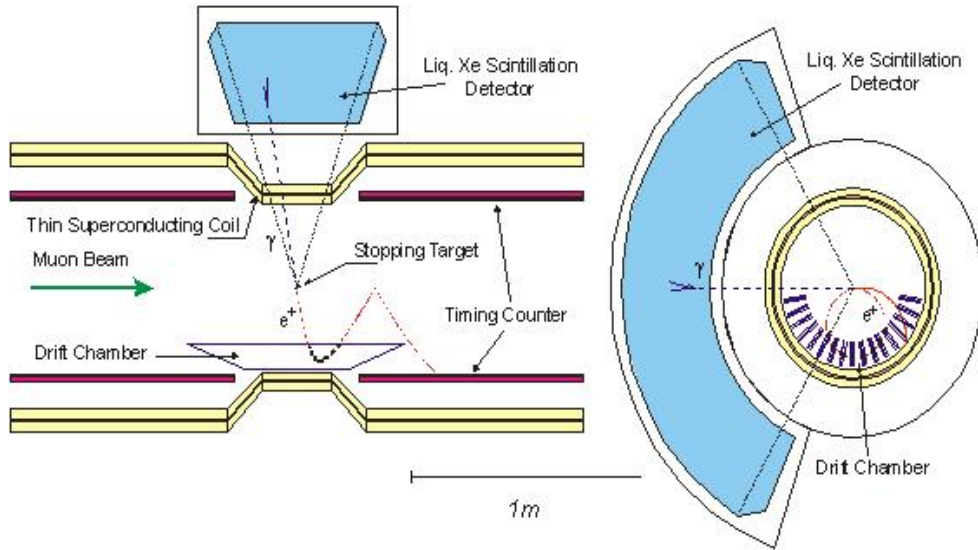


Figure 2 A sketch of the MEG detector

Another channel for CLFV investigation which is not limited by accidental background and can therefore be used to improve the sensitivity to CLFV is muon to electron (μe) conversion in nuclei. The ratio of the rate for this process with respect to $\mu \rightarrow e\gamma$ has been calculated by several authors, for various nuclei, under assumptions on the relevant matrix elements which are valid in many SUSY models (see Figure 3, from [Kitao02]).

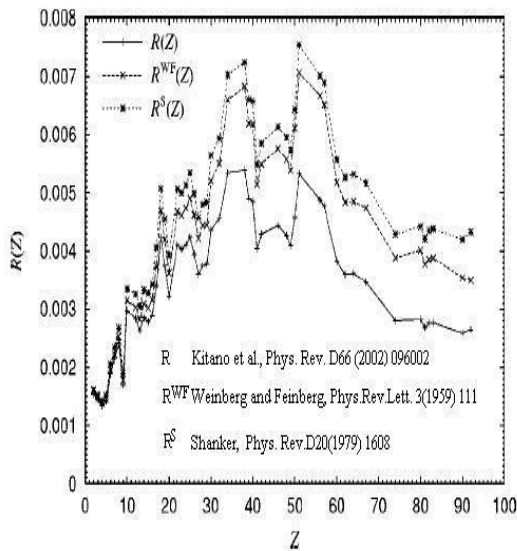


Figure 3 Computed ratio of $BR(\mu e)/BR(\mu \rightarrow e\gamma)$.

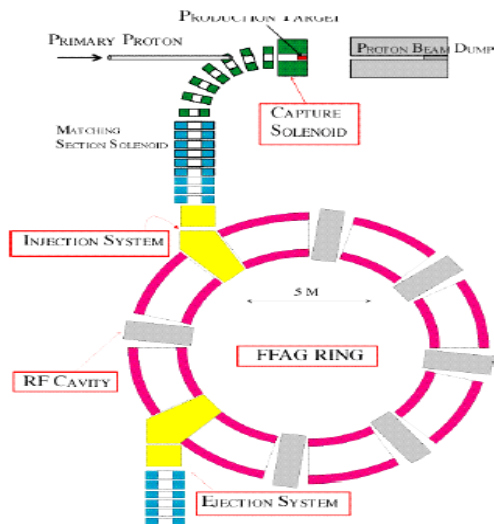


Fig. 4 Muon production in the PRISM/PRIME project at J-PARC

Experimentally, negative muons are brought to stop in a thin target and are subsequently captured around a nucleus. The energy of a possible converted electron would be equal to the rest muon mass minus the muon binding energy EB . Two main sources of background are: i) beam correlated background due mainly to radiative pion capture followed by $\gamma \rightarrow e^+e^-$ conversions and ii) electrons from muon decay in orbit (DIO). The first source of background can be controlled by improving the

muon beam quality, the second one is intrinsic; DIO electrons spectrum extends up to the energy region of electrons from μe conversion but with a spectrum proportional to $(m_\mu - EB - E_e)^5$. An excellent electron momentum resolution is fundamental in order to keep this background under control.

In the PRISM/PRIME (Fig. 4) project at J-PARC a pulsed proton beam is used to produce low energy pions that are captured by placing the target inside a superconducting solenoid magnet. The pulsed structure of the beam helps in reducing the beam correlated background. The beam is then transported in a circular system of magnets and RF cavities (FFAG ring) which acts as a pion decay section (increasing beam cleaning) and reduces the muon energy spread. The features of this beam would be an extremely high intensity ($10^{12}/s$) of very clean muons of low momentum ($\approx 70\text{MeV}/c$) with a narrow energy spread (few % FWHM). The last feature is essential to stop enough muons in thin targets. If the electron momentum resolution will be kept below $350\text{KeV}/c$ (FWHM) the experiment will be sensitive to μe conversion down to $BR \leq 10^{-18}$.

If CLFV were discovered, the angular distribution of electrons from the CLFV decay of polarized muons could be used to discriminate among the different SUSY-GUT SU(5), SUSY-GUT SO(10), SUSY-Seesaw models or others.

It is finally important to stress the complementarity between SUSY searches at LHC and CLFV searches, as shown in Fig. 5, where the different sensitivity to possible neutralino masses for these two classes of experiments, for some parameters of the SUSY model investigated [Masiero04], is shown.

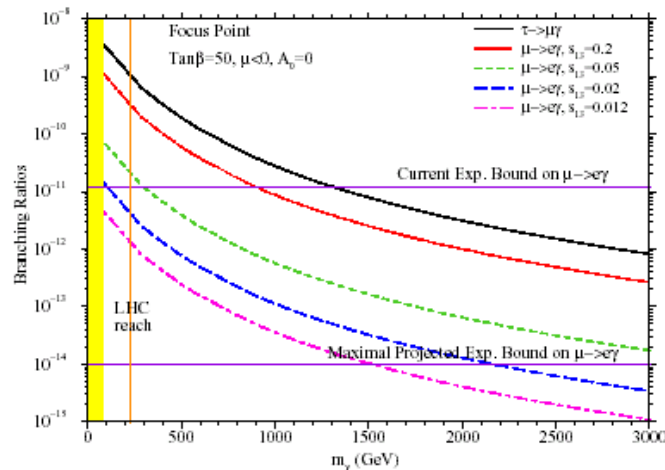


Fig. 5: Complementarity between SUSY search at LHC and CLFV searches

References

- [Aysto01] J. Aysto et al., ‘Physics with Low-Energy Muons at a Neutrino Factory Complex’, CERN-TH/2001-231, in ‘ECFA/CERN Studies of a European Neutrino Factory Complex’ Blondel, A (ed.) et al. CERN-2004-002.- ECFA-04-230, p259.
- [E821] H. N. Brown et al., Phys. Rev. Lett. 92 (2004) 161802.
- [JPARC] J-PARC Letter of intent L17: An improved Muon (g-2) Experiment at J-PARC, contact person: B.L. Roberts.
- [Bailey79v] J. Bailey et al., Nucl. Phys. B150,1 (1979).
- [Farley04] F.J.M. Farley et al., Phys. Rev. Lett. 92, 052001 (2004).
- [Kitao02] R. Kitano et al., Phys. Rev. D66 (2002) 096002
- [Masiero04] A. Masiero et al., hep-ph/0401138

Betabeams

1. Introduction

The beta-beam concept for the generation of an electron (anti-)neutrino beam was proposed [zuc02] in 2002. A first study [aut02,lin04] of the possibility of using the existing CERN machines for the acceleration for radioactive ions to a relativistic gamma of roughly 100, for later storage in a new decay ring of approximately the size of SPS, was made in 2002. The results from this very first short study were very encouraging, but as no resources could be allocated at CERN for this work, it was not continued. In 2004 it was decided to incorporate a design study for the beta-beam within the EURISOL DS proposal. EURISOL [eurisol] is a project name for a next-generation radioactive beam facility based on the ISOL method [isol] for the production of intense radioactive beams for nuclear physics, astrophysics and other applications. The proposal was accepted with the beta-beam task as an integral part. The design study officially started 1 February 2005 and will run for 4 years resulting in a conceptual design report for a beta-beam facility as one potential user of EURISOL.

2. First study

The first study of the feasibility of using the existing CERN accelerator complex for a beta-beam facility was made over a few months with very limited manpower (see figure 1). The main objectives were twofold: i) to identify a possible scenario for bunching, acceleration and storage in a few very short bunches of a sufficient amount of radioactive ions for a beta-beam and ii) to identify possible bottlenecks in the proposed scheme. A main objection raised early on concerned the possible activation of the accelerators. Consequently, some time was spent to simulate the activation problem in the decay ring and to calculate the average losses in the accelerator chain. The overall result was encouraging but unfortunately no further work was approved due to other commitments for the accelerator departments at CERN. The study proposed to use a thick ISOL target for production of ^6He and ^{18}Ne as both isotopes can be produced in large quantities and are easy to handle. Neither of the isotopes have any long-lived daughter products that could create a problem in the low-energy part of the facility. Several iterations were required for the “bunching” but eventually a high frequency (60 GHz) ECR source was identified as a possible highly efficient tool to create sufficiently short bunches after the target for multi-turn injection into a synchrotron. For the first stage of acceleration, it was proposed to use the 100 MeV/u linac of the EURISOL facility. Further acceleration was to be done with a new rapid cycling synchrotron (RCS), the PS and finally the SPS. A new injection and stacking method was proposed to keep the duty factor of the decay ring low. The method makes use of a dispersion orbit in the decay ring to avoid that the injection elements interfere with the circulating beam, bunch rotation to bring the fresh bunches to the central orbit and asymmetric bunch merging to take the newly injected ions into the centre of the circulating bunch [han03,han05]. The maximum gamma of 150 that can be reached for fully stripped ^6He ions in the SPS, was initially chosen for the coasting beam in the decay ring but later revised to lower values taking physics reach considerations into account. The main bottlenecks in the scenario chosen for the first study were shown to be the tune shift at PS and SPS injection, the activation of the PS ring.

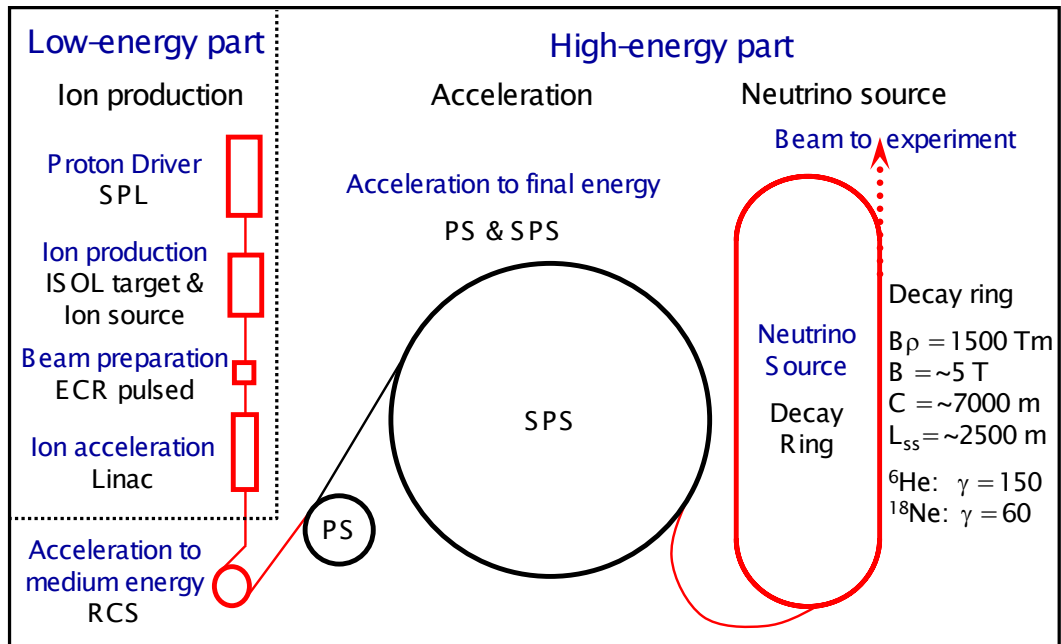


Figure 1: CERN Betabeam conceptual design

3. The beta-beam task in EURISOL

The beta-beam task within EURISOL has six participating institutes and receives close to 1 M€ over four years from the EU. Together with the contribution from the participating institutes, a total of almost 35 manyears is available for the study. The study officially started 1 February 2005 and the short term objective is to freeze a parameter list for the end of 2005. The deliverable of the study is a conceptual design report (CDR) for a beta-beam facility at CERN by the end of 2008.

3.1 Version 1

In the first study no consideration was made of the charge state distribution after the ECR. While it seems feasible to extract most He ions in the fully stripped charge state 2^+ , it is unlikely that more than a maximum of 30% of Ne ions can be extracted in one single charge state. Furthermore, the multi-turn injection into the RCS is typically done with an efficiency of 50%. Revising the first study with these points taken into consideration, the annual rate of neutrinos from ^{18}Ne was found to be drastically reduced [version1], while the annual rate of anti-neutrino from ^6He was acceptable.

Table 1: A comparison of three different scenarios for a CERN beta-beam facility

Publication		NuFact02	Version 1	Version 2
Date		2002	April 05	Aug 05
Target	rate to ECR [s^{-1}]	2×10^{13}	2×10^{13}	2×10^{13}
		(0.8×10^{12})	(0.8×10^{12})	(0.8×10^{12})
ECR	Type	Compact	60 GHz pulsed	60 GHz pulsed
	T ejection [keV/u]	20 keV/amu	U=50kV	U=50kV
	Efficiency			
Post accel.		Cyclotron	LINAC	LINAC
	T ejection [MeV/u]	50	100	100
RCS	Rep. rate [Hz]	16	16	10
	T ejection [MeV/u]	300	500 (1124)	500 (1124)
PS	Number of RCS bunches/cycle	16	16	20
	T ejection [GeV/u]	3.5 (7.8)	7.8 (14)	7.8 (14)
	Number of ejected bunches	8 (top merge)	8 (top merge)	20
SPS	T ejection [GeV/u]	139 (55)	92.5	92.5
	γ	150 (60)	100 (100)	100 (100)
Decay ring	Cycle time [s]	8	6 (3.6)	6 (3.6)
	Imax, stored [ions]	2.02×10^{14}	5.88×10^{13}	9.7×10^{13}
		(9.11×10^{12})	(1.19×10^{12})	(3.11×10^{12})
	Neutrino flux (ν /year)	Not given	1.76×10^{18}	2.9×10^{18}
			(0.02×10^{18})	(0.046×10^{18})
	Duty factor	2×10^{-3}	2×10^{-3}	$4.5(3.9) \times 10^{-3}$

Values given for the two baseline isotopes as ${}^6\text{He}({}^{18}\text{Ne})$

3.2 Version 2

The second version [version2] was primarily motivated by the need to deal with space charge issues in the PS and SPS. The first study had a duty factor of the neutrino beam of 2×10^{-3} , with a moderate increase the known space charge induced limitations of PS and SPS were shown to be respected for an annual rate of 10^{18} (anti-)neutrinos per year. However, the second version still fails to deliver this rate for neutrinos. In table 1 the main differences between the different versions are listed together with the resulting annual rate. The increase of the number of bunches also resolves a problem with a RF incompatibility at transfer between the 40 Mhz Rf system and the 200 MHz RF system during acceleration.

4. Technical challenges

The main challenge for the beta-beam study is to reach the required annual rate for a physics reach that would make the beta-beam an attractive option for neutrino physics in 10-15 years time. Detailed physics studies available in these proceedings show that a beta-beam facility must deliver at least 10^{18} (anti-)neutrinos at the end of the straight section per year in the direction of a MegaTon detector to be of any interest in this time perspective. At a higher rate of some 10^{19} (anti-)neutrinos per year the detector size could be greatly reduced and still leaving the facility as a highly competitive alternative to even the neutrino factory.

The achievable production rate of the isotopes of interest for a beta-beam facility was a major issue in the first study at CERN. The numbers presented in ref. [aut02] were compiled from general parameters for beam current and targets taken from the EU-supported EURISOL RTD project [eurisolrtd]. The possibility of increasing these numbers has been discussed ever since then. Different options have been proposed, such as multiple target of MgO in series for the production of ${}^{18}\text{Ne}$. However, the production of the baseline isotopes will again be carefully

investigated within the EURISOL DS [eurisol] target tasks. Any improvements will translate linearly into annual rate making it the most straightforward way to increase the flux of (anti-neutrinos). In table 2 the required production rate for version 2 of the beta-beam facility with an annual rate of 1.1×10^{18} anti-neutrinos and 2.9×10^{18} [bou03] neutrinos are given.

Table 2: The assumed possible production [aut02] (for atoms at the entrance to the ECR) compared to a theoretically required production [version2].

	Nominal production rate [ions/s]	Required production rate [ions/s]	Missing factor
6He	2×10^{13}	2×10^{13}	1
18Ne	8×10^{11}	1.9×10^{13}	24

Tests with a 28 GHz ECR source at LPSC in Grenoble [sor04] have demonstrated that short pulses of noble gas in a high charge state can be extracted through the so-called pre-glow effect. Theoretical estimates show that the increased plasma density at 60 GHz could produce an intense pulsed beam of noble gases suitable for a multi-turn injection into a synchrotron. Such a source, at a 10 Hz repetition rate, could feasibly generate 10-20 microsecond long pulses of up to a few 10^{12} charges per pulse. A similar set-up in Louvain-La-Neuve operates with high efficiency for the production of a radioactive ion beam where the ions are produced on-line in a thick target [lln]. The He ions would, in the proposed 60 GHz ECR source, be extracted almost exclusively in the highest charge state (2^+) but the Ne ions would be extracted as a spectrum of several charge states with a maximum of 30% in one single state.

A high energy ion beam requires acceleration in large synchrotrons. Such synchrotrons are costly and, consequently, the re-use of existing accelerators at e.g. CERN is essential to keep the cost down. The PS and SPS at CERN were built for fixed target physics with protons and have later been adapted to accelerate ions for the CERN heavy-ion physics programme. The aperture and space charge limitations of these two accelerators have been carefully studied over many years of operation. This is a great help for the study as it sets strict limits for the number of charges per bunch that can be injected and accelerated, but it also shows that neither of the two machines are well adapted for very high intensity ion beams. Already in the first study a modification of the RF system was proposed to reduce the tune shift at injection in the SPS.

The activation of magnets and tunnels is a major concern for all accelerators and in particular for the beta-beam where the ions decay (and are lost) during acceleration. The lost ions will be distributed all along the circumference making it very difficult to construct efficient beam collimation and magnet protection systems. Following the first study a first attempt was made [mag02] to estimate the decay losses. The losses in the PS for He are serious enough to make further hands on maintenance of the machine difficult while the losses in the SPS are comparable to losses already experienced with operational beams. The losses in the decay ring were simulated and shown to be important but not a cause of major concern as the subsequent secondary production of radioactivity in the surrounding rock were shown to be well below national limits.

Synchrotrons are highly efficient considering space and number of RF cavities compared to linacs, but with the inherent limitation of requiring a long repetition time between macropulses as the synchrotron, once “filled”, has to ramp the magnetic field to accelerate. This produces long gaps between injections, of ions from the source. For the beta-beam

baseline scenario, the repetition time between injections can be several seconds and, as the injection is only sustained for a second a large fraction of the ions produced are already lost at the exit of the ion source.

A further challenge is caused by the rather low energies reachable with the SPS at CERN as this implies a low duty factor ($< 2 \times 10^{-3}$) of the neutrino beam to permit a clean discrimination between atmospheric background and signal in the detector. This requires that the ions are kept in a few short bunches in the decay ring. The proposed injection scheme in the baseline scenario results in an acceptable duty factor, but due to longitudinal aperture limitations, it quickly saturates and prevents a further increase of the total number of stored ions. The long-term stability of the short high-intensity ion bunches in the decay ring due to intra-beam scattering is another point of concern which has to be addressed in the next stage of the study.

5. Optimization of the beta-beam baseline

A powerful method to understand the functional dependence of the many parameters influencing the figure of merit for a certain facility is available with modern analytical calculation software. The method requires that a symbolic analytical description is produced of the full accelerator chain and such a description has been done for the proposed beta-beam facility at CERN using Mathematica [wolfram,version1,version2]. The results can be visualized as e.g. two-dimensional plots for the figure of merit as function of the different machine parameters. However, it is important to note that while such plots will help to identify the right parameter space for a requested machine design, they do not themselves guarantee that a realizable technical solution can be found. The work presented here has been done respecting the constraint of using the existing CERN accelerators, PS and SPS. In addition, the current decay ring design [pay05] is used even though it has been assumed that the Lorenz gamma of the stored ions can be changed without major changes to it. For the stacking we have assumed that 15 ^6He bunches and 20 ^{18}Ne bunches can be merged in the decay ring without major losses. The situation is slightly better for ^{18}Ne , which due to a more advantageous charge-to-mass ratio, will see an almost three times larger longitudinal acceptance of the decay ring. Theoretically this would permit up to 45 merges for ^{18}Ne , but known limits in low level RF beam control precision restricts it to a maximum of 20 merges. All parameters for the EURISOL DS baseline of May 2005 are documented in the appendixes of [version1,version2].

5.1 Annual rate

The figure of merit for the beta-beam is the rate (R) of (anti-)neutrinos at the end of one of the straight section over a given period and it can be expressed as,

$$R = \frac{I_{in} \times f}{T_{rep}} \times \left(1 - 2^{-\frac{mr \times T_{rep}}{\gamma_{top} t_{half}}} \right) \times T_{run} \quad (1)$$

where the mr is the number of merges that can be done in the decay ring without major losses from the merging process itself, T_{rep} is the repetition period for the fills in the decay ring, γ_{top} the gamma factor of the decay ring, t_{half} the half life at rest for the ions, I_{in} the total number of

ions injected into the decay ring for each fill, f the fraction of the decay ring length for the straight section generating the neutrino beam and T_{run} the length in seconds of the run. The first term gives the rate per chosen period for the ideal decay ring in which an endless number of merges can be accepted, the second term gives the limit set by the restriction on maximum number of merges that can be accepted due to longitudinal emittance limitations and the third term is by convention chosen as the length of a “snow mass year” [snowmass] which is used as an international standard to calculate running time of high energy physics experiments and it is 10^7 seconds long.

5.2 Gamma dependence of the CERN baseline

The gamma dependence of the rate is mainly due to the gamma dependence of the acceleration time in the SPS, the lifetime of the ions in the decay ring and of the available longitudinal acceptance in the decay ring. The acceleration time in the SPS has a minimum length depending on hardware limitations and also increases in steps of 1.2 seconds due to the basic timing period of the CERN accelerator complex. For a given radio-frequency (RF) voltage, the longitudinal acceptance will to first order scale as the square root of gamma. In figure 2 dependence of ${}^6\text{He}$ is shown. The maximum gamma that can be reached with the CERN SPS for ${}^6\text{He}$ is 150 and for ${}^{18}\text{Ne}$ 250.

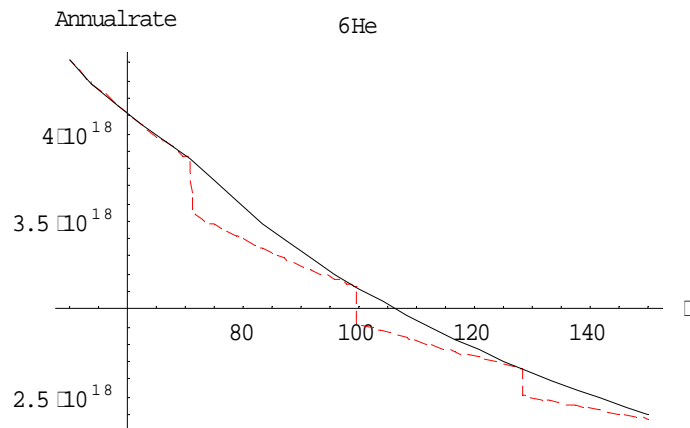


Figure 2: The annual rate of anti-neutrinos from ${}^6\text{He}$ as a function of γ . The red dashed line shows the annual rate with the given source rate respecting a basic period of 1.2 seconds of the CERN accelerator complex. The solid line shows the same dependence but with a “smooth” choice of acceleration time for PS and SPS.

5.3 Duty factor dependence

The available longitudinal acceptance for stacking in the decay ring can be increased by increasing the number of bunches in the decay ring. This will increase the duty factor which is the total length in time of all bunches in the decay ring divided by the revolution time. The present limits of 15 merges for ${}^6\text{He}$ and 20 merges for ${}^{18}\text{Ne}$ will truncate the stacking well before decay rate equals the stacking rate. In the formula for the annual rate (1) the second term is the result of this truncation. The duty cycle in the beta-beam baseline is 4.5×10^{-3} for 20 bunches ${}^6\text{He}$ and 3.9×10^{-3} for 20 bunches of ${}^{18}\text{Ne}$ at a gamma of 100. Figure 3 shows the annual rate for ${}^{18}\text{Ne}$ as a function of number of bunches in the decay ring. [version1,version2].

The constraint on the duty-cycle is set by the atmospheric background and it becomes less severe at higher energies of the neutrinos. A beta-beam facility operating at a higher gamma

than the nominal value of 100 could accept a larger duty factor. Consequently, to fully understand the evolution of the annual rate as a function of gamma the possibility to increase the duty factor must be fully explored.

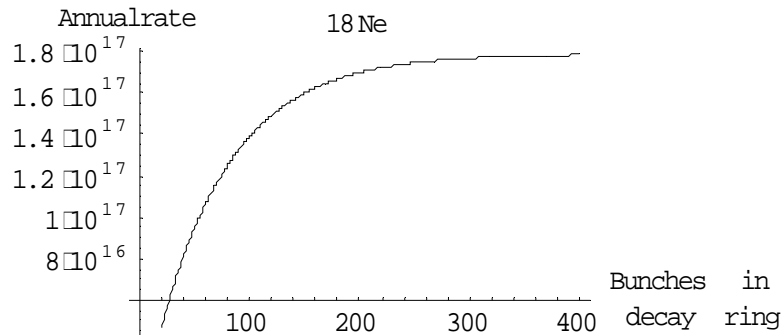


Figure 3: The annual rate of anti-neutrinos from the decay of ^{18}Ne as a function of number of bunches. The curve starts to the left with 8 bunches at a duty factor of 2×10^{-3} and saturates when the decay rate equals the stacking rate.

5.4 Accumulation

In the beta-beam baseline the ions are produced for one second and thereafter accelerated during several seconds to a γ of 100. In theory, ions could be produced continuously as long as they can be stored while waiting for acceleration. The optimum storage time determined by the life-time of the ion and the required acceleration time as can be seen in the curve in figure 4. In these calculations the accumulation is done before the PS in a storage ring fitted with a cooling system for an efficient storage through multiple injection cycles [rnbstorage].

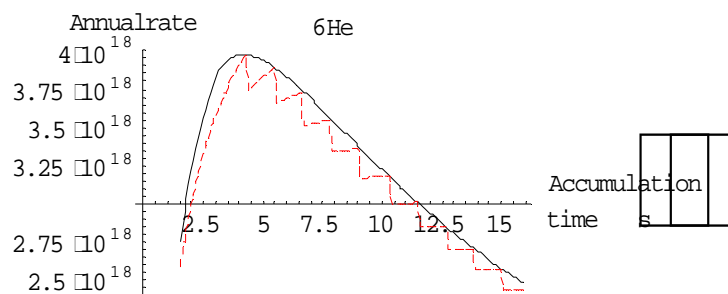


Figure 4: The annual rate of anti-neutrinos from ^6He as a function of the accumulation time in a storage ring before the PS. The red dashed line shows the annual rate respecting a basic period of 1.2 seconds of the CERN accelerator complex. The solid line shows the same dependence but with a “smooth” choice of acceleration time for PS and SPS. All other parameters have been taken from the baseline design

6. Conclusions

The conceptual design of the first study of the beta-beam facility has been taken a step further with the start of the EURISOL design study's beta-beam task. The first objective is to establish the annual rate of (anti-)neutrinos from a given production rate [aut02] with the re-use of the PS and SPS for acceleration. The next step, which already started, is to optimize the many machine parameters in the baseline to establish the maximum achievable rate. The work done so far shows that: i) the increase of production rate translates linearly into an increase of the rate [lin05], ii) an increase of the duty cycle permits more particles to be stored in the decay ring and iii) accumulation can potentially be a powerful tool to make better use of the ions produced. The gamma dependence of the rate is often forgotten in physics reach calculations, but it should be noted that it is linked to longitudinal acceptance and duty factor and has to be calculated specifically for each choice of machine parameters.

A green field study to establish the ultimate limit of the beta-beam concept is beyond the scope of the on-going design study. However, before a final decision on the construction of the next generation neutrino source, such a study should be done.

References

- [aut02] B.Autin, M. Benedikt, M. Grieser, S. Hancock, H. Haseroth, A. Jansson, U. Köster, M. Lindroos, S. Russenschuck and F. Wenander, "The acceleration and storage of radioactive ions for a neutrino factory", CERN/PS 2002-078 (OP), Nufact Note 121, Proceedings of Nufact 02, London, UK, 2002, J. Phys. G: Nucl. Part. Phys. 29 (2003) 1785-1795
- [bou03] J. Bouchez, M. Lindroos and M. Mezzetto, "Beta-beams: Present design and expected performance", In the proceedings of Nufact 03, New York, 2003
- [eurisol] <http://eurisol.org>
- [eurisolrtd] "The EURISOL report", Edt. J. Cornell, GANIL, Caen, 2003, European commission contract No. HPRI-CT-1999-500001
- [han03] M. Benedikt, S. Hancock and J.-L. Vallet, "A proof of principle of asymmetric bunch pair merging", AB-note-2003-080 MD
- [han05] M. Benedikt and S. Hancock, Nucl. Instr. and Meth. A550(2005)1-5
- [isol] H.R. Ravn and B.W. Allardyce, "On-Line Mass Separators", in Treatise on Heavy-Ion Science, Edt. D. A. Bromley, Plenum Press, New York, 1989, ISBN 0-306-42949-7
- [lin04] M. Lindroos and T. Nilsson, "The acceleration and storage of radioactive ions for a beta-beam facility", Proceedings of RNB6, Argonne, USA, 2003, Nucl. Phys. A 746(2004)316c-322c
- [lln] M. Gaelens, M. Cogneau, M. Loiselet and G. Ryckewaert, "Post-acceleration of ^7Be at the Louvain-la-Neuve radioactive ion beam facility", Nucl. Instr. Meth. B 204(2003)48-52
- [mag02] M. Magistris and M. Silari, "Parameters of radiological interest for a beta-beam decay ring", TIS-2003-017-RP-TN
- [pay05] A.Chance and J. Payet, "Storage ring for a beta-beam facility", Beta-beam task meeting, July 2005
- [rnbstorage] T. Fritioff, M. Grieser, M. Lindroos, A. Trzcinska, Nucl. Phys. A 746(2004)384c-388c
- [snowmass] <http://snowmass2001.org/>

- [sor04] P. Sortais, J.-L. Bouly, J.-C. Curdy, T. Lamy, P. Sole, T. Thuillier, J.-L. Vieux-Rochaz, and D. Voulot, "ECRIS development for stable and radioactive pulsed beams", *Rev. Sci. Instrum.* 75 (2004) 1610-1612
- [version1] M. Benedikt, S. Hancock, M. Lindroos, "Parameter and Intensity Values, Version 1, April 2005", EURISOL DS/TASK12/TN-05-01
- [version2] M. Benedikt, A. Fabich, S. Hancock, M. Lindroos, "Parameter and Intensity Values, Version 2, July 2005", EURISOL DS/TASK12/TN-05-03
- [wolfram] © Wolfram Research, Inc., <http://www.wolfram.com>
- [zuc02] P. Zucchelli, "A novel concept for a neutrino factory: the beta-beam", *Phys. Let. B*, 532 (2002) 166-172

Physics with a first very low energy Betabeam

We describe the importance of having low-energy (10-100 MeV) neutrino beams produced through the decay of boosted radioactive ions (“beta-beams”). We focus on the interest for neutrino-nucleus interaction studies and their impact for astrophysics, nuclear and particle physics. In particular, we discuss the relation to neutrinoless double-beta decay. Finally, we mention the status as far as the feasibility of low-energy beta-beams is concerned.

1 Introduction

Nuclei are a wonderful laboratory for searches on fundamental issues, such as the knowledge of the neutrino mass scale, or of the Majorana versus Dirac nature of neutrinos. Nuclei can also be a beautiful tool for the search of new physics. The original idea of “beta-beams”, first proposed by Zucchelli [zu02], enter in this category. Beta-beams use the beta-decay of boosted radioactive ions to produce well known electron (anti)neutrino beams, while the conventional way exploits the decay of pions and muons. This simple but intriguing idea has opened new strategies, thanks to the future radioactive ion beams, at present under study, in various nuclear physics laboratories. In fact, the planned intensities of 10^{11-13} ions/s can actually render neutrino accelerator experiments using ions, feasible.

In the original paper [zu02], a new facility is described, based on the beta-beam method, the central motivation being the search for CP violation in the lepton sector — the Maki-Nakagawa-Sakata-Pontecorvo (MNSP) matrix, relating the neutrino flavor and mass basis, might indeed be complex. With this aim the ions would be accelerated to 60-100 GeV/A (or $\gamma=60-100$, where γ is the Lorentz factor), requiring accelerator infrastructure like the PS and SPS at CERN, as well as a large storage ring pointing to an (enlarged) Frejus Underground Laboratory, where a big detector would be located.

Very soon the interest of this new concept for the production of low-energy neutrino beams has been recognized [vol04]. Here the ions are boosted to a much lower γ , i.e. 5-15. High energy scenarios have been proposed [jj04]. Afterwards, requiring different (or revised) accelerator infrastructures to boost the ions at very high γ ($\gamma \gg 100$). (Note that for this reason the original scenario [zu02] is sometimes referred to in the literature as “standard”, or misleadingly “low-energy”.) Detailed works exist at present both on the feasibility [feas] as well as on the physics potential of the standard [stand] scenario, contributing to determining the conditions for the best CP violation sensitivity, in possible future searches. A feasibility study is now ongoing within the Eurisol Design Study [eurisol]. Here we will focus on the physics potential of low-energy beta-beams.

2 Low energy beta-beams

2.1 Physics Motivations

The idea of establishing a facility producing low-energy neutrino beams, based on beta-beams, has been proposed in [vol04]. This opens new opportunities, compared to the original scenario. First one might use the ion decay at rest as an intense neutrino source in order to explore neutrino properties that are still poorly known, such as the neutrino magnetic moment

[mc04]. In fact direct measurements achieving improved limits are precious, since the observation of a large magnetic moment points to physics beyond the Standard Model.

The interest of low-energy beta-beams in the tens of MeV, to perform neutrino-nucleus interaction studies, has been discussed in [vol04,ser04]. At present, there is a limited number of measurements available in this domain (essentially on three light nuclei), so that theoretical predictions are of absolute necessity. Getting accurate predictions can be a challenging task, as the discrepancies on the ^{12}C [c12] and ^{208}Pb [Pb] cross sections have been demonstrating. Neutrino-nucleus applications are numerous and span from a better knowledge of neutrino detector response using nuclei, like supernova observatories or in oscillation experiments, to nuclear astrophysics, for the understanding of processes like the nucleosynthesis of heavy elements. (More information can be found e.g. in [ku94,review].)

In [vol04] we have pointed out that performing neutrino-nucleus interaction studies on various nuclei would improve our present knowledge of the "isospin" and "spin-isospin" nuclear response (the nuclear transitions involved in charged-current reactions are in fact due to the isospin, like e.g. t_{\pm} , and spin-isospin, like e.g. σt , operators). A well known example is given by the super-allowed Fermi transitions (due to the isospin operator), which are essential for determining the unitarity of the Cabibbo-Kobayashi-Maskawa (CKM) matrix, the analog of the MNSP matrix in the quark sector. Another (less known but still intriguing) example is furnished by the so-called Gamow-Teller transitions (these are due to the spin-isospin operator) in mirror nuclei, which can be used to observe second class currents, if any. These terms transform in an opposite way under the G -parity transformation — the product of charge-conjugation and of a rotation in isospin space — as the usual vector and axial-vector terms [vol03], and are not present in the Standard Model. In [vol04] we have pointed out that spin-isospin and isospin states of higher multipolarity (than those just mentioned) contribute significantly to the neutrino-nucleus cross sections, as the energy of the neutrino increases. Such contributions are larger when the nucleus is heavier. Since low-energy beta-beams have the specificity that the average energy can be increased by increasing the Lorentz boost of the ions (more precisely $\langle E\nu \rangle \approx 2 \gamma Q_{\beta}$), they appear as an appropriate tool for the study of these states. Apart from their intrinsic interest, neutrino-nucleus interaction measurements would put theoretical predictions for the extrapolation to exotic nuclei useful for astrophysical applications on really solid grounds. They are also important for the open question of the neutrino nature.

One of the crucial issues in neutrino physics is to know if neutrinos are Dirac or Majorana particles. The answer to this question can be furnished for example by the observation of neutrinoless double beta decay in nuclei, since this lepton violating process can be due to the exchange of a Majorana neutrino. While the present limit is of about 0.2 eV [ex], future experiments aim at the challenging 50 meV energy range. However, it has been longly debated that the theoretical situation, as far as the half-life predictions are concerned, should be clarified: different calculations present significant variations for the same candidate emitters. Reducing these differences certainly represents an important theoretical challenge for the future, and one might hope that dedicated experiments will help making a step forward [zu05]. One way to constrain such calculations is by measuring related processes, such as beta-decay [su05], muon capture [kor03], charge-exchange reactions [jo] and double-beta decay with the emission of two neutrinos [vo03] (the latter process is allowed within the Standard Model and does not tell us anything about the neutrino nature). Such a procedure has been used since a long time. However each of these processes bring part of the necessary information only.

Recently we have been showing that there is a very close connection between neutrinoless double-beta decay and neutrino-nucleus interactions [vol05]. In fact, by rewriting the neutrino exchange potential in momentum space and by using a multiple decomposition, the two-body transition operators, involved in the former, can be rewritten as a product of one-body operators, which are essentially the same as the ones involved in neutrino-nucleus interactions. (Note, however, that there keep being some differences like for example short range correlations which can play a role in the two-body process, but not in the one-body one.) Therefore, besides the above-mentioned processes an improved knowledge of the nuclear response through either low-energy beta-beams or conventional sources (decay of muons at rest) could help constraining the neutrinoless half-life predictions as well. Figures 1 and 2 show the contribution of different states for two impinging neutrino energies on ^{48}Ca taken just as an illustrative example.

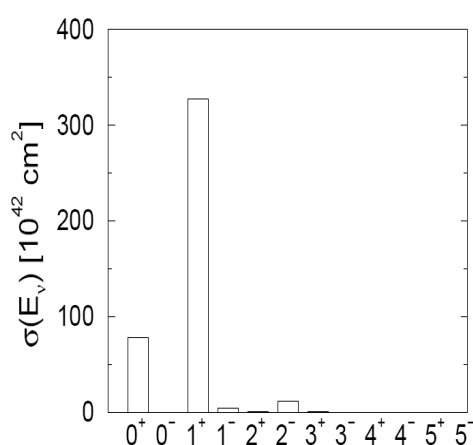


Figure 1 Contribution of the states of different multipolarity to the total charged-current $\nu_e - ^{48}\text{Ca}$ cross section for neutrino energy $E_\nu=30$ MeV. The histograms show the contribution of the Fermi ($J^\pi=0^+$), the Gamow-Teller (1^+) and the spin-dipole ($0, 1, 2^-$) states and all higher multipoles up to 5 [vol05].

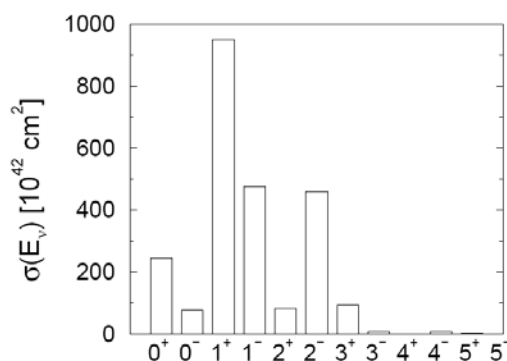


Figure 2 One can see that the Gamow-Teller transition is giving the dominant contribution at low neutrino energies, while many other states become important when the neutrino impinging energy increases. These states are an essential part of the neutrinoless double-beta decay half-lives as well [vol04].

2.2 A small storage ring

The main aim of the work in Ref. [ser04] has been to calculate exactly: i) the neutrino-nucleus interactions rates expected at a low-energy beta-beam facility, by using parameters from the first feasibility study [feas]; ii) to study how these scale by changing the geometry of the storage ring. In particular, two sizes have been considered: a small one, i.e. 150 m straight sections and 450 total length, like the one planned for the future GSI facility [gsi]; a large one, having 2.5 km straight sections and 7 km total length, such as the one considered in the original beta-beam baseline scenario [zu02]. Table 1 shows the events for deuteron, oxygen, iron and lead, taken as typical examples, the detector being located at 10 m from the storage ring.

Table 1: Neutrino-nucleus interaction rates (events/year) at a low-energy beta-beam facility [ser04]: Rates on deuteron, oxygen, iron and lead are shown as examples. The rates are obtained with $\gamma=14$ as boost of the parent ion. The neutrino-nucleus cross sections are taken from referred references. The detectors are located at 10 meters from the storage ring and have cylindrical shapes ($R=1.5$ m and $h=4.5$ m for deuteron, iron and lead, $R=4.5$ m and $h=15$ m for oxygen, where R is the radius and h is the depth of the detector). Their mass is indicated in the second column. Rates obtained for two different storage ring sizes are presented: the small ring has 150 m straight sections and 450 total length, while the large ring has 2.5 km straight sections and 7 km total length. Here 1 year = 3.2×10^7 s.

Reaction	Ref.	Mass (tons)	Small Ring	Large Ring
$\nu+D$	[11]	35	2363	180
$\bar{\nu}+D$	[11]	35	25779	1956
$\nu+^{16}O$	[23]	952	6054	734
$\bar{\nu}+^{16}O$	[23]	952	82645	9453
$\nu+^{56}Fe$	[24]	250	20768	1611
$\nu+^{208}Pb$	[25]	360	103707	7922

One can see that interesting interaction rates can be obtained on one hand and that clearly a small devoted storage ring is more appropriate for such studies on the other hand. The physical reason is simple. Since the emittance of the neutrino fluxes is inversely proportional to the γ of the ions, only the ions which decay close to the detector contribute significantly to the number of events, while those who decay far away see the detector under a too small opening angle. The complementarity between a low-energy beta-beam and conventional source is discussed in [mcl04].

3 Conclusions

The use of the beta-beam concept to produce neutrino beams in the tens of MeV energy range is very appealing. If both electron-neutrino and electron (anti)neutrino beams of sufficiently high intensities can be achieved, low-energy beta-beams can offer a flexible tool, where the average neutrino energy can be varied by varying the γ of the ions. The studies realized so far indicate clearly that a small devoted storage ring is more appropriate to obtain such beams in particular for performing neutrino-nucleus interaction studies, a promising axis of research. We have particularly discussed the interest of such measurements for a better knowledge of the nuclear response relevant for neutrinoless double-beta decay searches. The feasibility

study of the small storage ring is now ongoing within the Eurisol Design Study. Several issues need to be addressed (e.g. stacking ion method, duty factor). The realization of low-energy beta-beams would be a *proof-of-principle* that the beta-beam concept works.

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References

- [c12] E. Kolbe et al, Phys. Rev. C 52 (1995) 3437; S.K. Singh, N.C. Mukhopadhyay and E. Oset, Phys. Rev. C 57 (1998) 2687; E. Kolbe, K. Langanke and P. Vogel, Nucl. Phys. A 613 (1997) 382; A.C. Hayes and I.S. Towner, Phys. Rev. C 61 (2000) 044603; C. Volpe et al, Phys. Rev. C 62 (2000) 015501; N. Jachowicz et al, Phys. Rev. C 65 (2002) 025501; A. Brotugno and G. Co, nucl-th/0506028.
- [eurisol] See <http://www.ganil.fr/eurisol/>.
- [ex] L. Baudis et al, Phys. Rev. Lett. 83(1999)41; H.V. Klapdor-Kleingrothaus et al, Eur. Phys. J. A 12 (2001) 147; C.E. Aalseth, Phys. Rev. D 65, (2002).
- [feas] B. Autin et al, J. Phys. G 29 (2003) 1785 [physics/0306106]. See also <http://beta-beam.web.cern.ch/beta-beam/> and the contribution of M. Lindroos to this volume.
- [gsi] See <http://www.gsi.de/>.
- [jj04] J. Burguet-Castell et al, Nucl. Phys. B 695 (2004) 217 [hep-ph/0312068].
- [jo] J. Bernabeu et al, IPNO/TH-88-58, FTUV-88/20; H. Akimune et al, Phys. Lett. B 394, 23 (1997); S. Rakers et al, Phys. Rev. C 70, 054302 (2004) and 71, 054313 (2005).
- [kor03] M. Kortelainen and J. Suhonen, Europhys. Lett. 58, 666 (2002); M. Kortelainen and J. Suhonen, Phys. Atom. Nucl. 67, 1202 (2004).
- [ku94] K. Kubodera and S. Nozawa, Int. J. Mod. Phys. E 3 (1994) 101.
- [mc04] G.C. McLaughlin and C. Volpe, Phys. Lett. B 591 (2004) 229 [hep-ph/0312156].
- [mcl04] G.C. McLaughlin, Phys. Rev. C 70 (2004) 045804 [nucl-th/0404002].
- [nuFe] E. Kolbe and K. Langanke, Phys. Rev. C 63 (2001) 025802.
- [nuO] E. Kolbe, K. Langanke, P. Vogel, Phys. Rev. D 66 (2002) 013007; W.C. Haxton, Phys. Rev. D 36 (1987) 2283.
- [nuPb] J. Engel, G.C. McLaughlin, C. Volpe, Phys. Rev. D 67 (2003) 013005.
- [Pb] G.M. Fuller, W.C. Haxton and G.C. McLaughlin, Phys. Rev. D 59 (1999) 085005; E. Kolbe and K. Langanke, Phys. Rev. C 63 (2001) 025802; C. Volpe et al, Phys. Rev. C 65 (2002) 044603; N. Jachowicz, K. Heyde and J. Ryckebush, Phys. Rev. C 66 (2002) 055501; M. Sajjad Athar et al, nucl-th/0506046.
- [review] A.B. Balantekin, Prog. Theor. Phys. Suppl. 146 (2003) 227 [nucl-th/0201037]; A.B. Balantekin and G.M. Fuller, J. Phys. G 29 (2003) 2513 [astro-ph/0309519]; C. Volpe, Nucl. Phys. B 143 (2005) 43 [hep-ph/0409249]; E. Kolbe et al, J. Phys. G 29 (2003) 2569, and references therein.
- [ser04] J. Serreau and C. Volpe, Phys. Rev. C 70 (2004) 055502 [hep-ph/0403293].

- [stand] Mezzetto, J. Phys. G29 (2003) 1771 [hep-ex/0302007]; J. Bouchez, M. Lindroos, M. Mezzetto, AIP Conf. Proc. 721, (2004) 37 [hep-ex/0310059]; M. Mezzetto, Nucl. Phys. Proc. Suppl. 143 (2005) 309 [hep-ex/0410083] ; A. Donini et al, Phys. Lett. B 621 (2005) 276 [hep-ph/0411402]; C. Volpe, Nucl. Phys. A 752 (2005) 38 [hep-ph/0409357]; J. Burguet-Castell et al, Nucl. Phys. B 725, 306 (2005) [hep-ph/0503021]; J. Bernabeu, J. Burguet-Castell, C. Espinoza, M. Lindroos, hep-ph/0505054; P. Huber, M. Lindner, M. Rolinec, W. Winter, [hep-ph/0506237]; see also contributions of J. Burguet-Castell and M. Mezzetto to this volume.
- [su05] K. Muto, E. Bender and H.V. Klapdor, Z. Phys. A 334, 177 (1989); M. Aunola, J. Suhonen, Nucl. Phys. A 602, 133 (1996).
- [vo03] V.A. Rodin, A. Faessler, F. Simkovic, P. Vogel, Phys. Rev. C 68, 044302 (2003).
- [vol03] N.A. Smirnova and C. Volpe, Nucl. Phys. A 714 (2003) 441, and references therein.
- [vol04] C. Volpe, Jour. Phys. G 30 (2004) L1 [hep-ph/0303222].
- [vol05] C. Volpe, J. Phys. G 31 (2005) 903 [hep-ph/0501233].
- [zu02] P. Zucchelli, Phys. Lett. B 532 (2002) 166.
- [zu05] K. Zuber et al, article in preparation.

R&D on Neutrino Detectors: status and priorities

The challenge of very large detectors, assembling enough detector mass adequately instrumented for effective detection of neutrinos events, accompanies neutrino physics since its early days. It has never been easy and has to deal today with a much larger scale. Several detectors options are being considered, in conjunction to different neutrino sources. The R&D in progress must be given all possible support .

1. Low-Z Tracking Calorimetry

Low-Z Tracking Calorimetry optimizes the detection of electrons in the final state by using a fine sampling in terms of radiation lengths, leading to the choice for a low-Z passive material. This is the technique used for the study of $\nu_\mu e$ scattering in the CHARM II experiment at CERN [3]. Here we discuss the status of the design of the NOvA experiment proposed for the detection of $\nu_\mu - \nu_e$ oscillations in the off-axis NuMI beam at Fermilab [4], with the observation of θ_{13} as its prime aim.

The NOvA detector foresees a mass one order of magnitude larger than MINOS [5] and at the same time, in relation to its aim, a finer sampling ($\Delta X_0 < 0.3$ to be compared to 1.5 with MINOS, which has iron as passive material). The principal technical issue is to improve the performance and substantially reduce the unitary cost of the trackers. The main technological innovation consists in the use of liquid scintillator read by Avalanche Photodiodes (APDs), instead of plastic scintillator read by multianode PMTs.

The detector is a "Totally Active" Scintillator Detector (TASD), with a 30 kton mass, of which 24 kton are of liquid scintillator and the rest of PVC. In TASD, the scintillator modules have cells along the beam 6.0 cm long and 3.9 cm width. TASD consists of a single block with overall dimensions 15.7x15.7x132 m³. Lacking the particle board, the PVC must provide a self-supporting structure for a detector as high as a five-storey building. Since last year, progress has been made also in the mechanical design and in the assembling methods.

The electromagnetical energy resolution is $\Delta E/E \sim 10\%/\sqrt{E(\text{GeV})}$, the almost continuous pulse height information along the track helps in e/π_0 discrimination.

The use of APDs results in a considerably lower cost than with PMTs. 2x16 pixel APDs are commercially produced in large quantities and already foreseen for the CMS experiment at LHC. The high quantum efficiency, about 85%, allows to have longer strips and less readout channels.

The NOvA design is in constant progress. If funding would begin in late 2006, the NOvA detector could be ready in 2011. It is worth noting that the progress with the development of the trackers is potentially useful also for magnetised iron spectrometers for neutrino factories or colliders.

2. Magnetised Iron Spectrometers

This technique is conventional, but the mass to be considered is one order of magnitude larger than for present magnetised iron spectrometers, like MINOS.

Recent studies indicate that a magnetised iron toroidal spectrometer of the required mass is feasible [15]. On one hand, the design of toroids with radius up to 10 m can be extrapolated from MINOS, with thicker plates for larger planes. On the other hand, the No_A liquid scintillator technology with APD readout allows to have transverse dimensions twice as large than in MINOS. Such a detector concept permits direct use of the experience with MINOS.

The India Neutrino Observatory (INO) [16] foresees a dipole magnet equipped with RPCs, as in MONOLITH [17]. The INO basic motivation is the study of atmospheric neutrinos like with MONOLITH; a future use in a very long baseline experiment with a ν -factory is envisaged. The investigations started with MONOLITH on the detector performance for design optimization have to be continued, in comparison with other detectors.

A conceptual spectrometer based on a 40 kton iron solenoid magnetised at 1 T by a superconducting coil, with embedded solid scintillator rods as the active detectors, has been presented in Ref. [14]. Considerable more work is required to define its features and assess its practical feasibility.

In general, practical problems (mechanics, magnet design, etc.) must be thoroughly addressed. In addition, as already mentioned more simulation work is required in order to understand and optimize the performance by a proper choice of the main detector parameters.

3. Water \checkmark Cerenkov

Water \checkmark Cerenkovs can provide a very large target mass and, if the photo-sensors have a sufficient density, a sensitivity down to the low energies of solar neutrinos. Its capabilities concern ν astrophysics, ν oscillations and proton decay. Above a few GeV, DIS dominates over QE scattering and leads to frequent multi-ring events more complicated to reconstruct. A similar limitation in energy comes from difficulties in the e/π_0 discrimination at high energies. The technique is not suitable at the high energies of ν -factories, where, in addition, a muon charge measurement is needed. One should also remark that the low neutrino cross-section at low energies reduces the advantage given by the very large mass which can be assembled.

The detectors presently under study represent the third generation of successful detectors, with in each stage an increase by one order of magnitude in mass. The performance of Super-KamiokaNDE has been widely simulated and observed, providing a basis for a mass extrapolation by one order of magnitude. The performance as well as the limitations are well known, also from K2K and related tests.

Two detector designs are being carried out, namely Hyper-KamiokaNDE [6] and UNO [7]. The design of a detector to be located at Frejus has been also initiated.

Hyper-KamiokaNDE foresees two 500 kton modules placed sideways, each consisting of five 50 m long optical compartments. The cost is higher than for a single module, but maintenance with one module always alive and a staging in the detector construction become possible. The present design foresees about 200,000 20" PMTs, to be compared with 11,146 in Super-KamiokaNDE. The detector could be constructed in about 10 years, starting after a few years of T2K operation.

The UNO design provides a 650 kton mass subdivided into three optical compartments with different photo-sensor coverage. The central one has a 40% coverage as in Super-

KamiokaNDE, allowing to pursue solar ν studies. The side compartments have 10% coverage. The number of 20" PMTs is two times smaller than with 40% coverage for the full detector, but still amounts to 56,650. The question arises as to whether this subdivision, with its non-uniformity given by the lower coverage in 2/3 of the detector, is the optimal solution to reduce the global cost.

By giving appropriate aspect ratio and shape to the cavern, its realization does not seem a problem.

A large fraction of the total detector cost, reaching about 1/2 or more if PMTs are used, comes from the photo-sensors. The present cost of 100,000-200,000 20" PMTs is hard to cover. Moreover, their production would take about 8 years, leading also to storage problems. The main issue is thus the development and acquisition of photo-sensors at acceptable cost and production rate, as well as an improvement of their performance. A better time resolution would improve neutrino vertexing and single photon sensitivity would give better ring reconstruction. A strong collaboration with industry is essential, as for the development of 20" PMTs by Hamamatsu for KamiokaNDE and Super-KamiokaNDE.

The Hamamatsu 20" glass bulbs are manually blown by specially trained people. Automatic manufacturing does not seem a practical solution to reduce the cost and speed-up the production rate, as the required quantity is still small compared to commercial PMTs. The question is whether a size smaller than 20", with an appropriate coverage, is more practical and cost effective. With a smaller size, automatic bulb manufacturing is eased and the risk of implosion decreased, with a possible saving in the plastic protection to damp implosions. For R&D on PMTs, collaborations have been established with industries also in Europe (Photonis) and USA.

To explore alternatives to PMTs, studies on new photo-sensors have been launched. In addition to reduce cost, while improving production rate and performance, it is essential to achieve the long term stability and reliability which is proven for PMTs.

Hybrid Photo-Detectors (HPD) are being developed by Hamamatsu, in collaboration with ICRR of Tokyo University. The HPD glass envelope is internally coated with a photo-cathode and a light reflector. Electrons are accelerated by a very high voltage towards an Avalanche Diode (AD). The strong electron bombardment results in a high gain (about 4500 for 20 kV voltage) in this first stage of amplification. It gives a remarkable single photon sensitivity and makes ineffective the AD thermally generated noise. The gain is lower than with PMTs, hence stable and highly reliable amplifiers are needed. The degradation in time resolution given by the transit time spread through the dynode chain is avoided, so that a 1 ns or so time resolution can be achieved, to be compared with the 2.3 ns of the 20" PMTs. The cost reduction with respect to PMTs essentially comes from the use of solid state devices like the AD, avoiding the complicated PMT dynode structure.

The principle has been proved with a 5" HPD prototype. Successful results from tests of an 13" prototype operated with 12 kV are now available, showing a $3 \cdot 10^4$ gain, good single photon sensitivity, 0.8 ns time resolution and a satisfactory gain and timing uniformity over the photo-cathode area. The next step will be the operation at a voltage up to 20 kV, giving a higher gain and a wider effective area of the photo-cathode. The development of HPD has been initiated also in Europe, in collaboration with Photonis.

4. Liquid Argon Time Projection Chamber

The Liquid Argon Time Projection Chamber (LAr TPC) provides an excellent imaging device and a dense neutrino target. It is a true "electronic bubble chamber" with a much larger mass (3 ton for Gargamelle, up to 100 kton now envisaged for LAr TPCs). The "state of the art" is given by the 300 ton T300 ICARUS [8] prototype module, tested at ground level in Pavia but not yet used in an experiment.

Two mass scales are foreseen for future experiments. For close ν detectors in Super-Beams, a mass of the order of 100 ton is envisaged. Detectors with 50-100 kton masses are proposed for ν oscillation, ν astrophysics and proton decay [9][10][11], implying a step in mass by more than two orders of magnitude with respect to the T300 ICARUS module.

Cryogenic insulation imposes a minimal surface/volume ratio. The modular ICARUS approach has thus to be abandoned for a single very large cryogenic module with about 1:1 aspect ratio.

To limit the number of readout channels, drift lengths have to be longer than the 1.5 m of the T300 module. The LAr TPC envisaged for the off-axis NuMI beam [9] [10] foresees 3 m drift lengths with readout, conceptually as in ICARUS. In another approach [11], the very strong attenuation over a much longer drift length (20 m) is compensated by the so-called "Double Phase" amplification and readout [12], tested on the ICARUS 50 litre chamber [13]. In both cases, the signal attenuation imposes a liquid Argon contamination by electronegative elements at the 0.1 ppb level.

In Ref. [11] a 100 kton detector consisting of a single module both for cryogenics and readout has been proposed. The 20 m drift in a field raised to 1 kV/cm results in a 10 ms drift time. With a 2 ms electron live-time in liquid Argon, the 6000 electrons/mm signal is attenuated by $e^{-t/\tau} \sim 1/150$ and becomes too low for a readout as in ICARUS. In the Double Phase readout, electrons are extracted from the liquid by a grid and collected in the gas phase using gas chamber techniques. The $\sigma \sim 3$ mm spread from the diffusion in the 20 m drift gives an intrinsic limit to the readout granularity.

The design of the single large cryostat [11] can benefit from the techniques developed for transportation and storage of large quantities of liquefied natural gas, kept at boiling temperature by letting it evaporate. A large cryogenic plant is needed for the initial filling and for the continuous refilling to compensate the evaporation.

The R&D for the detector of Ref. [11] foresees: drift under 3 atm pressure as at the tank bottom; charge extraction and amplification; imaging devices; cryogenics and cryostat design, in collaboration with industry; a column-like prototype with 5 m drift and double-phase readout, with a 20 m drift simulated by a reduced E field and by a lower liquid Argon purity; test of a prototype in a magnetic field (as for use in a ν -Factory); underground safety issues.

Tests on a 10 litre LAr TPC inserted in a 0.55 T magnetic field have been performed [18], together with a tentative layout for the implementation of a large superconducting solenoidal coil into the design of Ref. [11].

The experience accumulated in two decades with ICARUS is very important. However, a substantial R&D is required, at an extent which depends on the detector design and on the

(underground) location. In the design of very large detectors, with or without magnetic field, one has to proceed from concepts or conceptual designs to a practical design.

5. The Emulsion Cloud Chamber

The use of Emulsion Cloud Chamber (ECC) has been proposed to detect "silver events" from ν_e - ν_τ oscillations at a ν -Factory, in order to complement the ν_e - ν_μ "golden events" in resolving θ_{13} - δ ambiguities [19].

The ECC of the OPERA experiment [20] in the CNGS beam consists of a multiple sandwich of lead and nuclear emulsion sheets. The production and decay of τ leptons is expected to be observed with a very low background thanks to the emulsion sub- μm resolution. The 1.8 kton OPERA target is built up with about 200,000 lead-emulsion "bricks".

At the ν -Factory, a 4 kton mass is envisaged. By scanning only "wrong sign" muons as those coming from τ decays, the scanning load is expected to be comparable with that in OPERA. An increase in the speed of the automatic microscopes is foreseeable. The OPERA Collaboration is working to provide a milestone in the application of the ECC technique.

REFERENCES

1. J. Burguet-Castell et al., hep-ph/0312068 (2003)
2. P. Lipari, Int. Conf. NuInt01, Nucl. Phys. Proc. Suppl. 112 (2002) 274
3. D. Geiregat et al., Nucl. Instr. and Meth., A 325 (1993) 92
4. NO_A Coll., Proposal P929 (2004)
5. MINOS Coll., Proposal P-875, Fermilab Report NuMI-L-63 (1995)
6. K. Nakamura, Int. Journal of Mod. Phys. A 18 (2003) 4053, and references therein
7. UNO Proto Coll., SBHEP-01-03 (2000) Waseda Univ. (1992)
8. ICARUS Coll., LNGS-P28/2001 and ICARUS-TM/2001-03, hep-ex/0103008
9. FLARE Collaboration, Letter of Intent hep-ex/0408121 (2004)
10. D. Cline et al., Nucl. Instr. and Meth. A503 (2003) 136
11. A. Rubbia, Proc. of the II Int. Workshop on Neutrinos in Venice, Venice (2003)
12. B. A. Dolgoshein et al., Fiz. El. Chast. Atom. Yad. 4 (1973) 167
13. D. Cline et al., Astroparticle Physics 12 (2000) 373
14. A. Cervera et al., Nucl. Instr. and Meth. A451 (2000) 123
15. J. Nelson, proceedings of Nufact 05, LNF (Italy), 21-26 June 2005.
16. N.K. Mondal, proceedings of Nufact 05, LNF (Italy), 21-26 June 2005.
17. P. Antonioli, for the MONOLITH Coll., Nucl. Phys. (Proc. Suppl.) 100 (2001) 142
18. A. Rubbia, proceedings of Nufact 05, LNF (Italy), 21-26 June 2005.
19. A. Donini, D. Meloni and P. Migliozzi, hep-ph/0206034
20. OPERA Coll., CERN SPSC-2000-028, LNGS P25/2000