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Performances of the CsI(Tl) detector element of the GLAST calorimeter

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Abstract

The CsI(Tl) Detector Element (CDE) is the unit of detection of the calorimeter subsystem of the Large Area Telescope (LAT) of the GLAST satellite (IEEE Nuclear Science Symposium, 2002), which will be launched in September 2006. The LAT consists of 4×4 identical towers, each one including a strip silicon tracker on the top and a CsI(Tl) calorimeter on the bottom. A segmented plastic scintillator anticoincidence detector surrounds the 16 towers. The calorimeter modules, 8 layers of 12 CDEs each, have a hodoscope's shape to allow the measurement of the gamma-ray direction.

The CDE main characteristics are its Light Yield, the energy resolution and the tapering, ratio of left/right signals used for the determination of the crossing position of the particle in the crystal.

The CDE have to operate in space at about -10°C . We will describe the CDE, the optimization of its main parameter, the L.Y., and finally its variation with temperature.

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1. Introduction

The GLAST Large Area Telescope (LAT) Calorimeter [1] is characterized by its huge energy range from 20 MeV up to 300 GeV/c, and also by its tracking capability, especially for high-energy gamma. To cope with these requirements, the calorimeter is segmented into 16 identical modules forming a hodoscope of 1536 identical CsI(Tl) Detector Elements (CDEs). The Scintillation light is detected at each end of the crystal by two PIN photodiodes with an active area ratio of about 6,

the smaller (PIN A) for the highest energy deposition and the larger (PIN B) for the lowest. To increase position detection capability along the crystal, the left/right signal difference (named: tapering) is improved by depolishing 2 of the 4 long sides of the crystal.

2. CDE description and main parameters

The CDE is the element of detection of the calorimeter and it can be independently fully characterized. It is composed of a CsI(Tl) crystal ($326 \times 19.9 \times 26.7 \text{ mm}^3$, from Amcrys) with two custom dual PIN photodiodes (DPD, S8576 from

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1 Hamamatsu) bonded at each ends (0.7 mm of
 3 wrapped in a very reflective material: VM2000,
 5 multilayer of Mylar of 2 different indices, total
 7 thickness 65 μm from 3M Company, with a end
 9 capsule (Delrin endcap) at each ends.

11 The main parameter of the CDE is the Light
 13 Yield of each PIN diodes, the number of collected
 15 electrons by deposited MeV in the crystal, the
 17 integration constant being 3 μs . Two other para-
 19 meters also characterize the CDE: the tapering and
 the energy resolution. The tapering is the L.Y. of
 the PIN diode when particles cross the crystal at
 2 cm from its opposite end over the L.Y. when they
 cross at 2 cm from its side. The energy resolution is
 calculated with particles crossing the crystal at its
 center, it is deduced from the normalized difference
 between the left and the right CDE PIN
 diodes.

21 The cosmic test bench uses cosmic muons
 23 depositing about 10.2 MeV in the 2 cm of CsI(Tl).
 25 Two plastic scintillator hodoscopes perpendicular
 27 to the CDE, up and down, give the muon crossing
 29 position. The coincidence between at least one
 31 plastic of the bottom and the top hodoscopes gives
 the trigger to the data acquisition system (NIM,
 CAMAC with Labview). The PIN diode signal is
 first pre-amplified (eV 5093, 3.6 mV/fC) then
 shaped (3 μs) and amplified and finally converted
 by a peak sensing ADC. All events are registered
 and analyzed with a C++ program base upon
 ROOT program (<http://root.cern.ch/>); Figs. 1–3

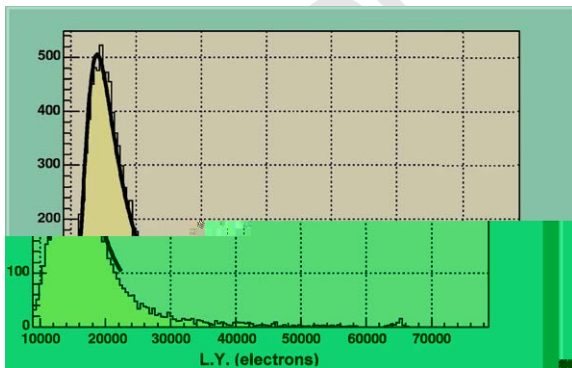


Fig. 1. Typical L.Y. spectrum obtained with cosmic muon crossing the CDE, with Landau fit.

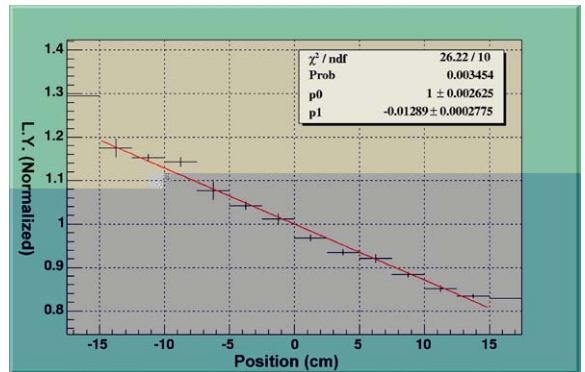


Fig. 2. Typical Tapering curve: normalized L.Y. vs muons crossing position along the CDE, with a linear fit.

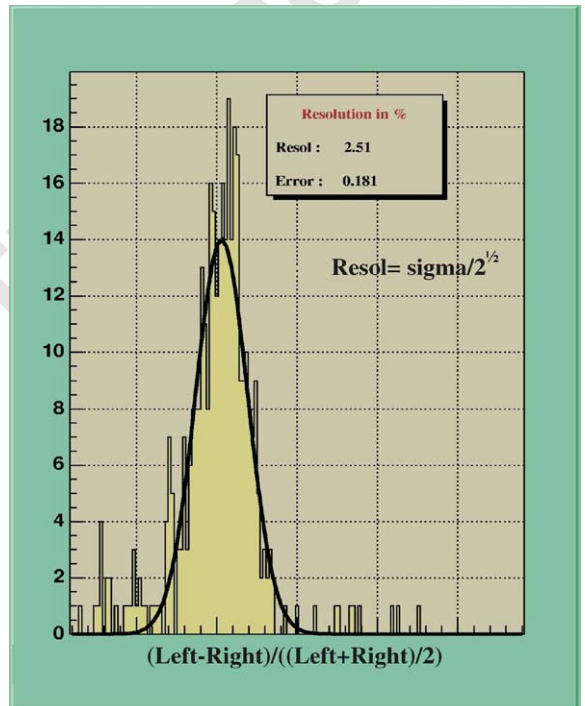


Fig. 3. Typical energy resolution spectrum, with Gaussian fit.

give typical curves or histogram from which we
 extract the main parameters.

We have assembled and tested 14 CDEs for their
 insertion in the “engineer module” at the Naval
 Research Laboratory (NRL, USA); both for
 integration tests with the front-end electronic and

1 Table 1 49
CDE L.Y., energy resolution and tapering requirements and measurements of the first 14 elements delivered to NRL

3 Parameter	Unit	PIN A (25 mm ²)		PIN B (152 mm ²)	
		Requier.	Meas.	Requier.	Meas.
5 Light Yield	(e/MeV)	> 1100	1170–1400	> 7000	7700–8600
7 Energy Resol.	(%)	No	3.6–5.3	< 8	1.5–2.3
Tapering	(%)	60–80	69–82	60–80	65–80

9 57

11 environmental tests. Table 1 gives their required
and measured performances.

15 3. Light yield optimization

17 The L.Y. is very dependent upon wrapping
19 material, and the tapering is anti-correlated to it.
First the base line material for the wrapping was
21 Tyvek plus Aluminum, which was used for the
crystal characterizations and tests.

23 To increase the L.Y. different reflective materi-
als have been tested and compared to a novel one:
25 a multilayer of Mylar of different indices named
VM2000 (Visual Mirror 2000 from 3M). Table 2
27 summarizes the comparison of VM2000 with:
Tyvek plus Aluminum, Millipore plus Aluminum,
29 Aluminum and finally Tedlar, which is a black
material giving the limit value when all the light
31 coming out of the crystal is lost. It shows the very
good performances of the VM2000 up to 40%
33 better than Tyvek plus Aluminum.

The VM2000 wrapping material covers only the
35 four long sides of the crystal and the PIN
photodiodes cover 335 mm² of the 472 mm² of
37 the crystal end. An endcap cover about 68 mm² of
the 137 mm² remaining area. We have tested four
39 different configurations for the crystal end: with
no endcap (137 mm² of bare crystal), with the
41 Delrin endcap, with VM2000 on the endcap and
finally with VM2000 covering the whole bare
43 crystal area. Table 3 summarizes the results and
shows an increase of the L.Y. of 15% with the
45 endcap and up to 8% more when adding VM2000.
Nevertheless, as our L.Y. CDE performance is
47 well over the requirement, for practical reason, we
decided to keep the Delrin endcap as our baseline.

Table 2

Wrapping material comparison performed with the largest PIN
photodiode, normalized to VM2000

Material	Normalized L.Y.
VM2000	1.00
Tyvek + Alu.	0.71
Millipore + Alu.	0.62
Aluminum	0.33
Tedlar (black)	0.18

61 63 65 67 69

Table 3

End crystal wrapping study: comparison between different end
crystal configurations of the L.Y. performance normalized to
VM2000 and remaining bare crystal area

Material	Bare Area (mm ²)	Normalized L.Y.
No Endcap	137	0.87
Endcap	69	1.00
Endcap with VM2000	69	1.04
VM2000	0	1.08

71 73 75 77 79 81

4. CDE operation from -20°C to 30°C

83 All tests on the CDE are performed at room
85 temperature (about 20°C) although the space
operating one will be around -10°C . As the
87 CsI(Tl) L.Y. decreases with the temperature, we
89 have to measure this dependency.

We have used a climate test cabinet where we
91 have put a CDE with its laboratory front-end
93 electronic (eV 5093 preamplifier) and a reference
PIN photodiode.

The L.Y. is mainly the convolution of: the
95 crystal scintillation and transmission, the PIN

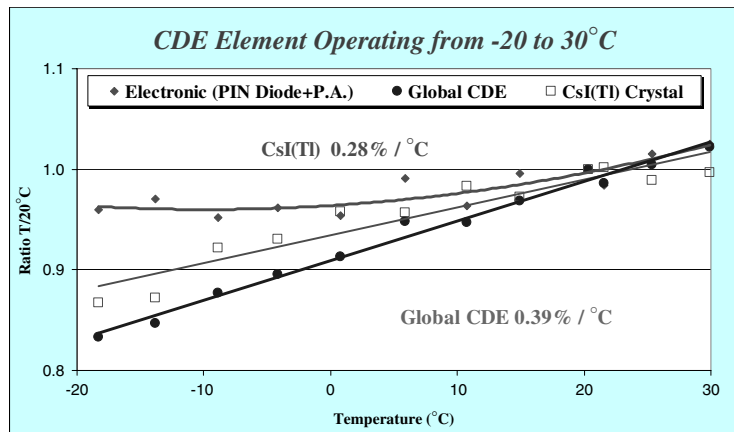


Fig. 4. Variation with temperature of the electronic, the global CDE and the CsI(Tl) crystal (de-convolution of the two formers).

diode photosensitivity and electron–hole collection, and the preamplifier gain.

To de-convolute these different contributions, we have used: a gamma-ray (AmBe, 4.43 MeV) source on the CDE, an optical fiber linked to a pulsed green LED and a X-ray source (Co57, 122 KeV) on the reference PIN photodiode, and finally a pulse generator linked to the dedicated input of the eV 5093 pre-amplifier.

We have first decreased the temperature from 30°C to –20°C by step of 10°C, and then increased it from –15°C up to 25°C.

The normalized values with the pulse generator, the green LED and the X-ray source on the reference PIN diode show the same variation. So we could conclude that the dominant contribution is the pre-amplifier gain. On the graph Fig. 4, we have put the mean value of this curve (labeled “electronic”). The normalized value with the gamma-ray source on the CDE (labeled “global CDE”) shows linear dependence with temperature with a slope of 0.39% per degree Celsius. If we de-convolute the electronic contribution we get the CsI(Tl) dependence with temperature (labeled “CsI(Tl) crystal”). The linear fit gives a slope of 0.28% per degree Celsius.

5. Conclusion

For the GLAST LAT calorimeter, we have chosen to use this very good reflective material,

VM2000 and a white Delrin endcap closing each crystal ends. In this condition we get a L.Y. well over the requirements, about 8200 electron by MeV with the larger PIN diode covering 32% of the crystal end. We have confirmed the temperature dependency of the Light Yield. The L.Y. lost when the calorimeter will operate in space at about –10°C, will be of the order of 8%.

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