# Extraction of the Compton Form Factor $\mathcal{H}$ from Deeply Virtual Compton Scattering measurements at Jefferson Lab 

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#### Abstract

In the framework of Generalised Parton Distributions, we study the helicity-dependent and independent cross sections measured in Hall A and the beam spin asymmetries measured in Hall B at Jefferson Laboratory. We perform a global fit of these data and fits on each kinematic bin. We extract the real and imaginary parts of the Compton Form Factor $\mathcal{H}$ under the main hypothesis of dominance of the Generalised Parton Distribution $H$ and twist 2 accuracy. We discuss our results and compare to previous extractions as well as to the VGG model. We pay extra attention to the estimation of errors on the extraction of $\mathcal{H}$.


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## Introduction

Since it was realised that Generalised Parton Distributions (GPD) were reachable through an harmonic analysis of the Deeply Virtual Compton Scattering (DVCS) process [1, 2], the study of the connection between GPDs and DVCS has been a very active field of research, concerning theoretical developments (see the reviews [3, 4, [5, 6]) as well as experimental ones [7, 8, 9, 10, 11, 12, 13, 14, 15]. After the first dedicated experiments and at the beginning of the experimental GPD program, it is already worth trying to extract GPDs from measurements.

The present work addresses this question, and illustrates it with recent JLab data, namely beam spin asymmetries (BSA) [14] and helicity-dependent and independent cross sections [12]. These data offer a large kinematic coverage and a fine kinematic binning, which are interesting features for our purpose. However, the methods we use and the conclusions we come to are presumably not restricted to JLab kinematics, and may be of interest for other experimental set-up. Furthermore, this is one of the first global fits of these measurements, and such fits are necessary to the completion of the experimental GPD program.

The harmonic analysis of $e p \rightarrow e p \gamma$ cross-sections has so far relied on the 2002 work of A.V. Belitsky, D. Müller and A. Kirchner [16]. In this formalism, the interference between the Bethe-Heitler (BH) and DVCS processes was treated with a leading order approximation of the BH part. This assumption was removed by A.V. Belitsky and D. Müller in [17] in the case of a spinless target, and by P.A.M. Guichon and M. Vanderhaegen in the case of a proton target 18. In all the following, we will use the expressions from the latter.

The first section of this paper describes JLab data, the difference of the evaluation of the $e p \rightarrow e p \gamma$ cross sections by [16] and [18], and outlines our hypothesis. In

[^0]the second part we explain our fitting strategy, which is twofold : we perform a global fit of the selected JLab data and fits on each kinematic bin (which we will be refering to as local fits). In the third section we systematically compare the results of global and local fits, and keep the output of the global fit as our best solution. We then discuss the extracted values of $\mathcal{H}$ and compare them to previous studies.

## I. PRELIMINARY ANALYSIS

In this study, $x_{B}$ is the standard Bjorken variable, $Q^{2}$ the virtuality of the initial photon and $t$ the square momentum transfer between initial and final protons.

## A. Description of the selected JLab data

C. Muñoz-Camacho et al. 12] published helicitydependent and helicity-independent cross sections for $x_{B}$ $=0.36$ and $t$ between -0.33 and $-0.17 \mathrm{GeV}^{2}$. Helicityindependent cross sections are given at $Q^{2}=2.3 \mathrm{GeV}^{2}$ only while helicity-dependent cross sections are measured for $Q^{2}$ between 1.5 and $2.3 \mathrm{GeV}^{2}$. Data come in 12 bins ( 3 values of $Q^{2}$ and 4 values of $t$ ). Helicity-dependent cross sections are given with relatively large uncertainties (from $20 \%$ to $100 \%$, depending on the bins and on the value of the angle between leptonic and hadronic planes) while helicity-independent cross sections are usually accurate at the $5 \%$ level.
F.-X. Girod et al. [14] released BSAs over a wide kinematic range $\left(x_{B}\right.$ from 0.11 to $0.58, Q^{2}$ from 1 to 4.8 $\mathrm{GeV}^{2}$ and $t$ from -1.8 to $-0.09 \mathrm{GeV}^{2}$, described by 62 multi-dimensional bins ( 5 values of $t, 5$ values of $x_{B}$ and 4 values of $Q^{2}$ ). Their accuracy is $25 \%$ on average, ranging from 5 to $100 \%$ depending on the bins and on the value of the angle between leptonic and hadronic planes.

These two sets of measurements have kinematic configurations in common, allowing us to perform crosschecks by evaluating a BSA through the ratio of helicitydependent to helicity-independent cross sections. This
simple exercise shows no unexpected discrepancy between Halls A and B measurements. This implies that both sets of data are consistent and can be used in a global fit.

## B. DVCS at leading twist

Four GPDs $H, E, \tilde{H}$ and $\tilde{E}$ appear at twist 2, but the cross sections depend on the Compton Form Factors (CFF). The convention of [16] is used to define the CFFs :

$$
\begin{gather*}
\mathcal{F}=\int_{-1}^{+1} d x F(x, \xi, t)\left(\frac{1}{\xi-x-i \epsilon}-\frac{1}{\xi+x-i \epsilon}\right)(1  \tag{1}\\
(F=H \text { or } E) \\
\tilde{\mathcal{F}}=\int_{-1}^{+1} d x \tilde{F}(x, \xi, t)\left(\frac{1}{\xi-x-i \epsilon}+\frac{1}{\xi+x-i \epsilon}\right)(2  \tag{2}\\
(\tilde{F}=\tilde{H} \text { or } \tilde{E})
\end{gather*}
$$

and $\xi=x_{B} \frac{1+\frac{t}{2 Q^{2}}}{2-x_{B}+\frac{x_{B} t}{Q^{2}}}$ is the generalised Bjorken variable [16, 18. The complex integration kernel yields a real and an imaginary part to the CFFs :

$$
\begin{align*}
& R e \mathcal{F}= \mathcal{P} \int_{-1}^{+1} d x F(x, \xi, t)\left(\frac{1}{\xi-x}-\frac{1}{\xi+x}\right)  \tag{3}\\
&(F=H \text { or } E) \\
& \operatorname{Im\mathcal {F}}=\pi(F(\xi, \xi, t)-F(-\xi, \xi, t))  \tag{4}\\
&(F=H \text { or } E)
\end{align*}
$$

and

$$
\begin{gather*}
\operatorname{Re} \tilde{\mathcal{F}}=\mathcal{P} \int_{-1}^{+1} d x F(x, \xi, t)\left(\frac{1}{\xi-x}+\frac{1}{\xi+x}\right)  \tag{5}\\
\quad(\tilde{F}=\tilde{H} \text { or } \tilde{E}) \\
\operatorname{Im} \tilde{\mathcal{F}}=\pi(F(\xi, \xi, t)+F(-\xi, \xi, t))  \tag{6}\\
\quad(\tilde{F}=\tilde{H} \text { or } \tilde{E})
\end{gather*}
$$

where the symbol $\mathcal{P}$ denotes the principal value of the integral.

The real and imaginary parts of a CFF are related by dispersion relations due to analyticity properties (see [19, 20, 21, 22]). However the unknown subtraction (the D-term [23]) and the limited kinematic range of our data make this constraint rather weak. In this work we consider the real and imaginary parts as independent (local fits) except in the case where we use an explicit parametrisation of the GPD $H$ to compute the CFF (global fit).

## C. The formalism of P.A.M. Guichon and M. Vanderhaeghen and its consequences

The recent work of P.A.M. Guichon and M. Vanderhaeghen (GV) gives analytic expression for the $e p \rightarrow$
$e p \gamma$ cross section ; these formulae are embodied in a Mathematica package [18, which we used to build a C++/ROOT library.

The $e p \rightarrow e p \gamma$ cross section is classically divided into three parts, namely BH, VCS and interference between the BH and VCS cross sections. The contribution of the BH amplitude to the interference is treated exactly. This introduces some differences with respect to the wellknown BMK expressions and two new important qualitative features to the discussion of the measurements under scrutiny.

Concerning Hall A measurements, the BMK formalism restricted to twist 2 asserts that helicity-dependent cross sections write :

$$
\begin{array}{r}
\frac{1}{2}\left[\frac{d^{4} \sigma^{+}}{d^{4} \Phi}-\frac{d^{4} \sigma^{-}}{d^{4} \Phi}\right]=C_{1} \sin \phi \operatorname{Im}\left(\mathcal{H}+\frac{x_{B}}{2-x_{B}}(1\right. \\
\left.\left.+\frac{F_{2}}{F_{1}}\right) \tilde{\mathcal{H}}-\frac{t}{4 M^{2}} \frac{F_{2}}{F_{1}} \mathcal{E}\right) \tag{7}
\end{array}
$$

where $d^{4} \Phi=d Q^{2} d x_{B} d t d \phi, F_{1}$ and $F_{2}$ are the Dirac and Pauli form factor, $M$ the proton mass and $C_{1}$ is a constant irrelevant for our purpose. Since the $Q^{2}$ dependence of this cross section is factorised, this expression allows a study of scaling without having to disentangle the different CFFs as in 12 .
With the exact equations of [18], an helicity-dependent cross section does not have such a simple form. It writes :

$$
\begin{gather*}
\frac{1}{2}\left[\frac{d^{4} \sigma^{+}}{d^{4} \Phi}-\frac{d^{4} \sigma^{-}}{d^{4} \Phi}\right]=C_{2} \\
\sin \phi \operatorname{Im}\left(\mathcal{H}+c_{\mathcal{E}} \mathcal{E}+c_{\tilde{\mathcal{H}}} \tilde{\mathcal{H}}\right.  \tag{8}\\
\left.+c_{\tilde{\mathcal{E}}} \tilde{\mathcal{E}}\right)+\ldots
\end{gather*}
$$

where $C_{2}$ is a constant irrelevant for our purpose. The dots stand for power-suppressed contributions. The test of scaling is more involved now since the coefficients $c_{\mathcal{E}}$, $c_{\tilde{\mathcal{H}}}$ and $c_{\tilde{\mathcal{E}}}$ do depend on $Q^{2}$. At given $x_{B}$ and $t$, the coefficients of Eq. (8) $c_{\mathcal{E}}, c_{\tilde{\mathcal{H}}}$ and $c_{\tilde{\mathcal{E}}}$ vary respectively at most by $20 \%, 20 \%$ and $30 \%$. If we only fit the combination of CFFs appearing in $\operatorname{Im}(\ldots)$ in Eq. (8), the kinematic $Q^{2}$-dependence of $c_{\mathcal{E}}, c_{\tilde{\mathcal{H}}}$ and $c_{\tilde{\mathcal{E}}}$ may appear as a scaling deviation of the same magnitude.
Differences also arise in the expression of a BSA. Its dependence on the angle $\phi$ between the leptonic and hadronic planes takes the following form :

$$
\begin{equation*}
\mathrm{BSA}=\frac{a \sin \phi+b \sin 2 \phi}{1+c \cos \phi+d \cos 2 \phi+e \cos 3 \phi} \tag{9}
\end{equation*}
$$

where $a=\mathcal{O}\left(Q^{-1}\right), b=\mathcal{O}\left(Q^{-4}\right), c=\mathcal{O}\left(Q^{-1}\right) d=$ $\mathcal{O}\left(Q^{-2}\right)$ and $e=\mathcal{O}\left(Q^{-5}\right)$ are real numbers. In the BMK picture, the coefficients $b, d$ and $e$ are higher-twist contributions. As a straightforward consequence, we see that the $90^{\circ}$ asymmetry is no longer proportional to the imaginary part of a linear combination of CFFs.

Moreover the coefficient $c$ of Eq. (9) now depends on the imaginary part of CFFs, and not only on the real
part as in [16]. This precludes a clean separation of the real and imaginary parts of

$$
F_{1} \mathcal{H}+\frac{x_{B}}{2-x_{B}}\left(F_{1}+F_{2}\right) \tilde{\mathcal{H}}-\frac{t}{4 M^{2}} F_{2} \mathcal{E}
$$

through the distinct measurements of the helicitydependent and helicity-independent cross sections as was done previously [12].

## D. Main assumptions

P.A.M. Guichon and M. Vanderhaeghen provide exact analytical expressions of the $e p \rightarrow e p \gamma$ observables as functions of the DVCS amplitudes. The latter can be written in terms of GPDs through the usual twist expansion [1. In this work we restrict ourselves to the twist 2 approximation. This is a reasonable assumption, since C. Muñoz-Camacho et al. [12] claimed the observation of early $Q^{2}$-scaling.

As we are considering experiments on a proton target, we neglect $E, \tilde{H}$ and $\tilde{E}$ ( $H$-dominance) for the following reasons. Firstly, this is supported by kinematics : in Eq. (8), the coefficient $c_{\mathcal{E}}$ varies between 0.05 and 0.28 , $c_{\tilde{\mathcal{H}}}$ between 0.3 and 1., and $c_{\tilde{\mathcal{E}}}$ between $2.10^{-5}$ and 0.015 for the 52 kinematic configurations of Hall B data satisfying $\frac{|t|}{Q^{2}}<\frac{1}{2}$. Secondly, for small $t$ and $\xi$, we expect $\frac{\tilde{H}}{H}$ to be close to $\frac{\Delta q}{q}$ i.e. $\frac{1}{4}$. Thirdly, we can check, for instance thanks to the VGG model [24, 26, 27, 28], that the relative sizes of $\operatorname{Im\mathcal {E}}$ and $\operatorname{Im} \tilde{\mathcal{H}}$ to $\operatorname{Im\mathcal {H}}$ are similar : $\frac{I m \mathcal{E}}{I m \mathcal{H}}$ varies between 0.21 and 0.92 , and $\frac{I m \mathcal{H}}{I m \mathcal{H}}$ between 0.13 and 0.91 for the same set of 52 Hall B kinematic configurations. This indicates that the hierarchy between the kinematic coefficients reflects the hierarchy of contributions to the interference.

Thus assuming $H$-dominance, we may hope to extract information on $H$ from BSAs or helicity-dependent cross sections with a systematic error of $20 \%$ to $50 \%$ [44, this approximation being better at small $t$. The advantage of this approach is the dramatic decrease of the number of degrees of freedom involved in fits. M. Guidal indeed showed in a recent work [29] in the same kinematic region that it is not possible to extract sensible information about the real and imaginary parts of $\mathcal{H}, \mathcal{E}, \tilde{\mathcal{H}}$ and $\tilde{\mathcal{E}}$ by direct fits of helicity-dependent and independent cross sections. More specifically, keeping only the dominant coefficients in Eq. (9) gives the minimal functional form :

$$
\begin{equation*}
\mathrm{BSA} \simeq \frac{a \sin \phi}{1+c \cos \phi} \tag{10}
\end{equation*}
$$

A direct fit of BSAs to this reduced expression on each $\left(x_{B}, Q^{2}, t\right)$-bin, along the lines of [14], shows that the coefficient $c$ is compatible with 0 (while with a marked trend to negative values) within error bars for 25 bins over the 52 bins for which $\frac{|t|}{Q^{2}}$ is less than $\frac{1}{2}$. Extracting the real part of the CFFs contained in the coefficient $c$ of Eq. 10) is thus a demanding task.

## II. FITTING STRATEGIES

The possibility to study GPDs in DVCS rests on factorisation theorems [30, which require a small value of $\frac{|t|}{Q^{2}}$. In the following, we restrict ourselves to kinematic configurations for which $\frac{|t|}{Q^{2}}<\frac{1}{2}$.

## A. Local fits

There are at most $12 \phi$-bins in each Hall B $\left(x_{B}, Q^{2}, t\right)$ bin, and $24 \phi$-bins in each Hall A $\left(x_{B}, Q^{2}, t\right)$-bin. $R e \mathcal{H}$ and $\operatorname{ImH}$ are the free parameters of the fits.

We estimate the systematic errors associated to our $H$-dominance hypothesis by first fitting data setting the subdominant GPDs to 0 , then fitting the same data setting the subdominant GPDs to their VGG value, and computing the difference.

## B. Global fits

Turning to global fits will help to decrease the statistical uncertainties on the fitted parameters. Moreover, we will benefit from the wide kinematic coverage of Hall B data, and of the accuracy of Hall A measurements in the same fit.

Since we are interested only in extracting values of CFFs, we will not try to extrapolate outside the kinematic region of the measurements we consider. This allows us to use a polynomial parametrisation to perform the fits. The forthcoming difficulty will be the evaluation of the systematic uncertainty related to that phenomenological choice.

## 1. A parametrisation of $H$ from the dual model

The singlet combination $H_{+}$is :

$$
\begin{equation*}
H_{+}\left(x, \xi, t, Q^{2}\right)=H\left(x, \xi, t, Q^{2}\right)-H\left(-x, \xi, t, Q^{2}\right) \tag{11}
\end{equation*}
$$

This is the quantity which is accessible through DVCS.
In the framework of the dual model for a spin $\frac{1}{2}$ target, and assuming $H$-dominance, $H_{+}$can be formally expanded according to [31]:

$$
\begin{align*}
H_{+}\left(x, \xi, t, Q^{2}\right)= & 2 \sum_{n=0}^{\infty} \sum_{l=0}^{n+1} B_{n l}\left(t, Q^{2}\right) \theta\left(1-\frac{x^{2}}{\xi^{2}}\right) \\
& \times\left(1-\frac{x^{2}}{\xi^{2}}\right) C_{2 n+1}^{\frac{3}{2}}\left(\frac{x}{\xi}\right) P_{2 l}\left(\frac{1}{\xi}\right) \tag{12}
\end{align*}
$$

This formal expansion can be resummed as a Gegenbauer
polynomial expansion [32, 33, 34, 35] :

$$
\begin{equation*}
H_{+}\left(x, \xi, t, Q^{2}\right)=2\left(1-x^{2}\right) \sum_{n=0}^{\infty} A_{n}\left(\xi, t, Q^{2}\right) C_{2 n+1}^{3 / 2}(x) \tag{13}
\end{equation*}
$$

The coefficients $A_{n}$ are defined by :

$$
\begin{align*}
A_{n}\left(\xi, t, Q^{2}\right)= & -\frac{4 n+5}{(n+1)(2 n+3)} \sum_{p=0}^{n} \xi R_{n p}(\xi) \\
& \times \frac{(p+1)(2 p+3)}{4 p+5} \sum_{l=0}^{p+1} B_{p l}\left(t, Q^{2}\right) P_{2 l}\left(\frac{1}{\xi}\right) \tag{14}
\end{align*}
$$

where $P_{2 l}$ is a Legendre polynomial and $R_{n p}(\xi)$ is a polynomial the degree of which is $2 n+1$ :

$$
\begin{align*}
R_{n p}(\xi)= & (-1)^{(n+p+1)} \frac{\Gamma\left(\frac{5}{2}+n+p\right)}{\Gamma(n-p+1) \Gamma\left(\frac{5}{2}+2 p\right)} \xi^{(2 p+1)} \\
& \times_{2} F_{1}\left(p-n, \frac{5}{2}+n+p, \frac{7}{2}+2 p, \xi^{2}\right) \tag{15}
\end{align*}
$$

with ${ }_{2} F_{1}$ the Gauss hypergeometric function.
The $Q^{2}$-evolution of $B_{p l}\left(t, Q^{2}\right)$ is given at leading order in 33:

$$
\begin{equation*}
B_{p l}\left(t, Q^{2}\right)=B_{p l}\left(t, Q_{0}^{2}\right)\left(\frac{\ln \frac{Q_{0}^{2}}{\Lambda^{2}}}{\ln \frac{Q^{2}}{\Lambda^{2}}}\right)^{\frac{\gamma_{p}}{\beta_{0}}} \tag{16}
\end{equation*}
$$

where $\beta_{0}=11-\frac{2}{3} n_{f}$ and, for $0 \leq p \leq N_{\max }-1$ [35] :

$$
\begin{equation*}
\gamma_{p}=\frac{4}{3}\left(3+\frac{1}{(p+1)(2 p+1)}-4\left(\Psi(2 p+2)+\gamma_{E}\right)\right) \tag{17}
\end{equation*}
$$

where we note $\Psi$ the Digamma function and $\gamma_{E}$ the Euler-Mascheroni constant. For the sake of clarity, note that the coefficients $B_{n l}, A_{n}, R_{n p}$ and $\gamma_{n}$ are redefined here. In fact they should have been labelled $B_{2 n+1} 2 l$, $A_{2 n+1}, R_{2 n+1}{ }_{2 p+1}$ and $\gamma_{2 n+1}$ to ensure a complete coherence with the traditional notations of the dual model. However since Eq. (12), Eq. (13), Eq. (14), Eq. (15), Eq. (16) and Eq. (17) are just intermediate steps in our extraction of the CFF $\mathcal{H}$ we will not try to compare the values of $B_{n l}, A_{n}, R_{n p}$ and $\gamma_{n}$ to those obtained in dual model studies.

We use $\Lambda=373 \mathrm{MeV}$ in the $\overline{\mathrm{MS}}$ scheme with 3 flavours of quarks. We obtained this value after a running of the strong coupling constant computed at four loops [36], starting from the 2008 world-averaged value of $\alpha_{S}\left(M_{Z}\right)$ and crossing each quark threshold at (twice) its 2008 averaged mass 37. This evaluation is in good agreement with a recent textbook one [38]. The reference scale $Q_{0}$ has been set to $Q_{0}^{2}=3 \mathrm{GeV}^{2}$.

## 2. Iterative fitting procedure

In practice we truncate Eq. at some maximum value $N_{\max }$ of $n$ and we assume the following form for
the coefficients $B_{p l}$ :

$$
\begin{equation*}
B_{p l}\left(t, Q_{0}^{2}\right)=\frac{a_{p l}}{1+b_{p l}\left(t-t_{0}\right)^{2}} \tag{18}
\end{equation*}
$$

with $t_{0}$ a constant and $a_{p l}$ and $b_{p l}$ the free parameters. Their number is $N_{\max } *\left(N_{\max }+3\right)$. Due to the truncation at $n=N_{\max }$ the representation of the GPD that we get from a fit of $\left(a_{p l}, b_{p l}\right)$ can hardly be trusted outside the domain of the fit. We take it as a smooth parametrisation of the data.

The selected JLab data consist in 1001 measurements with $\frac{|t|}{Q^{2}}<\frac{1}{2}$. We fitted them with $N_{\max }=2,3,4$ which corresponds to 10,18 and 28 -parameter fits (performed with Minuit 39]). To constrain the polynomial oscillations in $\xi$, we adopt an iterative fitting procedure. We simplify the problem by first working with bins for which we can neglect the $t$-dependence. Indeed, $40 \%$ of the data with $\frac{|t|}{Q^{2}}<\frac{1}{2}$ satisfy $0.2 \mathrm{GeV}^{2} \leq-t \leq 0.4 \mathrm{GeV}^{2}$. We choose $t_{0}=-0.28 \mathrm{GeV}^{2}$ in Eq. (18) and extract a first value of the parameters $a_{p l}$. We then add bins with $-t$ between 0.09 and $0.2 \mathrm{GeV}^{2}$, and between 0.6 and $1.0 \mathrm{GeV}^{2}$, fitting both $a_{p l}$ and $b_{p l}$, initialising the fit at the values of $a_{p l}$ fitted at the previous iteration, and $b_{p l}$ at 0 . We then add the two last $t$-bins to the fit, using at each step the previous extractions of $a_{p l}$ and $b_{p l}$.

## 3. Systematic uncertainties

The fits with $N_{\max }=2,3,4$ are qualitatively similar. Their $\chi^{2} /$ d.o.f. are respectively $1.73,1.61$ and 1.78 . The comparison of the values of the CFFs derived from these fits gives an estimate of the systematic error on $\mathcal{H}$ induced by the truncation. This truncation error is precisely defined at the beginning of the next section when discussing the results of these fits.
Since we assume $H$-dominance, we must take into account the systematic error linked to the neglect of subdominant GPDs. We proceed as in the case of local fits : we fit the data with the subdominant CFFs set to 0 or to their VGG value, and take the difference as an estimate of the systematic uncertainty.

## III. RESULTS

## A. Extraction of $\operatorname{ImH}$ and $\operatorname{ReH}$

The Fig. 1 and Fig. 22 display the effect of the truncation of the series in Eq. (13). When $\xi$ is small, $N_{\max }=2$ is not enough to describe the BSAs. On the contrary, when $\xi$ is large, we cannot control the fit with $N_{\max }=4$. This comes from the fact that statistical errors on BSAs are getting larger when $\xi$ grows. The fit with $N_{\max }=3$ is always good, and close to the local fits, which are optimal by construction. We also see that the VGG model overestimates the data, which is a known feature [12. It


FIG. 1: The $x_{B}$-range is divided into 5 bins. The BSAs are displayed at the lowest $Q^{2}$ and for the value of $t$ used at the first iteration of the fit : $x_{B}=0.13$ and $Q^{2}=1.17 \mathrm{GeV}^{2}$ (a), $x_{B}=0.18$ and $Q^{2}=1.37 \mathrm{GeV}^{2}$ (b), $x_{B}=0.25$ and $Q^{2}=1.69 \mathrm{GeV}^{2}$ (c), $x_{B}=0.34$ and $Q^{2}=1.99 \mathrm{GeV}^{2}$ (d) and $x_{B}=0.44$ and $Q^{2}=2.70 \mathrm{GeV}^{2}(\mathrm{e})$. In all cases $t=$ $-0.30 \mathrm{GeV}^{2}$. The results of the global fits with $N_{\max }=2$ (red, dotted), $N_{\text {max }}=3$ (green, dot-dashed), $N_{\max }=4$ (blue, dot-dot-dashed), the result of the local fits (magenta, plain) and the prediction of the VGG code (cyan, dashed) are plot for $\phi$ between 0 and 360 degrees and values of BSA between -0.35 and +0.35 .
presumably stems from an overestimation of the imaginary parts of CFFs by VGG. At last, global and local fits to helicity-dependent and helicity-independent cross sections are all good, and almost indistinguishable.

We thus choose the fit with $N_{\text {max }}=3$ as our nominal solution. The systematic uncertainty on $\operatorname{ImH}$ and $R e \mathcal{H}$ linked to the truncation of Eq. 133 is estimated as the maximum of the (absolute values of the) difference between the results for the nominal solution and for the 2 other fits.

The Fig. 3. Fig. 4. Fig. 5 and Fig. 6 display our results for $\operatorname{ImH}$ and $R e \mathcal{H}$ respectively. Both local fits and global


FIG. 2: Helicity-dependent (up) and helicity-independent (down) cross sections at $x_{B}=0.36, Q^{2}=2.3 \mathrm{GeV}^{2}$ and $t=-0.17 \mathrm{GeV}^{2}(\mathrm{a}), t=-0.23 \mathrm{GeV}^{2}(\mathrm{~b}), t=-0.28 \mathrm{GeV}^{2}$ (c) and $t=-0.33 \mathrm{GeV}^{2}(\mathrm{~d})$. The results of the global fits with $N_{\max }=2$ (red, dotted), $N_{\max }=3$ (green, dot-dashed), $N_{\max }=4$ (blue, dot-dot-dashed), the result of the local fits (magenta, plain) and the prediction of the VGG code (cyan, dashed) are plot for $\phi$ between 0 and 360 degrees. Helicity-dependent cross sections range between -0.025 and $+0.025 \mathrm{nb} / \mathrm{GeV}^{4}$, and helicity-independent cross sections between 0.015 and $0.105 \mathrm{nb} / \mathrm{GeV}^{4}$.
fit give results with comparable accuracy for $\operatorname{Im\mathcal {H}}$, but as expected the results of the global fits are smoother. This is especially true concerning $R e \mathcal{H}$ : in this case the local fits suffer from large fluctuations of $R e \mathcal{H}$ with values which fall outside the plot range. However, we could not reliably extract values of the CFFs for the larger values of $\xi$ with the global fit. This is reminiscent of the difficulty in controlling the oscillating behaviour of the polynomial expansion displayed in Fig. 1.

The results for local and global fits are almost always compatible, which is a strong consistency check. Both rely on the assumptions of twist 2 accuracy and of H -dominance. On one hand, local fits suffer from numerical fluctuations (the 2-parameter local fits are not constrained enough on some bins) but are almost model-independent. On the other hand, global fits are smoother, but suffer from oscillations. That both methods give the same results indicates that fluctuations and oscillations are reasonably controlled in the bins for which results are displayed. Since, in both cases, the total error bars have the same size, we conclude that our estimation of systematic uncertainties due to the truncation of the series Eq. 13) is realistic.

All fits keep data satisfying $\frac{|t|}{Q^{2}}<\frac{1}{2}$. For local fits, changing the maximal value of $\frac{|t|}{Q^{2}}$ amounts to dropping points. For global fits, the whole results may be changed, but the good agreement between the results of both types of fits, and the slow $Q^{2}$-evolution of the extracted CFFs, indicate that this restricted kinematic region is suitable for an analysis in the GPD framework.


FIG. 3: $Q^{2}$-behaviour ( $1<Q^{2}<4 \mathrm{GeV}^{2}$ ) of the extracted values of $\operatorname{Im} \mathcal{H}$ of local fits (left) and global fit (right) on Hall B kinematics : $0.09<-t<0.2 \mathrm{GeV}^{2}$ (a), $0.2<-t<0.4 \mathrm{GeV}^{2}$ (b), $0.4<-t<0.6 \mathrm{GeV}^{2}$ (c), and $0.6<-t<1 . \mathrm{GeV}^{2}$ (d). The error bars include both statistics and systematics. $\operatorname{ImH}$ ranges between 0 and 10. The black full circles correspond to $x_{B}=0.125$, red squares to $x_{B}=0.175$, green up triangles to $x_{B}=0.250$, blue down triangles to $x_{B}=0.360$ and magenta open circles to $x_{B}=0.491$.

Nevertheless, we observe a sizeable scaling deviation on $\operatorname{Im} \mathcal{H}$ extracted from Hall A data for $t=-0.17 \mathrm{GeV}^{2}$. The choice of the expression of $\xi$ is also related to the issue of scaling. Changing the exact expression of $\xi$ to its asymptotic form $\frac{x_{B}}{2-x_{B}}$ induces differences on the extraction of CFFs, the amplitudes of which depend on $x_{B}$ and $t$. Deviations are noticeable but results with both expressions of $\xi$ are compatible within error bars. This indicates an effect of higher-order power corrections. In view of the whole set of data, the conclusion of early $Q^{2}$-scaling 12 presumably still holds but a higher-twist study is needed to make it final.

Tab. $\mathbb{\square}$ and Tab. II summarize our results. Our error bars are dominated by systematic effects. Typically we obtain a relative accuracy of 20 to $50 \%$ on $\operatorname{Im\mathcal {H}}$, which is


FIG. 4: $Q^{2}$-behaviour ( $1<Q^{2}<4 \mathrm{GeV}^{2}$ ) of the extracted values of $R e \mathcal{H}$ of local fits (left) and global fit (right) on Hall B kinematics : $0.09<-t<0.2 \mathrm{GeV}^{2}$ (a), $0.2<-t<0.4 \mathrm{GeV}^{2}$ (b), $0.4<-t<0.6 \mathrm{GeV}^{2}$ (c), and $0.6<-t<1 . \mathrm{GeV}^{2}$ (d). The error bars include both statistics and systematics. $\operatorname{Im} \mathcal{H}$ ranges between -7.5 and +7.5 . The black full circles correspond to $x_{B}=0.125$, red squares to $x_{B}=0.175$, green up triangles to $x_{B}=0.250$, blue down triangles to $x_{B}=0.360$ and magenta open circles to $x_{B}=0.491$.
quite satisfactory under the assumption of $H$-dominance and given the statistical accuracy of JLab data. On the contrary, $R e \mathcal{H}$ is still largely undetermined, and is never extracted with a precision better than $50 \%$. But the imaginary parts of CFF might be the quantities of prime importance as stressed in 40 : it may be possible to evaluate $\operatorname{Re} \mathcal{H}$ (for instance) from the knowledge of $\operatorname{ImH}$ on a wide kinematic range through the use of dispersion relations.

## B. Discussion

The Fig. 7 compares our results to a twist 2 modelindependent extraction [29] and an extraction with the

| Hall B kinematics |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $x_{B}$ | $t\left(\mathrm{GeV}^{2}\right)$ | $Q^{2}\left(\mathrm{GeV}^{2}\right)$ | $\operatorname{Im} \mathcal{H}$ | ReH |
| 0.1342 | -0.1337 | 1.1661 | 5.90 (0.01) (1.57) (1.01) | -3.16 (0.04) (1.70) (4.11) |
| 0.1763 | -0.1346 | 1.3651 | 6.91 (0.02) (0.67) (0.72) | -3.54 (0.03) (3.60) (2.66) |
| 0.1767 | -0.1376 | 1.5557 | 7.69 (0.02) (0.72) (0.72) | -4.92 (0.03) (3.80) (2.47) |
| 0.2350 | -0.1465 | 1.6453 | 5.43 (0.03) (0.34) (0.42) | -2.75 (0.04) (5.01) (1.41) |
| 0.2377 | -0.1448 | 1.8895 | 5.59 (0.03) (0.35) (0.46) | -3.17 (0.04) (4.33) (1.28) |
| 0.2460 | -0.1442 | 2.1641 | 5.28 (0.03) (1.08) (0.57) | 8) $(1.16)$ |
| 0.3205 | -0.1705 | 1.9424 | 3.33 (0.05) (1.14) (1.64) | -2.87 (0.04) (7.62) (0.94) |
| 0.3215 | -0.1719 | 2.2170 | 3.46 (0.05) (0.48) (1.74) | -2.74 (0.04) (6.66) (0.87) |
| 0.3213 | -0.1743 | 2.5078 | 3.55 (0.05) (0.12) (1.81) | -2.66 (0.04) (5.85) (0.82) |
| 0.3211 | -0.1753 | 2.7865 | 3.64 (0.05) (0.59) (1.89) | -2.58 (0.04) (5.21) (0.79) |
| 0.1341 | -0.2840 | 1.1678 | 5.37 (0.01) (0.60) (1.82) | -4.50 (0.03) (4.77) (5.48) |
| 0.1764 | -0.2798 | 1.3680 | 6.34 (0.01) (1.52) (1.54) | -4.21 (0.03) (3.09) (3.39) |
| 0.1772 | -0.2819 | 1.5653 | 6.99 (0.01) (1.97) (1.55) | -5.37 (0.03) (3.56) (3.26) |
| 0.2466 | -0.2842 | 1.6881 | 4.29 (0.02) (0.26) (0.46) | -2.41 (0.03) (1.68) (0.72) |
| 0.2 | -0.2809 | 1.9490 | 4.28 (0.02) (0.20) (0.43) | -2.70 (0.03) (1.45) (0.62) |
| 0.2525 | -0.2814 | 2.2131 | 4.06 (0.02) (0.65) (0.41) | -2.82 (0.03) (1.32) (0.50) |
| 0.3 | -0.3062 | 1.9930 | 3.10 (0.04) (0.66) (0.77) | -3.48 (0.03) (3.43) (0.06) |
| 0.3431 | -0.3012 | 2.3060 | 3.52 (0.04) (0.36) (0.82) | -3.31 (0.03) (2.56) (0.02) |
| 0.3 | -0.2966 | 2.6372 | 3.92 (0.05) (0.49) (0.84) | -3.10 (0.03) (1.84) (0.09) |
| 0.3480 | -0.2942 | 2.9706 | 4.62 (0.05) (0.36) (0.78) | -2.82 (0.03) (0.87) (0.15) |
| 0.1331 | -0.4929 | 1.1611 | 4.03 (0.01) (0.64) (2.01) | -4.27 (0.02) (4.81) (6.03) |
|  | -0.4910 | 1.3580 | 4.96 (0.01) (1.85) (2.03) | -3.98 (0.02) (3.24) (4.14) |
| 0.1765 | -0.4909 | 1.5611 | 5.47 (0.01) (2.37) (2.04) | -4.86 (0.02) (3.19) (3.99) |
| 0. | -0.4 | 1.7039 | 3.13 (0.02) (0.57) (0.84) | -1.91 (0.02) (0.47) (0.77) |
| 0.2 | -0.4 | 1.9485 | 3.26 (0.02) (0.54) (0.87) | -2.21 (0.02) (0.35) (0.83) |
| 0.2504 | -0.4883 | 2.2028 | 3.06 (0.02) (0.36) (0.81) | -2.30 (0.02) (0.62) (0.73) |
| 0.3443 | -0 | 2.0062 | 1.63 (0.03) (0.34) (0.26) | -3.81 (0.02) (2.02) (0.41) |
| 0.3501 | -0.4938 | 2.3282 | 1.95 (0.04) (0.37) (0.19) | -4.00 (0.02) (1.31) (0.28) |
| 0.3555 | -0.4889 | 2.6851 | 2.53 (0.04) (0.48) (0.02) | -4.06 (0.02) (0.43) (0.13) |
| 0.3600 | -0.4854 | 3.0455 | 3.25 (0.04) (1.06) (0.25) | -3.99 (0.02) (1.98) (0.03) |
| 0.1753 | -0.7741 | 1.5516 | 2.51 (0.01) (0.68) (0.63) | -1.60 (0.02) (4.81) (1.75) |
| 0.2493 | $-0.7731$ | 1.6847 | 1.31 (0.01) (0.46) (0.05) | -0.52 (0.02) (1.83) (0.25) |
| 0.2476 | -0.7694 | 1.9394 | 1.24 (0.02) (0.54) (0.03) | -0.73 (0.02) (1.62) (0.21) |
| 0.2494 | -0.7689 | 2.1990 | 0.99 (0.02) (0.71) (0.14) | -0.82 (0.02) (1.59) (0.13) |
| 0.3516 | -0.7752 | 2.0231 | 1.39 (0.03) (1.04) (0.68) | -4.28 (0.02) (1.45) (0.26) |
| 0.3597 | -0.7684 | 2.3596 | 2.09 (0.03) (0.63) (0.71) | -4.98 (0.02) (0.97) (0.49) |
| 0.3607 | -0.7623 | 2.7054 | 2.20 (0.03) (0.54) (0.56) | -5.18 (0.02) (1.60) (0.66) |
| 0.3582 | -0.7573 | 3.0357 | 1.83 (0.03) (0.04) (0.39) | -5.15 (0.02) (2.04) (0.76) |

TABLE I: Global fit extraction of $\operatorname{Im\mathcal {H}}$ and $\operatorname{Re\mathcal {H}}$ on Hall B kinematics. The horizontal lines gather bins with common $\left(x_{B}, t\right)$ and the horizontal double lines bins with common $t$. The errors are in parenthesis, the first one being statistical, the second and third systematic (respectively truncation and subdominant CFFs).

BMK formalism [12. Firstly, the use of the GV expressions creates important deviations to the latter extraction. Since the extracted combinations of GPDs are not the same, we will not make the argument more quantitative. Secondly, we obtained results in very good agree-
ment with 29, but with errors considerably smaller.
For a given $t$, and at $Q^{2}=3 \mathrm{GeV}^{2}$, the Fig. 8 displays the $x_{B}$-dependence of the CFFs we extracted, and compares them to the predictions of the VGG model. To draw the authorised region for $\operatorname{Im\mathcal {H}}$ and $\operatorname{Re} \mathcal{H}$ we had to estimate the systematic uncertainty induced by the ne-

| Hall A kinematics |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $x_{B}$ | $t\left(\mathrm{GeV}^{2}\right)$ | $Q^{2}\left(\mathrm{GeV}^{2}\right)$ | ImH | ReH |  |
| 0.3600 | -0.1700 | 1.5000 | $4.73(0.07)(0.53)(0.72)$ | $-1.35(0.03)(3.62)(0.47)$ |  |
| 0.3600 | -0.1700 | 1.9000 | $5.87(0.07)(0.12)(0.57)$ | $-0.39(0.03)(1.33)(0.38)$ |  |
| 0.3600 | -0.1700 | 2.3000 | $7.05(0.07)(0.31)(0.45)$ | $0.36(0.03)(0.51)(0.31)$ |  |
| 0.3600 | -0.2300 | 1.5000 | $5.03(0.06)(0.94)(0.82)$ | $-2.77(0.03)(2.92)(0.12)$ |  |
| 0.3600 | -0.2300 | 1.9000 | $5.83(0.06)(0.36)(0.74)$ | $-2.04(0.03)(1.08)(0.15)$ |  |
| 0.3600 | -0.2300 | 2.3000 | $6.67(0.06)(0.19)(0.64)$ | $-1.46(0.03)(0.32)(0.19)$ |  |
| 0.3600 | -0.2800 | 1.5000 | $4.75(0.05)(0.99)(0.69)$ | $-3.45(0.03)(2.51)(0.16)$ |  |
| 0.3600 | -0.2800 | 1.9000 | $5.32(0.05)(0.43)(0.64)$ | $-2.90(0.03)(0.97)(0.05)$ |  |
| 0.3600 | -0.2800 | 2.3000 | $5.94(0.05)(0.20)(0.55)$ | $-2.45(0.03)(0.20)(0.04)$ |  |
| 0.3600 | -0.3300 | 1.5000 | $4.30(0.04)(0.95)(0.51)$ | $-3.86(0.03)(2.20)(0.40)$ |  |
| 0.3600 | -0.3300 | 1.9000 | $4.68(0.05)(0.52)(0.48)$ | $-3.47(0.03)(0.91)(0.24)$ |  |
| 0.3600 | -0.3300 | 2.3000 | $5.11(0.05)(0.15)(0.38)$ | $-3.12(0.03)(0.15)(0.10)$ |  |

TABLE II: Global fit extraction of $\operatorname{Im} \mathcal{H}$ and $R e \mathcal{H}$ on Hall A kinematics. The horizontal lines gather bins with common $\left(x_{B}, t\right)$ and the horizontal double lines bins with common $t$. The errors are in parenthesis, the first one being statistical, the second and third systematic (respectively truncation and subdominant CFFs).
glect of subdominant CFFs for kinematic configurations for which we have no measurements. According to Tab. II and Tab. II this error weakly depends on $Q^{2}$ for fixed $t$ and $x_{B}$. We averaged it over each $\left(x_{B}, t\right)$-bin, and used it to draw the colour band of Fig. 8. Doing so, we enhanced the oscillating behaviour of the contours of the plot by adding discontinuities. Since the errors are mostly due to systematic effects, the points in the coloured domain have the same probability.

This being stated, we notice that the extracted values of $\operatorname{ImH} \mathcal{H}$ have an $x_{B}$-dependence similar to that predicted by the VGG model. We again observe that VGG tends to overestimate $\operatorname{ImH}$. The extracted CFFs are often but not always compatible with VGG. The errors on the extracted $\operatorname{Re} \mathcal{H}$ and $\operatorname{ImH}$ in the hatched zone may be underestimated. We already noticed that polynomial oscillations are more severe at large $x_{B}$, where the data are less accurate, as shown in Fig. 1 and Fig. 2. However, from Fig. 3 and Fig. 4 we know that the extraction is independent of the type (local or global) of the fit for $0.3 \leq x_{B} \leq 0.4$. Thus these data should not be rejected but it is necessary to stay cautious when interpreting them.

## Conclusions

Working at leading twist, and assuming $H$-dominance, we extracted $\operatorname{Im} \mathcal{H}$ and $\operatorname{Re} \mathcal{H}$ with two different methods. The local fits do not benefit from the wide kinematic coverage of Hall B data, and are bound to produce rather large statistical uncertainties when fitting two independent parameters from an asymmetry measured at (at most) twelve $\phi$-bins. On the opposite, the global fit dramatically decreases the systematic uncertainties and
uses the kinematic coverage to disentangle the contributions of the different CFFs at the expense of a fitting Ansatz. The good agreement between the results of both extractions is a strong consistency check. It demonstrates that the extraction of $\operatorname{ImH}$ and $\operatorname{Re\mathcal {H}}$ and the estimation of the systematics related to this fitting procedure produce relatively model-independent results, at least for $x_{B}$ not too large. This estimation of the systematics will be refined but these first results are already encouraging. One immediate advantage is the smoothness of the results. This comparison and the weak $Q^{2}$-dependency of the results also validates a posteriori the restriction to kinematic configurations with $\frac{-t}{Q^{2}}<\frac{1}{2}$.
It is one of the first global fits of measurements from different experiments. In spite of the difficulty of controlling polynomial oscillations, the (fitted) behaviour of $\operatorname{Im} \mathcal{H}$ and $\operatorname{Re} \mathcal{H}$ versus $x_{B}$ is similar to the VGG model prediction. The $t$-dependence of our global fit is also in fair agreement with that obtained in the first GPD analysis of Hall A data in the BMK framework. As expected, we find that the VGG model tends to overestimate the physical value of $\operatorname{ImH}$.

Our results are dominated by systematic uncertainties. Their origin is twofold. Firstly, we assumed $H$ dominance and neglected the contribution of $E, \tilde{E}$ and $\tilde{H}$. Secondly, the Gegenbauer and Legendre expansions induce oscillations in the partial sums of the series. We may hope to reduce these systematic uncertainties in the near future using additional BSA measurements 41] (unpolarised proton target) and 42] (longitudinally polarised proton target) which will put stronger constraints on the global fits. The extension of our procedure to smaller values of $x_{B}$ is in progress.

However, our total errors are already of reasonable size, since they are comparable to, or smaller than those com-


FIG. 5: $Q^{2}$-behaviour ( $1<Q^{2}<3 \mathrm{GeV}^{2}$ ) of the extracted values of $\operatorname{Im} \mathcal{H}$ of local fits (left) and global fit (right) on Hall A kinematics : $t=-0.17 \mathrm{GeV}^{2}$ (a), $t=-0.23 \mathrm{GeV}^{2}$ (b), $t=-0.28 \mathrm{GeV}^{2}$ (c) and $t=-0.33 \mathrm{GeV}^{2}$ (d). In all cases $x_{B}$ $=0.36$. The error bars include both statistical and systematic uncertainties. $\operatorname{ImH}$ ranges between 2.5 and 7.5 .
ing from previous extractions. We typically obtain a 20 to $50 \%$ accuracy on $\operatorname{ImH}$, which is already good regarding our hypothesis of $H$ dominance.

At last, C. Muñoz-Camacho et al. 12 performed the extraction of a combination of GPDs with the BMK formulae. They concluded that Hall A data indicates twist 2 dominance of DVCS through early $Q^{2}$-scaling. Using the new GV formalism, we tend to come to the same conclusion in this study, but we plan to work out a refined analysis at twist 3 in the Wandura-Wilczek approximation to make sure we distinguish between power law (higher twist) and logarithmic (evolution) behaviours.

Note added During the writing of this paper, K. Kumericki and D. Müller released a detailed modeldependent fit of Hermes and JLab DVCS measurements 43]. We will compare their results to ours (both present results and the on-going extension to smaller $x_{B}$ ) in a


FIG. 6: $t$-behaviour ( $0.1<-t<0.4 \mathrm{GeV}^{2}$ ) of the extracted values of $\operatorname{ImH}(u p)$ and $R e \mathcal{H}$ (down) of local fits (left) and global fit (right) on Hall A kinematics $\left(x_{B}=0.36\right.$ and $Q^{2}$ $\left.=2.3 \mathrm{GeV}^{2}\right)$. The error bars include both statistical and systematic uncertainties. $\operatorname{ImH}$ ranges between 4 and 8, and Re $\mathcal{H}$ between -4 and 1 .


FIG. 7: $\operatorname{Im} \mathcal{H}$ (up) and $\operatorname{Re} \mathcal{H}$ (down) vs $t(0.1<-t<$ $\left.0.4 \mathrm{GeV}^{2}\right)$ on Hall A kinematics $\left(x_{B}=0.36\right.$ and $Q^{2}=$ $2.3 \mathrm{GeV}^{2}$ ). $\quad \operatorname{Im} \mathcal{H}$ ranges between 3 and 8 , and $R e \mathcal{H}$ between -5 and 1. We compare our results (left column) to those of M. Guidal 29] (middle column) and C. MuñozCamacho et al. [12] (right column). In the latter column $\mathcal{H}+\frac{x_{B}}{2-x_{B}}\left(1+\frac{F_{2}}{F_{1}} \tilde{\mathcal{H}}\right)-\frac{t}{4 M^{2}} \frac{F_{2}}{F_{1}} \mathcal{E}$ is plot, and not $\mathcal{H}$. The error bars include both statistical and systematic uncertainties.


FIG. 8: Comparisons of the extracted $\operatorname{Im\mathcal {H}}$ (up) and $\operatorname{Re} \mathcal{H}$ (down) and the predictions of the VGG model (black circles) for $0.1<x_{B}<0.4$ and $0.09<-t<0.2 \mathrm{GeV}^{2}$ (a), $0.2<$ $-t<0.4 \mathrm{GeV}^{2}(\mathrm{~b}), 0.4<-t<0.6 \mathrm{GeV}^{2}(\mathrm{c})$, and $0.6<$ $-t<1 . \mathrm{GeV}^{2}(\mathrm{~d})$. $\operatorname{Im} \mathcal{H}$ ranges between 0 and 15 , and $\operatorname{Re} \mathcal{H}$ between -10 and +7.5 . The error bars include both statistical and systematics contributions. Since our errors are dominated by systematics, we can only say that the true value of $\operatorname{Im} \mathcal{H}$ and $\operatorname{Re} \mathcal{H}$ lie in the coloured bands. Moreover, we estimated the systematic errors due to the neglect of subdominant GPDs by averaging over the points where measurements are made, hence enhancing the oscillating behaviour of the fitting curve (see text for more explanations).
future study.

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## APPENDIX: GV EXPRESSIONS AT TWIST 2 ASSUMING $H$-DOMINANCE

We give the GV expressions used in the present work. Hence we restrict ourselves to the case of an unpolarised target and take into account only the GPD $H$. For the sake of consistency and for future references, we adopt the notations of 18.

## 1. Kinematics for $e p \rightarrow e p \gamma$

We note $k$ and $p$ the 4 -momentum of the initial lepton and proton, and $k^{\prime}, p^{\prime}$ and $q^{\prime}$ the 4 -momentum of the final lepton, proton and photon respectively. In the center of mass frame the coordinate system is chosen such that [24]:

$$
\begin{align*}
p= & \left(\sqrt{M^{2}+\vec{q}^{2}}, 0,0,-|\vec{q}|\right)  \tag{A.1}\\
q= & \left(q^{0}, 0,0,|\vec{q}|\right)  \tag{A.2}\\
k= & \left(\sqrt{m^{2}+\vec{k}^{2}},|\vec{k}| \sin \alpha \cos \phi\right. \\
& |\vec{k}| \sin \alpha \sin \phi,|\vec{k}| \cos \alpha)  \tag{A.3}\\
q^{\prime}= & \left(q^{\prime 0}, q^{\prime 0} \sin \theta, 0, q^{\prime 0} \cos \theta\right)  \tag{A.4}\\
k^{\prime}= & \left(\sqrt{m^{2}+\overrightarrow{k^{\prime}} 2},\left|\overrightarrow{k^{\prime}}\right| \sin \alpha^{\prime} \cos \phi\right. \\
& \left.\left|\overrightarrow{k^{\prime}}\right| \sin \alpha^{\prime} \sin \phi,\left|\overrightarrow{k^{\prime}}\right| \cos \alpha^{\prime}\right) \tag{A.5}
\end{align*}
$$

with $0 \leq \alpha, \alpha^{\prime}, \theta \leq \pi$ and $0 \leq \phi \leq 2 \pi . M$ (resp. $m$ ) is the hadron (resp. lepton) mass, $q_{\perp}^{\prime}=q^{\prime 0} \sin \theta$ and $s=(p+q)^{2}$. The parameter $\omega$ defined by :

$$
\begin{align*}
\cosh \omega & =\frac{-Q^{2}+4 E M x_{B}}{\sqrt{Q^{4}+4 M^{2} Q^{2} x_{B}^{2}}}  \tag{A.6}\\
\sinh \omega & =\sqrt{-1+\frac{\left(-Q^{2}+4 E M x_{B}\right)^{2}}{Q^{4}+4 M^{2} Q^{2} x_{B}^{2}}} \tag{A.7}
\end{align*}
$$

contains all the dependence on the beam energy $E$. We note $h_{e}$ the helicity of the incoming lepton.

The BMK and GV formalisms thus use different definitions for the angle $\phi$ between the hadronic and leptonic planes, and these definitions do not follow the Trento convention [25] chosen to describe the measurements : $\phi_{\text {Trento }}=\pi-\phi_{\text {BMK }}$ and $\phi_{\text {Trento }}=2 \pi-\phi$.

## 2. Bethe-Heitler process

The proton form factors are estimated according to the following expressions :

$$
\begin{align*}
G_{E} & =\frac{1}{\left(1-\frac{t}{M_{V}^{2}}\right)^{2}} \text { with } M_{V}^{2}=0.71 \mathrm{GeV}^{2}  \tag{A.8}\\
& =F_{1}+\frac{t}{4 M^{2}} F_{2}  \tag{A.9}\\
G_{M} & =\mu_{p} G_{E} \text { with } \mu_{p}=2.79285  \tag{A.10}\\
& =F_{1}+F_{2} \tag{A.11}
\end{align*}
$$

The amplitude $\mathcal{M}_{\mathrm{BH}}$ of the BH process writes:

$$
\begin{equation*}
\mathcal{M}_{\mathrm{BH}}=-\frac{\sigma_{\mathrm{BH}}[0,1]+\cosh 2 \omega \sigma_{\mathrm{BH}}[0,2]+\cos \phi \sinh 2 \omega \sigma_{\mathrm{BH}}[0,3]-\cos 2 \phi(\cosh 2 \omega-1) \sigma_{\mathrm{BH}}[0,4]}{4 t^{2}\left(\frac{1}{4}\left(Q^{2}+t\right)^{2}-\left(\frac{Q^{2}\left(Q^{2}+t\left(2 x_{B}-1\right) \cosh \omega\right.}{2 \sqrt{Q^{4}+4 M^{2} x_{B}^{2} Q^{2}}}-Q q_{\perp}^{\prime} \cos \phi \sinh \omega\right)^{2}\right)} \tag{A.12}
\end{equation*}
$$

where

$$
\begin{align*}
\sigma_{\mathrm{BH}}[0,1]= & \left.\frac{1}{( } M^{4}+2\left(Q^{2}-s\right) M^{2}+\left(Q^{2}+s\right)^{2}\right)\left(4 M^{2}-t\right)
\end{align*} \times\left(4 G_{M}^{2} t\left(6 t M^{8}-2\left(3 Q^{4}-11 t Q^{2}+6 t(2 s+t)\right) M^{6}\right)\right.
$$

$$
\begin{align*}
\sigma_{\mathrm{BH}}[0,2]= & \frac{1}{\left(M^{4}+2\left(Q^{2}-s\right) M^{2}+\left(Q^{2}+s\right)^{2}\right)\left(4 M^{2}-t\right)} \times\left(4 G _ { M } ^ { 2 } t \left(2 t M^{8}-2\left(Q^{4}-9 t Q^{2}+2 t(2 s+t)\right) M^{6}\right.\right. \\
& +\left(8 Q^{6}+(4 s+19 t) Q^{4}-2 t(15 s+7 t) Q^{2}+t\left(12 s^{2}+10 t s+t^{2}\right)\right) M^{4}-2\left(Q^{8}+(2 s+t) Q^{6}\right. \\
& \left.+\left(s^{2}-7 t s+5 t^{2}\right) Q^{4}-t\left(3 s^{2}+10 t s+t^{2}\right) Q^{2}+s t(2 s+t)^{2}\right) M^{2}+t\left(\left(Q^{4}+2 s Q^{2}+2 s^{2}\right)\right. \\
& \left.\left.\left(Q^{2}+s\right)^{2}+\left(Q^{4}-4 s Q^{2}+s^{2}\right) t^{2}+2 s\left(-2 Q^{4}-s Q^{2}+s^{2}\right) t\right)\right)-32 G_{E}^{2} M^{2}\left(t M^{8}\right. \\
& +\left(Q^{4}+t Q^{2}-4 s t\right) M^{6}+\left(-4 Q^{6}+(5 t-2 s) Q^{4}+(s-t) t Q^{2}+s t(6 s+t)\right) M^{4}+\left(\left(Q^{2}+s\right)^{2} Q^{4}\right. \\
& \left.-2\left(Q^{4}+s Q^{2}+s^{2}\right) t^{2}+\left(Q^{2}-s\right)\left(Q^{2}+s\right)\left(5 Q^{2}+4 s\right) t\right) M^{2} \\
& \left.\left.+\left(Q^{2}+s\right)^{2} t\left((s-t) Q^{2}+s(s+t)\right)\right)\right) \tag{A.14}
\end{align*}
$$

$\sigma_{\mathrm{BH}}[0,3]=\frac{1}{\sqrt{M^{4}+2\left(Q^{2}-s\right) M^{2}+\left(Q^{2}+s\right)^{2}}\left(4 M^{2}-t\right)} \times\left(16 G_{M}^{2} Q q_{\perp}^{\prime} t\left(\left(3 t-2 Q^{2}\right) M^{4}+\left(2 Q^{4}+(2 s+5 t) Q^{2}\right.\right.\right.$

$$
\left.-t(4 s+t)) M^{2}+t\left((s-t) Q^{2}+s(s+t)\right)\right)-64 G_{E}^{2} M^{2} Q q_{\perp}^{\prime}\left(M^{2}-Q^{2}-s\right)
$$

$$
\begin{equation*}
\left.\times\left(2 M^{2} Q^{2}-\left(M^{2}+Q^{2}+s\right) t\right)\right) \tag{A.15}
\end{equation*}
$$

$$
\begin{align*}
\sigma_{\mathrm{BH}}[0,4]= & \frac{1}{\left(M^{4}+2\left(Q^{2}-s\right) M^{2}+\left(Q^{2}+s\right)^{2}\right)\left(4 M^{2}-t\right)}\left(-64 G_{E}^{2} Q^{2}\left(t M^{4}+\left(Q^{4}+t Q^{2}-2 s t\right) M^{2}\right.\right. \\
& \left.\left.+s t\left(Q^{2}+s+t\right)\right) M^{4}-8 G_{M}^{2} Q^{2}\left(2 M^{2}-t\right) t\left(t M^{4}+\left(Q^{4}+t Q^{2}-2 s t\right) M^{2}+s t\left(Q^{2}+s+t\right)\right)\right) \tag{A.16}
\end{align*}
$$

## 3. Virtual Compton Scattering process

The amplitude $\mathcal{M}_{\mathrm{VCS}}$ of the VCS process writes :

$$
\begin{equation*}
\mathcal{M}_{\mathrm{VCS}}=\frac{\sigma_{\mathrm{VCS}}[0,1]+\cosh 2 \omega \sigma_{\mathrm{VCS}}[0,2]}{2 Q^{2}} \tag{A.17}
\end{equation*}
$$

where

$$
\begin{align*}
\sigma_{\mathrm{VCS}}[0,1] & =\frac{1}{4}\left(\frac{48\left(x_{B}-1\right)^{2} I m \mathcal{H}^{2}}{\left(1-x_{B}\right)\left(x_{B}-2\right)^{2}}+\frac{48\left(x_{B}-1\right)^{2} R e \mathcal{H}^{2}}{\left(1-x_{B}\right)\left(x_{B}-2\right)^{2}}\right)  \tag{A.18}\\
\sigma_{\mathrm{VCS}}[0,2] & =\frac{1}{4}\left(\frac{16\left(x_{B}-1\right)^{2} I m \mathcal{H}^{2}}{\left(1-x_{B}\right)\left(x_{B}-2\right)^{2}}+\frac{16\left(x_{B}-1\right)^{2} R e \mathcal{H}^{2}}{\left(1-x_{B}\right)\left(x_{B}-2\right)^{2}}\right) \tag{A.19}
\end{align*}
$$

## 4. Interference

The amplitude $\mathcal{M}_{\mathrm{I}}$ of the interference between the BH and VCS processes writes :

$$
\begin{align*}
\mathcal{M}_{\mathrm{I}}= & \frac{1}{Q^{2} t\left(\frac{1}{4}\left(Q^{2}+t\right)^{2}-\left(\frac{Q^{2}\left(Q^{2}+t\left(2 x_{B}-1\right)\right) \cosh \omega}{2 \sqrt{Q^{4}+4 M^{2} x_{B}^{2} Q^{2}}}-Q q_{\perp}^{\prime} \cos \phi \sinh \omega\right)^{2}\right)} \times\left(\cosh \omega \sigma_{\mathrm{I}}[0,1]+\cosh 3 \omega \sigma_{\mathrm{I}}[0,2]\right. \\
& \left.\left.+\cos \phi \sinh \omega \sigma_{\mathrm{I}}[0,3]+\sinh 3 \omega \sigma_{\mathrm{I}}[0,4]\right)+\cos 2 \phi \cosh \omega-\cosh 3 \omega\right) \sigma_{\mathrm{I}}[0,5] \\
& \left.+\frac{1}{3} \cos 3 \phi(3 \sinh \omega-\sinh 3 \omega) \sigma_{\mathrm{I}}[0,6]+h_{e} \sin \phi \sinh 2 \omega \sigma_{\mathrm{I}}[0,7]-h_{e}(\cosh 2 \omega-1) \sin 2 \phi \sigma_{\mathrm{I}}[0,8]\right) \quad(\mathrm{A} . \tag{A.20}
\end{align*}
$$

where

$$
\begin{align*}
& \sigma_{\mathrm{I}}[0,1]=-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{00} U_{40}^{r} Q^{4}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{01} U_{46}^{r} Q^{4}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{A.21}\\
& \sigma_{\mathrm{I}}[0,2]=-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{00} U_{2}^{r} Q^{4}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{01} U_{10}^{r} Q^{4}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{A.22}\\
& \sigma_{\mathrm{I}}[0,3]=-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{01} U_{47}^{r} Q^{5}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{00} U_{41}^{r} Q^{3}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{А.23}\\
& \sigma_{\mathrm{I}}[0,4]=-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{01} U_{49}^{r} Q^{5}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{00} U_{43}^{r} Q^{3}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{A.24}\\
& \sigma_{\mathrm{I}}[0,5]=-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{00} U_{3}^{r} Q^{2}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}-\frac{8\left(x_{B}-1\right) R e \mathcal{H} J_{01} U_{11}^{r} Q^{2}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{A.25}\\
& \sigma_{\mathrm{I}}[0,6]=-\frac{8 Q\left(x_{B}-1\right) R e \mathcal{H} J_{00} U_{42}^{r}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}-\frac{8 Q\left(x_{B}-1\right) R e \mathcal{H} J_{01} U_{48}^{r}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{A.26}\\
& \sigma_{\mathrm{I}}[0,7]=\frac{8\left(x_{B}-1\right) \operatorname{Im\mathcal {H}J_{01}U_{51}^{r}Q^{5}}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}+\frac{8\left(x_{B}-1\right) \operatorname{Im} \mathcal{H} J_{00} U_{45}^{r} Q^{3}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}  \tag{A.27}\\
& \sigma_{\mathrm{I}}[0,8]=\frac{8\left(x_{B}-1\right) \operatorname{Im} \mathcal{H} J_{00} U_{44}^{r} Q^{2}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)}+\frac{8\left(x_{B}-1\right) \operatorname{Im\mathcal {H}J_{01}U_{50}^{r}Q^{2}}}{\sqrt{1-x_{B}}\left(x_{B}-2\right)} \tag{A.28}
\end{align*}
$$

The helicity currents $J$ are :

$$
\begin{align*}
J_{00}= & -\frac{\mathrm{F} 2\left(2 M^{2}+Q^{2}+2 s\right) \sqrt{-s t} \cos \left(\frac{\theta}{2}\right)\left(\sqrt{p^{0}-M} \sqrt{p^{\prime 0}-M}-\sqrt{M+p^{0}} \sqrt{M+p^{\prime 0}}\right)}{M \sqrt{s} \sqrt{Q^{4}-4 s t}} \\
& -\frac{8 G_{M} s \sqrt{-t}\left(\sqrt{M+p^{0}} \sqrt{p^{\prime 0}-M}+\sqrt{p^{0}-M} \sqrt{M+p^{\prime 0}}\right)\left(\cos \left(\frac{\theta}{2}\right)\left(q^{3}-q^{\prime 3}\right)-q^{\prime 1} \sin \left(\frac{\theta}{2}\right)\right)}{Q^{2} \sqrt{Q^{4}-4 s t}}  \tag{A.29}\\
J_{01}= & \frac{2 \sqrt{2} \mathrm{~F} 2 \sqrt{s} \cos \left(\frac{\theta}{2}\right)\left(\sqrt{p^{0}-M} \sqrt{p^{\prime 0}-M}-\sqrt{M+p^{0}} \sqrt{M+p^{\prime 0}}\right) q^{3} q^{\prime 1}}{M \sqrt{Q^{4}-4 s t}} \\
& +\frac{2 \sqrt{2} G_{M} \sqrt{s}\left(\sqrt{M+p^{0}} \sqrt{p^{\prime 0}-M}+\sqrt{p^{0}-M} \sqrt{M+p^{\prime 0}}\right)\left(q^{3}+q^{\prime 0}\right) \sin \left(\frac{\theta}{2}\right)}{\sqrt{Q^{4}-4 s t}} \tag{A.30}
\end{align*}
$$

and the kinematic coefficients $U^{r}$ are :

$$
\begin{align*}
& U_{2}^{r}=\frac{\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{\prime 0}\right) Q^{2}-4 s t\right)}{64 Q^{2} \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}}\left(\left(Q^{4}+\left(t+\left(q^{0}-q^{\prime 0}\right)\left(q^{0}-q^{3}+2 q^{\prime 0}\right)\right) Q^{2}\right.\right. \\
& \left.\left.-4 q_{\perp}^{\prime 2} q^{0} q^{3}+t\left(q^{0}-q^{3}\right)\left(q^{0}+q^{0}\right)\right)\right)  \tag{A.31}\\
& U_{3}^{r}=\frac{\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{\prime 0}\right) Q^{2}-4 s t\right)}{64 \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}}\left(\left(Q^{4}+\left(t+\left(q^{0}-q^{\prime 0}\right)\left(q^{0}+q^{3}+2 q^{\prime 0}\right)\right) Q^{2}\right.\right. \\
& \left.\left.+4 q_{\perp}^{\prime 2} q^{0} q^{3}+t\left(q^{0}+q^{3}\right)\left(q^{0}+q^{0}\right)\right)\right)  \tag{A.32}\\
& U_{10}^{r}=-\frac{q_{\perp}^{\prime} \sqrt{s}\left(\left(t-\left(q^{0}-q^{3}-2 q^{\prime 0}\right)\left(q^{0}-q^{\prime 0}\right)\right) Q^{2}+t q^{0}\left(q^{3}+q^{\prime 0}\right)\right)}{4 \sqrt{2} Q^{2} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{А.33}\\
& U_{11}^{r}=\frac{q_{\perp}^{\prime} \sqrt{s}\left(\left(t-\left(q^{0}+q^{3}-2 q^{00}\right)\left(q^{0}-q^{00}\right)\right) Q^{2}+t q^{0}\left(q^{00}-q^{3}\right)\right)}{4 \sqrt{2} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{A.34}\\
& U_{40}^{r}=-\frac{\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{\prime 0}\right) Q^{2}-4 s t\right)}{64 Q^{2} \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}} \\
& \times\left(Q^{4}+\left(16 q_{\perp}^{\prime 2}+t+\left(q^{0}-q^{\prime 0}\right)\left(q^{0}-q^{3}+2 q^{\prime 0}\right)\right) Q^{2}+4 q_{\perp}^{\prime 2} q^{0}\left(4 q^{0}+3 q^{3}\right)+t\left(q^{0}-q^{3}\right)\left(q^{0}+q^{\prime 0}\right)\right)  \tag{A.35}\\
& U_{41}^{r}=-\frac{q_{\perp}^{\prime}\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{\prime 0}\right) Q^{2}-4 s t\right)}{128 Q^{2} \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}} \\
& \times\left(Q^{4}+\left(12\left(q^{0}\right)^{2}+16 q^{3} q^{0}-22 q^{\prime 0} q^{0}-7 t-26 q^{3} q^{\prime 0}\right) Q^{2}+4 t q^{0}\left(q^{0}+4 q^{3}\right)\right)  \tag{A.36}\\
& U_{42}^{r}=-\frac{3 Q^{2} q_{\perp}^{\prime}\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{0}\right) Q^{2}-4 s t\right)\left(Q^{2}+t+2\left(q^{0}-q^{3}\right) q^{\prime 0}\right)}{128 \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{А.37}\\
& U_{43}^{r}=-\frac{q_{\perp}^{\prime}\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{0}\right) Q^{2}-4 s t\right)\left(5 Q^{4}-\left(3 t+2\left(-6\left(q^{0}\right)^{2}+7 q^{0} q^{0}+q^{3} q^{0}\right)\right) Q^{2}+4 t\left(q^{0}\right)^{2}\right)}{128 Q^{2} \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{A.38}\\
& U_{44}^{r}=-\frac{\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{0}\right) Q^{2}-4 s t\right)}{32 \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}} \\
& \times\left(Q^{4}+\left(2 q_{\perp}^{\prime 2}+t+\left(q^{0}-q^{\prime 0}\right)\left(q^{0}+q^{3}+2 q^{\prime 0}\right)\right) Q^{2}+\left(q^{0}+q^{3}\right)\left(2 q^{0} q_{\perp}^{2}+t\left(q^{0}+q^{0}\right)\right)\right)  \tag{А.39}\\
& U_{45}^{r}=\frac{q_{\perp}^{\prime}\left(2 Q^{4}+q^{0}\left(3 q^{0}+q^{3}-2 q^{\prime 0}\right) Q^{2}+t q^{0}\left(q^{0}+q^{3}\right)\right)\left(Q^{4}+2 \sqrt{s}\left(q^{0}-q^{\prime 0}\right) Q^{2}-4 s t\right)}{16 Q^{2} \sqrt{s} \sqrt{-t} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{A.40}\\
& U_{46}^{r}=-\frac{q_{\perp}^{\prime} \sqrt{s}\left(\left(7 t-3\left(q^{0}-q^{3}-2 q^{\prime 0}\right)\left(q^{0}-q^{\prime 0}\right)\right) Q^{2}+t\left(4\left(q^{0}\right)^{2}+3\left(q^{3}+q^{\prime 0}\right) q^{0}+4 q^{3} q^{\prime 0}\right)\right)}{4 \sqrt{2} Q^{2} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{A.41}\\
& U_{47}^{r}=\frac{\sqrt{s}}{16 \sqrt{2} Q^{4} \sqrt{Q^{4}-4 s t} q^{3}}\left(\left(3 Q^{6}+\left(12 q_{\perp}^{\prime 2}+3\left(q^{0}\right)^{2}-22\left(q^{\prime 0}\right)^{2}+3 t+7 q^{0} q^{3}+35 q^{0} q^{\prime 0}-3 q^{3} q^{\prime 0}\right) Q^{4}\right.\right. \\
& +\left(2\left(22\left(q^{0}\right)^{2}+\left(15 q^{3}+16 q^{00}\right) q^{0}-7 q^{3} q^{\prime 0}\right) q_{\perp}^{\prime 2}+8\left(4 q^{0}+q^{3}\right) q^{\prime 0}\left(\left(q^{0}\right)^{2}-\left(q^{\prime 0}\right)^{2}\right)\right. \\
& \left.\left.\left.+t\left(3\left(q^{0}\right)^{2}+7\left(q^{3}+5 q^{\prime 0}\right) q^{0}+q^{00}\left(11 q^{3}+16 q^{\prime 0}\right)\right)\right) Q^{2}+8 q^{0}\left(4 q^{0}+q^{3}\right)\left(q^{0}+q^{\prime 0}\right)\left(q^{0} q_{\perp}^{2}+t q^{\prime 0}\right)\right)\right)  \tag{A.42}\\
& U_{48}^{r}=-\frac{3 Q^{2} \sqrt{s}}{16 \sqrt{2} \sqrt{Q^{4}-4 s t} q^{3}}\left(\left(Q^{4}+\left(4 q_{\perp}^{\prime 2}+t+\left(q^{0}-q^{00}\right)\left(q^{0}+q^{3}+2 q^{\prime 0}\right)\right) Q^{2}\right.\right. \\
& \left.\left.+t\left(q^{0}+q^{3}\right)\left(q^{0}+q^{\prime 0}\right)+2 q_{\perp}^{\prime 2}\left(2\left(q^{0}\right)^{2}+q^{3} q^{0}-q^{3} q^{\prime 0}\right)\right)\right)  \tag{A.43}\\
& U_{49}^{r}=-\frac{\sqrt{s}}{16 \sqrt{2} Q^{4} \sqrt{Q^{4}-4 s t} q^{3}}\left(\left(Q^{6}+\left(4 q_{\perp}^{\prime 2}+\left(q^{0}\right)^{2}-2\left(q^{\prime 0}\right)^{2}+t-3 q^{0} q^{3}+q^{0} q^{00}-q^{3} q^{0}\right) Q^{4}\right.\right. \\
& +\left(\left(4\left(q^{0}\right)^{2}-22 q^{3} q^{0}+6 q^{3} q^{\prime 0}\right) q_{\perp}^{\prime 2}+8 q^{3} q^{\prime 0}\left(\left(q^{\prime 0}\right)^{2}-\left(q^{0}\right)^{2}\right)+t\left(\left(q^{0}\right)^{2}+\left(q^{00}-3 q^{3}\right) q^{0}-7 q^{3} q^{\prime 0}\right)\right) Q^{2} \\
& \left.\left.-8 q^{0} q^{3}\left(q^{0}+q^{\prime 0}\right)\left(q^{0} q_{\perp}^{\prime 2}+t q^{\prime 0}\right)\right)\right)  \tag{А.44}\\
& U_{50}^{r}=\frac{q_{\perp}^{\prime} \sqrt{s}\left(\left(q^{0}+q^{3}-2 q^{\prime 0}\right) Q^{2}+t\left(q^{0}+q^{3}\right)\right)\left(q^{0}-q^{\prime 0}\right)}{4 \sqrt{2} \sqrt{Q^{4}-4 s t} q^{3}}  \tag{A.45}\\
& U_{51}^{r}=-\frac{\sqrt{s}}{2 \sqrt{2} Q^{4} \sqrt{Q^{4}-4 s t} q^{3}}\left(\left(Q^{6}+\left(4 q_{\perp}^{\prime 2}+\left(q^{0}\right)^{2}-3\left(q^{\prime 0}\right)^{2}+t+3 q^{0} q^{\prime 0}+q^{3} q^{\prime 0}\right) Q^{4}\right.\right. \\
& +\left(2 q^{0}\left(3 q^{0}+2 q^{3}+q^{\prime 0}\right) q_{\perp}^{\prime 2}+2\left(q^{0}+q^{3}\right) q^{00}\left(\left(q^{0}\right)^{2}-\left(q^{\prime 0}\right)^{2}\right)+t\left(\left(q^{0}\right)^{2}+3 q^{\prime 0} q^{0}+q^{\prime 0}\left(q^{3}+q^{\prime 0}\right)\right)\right) Q^{2} \\
& \left.\left.+2 q^{0}\left(q^{0}+q^{3}\right)\left(q^{0}+q^{\prime 0}\right)\left(q^{0} q_{\perp}^{\prime 2}+t q^{0}\right)\right)\right) \tag{А.46}
\end{align*}
$$

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