



COLLÈGE
DE FRANCE
— 1530 —



Solar record based on cosmogenic isotopes (^{10}Be & ^{14}C)

Edouard BARD

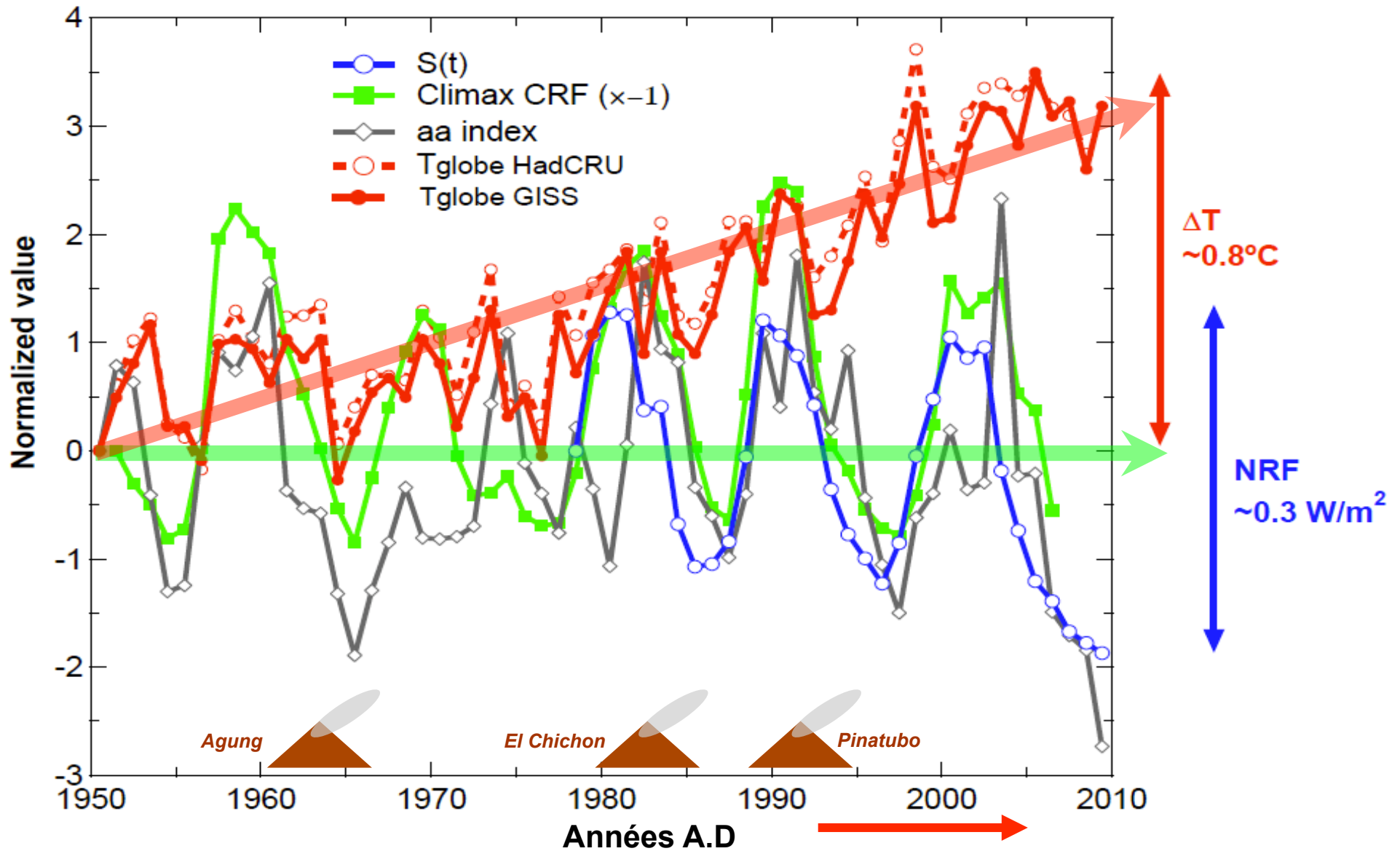
*Chaire de l'évolution du climat et de l'océan
du Collège de France
CEREGE, Aix-en-Provence*

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- LGGE: **J.R. Petit, O. Magand**

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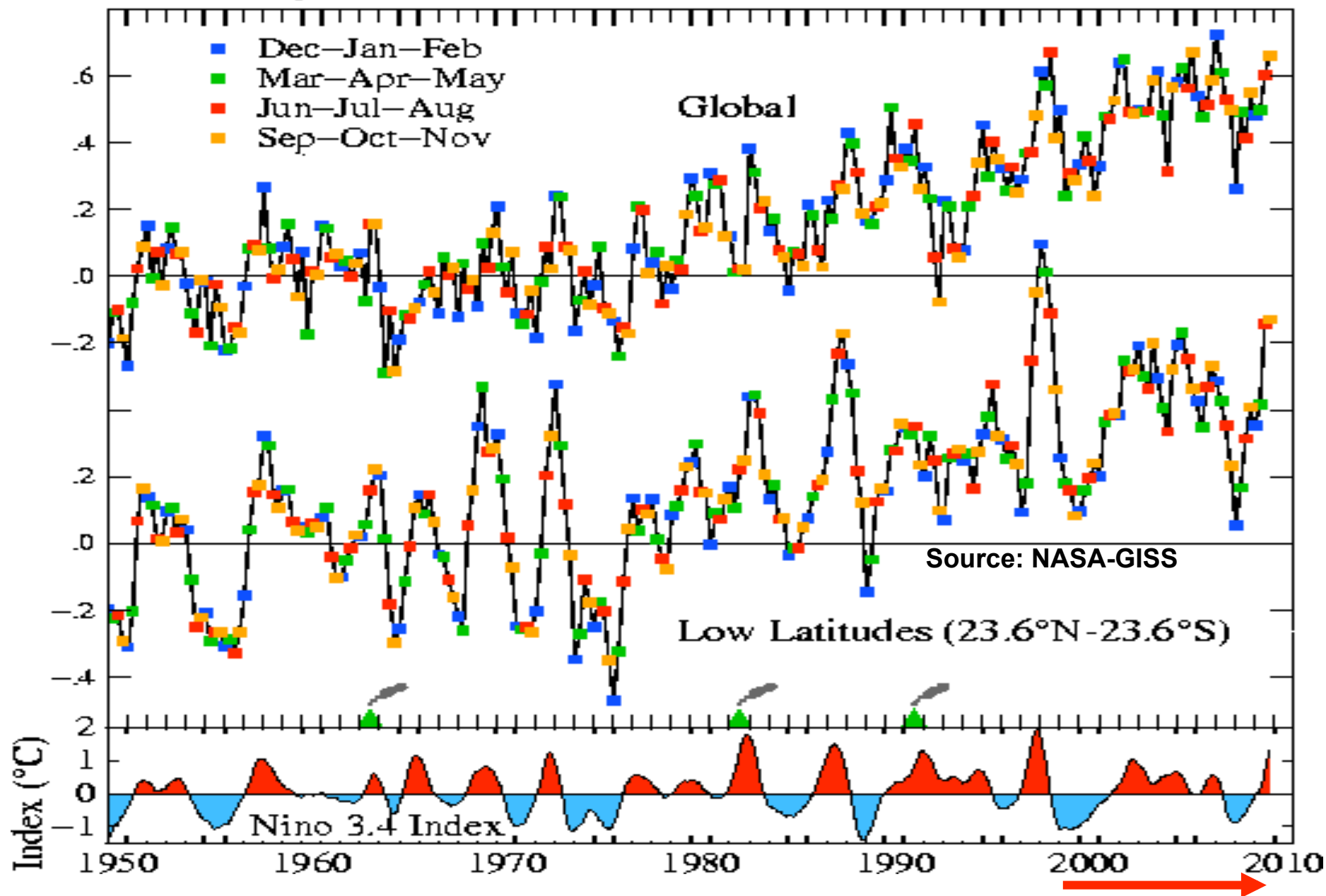
Over the past 60 years, the total irradiance and heliomagnetic modulation of cosmic rays doesn't seem to show a long-term trend that could have contributed to global warming



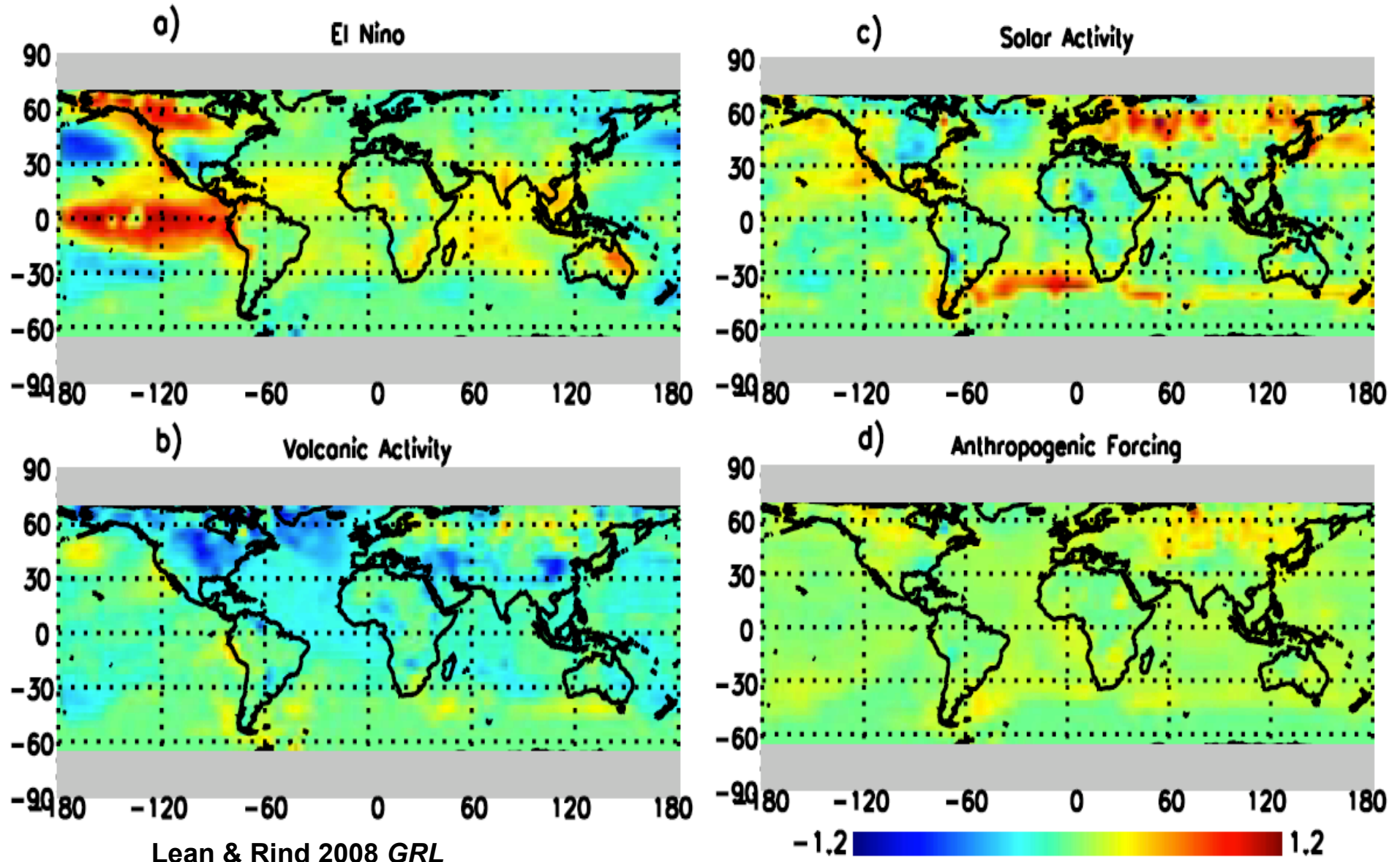
Bard & Delaygue 2008 *EPSL* (updated)

Beyond the annual mean T: seasonal and zonal variations

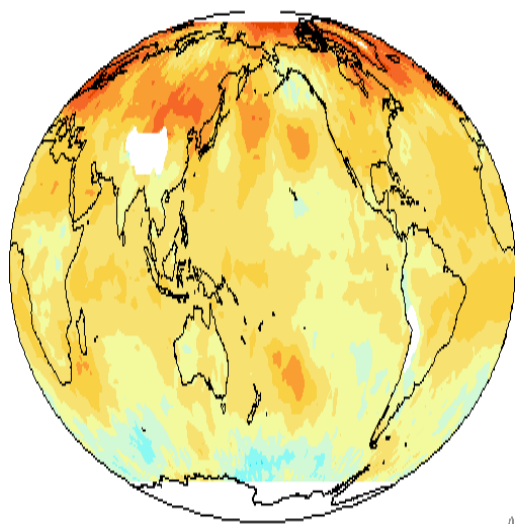
Temperature Anomaly ($^{\circ}\text{C}$): Seasonal Resolution



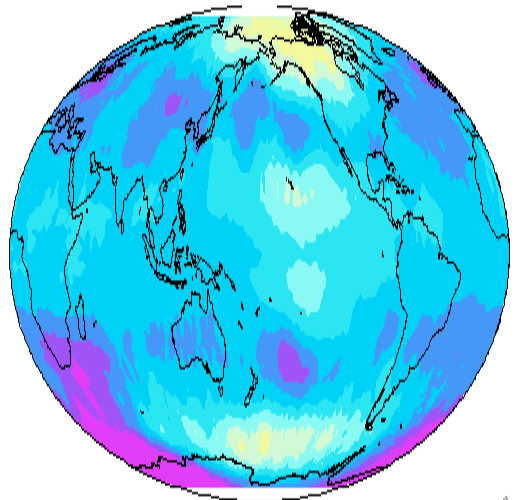
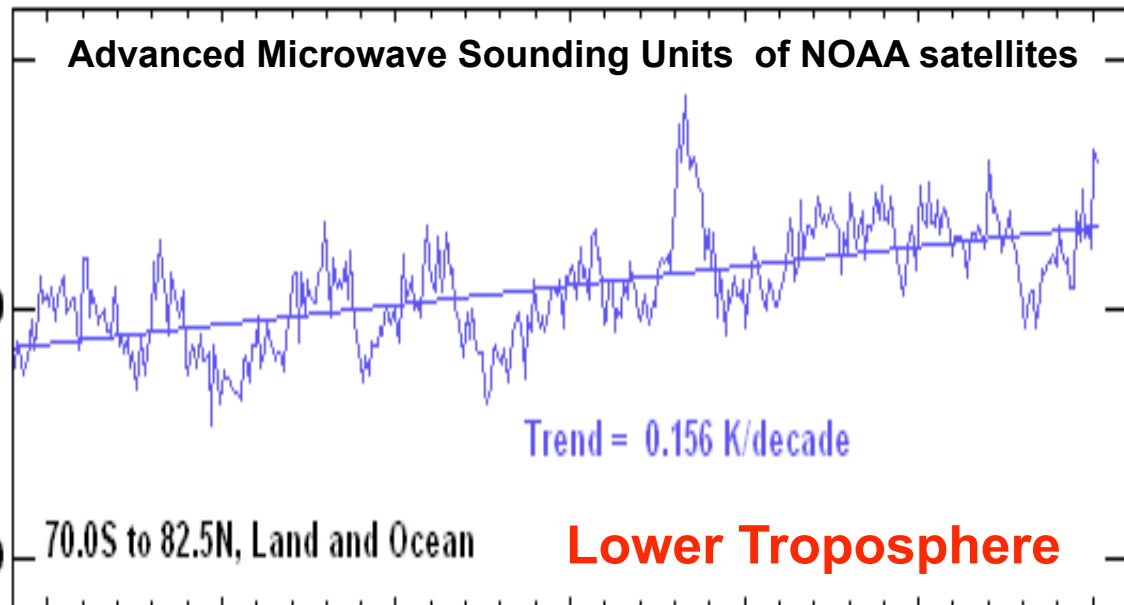
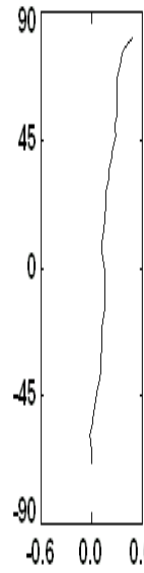
Beyond the annual mean temperature: multivariate spatial correlation analysis (1889-2006)



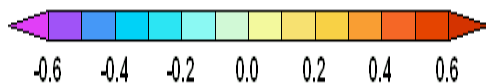
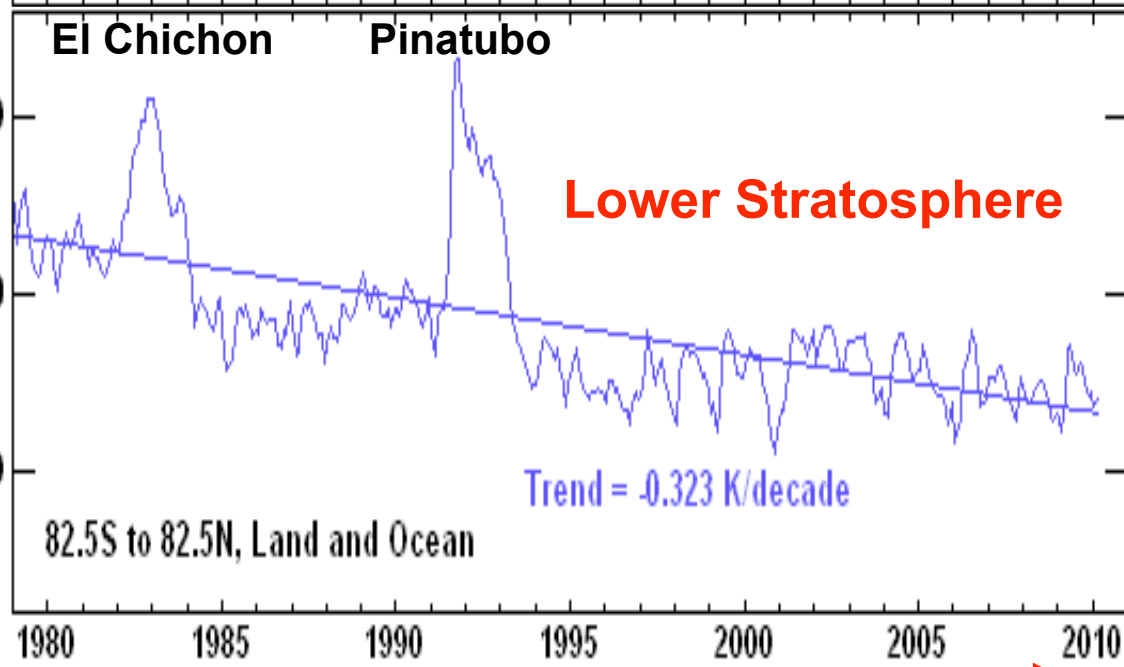
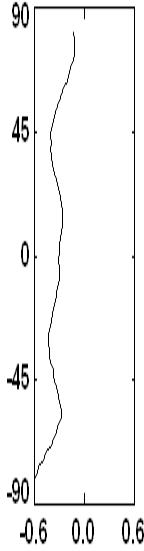
Beyond the annual mean temp : altitudinal variations



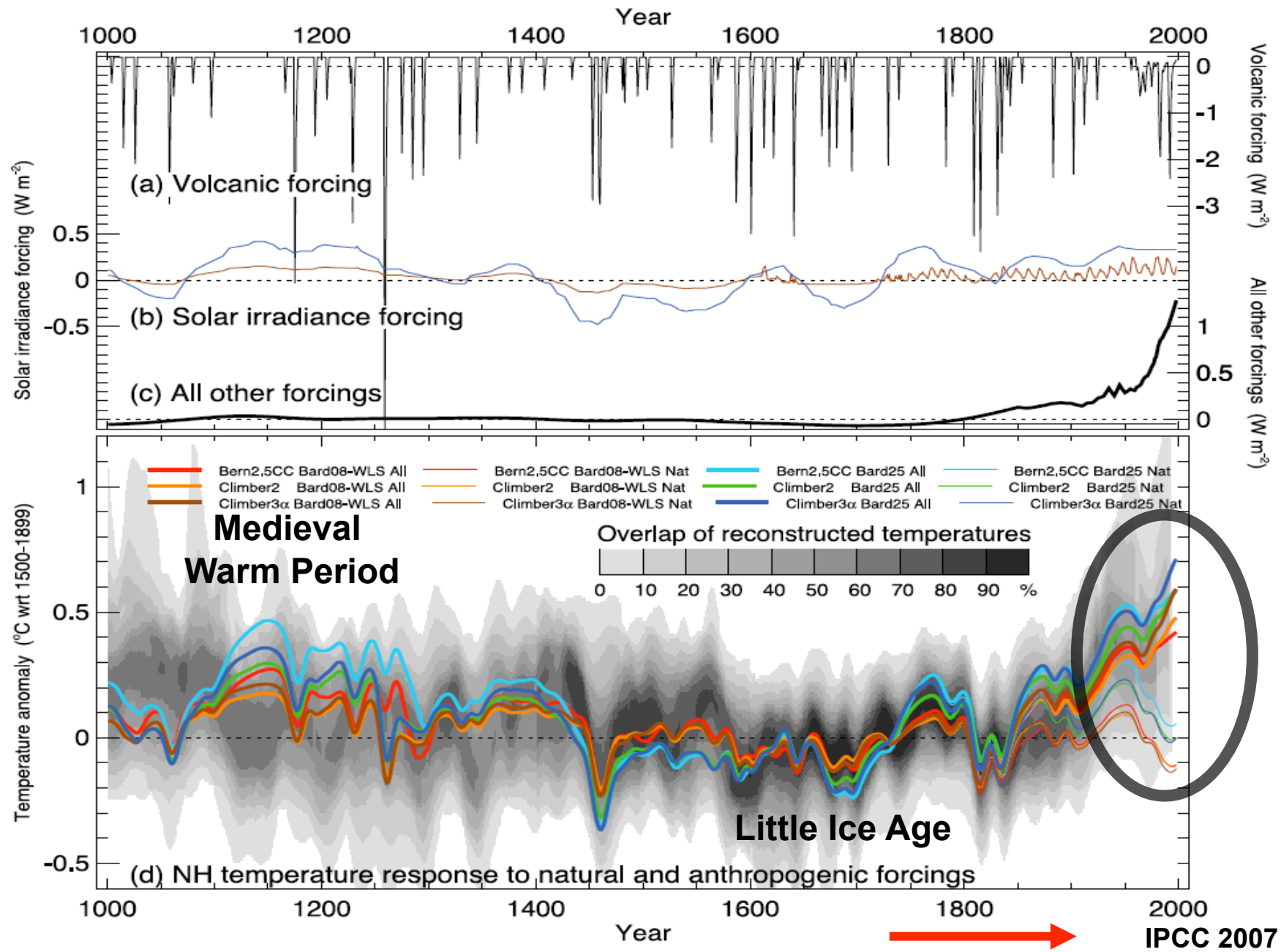
Remote Sensing Systems
www.remss.com



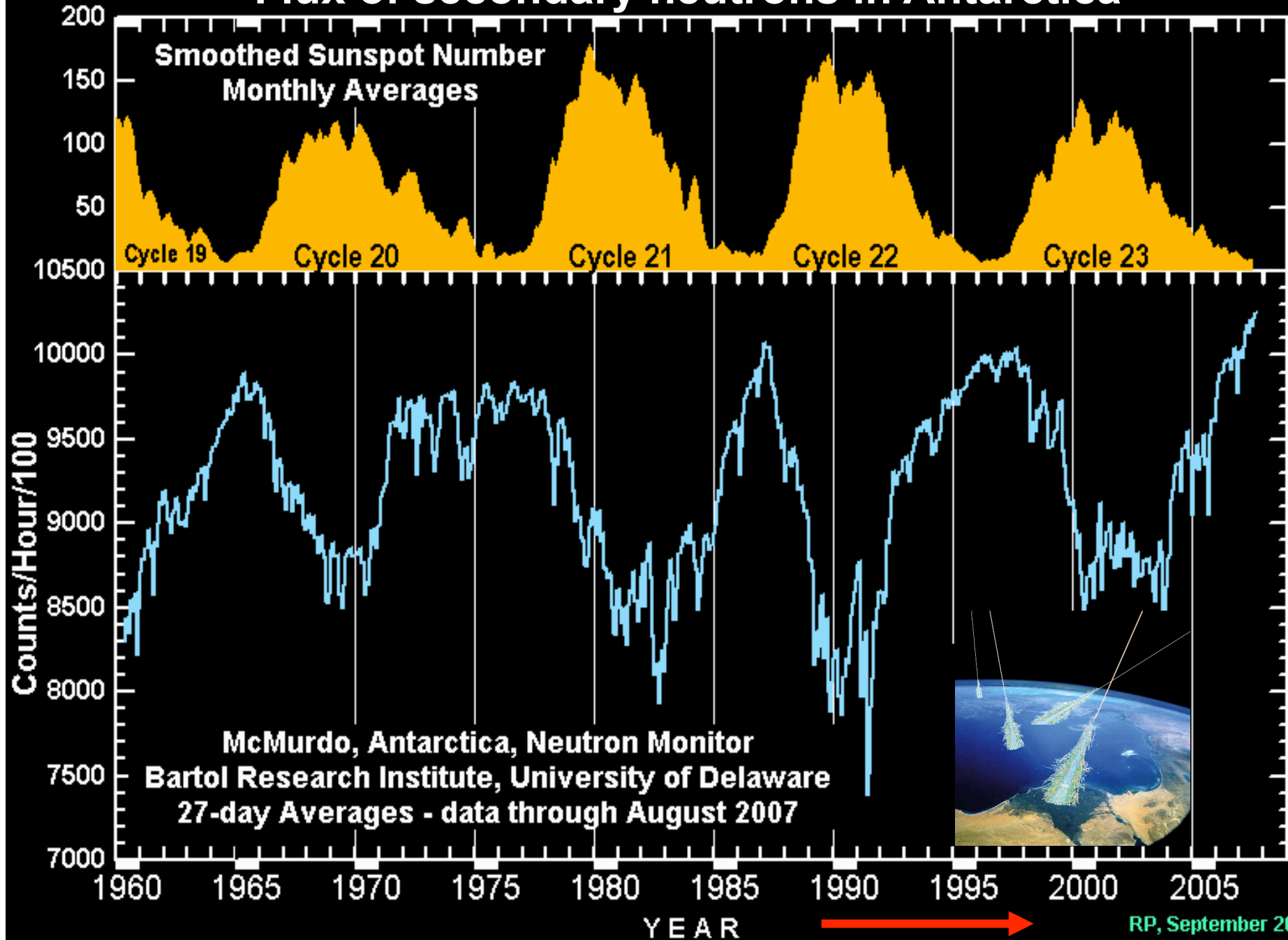
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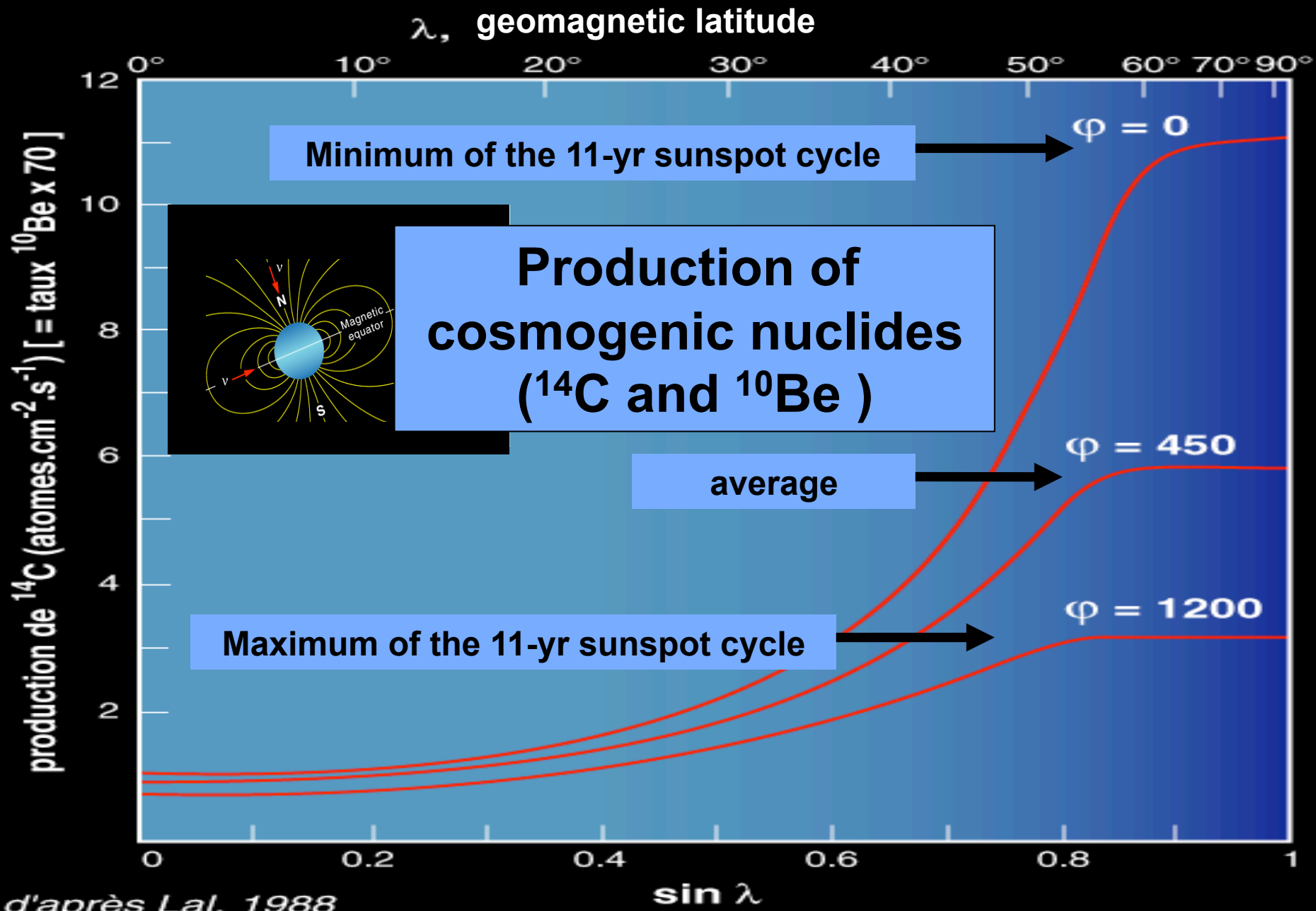
Model data comparison for the last millennium



Flux of secondary neutrons in Antarctica



Geo- and heliomagnetic modulation



^{10}Be production as a function of **altitude**,
the **solar modulation parameter Φ** and
the intensity of the **geomagnetic dipole M/M_0**

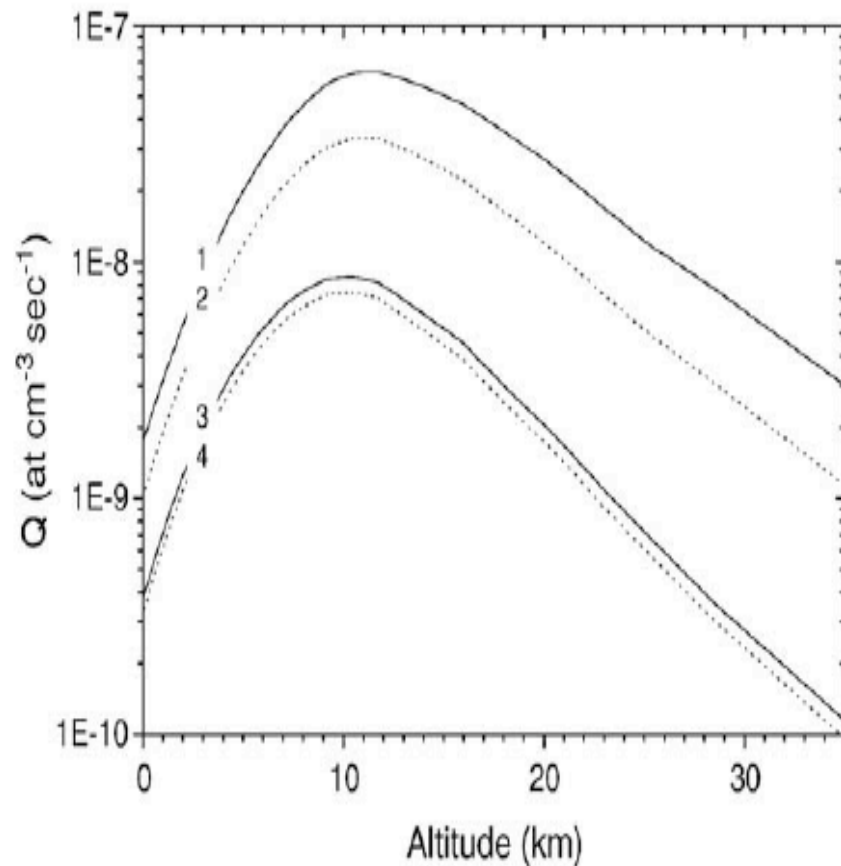


Fig. 2. Local production Q (atoms $\text{cm}^{-3} \text{s}^{-1}$) of ^{10}Be in the atmosphere as function of altitude. Different curves correspond to: 1 – solar minimum ($\phi = 300$ MV), polar region; 2 – solar maximum ($\phi = 1000$ MV), polar region; 3 – solar maximum, equator; and 4 – solar maximum, equator.

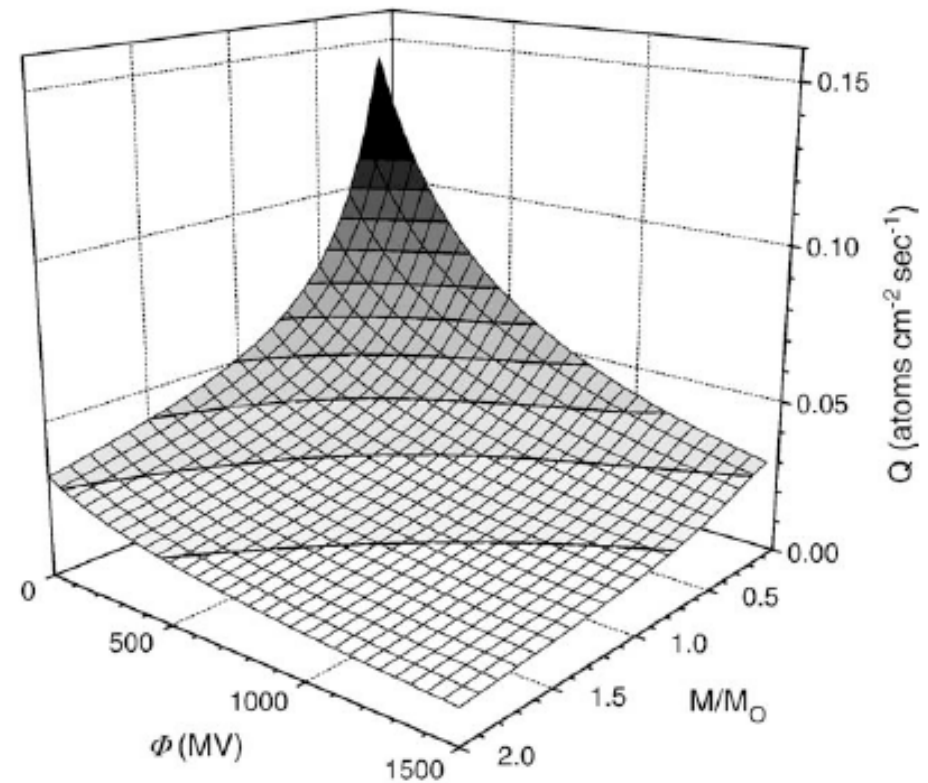
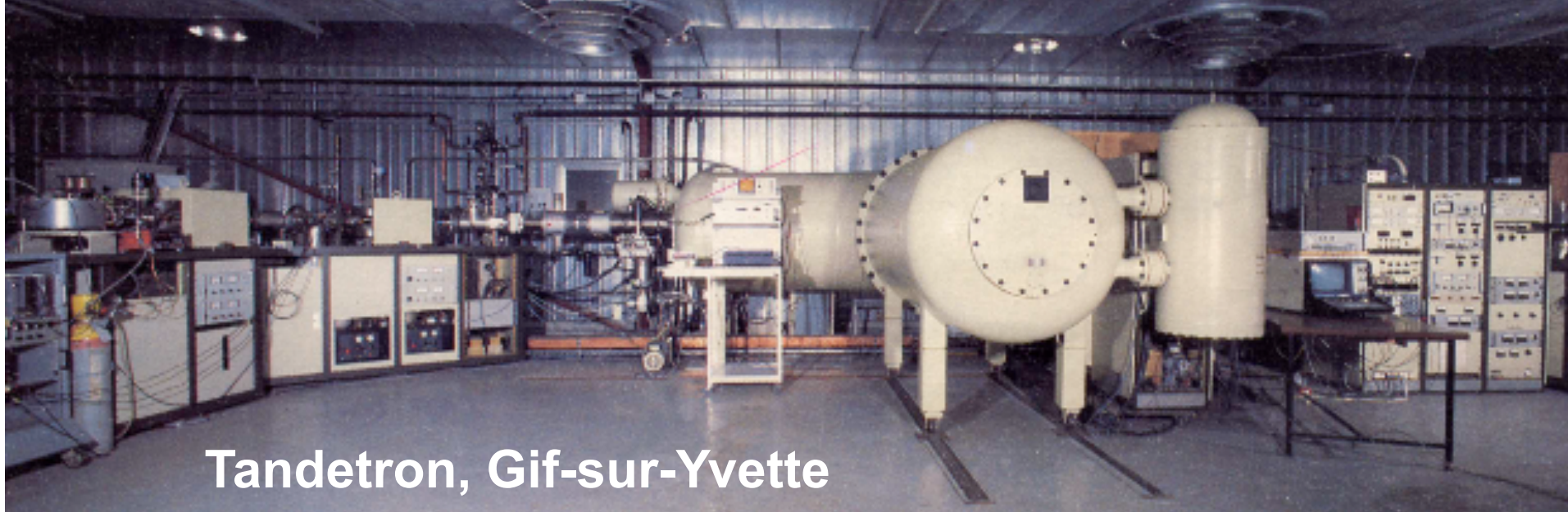


Fig. 1. Global production rate (atoms $\text{cm}^{-2} \text{s}^{-1}$) of ^{10}Be as function of the modulation potential ϕ and the magnetic dipole strength relative to the present one, M/M_0 .



Tandetron, Gif-sur-Yvette

**^{14}C , ^{10}Be and ^{36}Cl by
Accelerator Mass
Spectrometry**



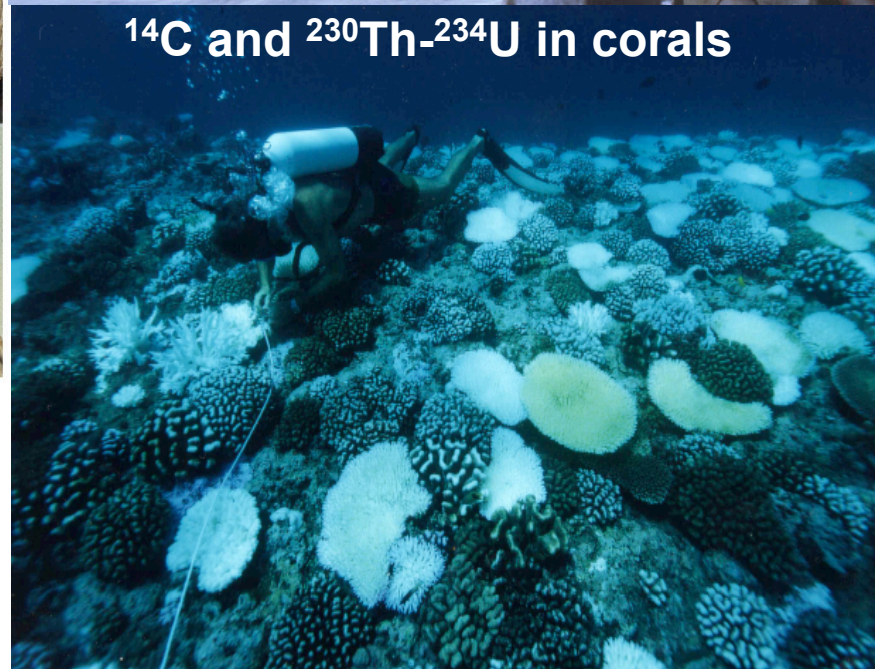
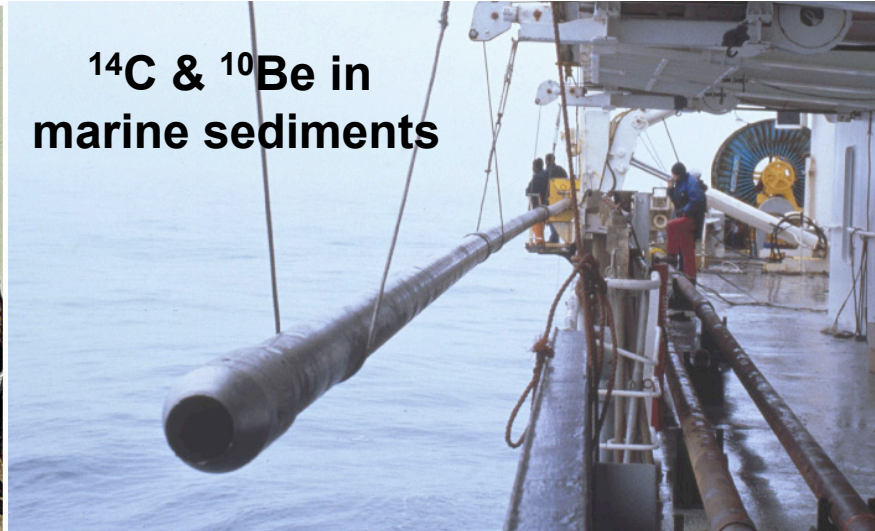
**ARTEMIS
Saclay**

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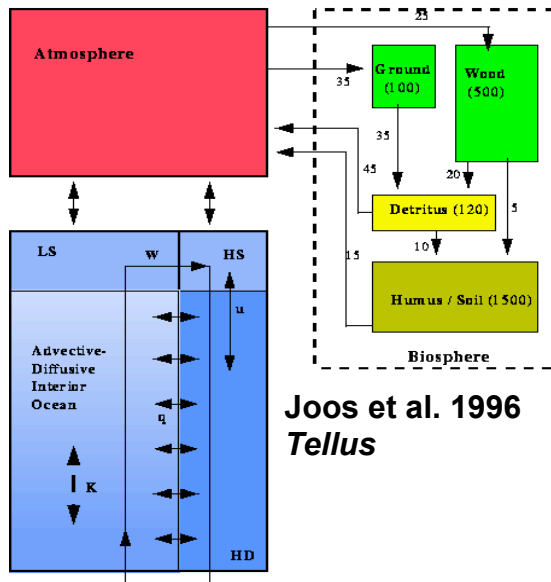
ASTER, Aix-en-Provence

Beyond the 17th century, the only way to study the solar activity is to measure cosmogenic nuclides in various “geological” archives

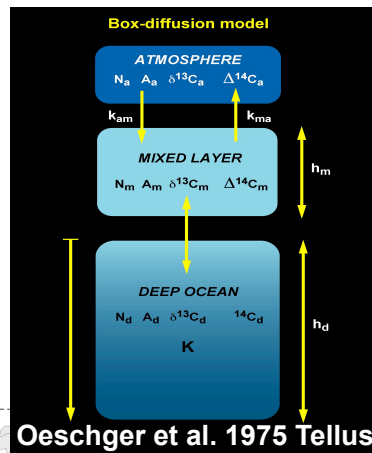
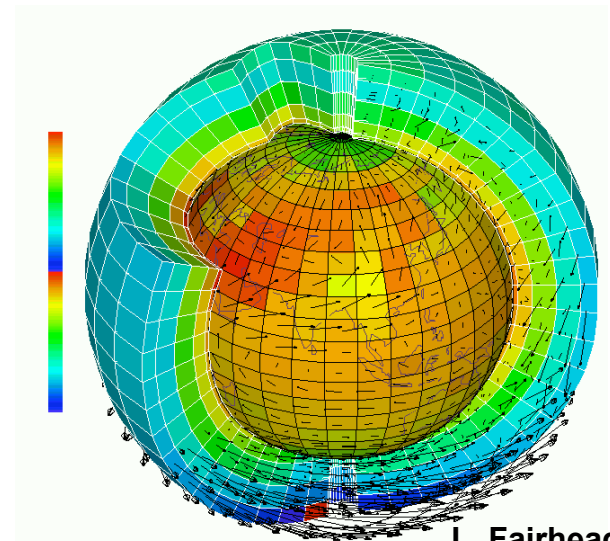
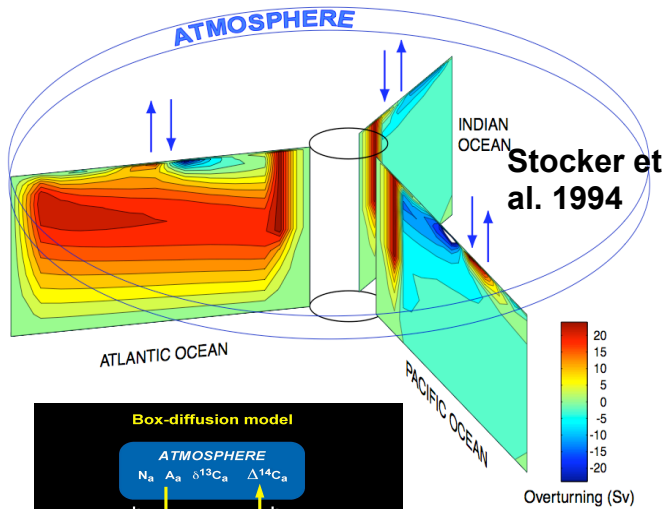


Comparing ^{14}C and ^{10}Be records requires the use of numerical models of various complexity

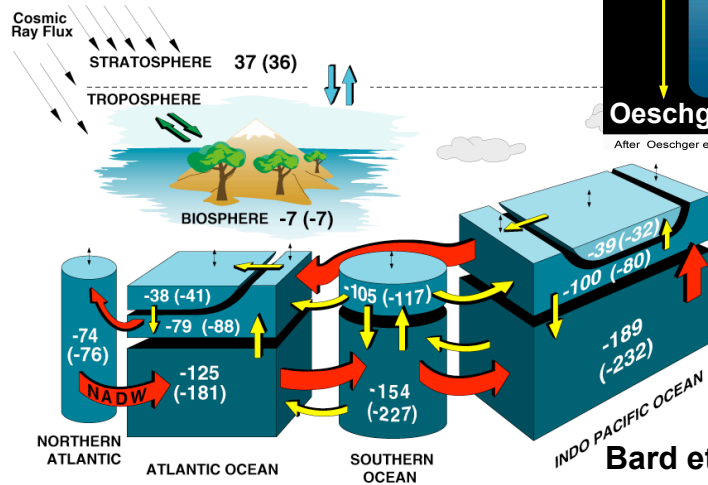
The Bern Carbon Cycle Model



Joos et al. 1996
Tellus

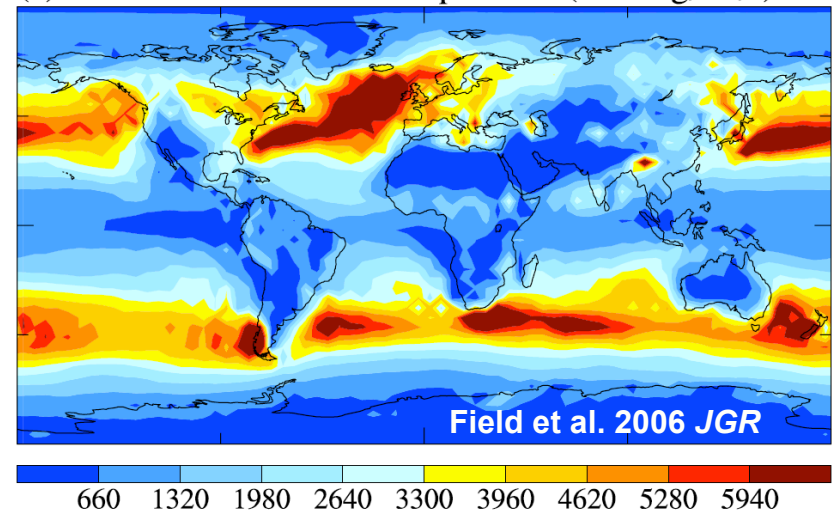


Oeschger et al. 1975 *Tellus*
After Oeschger et al. 1975

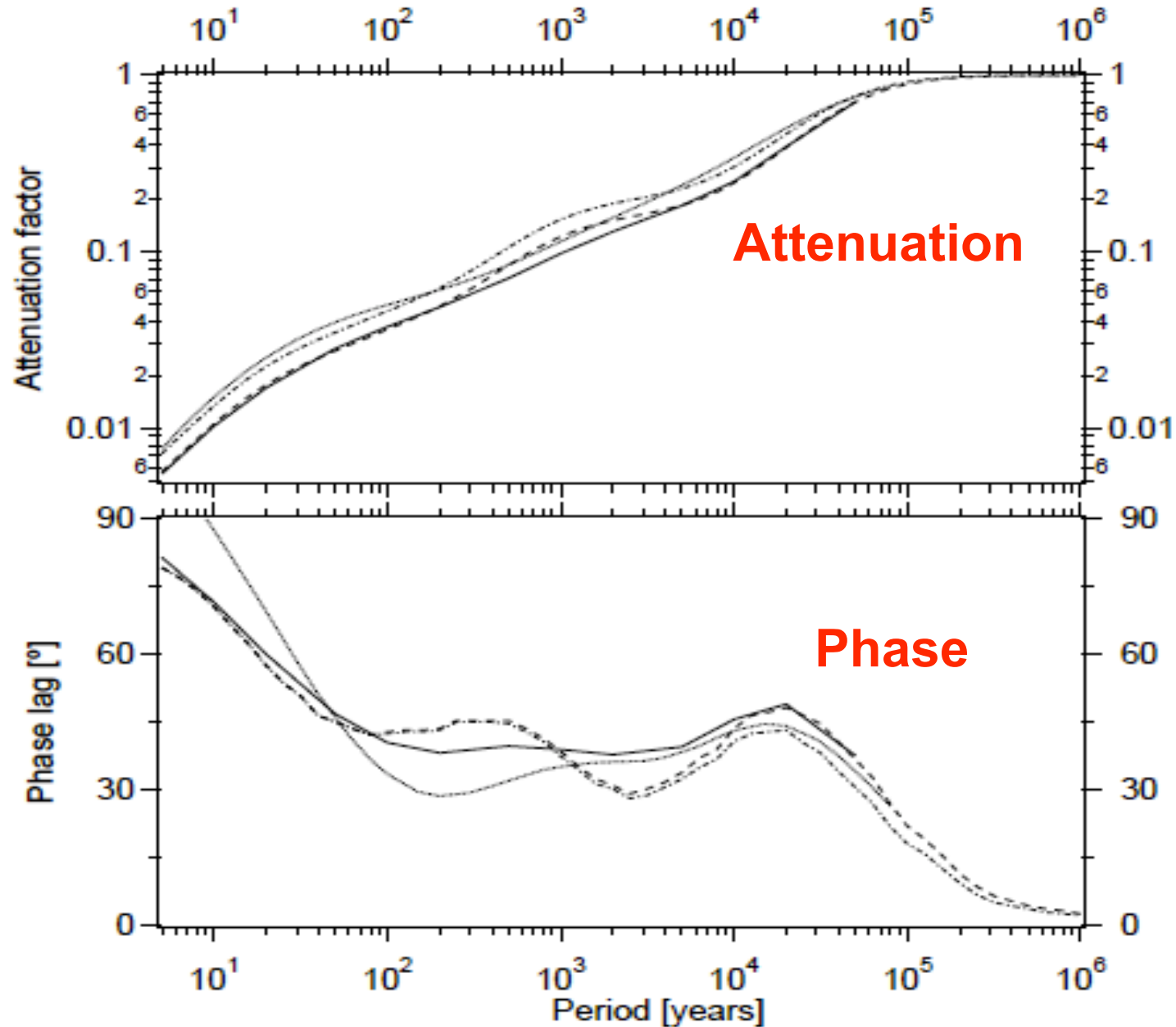


Bard et al. 1994 *EPSL*

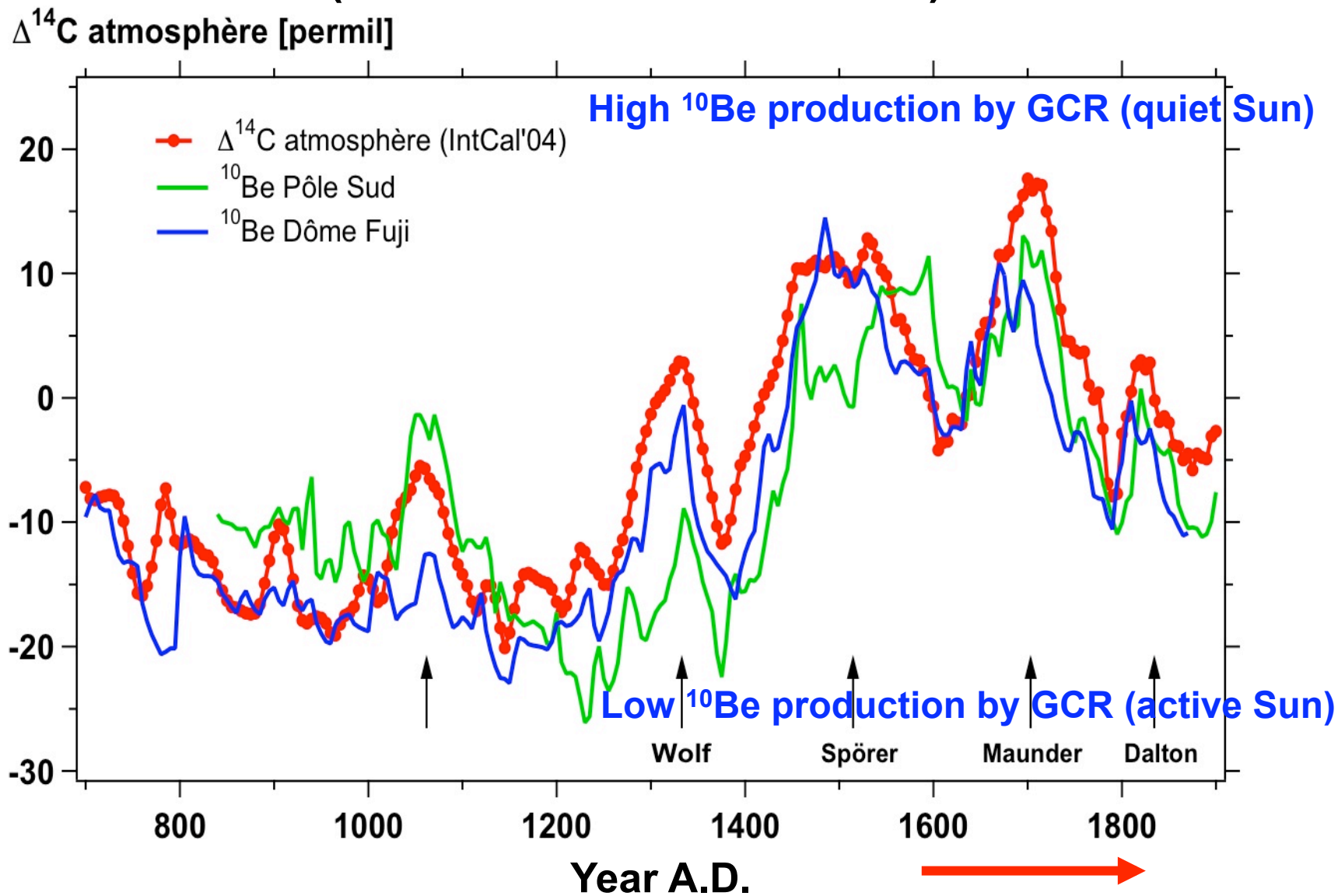
(a) Annual mean wet ^{10}Be deposition (10^{-27} kg/m²/s)



Geochemical cycles are low-pass filters for ^{14}C & ^{10}Be

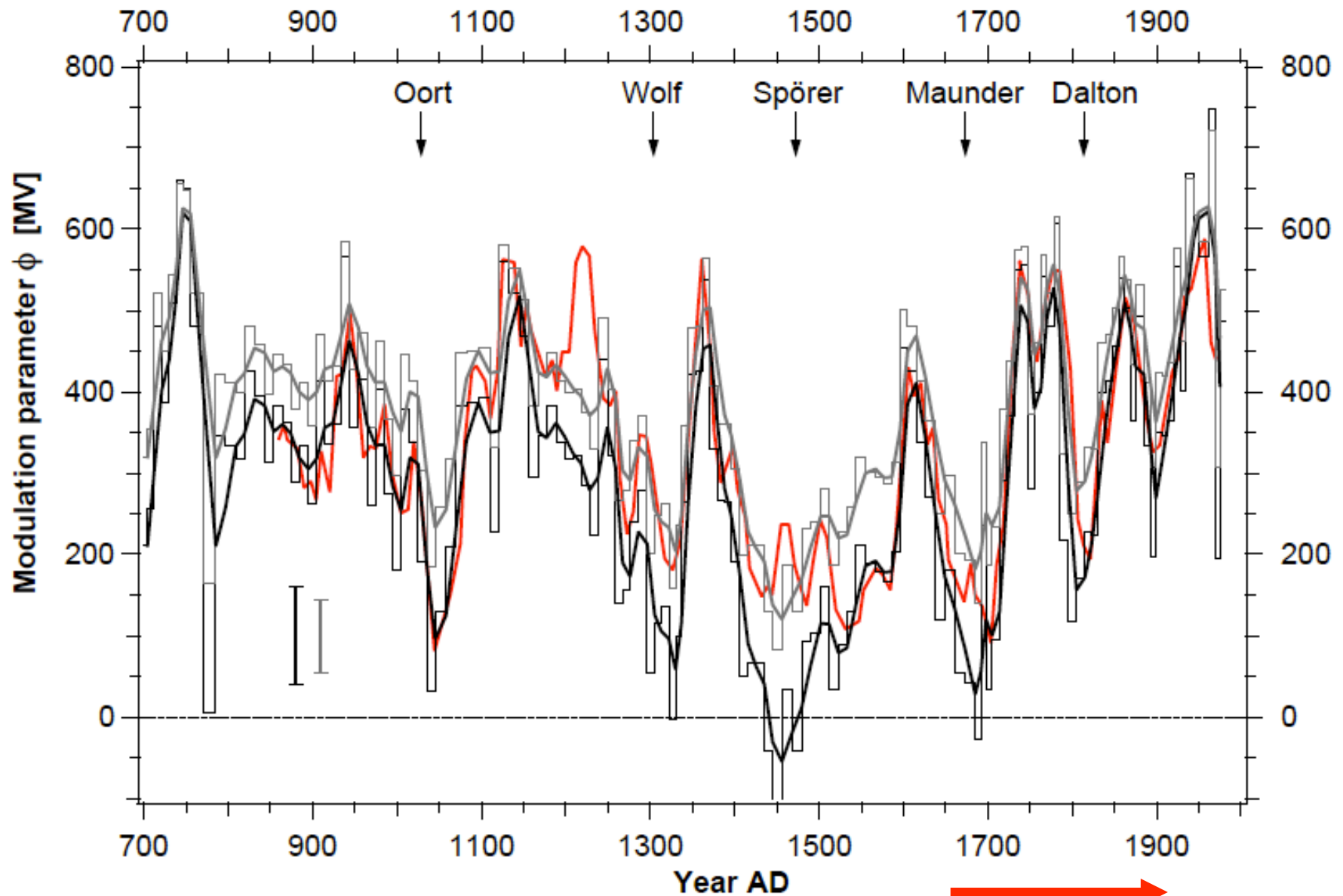


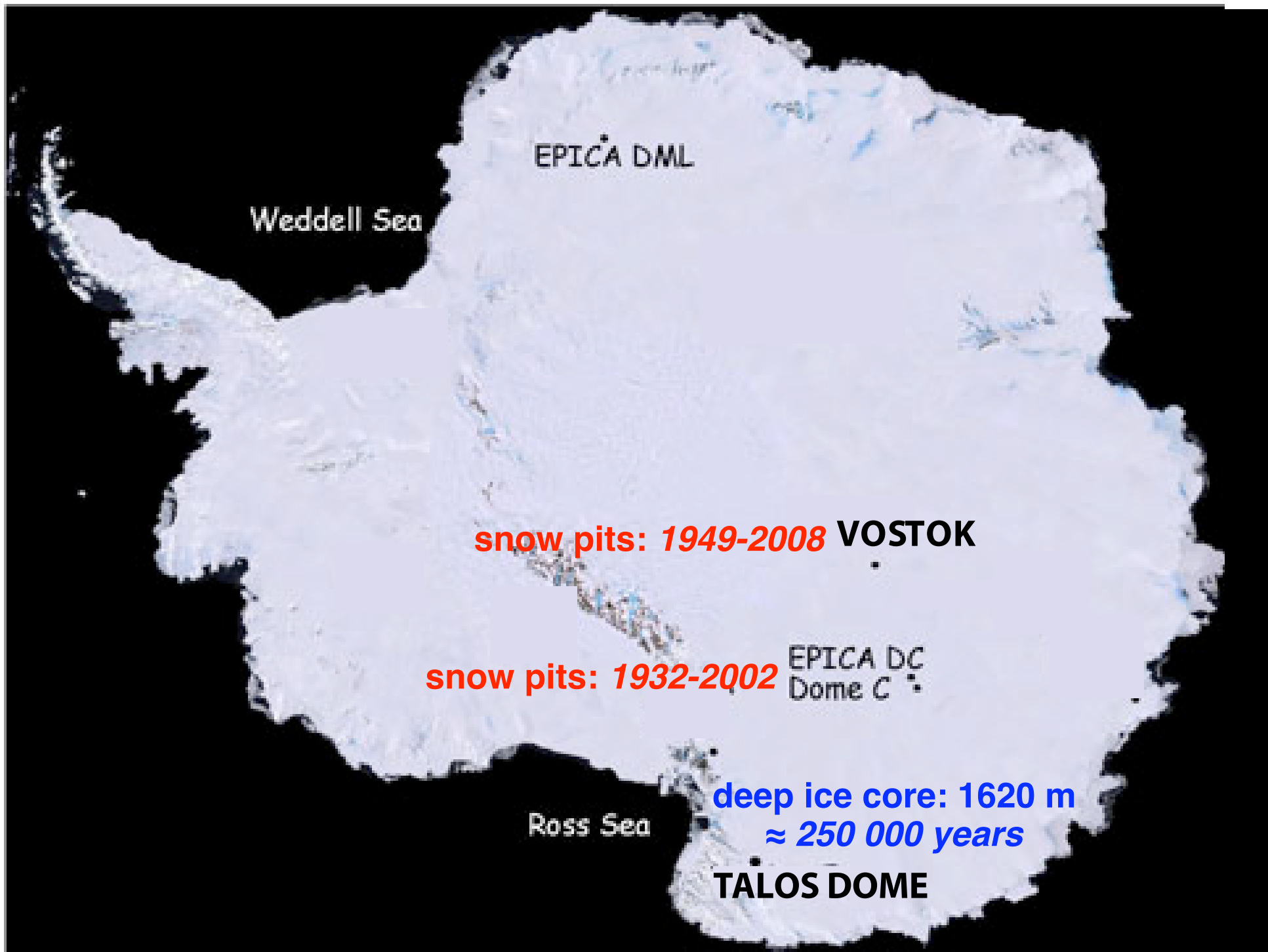
^{10}Be in ice cores from Antarctica and ^{14}C in tree-rings (all converted in ^{14}C units)



Bard et al. 1997 *EPSL*, 2000 *Tellus*, Horiuchi et al. 2008 *QG*, Delaygue & Bard 2010 *Clim. Dyn.*

An Antarctic stack of ^{10}Be -based solar activity for the past millennium





A new high-resolution ice-core allowing to scale the variability during the glacial period with that of the recent past (11-yr cycle, Maunder-type events ...)

**French-Italian base of Talos Dome Ross
Sea area (drilling in 2005-06)**

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Some conclusions about cosmogenic isotopes

- Time-series of the solar activity can be obtained with cosmogenic nuclides overlapping with other data over recent periods: TSI since 1978, geomagnetic index since 1868, sunspots since 1610, aurorae borealis since ca. 1500...
- The Sun varies over decades to millennia between active phases and solar minima (cycles of ca. 11, 90, 150 & 200 years, Maunder type minima),
- Overall, the Sun spent about a quarter of the last 7000 years in solar minima, often “deeper” than the Maunder Minimum.

Some remaining questions

- How was the Sun behaving during the (a) Maunder Minimum? Was the 11-yr cycle still present (some hint but still inconclusive) ? How rapid are the transitions in and out the (a) Maunder Minimum ?
- What is the best way to convert solar “proxy” data (^{10}Be , ^{14}C) in terms of irradiance and other effects relevant to Earth climate ?