The Reactor Antineutrino Anomaly and implications



Th. Lasserre (CEA-Saclay, Irfu APC & SPP)

New Reactor Antineutrino Spectra

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CEA / Irfu & IN2P3 / Subatech

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v spectrum emitted by a reactor

The prediction of reactor \mathbf{v} spectrum is the dominant source of systematic error for single detector reactor neutrino experiments





CEA DSM Irfu T. Lasserre

Complementary approaches to compute the v flux Integral Ab initio measurements Sum of all fission products' activities fission product inventory (t) fission rates (t) Fission Yields (JEFF, ENDF, JENDL) I ife time Sum of all β -branches Build total spectrum from **Reference** spectrum of each fission product sum of β -branches per isotope Complete β-decays schemes (ENSDF) • || | • β-strength (Greenwood et al.) electron Theory of β-• Total β spectrum per nucleus (Rudstam et al.) data decay • Masses (Q_b) Nuclear models … **Our mixed approach** Effective **Full ab-inito**

Unique reference to be met by any other measurement or calculation



ILL data: conversion to v spectra

- Fit e⁻ spectrum with a sum of 30 effective branches
- Conversion of the effective branches to v spectra



• All theory included in these effective branches but:

- What Z? : Mean fit on nuclear data Z=f(E0) $Z(E_0) \approx 49.5 - 0.7E_0 - 0.09E_0^2, Z \ge 34$
- What A_{CW} ? : effective correction on the v-spectra $DN_n^{C,W}(E_n) \approx 0.65 \times (E_n - 4MeV) \%$
- Conversion error from envelop of numerical studies





- MURE evolution code: core composition and off equilibrium effects
- BESTIOLE code: build up database of ~800 nuclei and 10000 β -branches



→ 95+/-5% of the spectrum reproduced but still not meeting required precision → Useful estimate of ²³⁸U spectrum which couldn't be measured @ ILL

→ Measurement at FRMII ongoing (N. Haag & K Schreckenbach)



- 1. SAME ILL e- data Anchorage
- 2. Ab-Initio: "true" distribution of β -branches reproduces >90% of ILL e⁻ data.
- 3. Old-procedure: five effective anchorage-branches to the remaining 10%.



- +3% normalization shift with respect to old v spectrum
- Similar result for all isotopes (²³⁵U, ²³⁹Pu, ²⁴¹Pu)
- Stringent Test Performed Origin of the bias identified



- Define "true" e⁻ and n spectra from reduced set of well-known branches from ENSDF nuclei data base.
- Apply exact same OLD conversion procedure to true e⁻ spectrum.
- 3. Compare the converted n spectrum to the true one.
- This technique gives a 3% bias compared to the true v spectrum



 \rightarrow **OLD** effective conversion method biases the predicted v spectrum at the level of -3% in normalization

Origin of the 3% shift



 E <4 MeV: deviation from effective linear A_{C,W} correction of ILL data

$$\Delta N_v^{C,W}(E_v) \approx 0.65 \times (E_v - 4\,MeV) \quad \%$$

• **E** >4 MeV: mean fit of $Z(E_0)$ doesn't take into account the very large dispersion of Z around the mean curve $Z(E_0) \approx 49.5 - 0.7E_0 - 0.09E_0^2, Z \ge 34$

Off-Equilibrium Effects MURE evolution code (IN2P3/Subatech)

- ILL electron reference spectra : 12 hours to 1.8 days irradiation time
- Neutrino reactor experiments irradiation time >> months

BUT 10% of fission products have a β-decay life-time long enough to keep accumulating after several days

- \rightarrow need a correction through simulation
- \rightarrow Not included prior to the CHOOZ experiment

Relative change of v spectrum w.r.t. infinite irradiation time

Correction included by default in our new reference model



The Reactor Antineutrino Anomaly

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CEA / Irfu

arXiv:1101.2755 [hep-ex], accepted for publication in PRD * corresponding author - Inverse Beta Decay: $\bar{\nu}_e + p \rightarrow e^+ + n$

- Theoretical predictions: our results agree with
 - Vogel 1984 (Phys Rev D29 p1918). Fayans 1985 (Sov J Nucl Phys 42)
 - Vogel-Beacom 1999: "supersedes" Vogel 84 (Phys Prev D60 053003)
 - Strumia-Vissani Phys. Lett. B564 (2003) 42-54

$$\sigma_{\rm V-A}(E_e) = \kappa \, p_e E_e (1 + \delta_{rec} + \delta_{wm} + \delta_{rad})$$

- The pre-factor κ (two pseudo-independent approaches)

$$\kappa = \frac{G_F^2 \cos^2(\theta_C)}{\pi} (1 + \Delta_{inner}^R)(1 + 3\lambda^2) = \frac{2\pi^2}{m_e^5 f^R \tau_n} \qquad \lambda = |\frac{g_A}{g_V}|^2$$

- κ ran down over the history, from 0.914 10⁻⁴² cm² in 1981
 - Vogel-Beacom 1999 : κ = 0.952 10⁻⁴² cm²
 - Our work is based on 2010 PDG $\tau_{\rm n}$: κ = 0.956 10^{-42} \, cm^2
 - But we anticipate 2011 κ =0.961 10⁻⁴² cm² (< τ_n > revision +0.5%)

Reactor Electron Antineutrino Detection

- Inverse Beta Decay: $\bar{\nu}_e + p \rightarrow e^+ + n$
 - Threshold: 1.806 MeV
- Anti-v_e interaction rate $n_{\nu} = \frac{1}{4\pi R^2} \frac{P_{\rm th}}{\langle E_f \rangle} N_p \varepsilon \sigma_f$
- Experimental cross section per fission: $\sigma_{\rm f}$

$$\sigma_f^{\text{meas.}} = \frac{4\pi R^2 n_{\nu}^{\text{meas.}}}{N_p \varepsilon} \frac{\langle E_f \rangle}{P_{\text{th}}}$$

- Predicted cross section per fission: σ_{pred}

$$\sigma_f^{\text{pred.}} = \int_0^\infty \phi_f^{\text{pred.}}(E_\nu) \sigma_{\text{V-A}}(E_\nu) dE_\nu$$

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Computing the expected rate/spectrum





Bugey-4 BenchmarkPhys Lett B 338(1994) 383

 $\cdot \tau_{n} = 887.4 \text{ s}$

- "old" spectra (30 effective branches)
- no off-equilibrium corrections

10 ⁻⁴³ cm ² / fission	²³⁵ U	²³⁹ Pu	²⁴¹ Pu
BUGEY-4	6.39±1.9%	4.18±2.4%	5.76±2.1%
This work	6.39±1.8%	4.19±2.3%	5.73±1.9%

Final agreement to better than 0.1% on best known ²³⁵U



- ν-flux: ²³⁵U +2.5%, ²³⁹Pu +3.1%, ²⁴¹Pu +3.7%, ²³⁸U +9.8% (σ_f^{pred} **7**)
- Off-equilibrium corrections now included $(\sigma_f^{\text{pred}} \nearrow)$
- Neutron lifetime decrease by a few % ($\sigma_{\rm f}^{\rm pred}$ **7**) $\sigma_{\rm V-A}(E_{\nu}) \propto 1/\tau_n$
- Slight evolution of the phase space factor ($\sigma_{f}^{pred} \rightarrow$)
- Slight evolution of the energy per fission per isotope ($\sigma_{f}^{pred} \rightarrow$)

• Burnup dependence:
$$\sigma_f^{pred} = \sum_k f_k \sigma_{f,k}^{pred} \quad (\sigma_f^{pred} \rightarrow)$$

_		old [3]	new	new/old
_	$\sigma^{pred}_{f,235U}$	$6.39{\pm}1.9\%$	$6.61{\pm}2.11\%$	+3.4%
New	$\sigma^{pred}_{f,239Pu}$	$4.19{\pm}2.4\%$	$4.34{\pm}2.45\%$	+3.6%
Results:	$\sigma^{pred}_{f,238_U}$	$9.21{\pm}10\%$	$10.10{\pm}8.15\%$	+9.6%
_	$\sigma_{f,^{241}Pu}^{ m pred}$	$5.73{\pm}2.1\%$	$5.97{\pm}2.15\%$	+4.2%

19 Experimental Results below 100m



Measured cross sections are taken at their face values

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ROVNO-88 (5 measurements, Sov Phys JETP67, 1988)

- Rovno, Russia, VVER, 1983-1986
- Technology
 - Integral detector with PE target containing ³He counters, only neutrons are detected
 - Liquid Scintillator detector
- Baselines
 - 18 m & 25 m
- Typical fuel composition: 60.7% ²³⁵U, 27.7% ²³⁹Pu, 7.4% ²³⁸U, 4.2% ²⁴¹Pu,
- Uncertainties:
 - statistics: < 0.9%</p>
 - systematics: 7-8%
- Correlated with:
 - Bugey-4
 - Rovno91 (integral measurement only),
 - with each other



ROVNO-91 (JETP Lett., 54, 1991, 253)

- Rovno, Russia, VVER, late 80's
- Technology:
 - Upgraded integral detector : water target containing ³He counters, only neutrons are detected
- Baselines
 - 18 m
- Fuel composition:
 - 61.4% ²³⁵U, 27.4% ²³⁹Pu, 7.4% ²³⁸U, 3.8% ²⁴¹Pu
- Uncertainties:
 - statistics: <1%</p>
 - systematics: 3.8%
- Correlated with:
 - Bugey-4 (same detector)



Bugey-4 (Phys. Lett. B338, 383, 1994)

- Bugey, France, PWR, early 1990s
- Technology:
 - Integral detector : water target containing ³He counters, only neutrons are detected
- Baseline
 - 15 m
- Fuel composition:

53.8% ²³⁵U, 32.8% ²³⁹Pu, 7.8% ²³⁸U, 5.6% ²⁴¹Pu

- Uncertainties:
 - statistics: 0.04%
 - systematics: 3% (most precise exp.)
- Correlated with:
 - ROVNO-91 (same detector)
 - ROVNO-88 (50% arb.)
- Experimental cross section used to normalize the CHOOZ experiment result





Bugey-3 (3 measurements, Nucl Phys B434, 504, 1995)

- Bugey, France, PWR, 80's
- Technology
 - Liquid scintillator segmented detectors doped with ⁶Li
- Fuel composition typical of PWR
 53.8% ²³⁵U, 32.8% ²³⁹Pu, 7.8% ²³⁸U 5.6% ²⁴¹Pu
- Baselines
 - 14m, 42m and 95m:
- Uncertainties:
 - statistics: 0.4%, 1.0%, 13.2%
 - systematics: 5.0%
- Correlated with
 - each other
- Stringent shape distortion analysis disfavoring sub-eV² oscillations





- Gösgen PWR, Switzerland, 1981-1984
- Technology:
 - liquid scintillator segmented detector + ³He counters for neutron capture
- Baselines:
 - **37.9m**, 45.9m, 64.7m
- 3 fuel compositions. Typical: 61.9% ²³⁵U, 27.2% ²³⁹Pu, 6.7% ²³⁸U, 4.2% ²⁴¹Pu
- Uncertainties:
 - statistics: 2.4%, 2.4%, 4.7%
 - systematics: 6.0%
- Correlated with
 - ILL (same detector)
 - each other





Detector assembly

30 liquid scintillator cells





- ILL, Research Reactor, Grenoble, 80-81
- Technology:
 - Liquid scintillator segmented detector + ³He counters for neutron capture
- Baselines
 - 8.76 (15) m
- Fuel composition:
 - almost pure ²³⁵U
- Uncertainties:
 - statistics: 3.5%
 - systematics: 8.9%
- Correlated with:
 - Goesgen
- Data reanalyzed in 1995 by sub-group of collaboration to correct 10% error in reactor opwer (underestimated for 10 years)



Krasnoyarsk (3 measurements, G.S. Vidyakin et al., JETP. 93, 1987)

- Krasnoyarsk research reactor, Russia
- Technology:
 - Integral detector filled with PE+ ³He counters
- Baselines:
 - 33m, 92m from 2 reactors (1987)
 - 57.3m from 2 reactors (1994)
- Fuel composition:
 - mainly ²³⁵U
- Uncertainties (33m, 57m, 92m):
 - statistics: 3.6%, 1%, 19.9%
 - systematics: 4.8% to 5.5% (corr)
- Correlated with:
 - each other





Cen Savannah River Plant (2 measurements, PRD53, 6054, 1996)

- Savannah River, USA, long standing program initiated by F. Reines. Only the last two results are included in our work.
- Technology:
 - Liquid scintillator doped with 0.5% Gd
- Baseline
 - 18.2m and 23.8 m
- Fuel composition:
 - Difference with pure ²³⁵U below 1.5%
- Uncertainties:
 - statistics: 0.6% and 1.0%: 3.7%
 - systematics:
- Correlated with:
 - each other,
 - but the two results are is slight tension





Technology							Baseline					
												<u>></u>
#	result	Det. type	$ au_n$ (s)	$^{235}\mathrm{U}$	²³⁹ Pu	$^{238}\mathrm{U}$	²⁴¹ Pu	old	new	$\operatorname{err}(\%)$	$\operatorname{corr}(\%)$	L(m)
1	Bugey-4	$^{3}\text{He}+\text{H}_{2}\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.942	3.0	3.0	15
2	ROVNO91	$^{3}\text{He}+\text{H}_{2}\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.988	0.946	4.8	4.8	15
4	Bugey-3-II	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.994	0.952	4.9	4.8	40
5	Bugey-3-III	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.915	0.876	14.1	4.8	95
6	Goesgen-I	3 He+LS	897	0.620	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	³ He+LS	897	0.584	0.298	0.068	0.050	1.045	0.992	6.5	6.0	45
8	Goesgen-II	3 He+LS	897	0.543	0.329	0.070	0.058	0.975	0.925	7.6	6.0	65
9	\mathbf{ILL}	³ He+LS	889	$\simeq 1$		—		0.832	0.802	9.5	6.0	9
10	Krasn. I	³ He+PE	899	$\simeq 1$				1.013	0.936	5.8	4.9	33
11	Krasn. II	$^{3}\text{He}+\text{PE}$	899	$\simeq 1$	—	—		1.031	0.953	20.3	4.9	92
12	Krasn. III	$^{3}\text{He}+\text{PE}$	899	$\simeq 1$		—		0.989	0.947	4.9	4.9	57
13	SRP I	Gd-LS	887	$\simeq 1$				0.987	0.952	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$		—		1.055	1.018	3.8	3.7	24
15	ROVNO88-1I	$^{3}\text{He}+\text{PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	3 He $+$ PE	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.2	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.2	25
19	ROVNO88-35	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

Neutron lifetime

#	result	Det. type	τ_n (s)	²³⁵ U	²³⁹ Pu	²³⁸ U	²⁴¹ Pu	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	³ He+H ₂ O	888.7	0.538	0.328	0.078	0.056	0.987	0.942	3.0	3.0	15
2	ROVNO91	$^{3}\text{He}+\text{H}_{2}\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
3	Bugey-3-I	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.988	0.946	4.8	4.8	15
4	Bugey-3-II	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.994	0.952	4.9	4.8	40
5	Bugey-3-III	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.915	0.876	14.1	4.8	95
6	Goesgen-I	³ He+LS	897	0.620	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
7	Goesgen-II	³ He+LS	897	0.584	0.298	0.068	0.050	1.045	0.992	6.5	6.0	45
8	Goesgen-II	³ He+LS	897	0.543	0.329	0.070	0.058	0.975	0.925	7.6	6.0	65
9	ILL	³ He+LS	889	$\simeq 1$	—		—	0.832	0.802	9.5	6.0	9
10	Krasn. I	³ He+PE	899	$\simeq 1$	—	—	_	1.013	0.936	5.8	4.9	33
11	Krasn. II	³ He+PE	899	$\simeq 1$	—		—	1.031	0.953	20.3	4.9	92
12	Krasn. III	³ He+PE	899	$\simeq 1$	—		—	0.989	0.947	4.9	4.9	57
13	SRP I	Gd-LS	887	$\simeq 1$	_	—	_	0.987	0.952	3.7	3.7	18
14	SRP II	Gd-LS	887	$\simeq 1$	—	-	—	1.055	1.018	3.8	3.7	24
15	ROVNO88-1I	³ He+PE	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
16	ROVNO88-2I	³ He+PE	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
17	ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.2	18
18	ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.2	25
19	ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

Averaged Fuel Composition

#	result	Det. type	τ_n (s)	285U	²³⁹ Pu	²³⁸ U	²⁴¹ Pu	old	new	err(%)	corr(%)	L(m)
1	Bugey-4	³ He+H ₂ O	888.7	0.538	0.328	0.078	0.056	0.987	0.942	3.0	3.0	15
2	ROVNO91	³ He+H ₂ O	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
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10	Krasn. I	³ He+PE	899	$\simeq 1$	—	—	_	1.013	0.936	5.8	4.9	33
11	Krasn. II	³ He+PE	899	$\simeq 1$	—		_	1.031	0.953	20.3	4.9	92
12	Krasn. III	³ He+PE	899	$\simeq 1$	—	-	—	0.989	0.947	4.9	4.9	57
13	SRP I	Gd-LS	887	$\simeq 1$	_	—	—	0.987	0.952	3.7	3.7	18
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19	ROVNO88-3S	Gd-LS	898.8	8,606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18

OBSERVED/PREDICTED ratios: OLD & NEW (this work)

-												
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19 Experimental Results Revisited (L<100m)

OBSERVED/PREDICTED ratios: OLD & NEW (this work)

result	Det. type	τ_n (s)	²³⁵ U	²³⁹ Pu	²³⁸ U	²⁴¹ Pu	old	new	err(%)	corr(%)	L(m)
Bugey-4	$^{3}\text{He}+\text{H}_{2}\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.942	3.0	3.0	15
ROVNO91	$^{3}\text{He}+\text{H}_{2}\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.940	3.9	3.0	18
Bugey-3-I	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.988	0.946	4.8	4.8	15
Bugey-3-II	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.994	0.952	4.9	4.8	40
Bugey-3-III	⁶ Li-LS	889	0.538	0.328	0.078	0.056	0.915	0.876	14.1	4.8	95
Goesgen-I	³ He+LS	897	0.620	0.274	0.074	0.042	1.018	0.966	6.5	6.0	38
Goesgen-II	³ He+LS	897	0.584	0.298	0.068	0.050	1.045	0.992	6.5	6.0	45
Goesgen-II	³ He+LS	897	0.543	0.329	0.070	0.058	0.975	0.925	7.6	6.0	65
ILL	³ He+LS	889	$\simeq 1$	—	-	—	0.832	0.802	9.5	6.0	9
Krasn. I	³ He+PE	899	$\simeq 1$	—	—	_	1.013	0.936	5.8	4.9	33
Krasn. II	³ He+PE	899	$\simeq 1$	—	_	_	1.031	0.953	20.3	4.9	92
Krasn. III	³ He+PE	899	$\simeq 1$	—	—	—	0.989	0.947	4.9	4.9	57
SRP I	Gd-LS	887	$\simeq 1$	—	—	—	0.987	0.952	3.7	3.7	18
SRP II	Gd-LS	887	$\simeq 1$	—	—	—	1.055	1.018	3.8	3.7	24
ROVNO88-1I	³ He+PE	898.8	0.607	0.277	0.074	0.042	0.969	0.917	6.9	6.9	18
ROVNO88-2I	³ He+PE	898.8	0.603	0.276	0.076	0.045	1.001	0.948	6.9	6.9	18
ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.972	7.8	7.2	18
ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.959	7.8	7.2	25
ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.938	7.2	7.2	18
	result Bugey-4 ROVNO91 Bugey-3-I Bugey-3-II Bugey-3-II Bugey-3-II Bugey-3-II Goesgen-I Goesgen-I Goesgen-II Goesgen-II Krasn. I Krasn. I Krasn. I Krasn. II Krasn. II Krasn. III SRP I SRP I SRP I SRP II ROVNO88-1I ROVNO88-1S ROVNO88-2S ROVNO88-3S	result Det. type Bugey-4 ³ He+H ₂ O ROVNO91 ³ He+H ₂ O Bugey-3-I ⁶ Li-LS Bugey-3-II ⁶ Li-LS Bugey-3-III ⁶ Li-LS Goesgen-I ³ He+LS Goesgen-II ³ He+LS Goesgen-II ³ He+LS Krasn. I ³ He+PE Krasn. II ³ He+PE Krasn. III ³ He+PE SRP I Gd-LS SRP II Gd-LS ROVNO88-1I ³ He+PE ROVNO88-1S Gd-LS ROVNO88-1S Gd-LS ROVNO88-2S Gd-LS ROVNO88-3S Gd-LS ROVNO88-3S Gd-LS	result Det. type τ_n (s) Bugey-4 ${}^{3}\text{He}+\text{H}_2\text{O}$ 888.7 ROVNO91 ${}^{3}\text{He}+\text{H}_2\text{O}$ 888.6 Bugey-3-I ${}^{6}\text{Li}\text{-LS}$ 889 Bugey-3-II ${}^{6}\text{Li}\text{-LS}$ 889 Bugey-3-III ${}^{6}\text{Li}\text{-LS}$ 889 Bugey-3-III ${}^{6}\text{Li}\text{-LS}$ 889 Goesgen-I ${}^{3}\text{He}+\text{LS}$ 897 Goesgen-II ${}^{3}\text{He}+\text{LS}$ 897 Goesgen-II ${}^{3}\text{He}+\text{LS}$ 897 ILL ${}^{3}\text{He}+\text{LS}$ 897 ILL ${}^{3}\text{He}+\text{PE}$ 899 Krasn. II ${}^{3}\text{He}+\text{PE}$ 899 Krasn. III ${}^{3}\text{He}+\text{PE}$ 899 SRP I Gd-LS 887 SRP II Gd-LS 887 ROVNO88-1I ${}^{3}\text{He}+\text{PE}$ 898.8 ROVNO88-1S Gd-LS 898.8 ROVNO88-2S Gd-LS 898.8 ROVNO88-3S Gd-LS 898.8 ROVNO88-	result Det. type τ_n (s) $^{2.5}$ U Bugey-4 3 He+H ₂ O 888.7 0.538 ROVNO91 3 He+H ₂ O 888.6 0.614 Bugey-3-I 6 Li-LS 889 0.538 Bugey-3-II 6 Li-LS 889 0.538 Bugey-3-III 6 Li-LS 889 0.538 Goesgen-I 3 He+LS 897 0.620 Goesgen-II 3 He+LS 897 0.543 ILL 3 He+LS 897 0.543 ILL 3 He+PE 899 $\simeq 1$ Krasn. II 3 He+PE 899 $\simeq 1$ SRP I Gd-LS 887 $\simeq 1$ ROVNO88-1I <	resultDet. type τ_n (s) 235 U 239 PuBugey-4 3 He+H ₂ O888.70.5380.328ROVNO91 3 He+H ₂ O888.60.6140.274Bugey-3-II 6 Li-LS8890.5380.328Bugey-3-III 6 Li-LS8890.5380.328Bugey-3-III 6 Li-LS8890.5380.328Bugey-3-III 6 Li-LS8890.5380.328Goesgen-I 3 He+LS8970.6200.274Goesgen-II 3 He+LS8970.5430.329ILL 3 He+LS8970.5430.329ILL 3 He+PE899 $\simeq 1$ Krasn. I 3 He+PE899 $\simeq 1$ Krasn. III 3 He+PE899 $\simeq 1$ SRP IGd-LS887 $\simeq 1$ SRP IIGd-LS887 $\simeq 1$ ROVNO88-1I 3 He+PE898.80.6030.276ROVNO88-1SGd-LS898.80.6060.277ROVNO88-2SGd-LS898.80.6060.274ROVNO88-3SGd-LS898.80.6060.274	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	resultDet. type τ_n (s) 235 U 235 U 238 U 241 PuoldBugey-4 3 He+H ₂ O888.70.5380.3280.0780.0560.987ROVNO91 3 He+H ₂ O888.60.6140.2740.0740.0380.985Bugey-3-I 6 Li-LS8890.5380.3280.0780.0560.994Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.994Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.994Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.994Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.915Goesgen-I 3 He+LS8970.6200.2740.0740.0421.018Goesgen-II 3 He+LS8970.5430.3290.0680.0501.045Goesgen-II 3 He+LS8970.5430.3290.0700.0580.975ILL 3 He+LS8970.5430.3290.0700.0580.975Krasn. II 3 He+PE899 $\simeq 1$ 1.031Krasn. III 3 He+PE899 $\simeq 1$ 0.989SRP IGd-LS887 $\simeq 1$ 0.987ROVNO88-1I 3 He+PE898.80.6070.2770.0740.0420.969ROVNO	resultDet. type τ_n (s) 233 U 239 Pu 238 U 241 PuoldnewBugey-4 3 He+H ₂ O888.70.5380.3280.0780.0560.9870.942ROVNO91 3 He+H ₂ O888.60.6140.2740.0740.0380.9850.940Bugey-3-II 6 Li-LS8890.5380.3280.0780.0560.9880.946Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.9940.952Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.9150.876Goesgen-I 3 He+LS8970.6200.2740.0740.0421.0180.966Goesgen-II 3 He+LS8970.5430.3290.0700.0580.9750.925ILL 3 He+LS8970.5430.3290.0700.0580.9750.925Krasn. I 3 He+PE899 $\simeq 1$ 1.0130.936Krasn. III 3 He+PE899 $\simeq 1$ 0.9870.952SRP IGd-LS887 $\simeq 1$ 0.9870.952SRP IIGd-LS887 $\simeq 1$ 0.9870.952SRP IIGd-LS887 $\simeq 1$ 1.0551.018ROVNO88-11 3 He+PE898.80.6070.2770.0740.0420.9690.917<	resultDet. type τ_n (s) 235 U 235 U 235 U 238 U 241 Puoldnewerr(%)Bugey-4 3 He+H ₂ O888.70.5380.3280.0780.0560.9870.9423.0ROVNO91 3 He+H ₂ O888.60.6140.2740.0740.0380.9850.9403.9Bugey-3-II 6 Li-LS8890.5380.3280.0780.0560.9880.9464.8Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.9940.9524.9Bugey-3-III 6 Li-LS8890.5380.3280.0780.0560.9150.87614.1Goesgen-I 3 He+LS8970.6200.2740.0740.0421.0180.9666.5Goesgen-II 3 He+LS8970.5840.2980.0680.0501.0450.9926.5Goesgen-II 3 He+LS8970.5430.3290.0700.0580.9750.9257.6ILL 3 He+LS8970.5430.3290.0700.0580.9750.9257.6ILL 3 He+LS899 $\simeq 1$ 0.8320.8029.5Krasn. II 3 He+PE899 $\simeq 1$ 1.0130.95320.3Krasn. III 3 He+PE899 $\simeq 1$ 0.9870.9523.7SRP IIGd-LS88	resultDet. type τ_n (s) 233 U 233 U 233 U 241 Puoldnewerr(%)corr(%)Bugey-4 ${}^{3}\text{He}+\text{H}_2\text{O}$ 888.70.5380.3280.0780.0560.9870.9423.03.0ROVNO91 ${}^{3}\text{He}+\text{H}_2\text{O}$ 888.60.6140.2740.0740.0380.9850.9403.93.0Bugey-3-II ${}^{6}\text{Li}\text{LS}$ 8890.5380.3280.0780.0560.9880.9464.84.8Bugey-3-III ${}^{6}\text{Li}\text{-LS}$ 8890.5380.3280.0780.0560.9940.9524.94.8Bugey-3-III ${}^{6}\text{Li}\text{-LS}$ 8890.5380.3280.0780.0560.9940.9524.94.8Bugey-3-III ${}^{6}\text{Li}\text{-LS}$ 8890.5380.3280.0760.0421.0180.9666.56.0Goesgen-I ${}^{3}\text{He}+\text{LS}$ 8970.6200.2740.0740.0421.0180.9666.56.0Goesgen-II ${}^{3}\text{He}+\text{LS}$ 8970.5430.3290.0700.0580.9750.9257.66.0ILL ${}^{3}\text{He}+\text{LS}$ 8970.5430.3290.0700.0580.9750.9257.66.0Krasn. I ${}^{3}\text{He}+\text{PE}$ 899 $\simeq 1$ 1.0130.95320.34.9Krasn. II ${}^{3}\text{He}+\text{PE}$ 899 $\simeq 1$ 0.9890.947



Our guiding principles: Be conservative - Be stable numerically (SRP case)

Reactor Antineutrino Sources

- 2% systematic on v-flux 100% correlated over ALL measurements
 - 1.8% corresponds to the normalization error on the ILL e- data
- Detector: Non-flux systematic error correlations across measurements:
 - Same experiment with same technology: 100% correlated
 - ILL shares 6% correlated error with Goesgen although detector slightly different. Rest of ILL error is uncorrelated.
 - Rovno88 integral measurements 100% corr. with Rovno 91 despite detector upgrade, but not with Rovno88 LS data
 - Rovno91 integral meas. 100% correlated with Bugey-4
 - Rovno88 integral meas. 50% correlated with Bugey-4

Experiments correlation matrix



- Main pink color comes from the 2% systematic on ILL β-spectra normalization uncertainty

The experiment block correlations come from identical detector, technology or neutrino source

The reactor antineutrino anomaly

0.6 0.7 0.8 0.9 1.1 1.2 1.3 1.4 1 t. ±0.008 ±0.068 ROVNO88_3S 0.938 18.2 m ROVNO88_2S ±0.009 ±0.075 -0.959 25.2 m ROVNO88_1S ±0.009 ±0.076 н 0.972 18.2 m ±0.009 ±0.065 **ROVNO88_21** 0.948 18.0 m ROVNO88_11 ±0.008 ±0.063 0.917 18.0 m ±0.010 ±0.038 SRP-II 23.8 m 1 1 A 1.019 SRP-I 18.2 m ±0.006 ±0.035 0.953 ±0.010 ±0.046 Krasnovarsk-III 0.954 ±0.190 ±0.053 Krasnovarsk-II 0.960 92.3 m • Krasnovarsk-I ±0.034 ±0.052 0.944 33.0 m ±0.059 ±0.048 ILL 8.76 m 0.801 ±0.043 ±0.055 Goesgen-III 0.924 65.0 m Goesgen-II 46.0 m ±0.024 ±0.059 0.991 -±0.023 ±0.058 Goesgen-I 0.966 38.0 m ±0.115 ±0.044 Bugey3 0.873 95.0 m ±0.009 ±0.047 Bugey3 0.948 -40.0 m Bugey-3/4 ±0.004 ±0.047 H 0.943 1 ROVNO91 ±0,023 ±0,028 0.940 ±0.000 ±0.028 Bugey-3/4 0.943 1 τ**"=885.7s** ±0.022 Average 0.943 0.6 0.7 0.8 0.9 1.1 1.2 1.3 1.4 1 $v_{
m Measured}$ / $v_{
m Expected}$

$$\chi^2 = \left(r - \overrightarrow{\mathbf{R}}\right)^T W^{-1} \left(r - \overrightarrow{\mathbf{R}}\right)$$

- Best fit : μ = 0.943
- Uncertainty : 0.023
- χ^2 = 19.6/19
- Deviation from unity
 - Naïve Gaussian : 99.3% C.L.
 - Toy MC: 98.6% C.L. (10⁶ trials)
- No hidden covariance
 - = 18% of Toy MC have χ^2_{min} <19.6

Are the ratios normally distributed?

- Our data points are ratios of Gaussians:
 - Numerator: measurement, Gaussian with stat & syst error, partially correlated
 - Denominator: common prediction, assumed to have Gaussian fluctuation of 2%
- Toy MC with correlated denominator with 2% fluctuation \rightarrow 10 6 events
 - Estimate weighted average R of 19 random points with correlations around 0.943.
 - P-value for (R >= 1): 1.4% (2.2σ)
 compared to naive Gaussian 2.4σ.
 - Our contours are reweighted by (2.2/2.4)²
 to take this slight non-normality into account

Hidden Covariance

• χ^2_{min} of data to straight line in the 18% quantile \rightarrow Data not incompatible with fluctuations





- 18/19 short baseline experiments <100m from a reactor observed a deficit of anti-v_e compared to the new prediction
- The effect is statistically significant at more 98.6%
- Effect partly due to re-evaluation of cross-section parameters, especially updated neutron lifetime, accounting for off equ. effect
- At least three alternatives:
 - Our conversion calculations are wrong. Anchorage at the ILL electron data is unchanged w.r old prediction
 - Bias in all short-baseline experiments near reactors : unlikely...
 - New physics at short baselines, explaining a deficit of anti-v_e:
 - Oscillation towards a 4th, sterile v ?
 - a 4th oscillation mode with θ_{new} and Δm^2_{new}





- Reactor at ILL with almost pure ²³⁵U, with compact core
- Detector 8.76(?) m from core. Any bias?
- Reanalysis in 1995 by part of the collaboration to account for overestimation of flux at ILL reactor by 10%... Affects the rate only



Large errors, but a striking pattern is seen by eye ?

Our ILL re-analysis (reproduce no-oscillation claim) CED

- 1981: Try to reproduce published contour
- 1995: Reproduce claim that global fit disfavors oscillation at 2σ
- How ? We add uncorrelated systematic in each bin until it's large enough Needed error : 11%, uncorrelated, in each bin.



Spectral shape analysis of Bugey-3



Combined Reactor Rate+Shape contours



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The Gallium Neutrino Anomaly

Based on PRD82 053005 (2010)

C. Giunti & M. Laveder



- 4 calibration runs with intense MCi neutrino sources:
 - 2 runs at Gallex with a ⁵¹Cr source (750 keV v_e emitter)
 - I run at SAGE with a ⁵¹Cr source
 - 1 run at SAGE with a 37 Ar source (810 keV v_e emitter)
 - All observed a deficit of neutrino interactions compared to the expected activity. Hint of oscillation ?
- Our analysis for Gallex & Sage:
 - Monte Carlo computing mean path lengths of neutrinos in Gallium tanks
 - **NEW** : Correlate the 2 Gallex runs together & the 2 SAGE runs together



The Gallium anomaly



- Effect reported in C. Giunti & M. Laveder in PRD82 053005 (2010)
- Significance reduced by additional correlations in our analysis
- No-oscillation hypothesis disfavored at 97.7% C.L.

 $\hat{\mathcal{A}}$





Implication for θ_{13}

Implication for θ_{13} at 1-2 km baselines

• The choice of normalization is crucial for reactor experiments looking for θ_{13} without near detector

 $\sigma_{f}^{pred,new}$: new prediction of the antineutrino fluxes



 σ_{f}^{ano} : experimental cross section (best fitted mean averaged)



The Normalization Dilemma

- Experiments with baselines > 500 m
- How do you normalize the expected flux, knowing the fuel composition?
- If near + far detector, not an issue anymore





CHOOZ reanalysis

- The choice of σ_f changes the limit on θ_{13}
- Chooz original choice was σ_f^{exp} from Bugey-4 with low error
- If $\sigma_{f}^{pred,new}$ is used, limit is worse by factor of 2
- If σ_f^{ano} is used with 2.7%, we obtain the original limit
 - \rightarrow But which error should we associate to σ_{f}^{ano} (burnup up error?)



Reanalysis of KamLAND's 2010 results

arXiv:1009.4771v2 [hep-ex]

Systematics

	Detector-related	(%)	Reactor-related (%)				
$\overline{\Delta m^2_{21}}$	Energy scale	1.8 / 1.8	$\overline{\nu}_e$ -spectra [<u>31</u>]	0.6/0.6			
Rate	Fiducial volume	1.8/2.5	$\overline{\nu}_e$ -spectra	2.4/2.4			
	Energy scale	1.1 / 1.3	Reactor power	2.1 / 2.1			
	$L_{cut}(E_{\rm p})$ eff.	0.7 / 0.8	Fuel composition	1.0 / 1.0			
	Cross section	0.2/0.2	Long-lived nuclei	0.3 / 0.4			
	Total	2.3/3.0	Total	3.3/3.4			

Reproduced KamLAND spectra within 1% in [1-6] MeV range





CEA DSM Irfu T. Lasserre

CHOOZ and KamLAND combined limit on θ_{13}



Our interpretation (different from Arxiv:1103:0734 for KamLAND-σ_f^{pred,new}, T. Schewtz's talk)

- No hint on θ_{13} >0 from reactor experiments : sin²(2 θ_{13})<0.11 (90%C.L., 1dof)
- CHOOZ 90 % CL limit stays identical to Eur. Phys. J. C27, 331-374 (2003)
- Multi-detector experiments are not affected

Need for new experimental inputs !



CEA DOMINU T. Lasserre

Conclusion and perspectives

New Reactor Antineutrino Anomaly Discovered

- Experimental bias to be deeply investigated
- New physics hypothesis tested: 4th neutrino
 - no-oscillation hypothesis disfavored at 99.8%

Clear experimental confirmation / infirmation is needed:

L/E ≈ few m/MeV or km/GeV

New Experiment at Reactor

- Short Baseline Shape + Rate Analysis
- Mci neutrino generator in/close to a large liquid scintillator

IikeSNO+, Borexino, KamLAND

- New neutrino beam experiment probing for electron GeV neutrino disappearance at 100 m & 1 km
 - C. Rubbias's proposal at CERN-PS
 - Fermilab workshop in May



NUCIFER in Saclay

- Osiris-Saclay: Core Size: 57x57x60 cm
- Nucifer Detector Size : 1.2x0.7m (850I)
- Baseline distribution
 - <L>=7.0 m, σ =0.3 m \rightarrow eV² oscillations are not washed out
- Folding Nucifer Geant4 Monte Carlo detector response
- Δm² = 2.4 eV² & sin²(2θ)=0.15
- No backgrounds. Thus to be taken with a grain of salt ...



Such pattern could not be seen at Bugey-3 (extended core & 14 m baselin

³⁷Ar Neutrino Generator Experiment

- A strong 1 Mci v source in the middle of a large LS detector
- Elastic scattering on electrons (few 10000 evts, 150 days, >250 keV)
- A good resolution in position (15cm) Low Backgrounds

