

A Decade of WHIM Searches: Where do we Stand and Where do we Go?

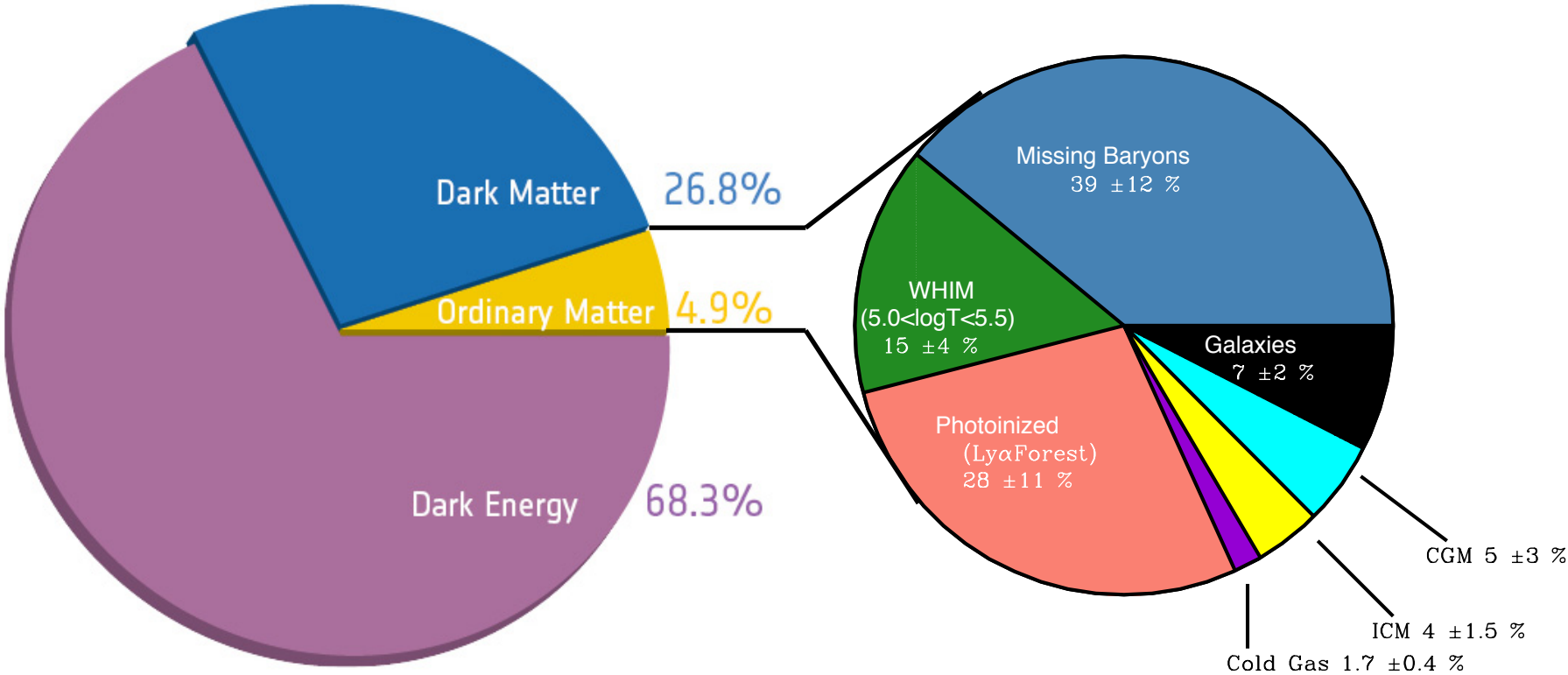
Nicastro, Krongold, Mathur and Elvis

<https://arxiv.org/abs/1611.03722>

Jean-Baptiste Melin

February 21st 2017

Baryon budget



Planck 2015
(observation @ z=1100)

This paper
(observation @ z=0)

Between $z=1100$ and $z=3$



Gunn Peterson relation for the Ly α forest

$$\tau = \frac{0.091}{\Gamma_{-12}} \left(\frac{\Omega_b h_{50}^2}{0.05} \right)^2 \left[\frac{5.2 H(0)}{h_{50} H(z)} \right] T_4^{-0.7} \left(\frac{\rho}{\bar{\rho}} \right)^\alpha \left(\frac{1+z}{1+2} \right)^6 \\ \times \left[1 + \frac{dv_{pec}}{H(z)dr} \right]^{-1}.$$

Rauch 1998

$\tau(z)$ from observations

Ω_b^2/Γ_{-12} from comparison with simulations

 $\Omega_b h^2 > 0.017$ (Rauch et al 1997).  No missing baryons @ $z=3$

Between $z=1100$ and $z=3$

Gunn Peterson relation for the Ly α forest

$$\tau = \frac{0.091}{\Gamma_{-12}} \left(\frac{\Omega_b h_{50}^2}{0.05} \right)^2 \left[\frac{5.2 H(0)}{h_{50} H(z)} \right] T_4^{-0.7} \left(\frac{\rho}{\bar{\rho}} \right)^\alpha \left(\frac{1+z}{1+2} \right)^6 \\ \times \left[1 + \frac{dv_{pec}}{H(z)dr} \right]^{-1}.$$

Rauch 1998

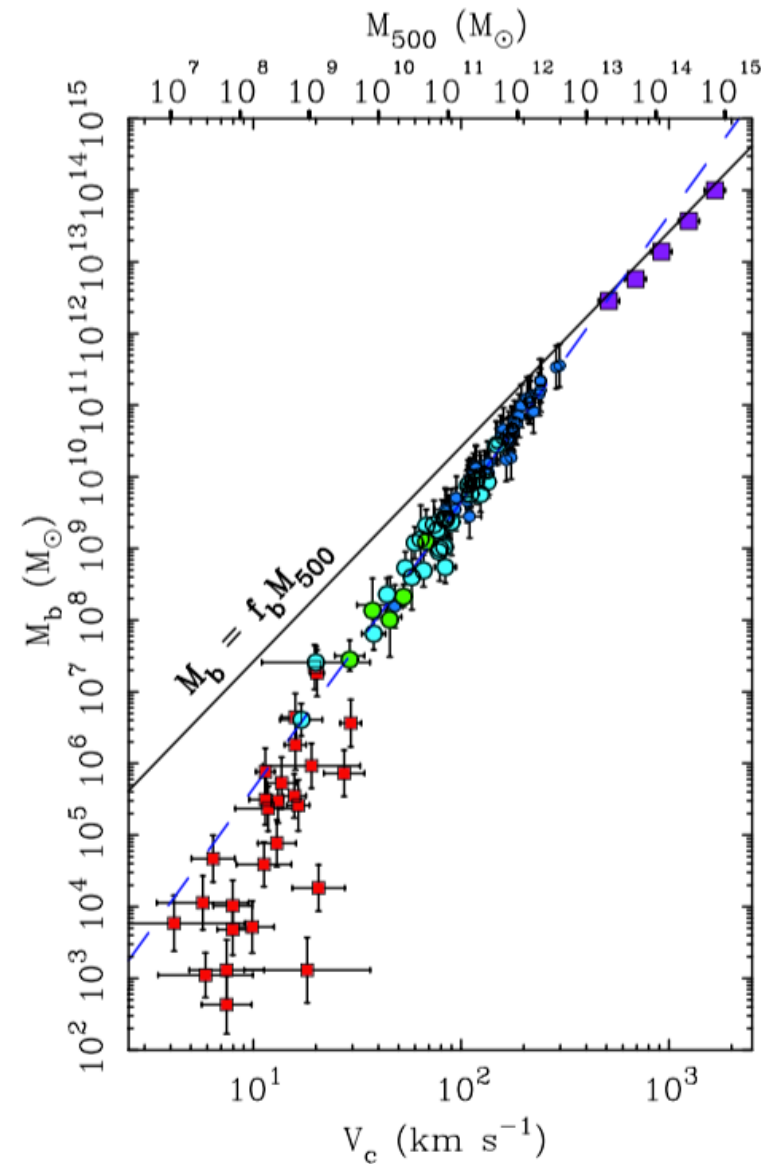
$\tau(z)$ from observations

Ω_b^2/Γ_{-12} from comparison with simulations

→ $\Omega_b h^2 > 0.017$ (Rauch et al 1997). **→** No missing baryons @ $z=3$

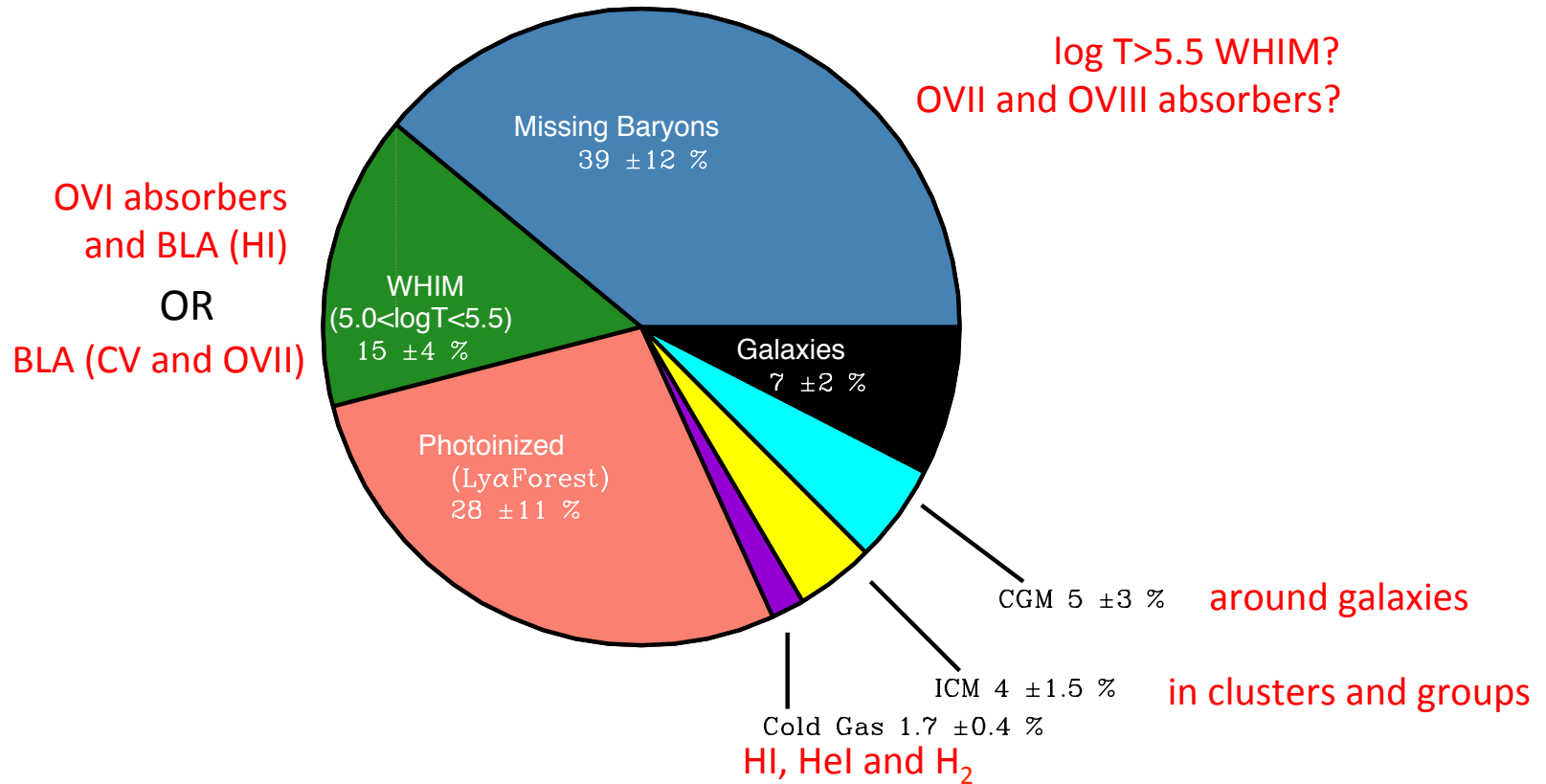
→ $\Omega_b h^2 = 0.02230 \pm 0.00014$ (Planck 2015) Note: $0.017/0.02230=0.76$

Missing baryons



McGaugh et al. 2010

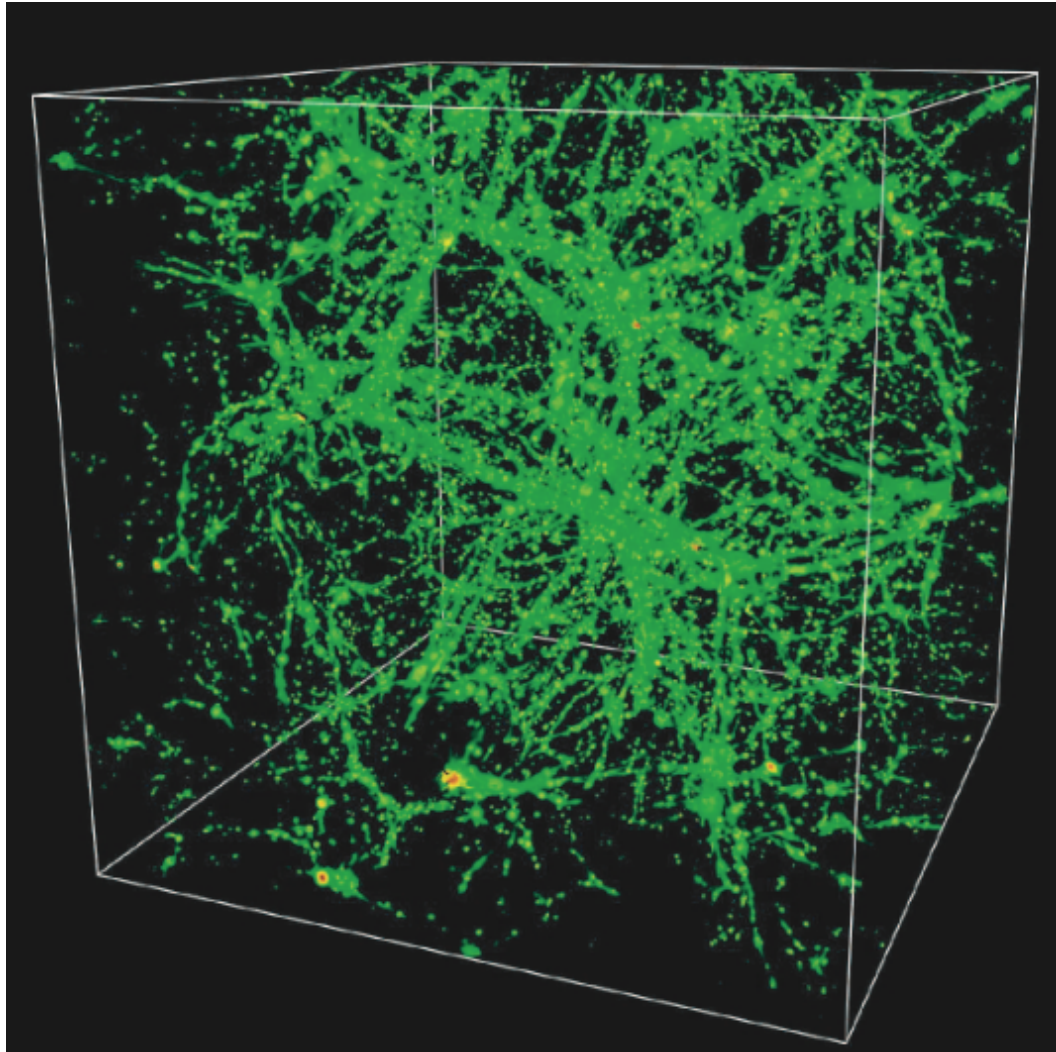
Between $z=3$ and $z=0$



This paper
(observation @ $z=0$)

See also Shull, Smith & Danforth 2006

WHIM



$$T = 10^5 - 10^7 \text{ K}$$
$$n_b = 10^{-6} - 10^{-4} \text{ cm}^{-3}$$

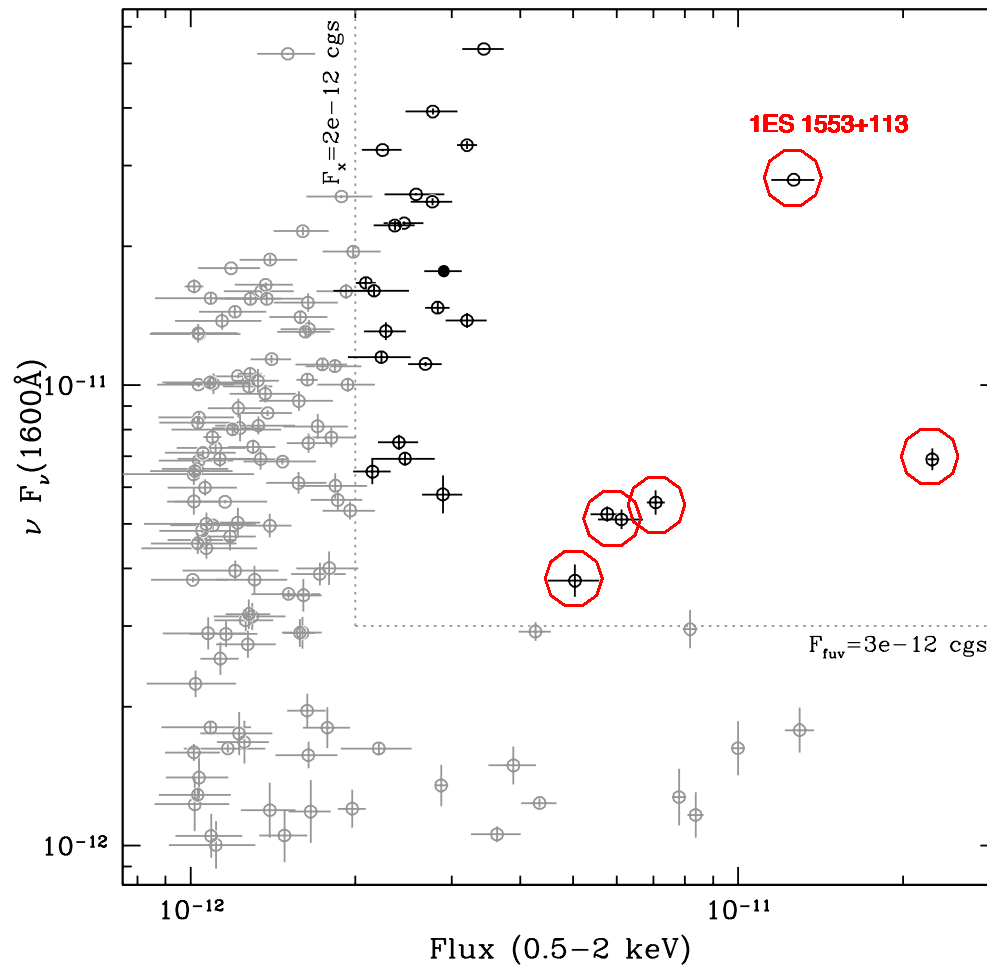


Absorption lines
in X-ray spectra
of distant astro sources

(*Chandra*, *XMM*)

Cen & Ostriker 2006

Targets



OVII He α absorbers


Only a few $\text{EW} \sim 3\text{m\AA}$
XMM spectro $\Delta\lambda \sim 70\text{m\AA}$

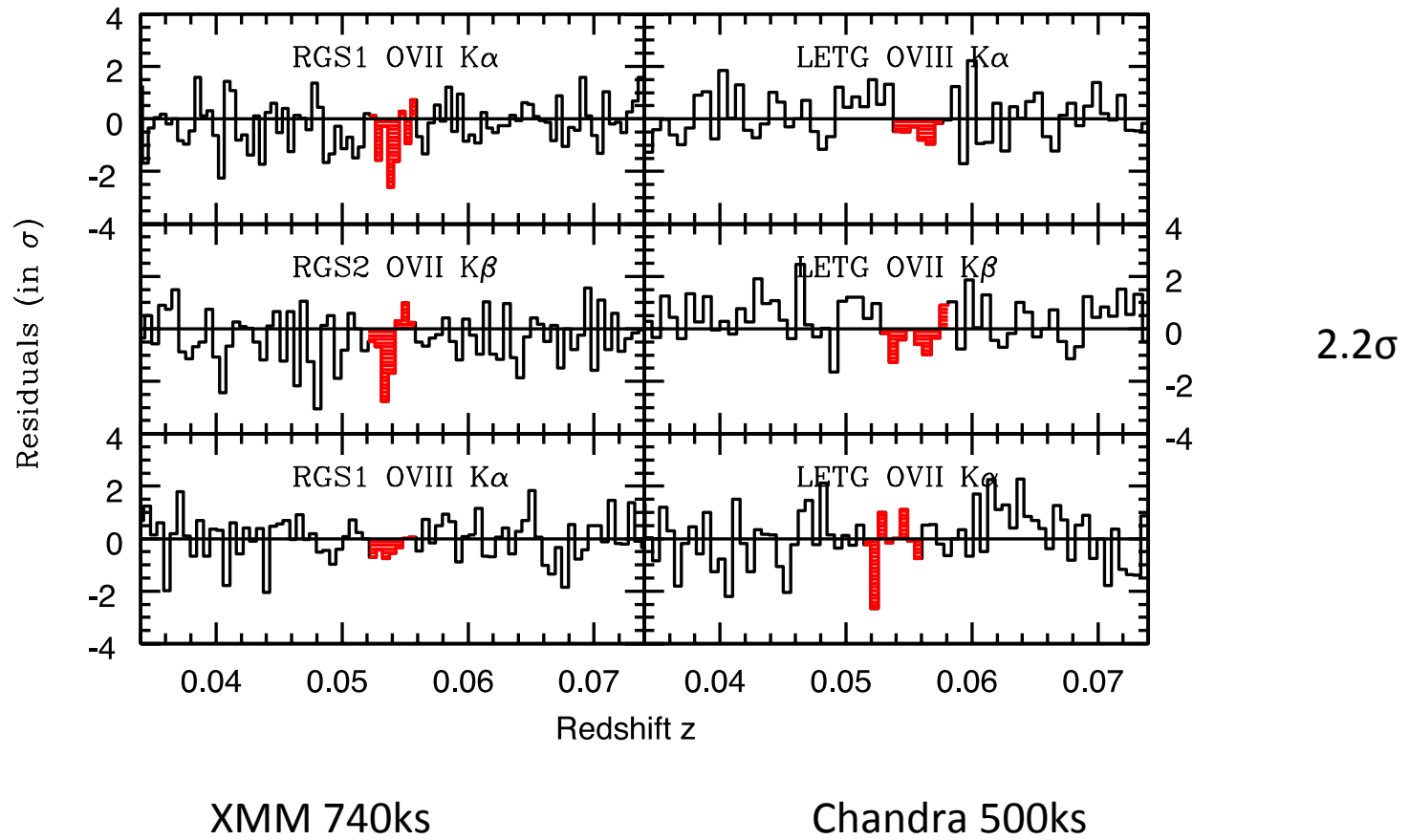
$\text{EW}/\Delta\lambda \sim 4\%$

syst. $\sim 2\%$

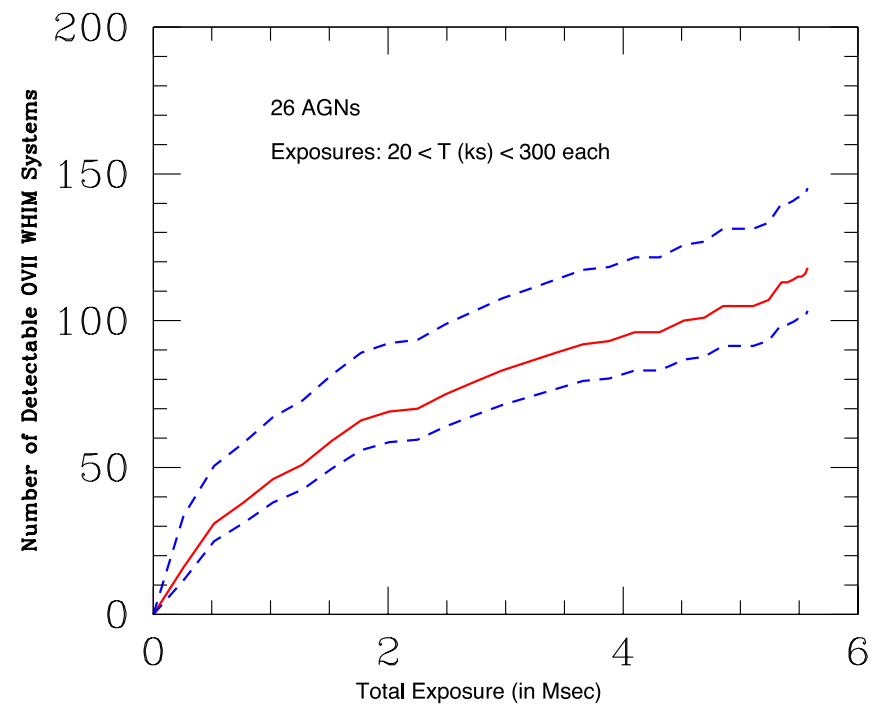
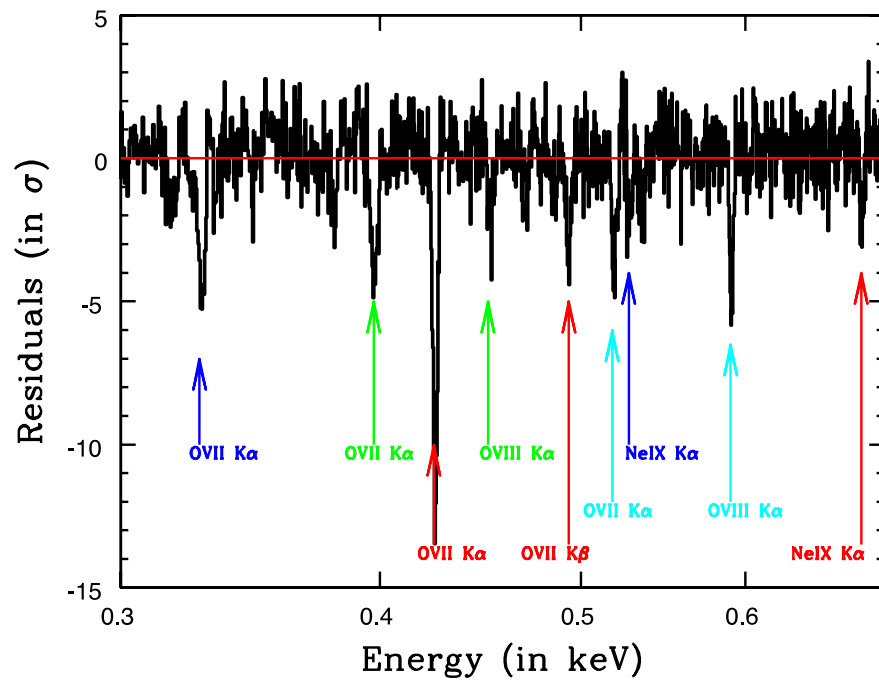
➡ Long observing time
➡ Very bright targets

Sensitivity

Blazar 1ES 1553+113  1.6 Ms observation with XMM



Future: Athena



Conclusions

- Firm detection of $\log T > 5.5$ WHIM not done yet despite strong efforts and controversies
- Second half of the observation of 1 ES 1553+113 this year
- But not sufficient (variation from line-of-sight to line-of-sight)
- XMM-Newton Legacy program (11Ms on 5 other targets) + HST (150 additional orbits)
- Future: HST and Athena