

# Study of final-state interactions of protons in neutrino-nucleus scattering with INCL+ABLA and NuWro cascade models

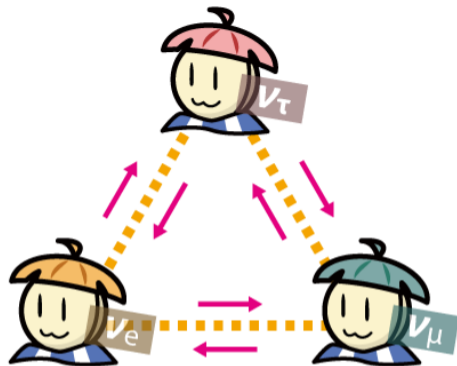
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March 13, 2022



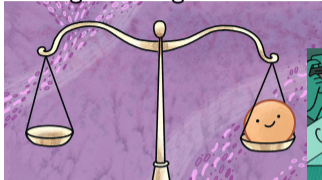
# Outlook

- 1 Introduction
- 2 Liège Intranuclear Cascade model (INCL)
- 3 Results
  - Neutrino energy reconstruction
  - Leading proton kinematics
  - Experimental observables

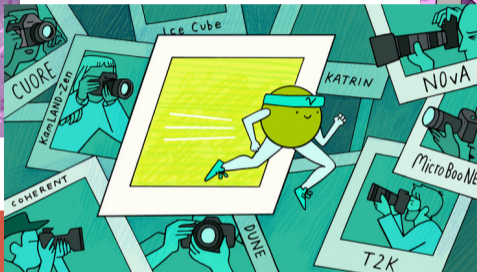
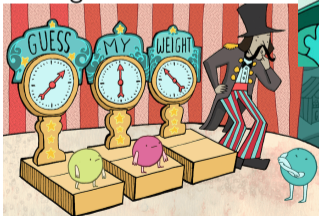


# Mysterious neutrinos

Mass generating mechanism?



$\nu$  weight?



Oscillation parameters?

Mass hierarchy?

CP violation in lepton sector?

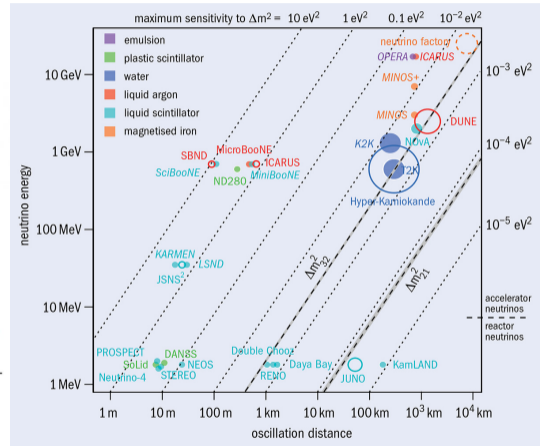
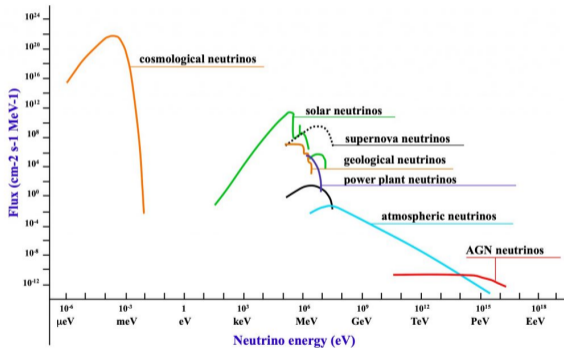
Sterile neutrinos?

...



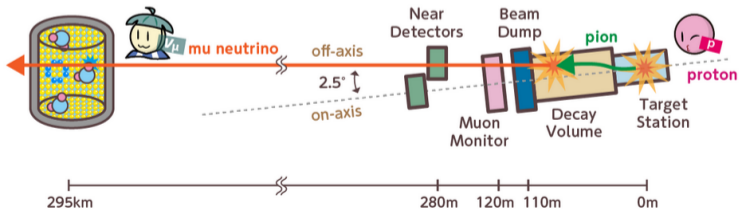
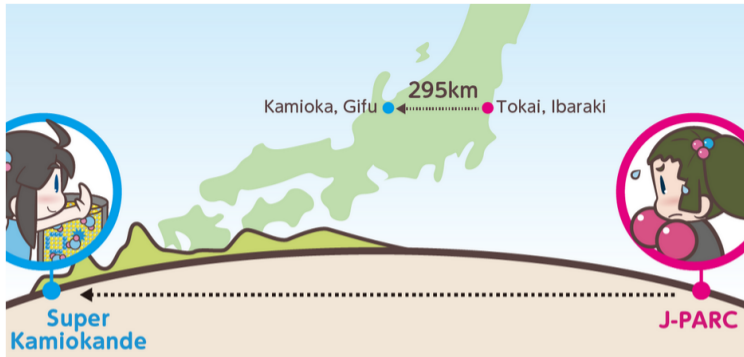
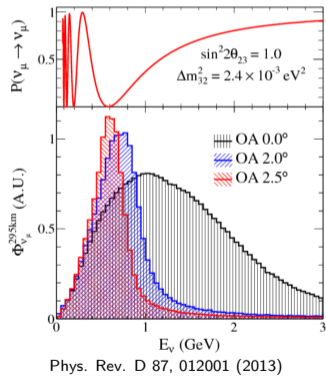
[symmetrymagazine.org](http://symmetrymagazine.org)

# Neutrino studies



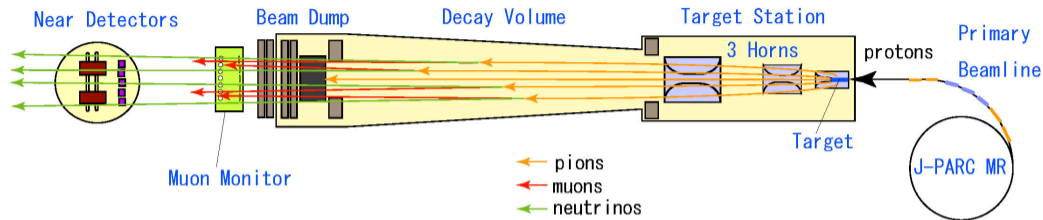
# T2K: design

The **off-axis** configuration allows a better **focus** of  $\nu$  energy at the **peak** of  $\nu$  oscillation probability.



Akimoto, Yuki @ higgstan

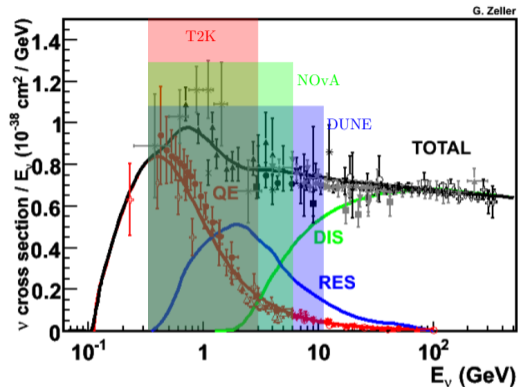
# T2K: $\nu$ rate measurement



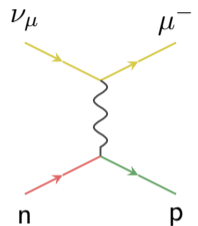
$$R_{FD}^{\nu'}(E_\nu) = \underbrace{\Phi^\nu(E_\nu)}_{\nu \text{ flux}} \otimes \underbrace{P_{osc}^{\nu \rightarrow \nu'}(E_\nu)}_{\text{oscillation probability}} \otimes \underbrace{\sigma^{\nu'}(E_\nu)}_{\nu \text{ cross-section}} \otimes \underbrace{\varepsilon(E_\nu)}_{\text{detector acceptance}}$$

In order to get accurate neutrino rate, we need to **measure** neutrino energy **precisely**.

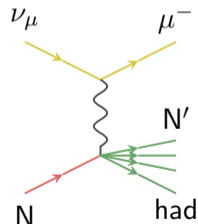
# $\nu$ energy reconstruction



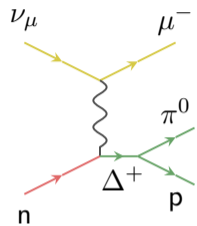
**CCQE** is the most important channel for T2K, and in this presentation we will focus on it



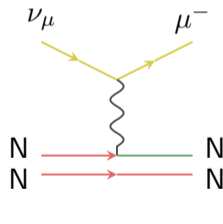
**CCQE**



**DIS**



**RES**



**2p2h**

## $\nu$ energy reconstruction

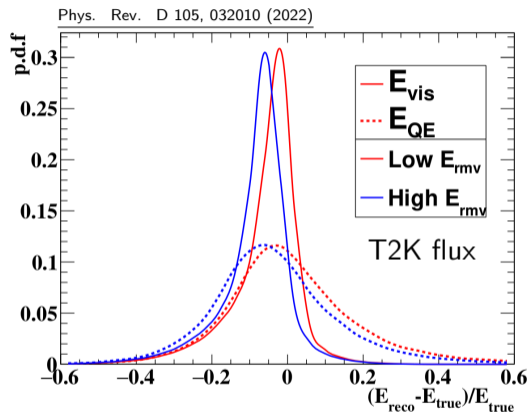
Energy reconstruction using only muon kinematics (works well for **quasi-elastic reaction**):

$$E_{\nu}^{QE} = \frac{m_p^2 - (m_n - E_B)^2 - m_{\mu}^2 + 2(m_n - E_B)E_{\mu}}{2((m_n - E_B) - E_{\mu} + p_{\mu} \cos \theta_{\mu})}$$

Energy reconstruction using **muon and kinetic energy of the nucleon**:

$$E_{\nu}^{vis} = E_{\mu} + T_N$$

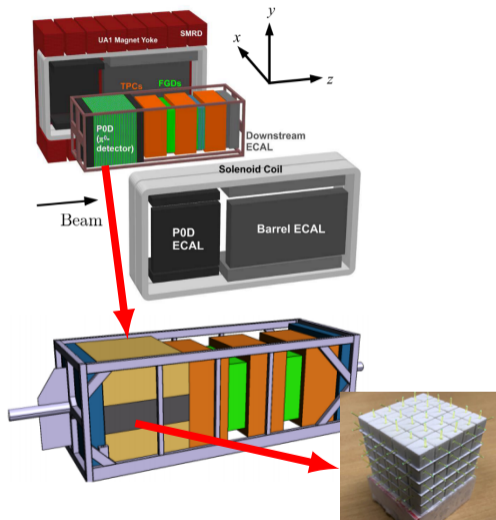
The community is moving from **inclusive** measurements (using only the outgoing lepton) to **exclusive** ones where also the **hardonic part**.



$E_{\nu}^{vis}$ , dashed line — QE formula  
solid line —  $\mu + N$  formula



# Upgrade of ND280 detector

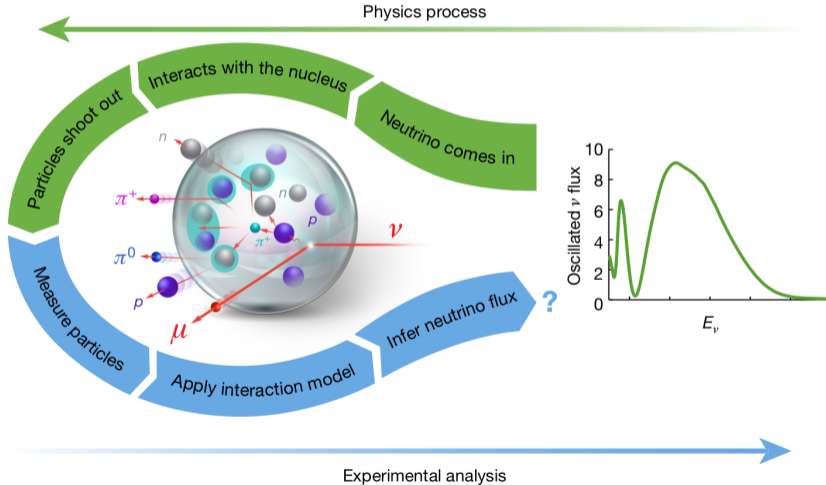


- better reconstruction efficiency of **high angle muons**: **angular acceptance** close to SK  $4\pi$  acceptance
- ability to detect **neutrons**
- lower **proton** momentum threshold
- $\nu$  **energy reconstruction** using also hadronic part

Upgrade will give us access to a **new phase space** with which we can constrain better the **nuclear models** used in the oscillation analysis. IRFU has a major contribution to the T2K upgrade: new HA-TPC with the resistive MicroMegas technology.

# Importance of nuclear effects

We need **not only** a better detector, but also better **modelling** of the neutrino-nucleus interactions, e.g. improved Monte-Carlo generators!



# My work

- I use INCL use for the first time to predict **exclusive final states of neutrino interactions**
- INCL model features various **novelties**, including the **production of nuclear clusters** (e.g., deuterons,  $\alpha$  particles)
- De-excitation process (that we model with ABLA) was **never modelled and studied** in neutrino physics before

## Study of final-state interactions of protons in neutrino-nucleus scattering with INCL and NuWro cascade models

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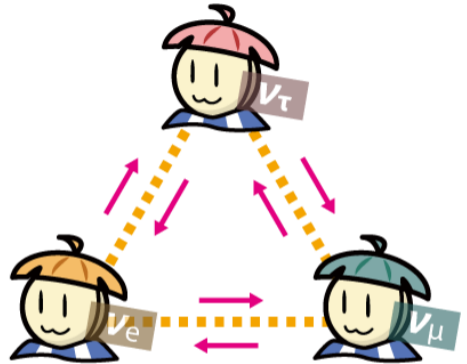
<sup>9</sup>Institute for Nuclear Research of the Russian Academy of Sciences, Moscow, Russia

## 1 Introduction

## 2 Liège Intranuclear Cascade model (INCL)

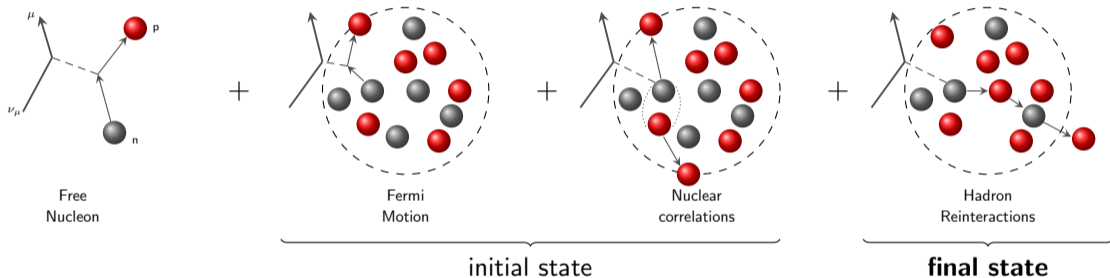
## 3 Results

- Neutrino energy reconstruction
- Leading proton kinematics
- Experimental observables



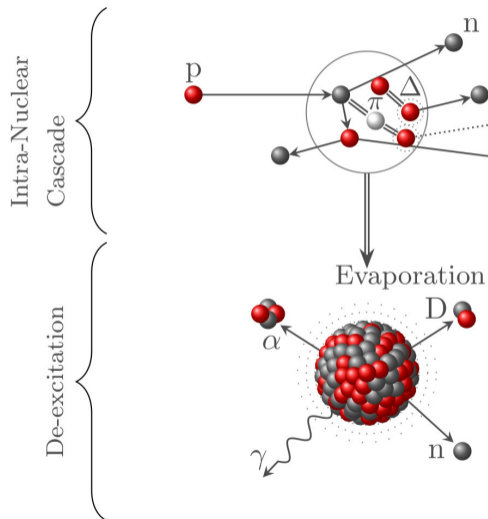
# Simulation of nuclear effects


Simulation allows us to **factorize** neutrino interaction and simulate in few stages:



I will focus on **CCQE**  $\nu$  reaction channel and the **Final State Interactions (FSI)** that are described by **cascade models**.

# Liège Intra Nuclear Cascade



**Projectiles:** baryons (nucleons,  $\Lambda$ ,  $\Sigma$ ), mesons (pions and Kaons) or light nuclei ( $A \leq 18$ ). **No neutrinos** yet! We use neutrino vertex from  **NuWro** (widely used  $\nu$ -nucleus MC generator).

**De-excitation:** ABLA, SMM, GEMINI

We will use **ABLA**: proved to work for the **light nuclei** (Phys. J. Plus 130, 153 (2015))

**Flexible tool:** has been implemented in GEANT4 and GENIE

# Cascade ingredients

## Potential

Each nucleon in the nucleus has its **position and momentum** and moves **freely** in a square potential well. Nuclear model is essentially **classical**, with some additional ingredients to mimic quantum effects.

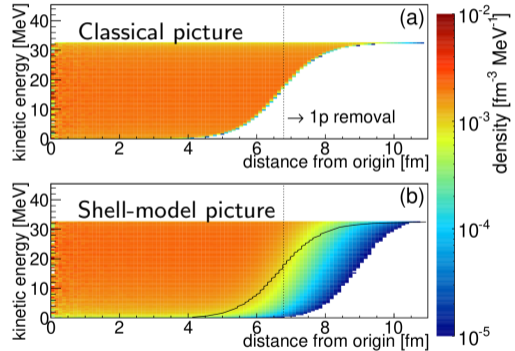
## Pauli Blocking

the phase-space below Fermi momentum is occupied and restricted

## Events inside cascade

- decay/collision
- reflection/transmission with probability to **leave the nucleus as a nuclear cluster**

Space-kinetic-energy density of protons in  $^{208}\text{Pb}$



Phys.Rev.C 91, 034602 (2021)

# Cascade ingredients

## Potential

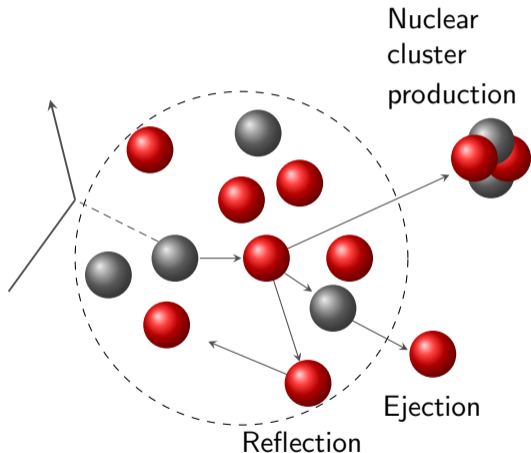
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## Pauli Blocking

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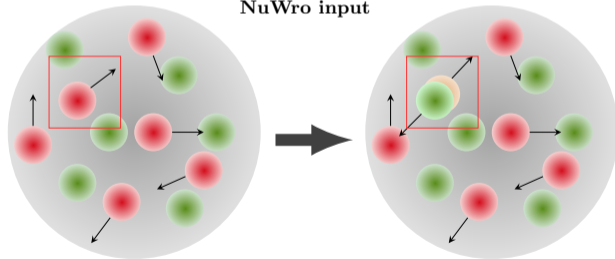
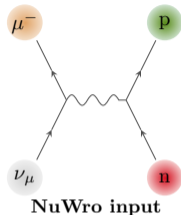
## Events inside cascade

- decay/collision
- reflection/transmission with probability to **leave the nucleus as a nuclear cluster**





# Using INCL with NuWro input



INCL nucleus

I use **NuWro sample** (one of the many generators on the market) to model  $\nu$  **CCQE** reaction on **carbon** target. I want to compare **FSI cascades** modelled by **INCL** and **NuWro**.

But there is no neutrino vertex implemented in INCL, so:

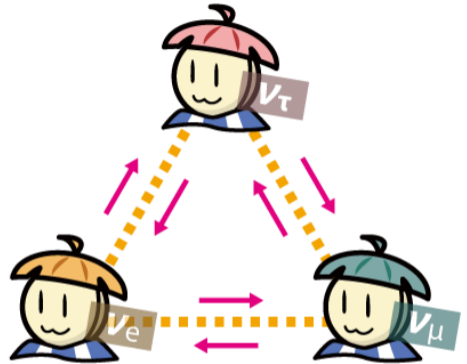
- I choose in INCL the neutron with the momentum closest to the NuWro neutron (on which  $\nu$  reacted)
- change this neutron to the reaction products:  $\mu$  and proton

## 1 Introduction

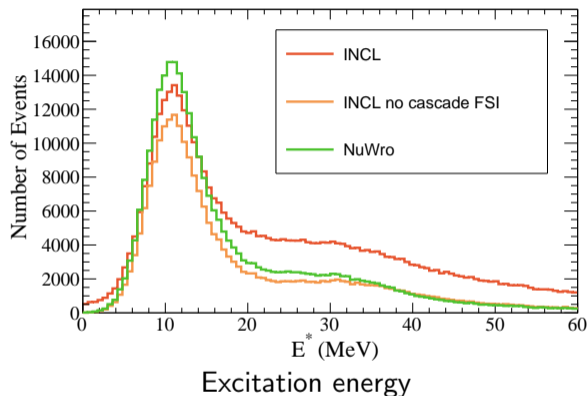
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# Excitation energy

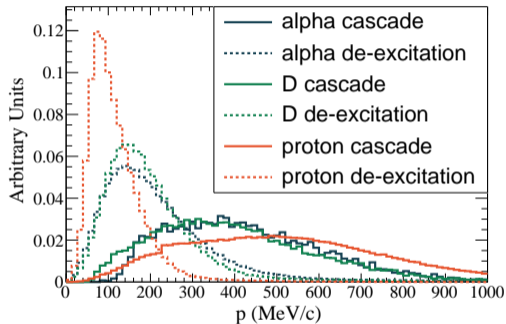
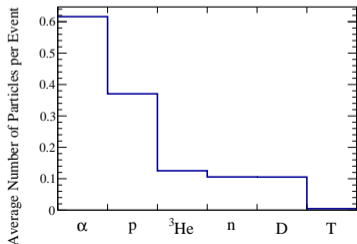
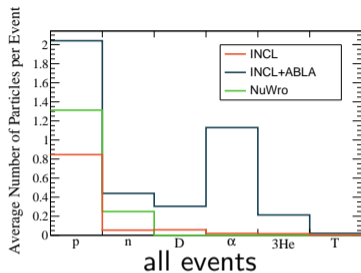


$$E = E_\nu + {}_6^{12}\text{M} - \sum_i E_i, \quad p = p_\nu - \sum_i p_i$$
$$E^* = \sqrt{E^2 - p^2} - M_{rem}$$

We have excitation energy even **without FSI** due to fundamental  $\nu$  interaction and it will be dealt with ABLA producing **de-excitation particles** ('binding energy' does not stay in the nucleus, it becomes observable in the final state)

In **presence of FSI** we produce additional excitation energy which is different for INCL and NuWro (INCL tend to have stronger FSI and produces more excitation in FSI than NuWro)

# Production of the nuclear clusters ( $\alpha$ , deuterons, tritons...)



Momentum of nuclear clusters produced during the cascade and de-excitation

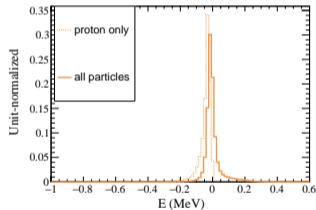
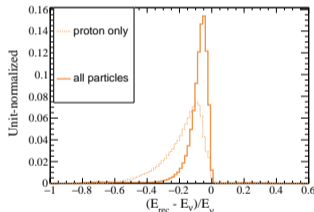
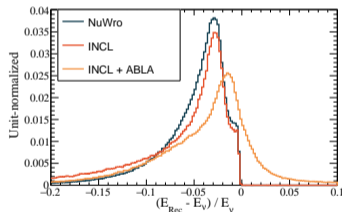
# Neutrino energy reconstruction

proton only:

$$E_{rec} = E_{\mu} + T_p$$

all particles:

$$E_{rec} = E_{\mu} + \sum_i T_i$$



"all particles" reconstruction

INCL+ABLA cascade FSI

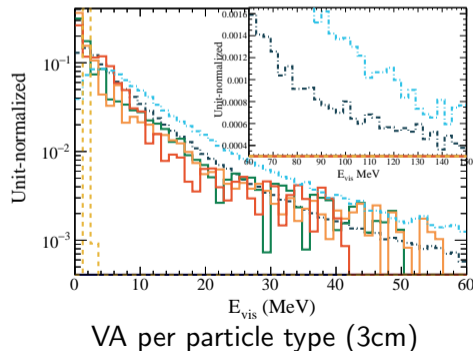
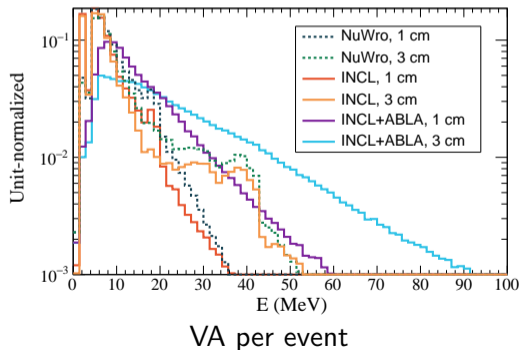
INCL+ABLA no cascade FSI

Explanation of  $E_{rec} > E_{\nu}$  in backup

# Vertex Activity

We define vertex activity as **visible energy deposited** (with Birks correction) in a 1(3) cm sphere **around** the neutrino interaction vertex. We distinguish **two types** of VA:

- **per event**: sum of energy deposits of all particles produced in a given event
- **per particle type**: energy deposit separately for different particle types

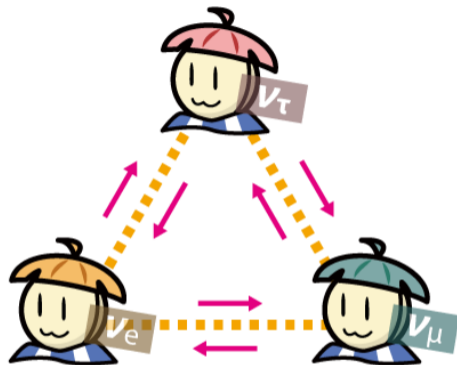


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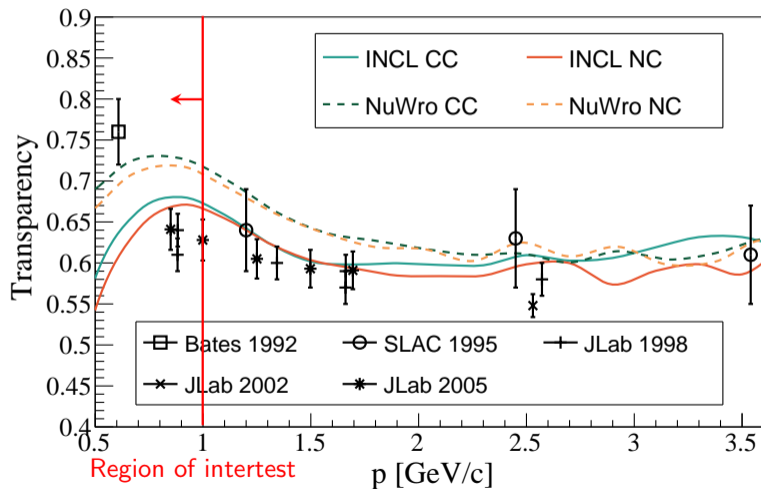
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# Nuclear transparency

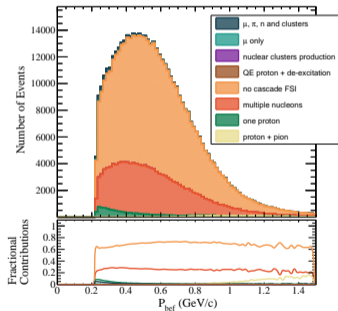


Here transparency is a probability for the **leading proton** to leave the nucleus "untouched".

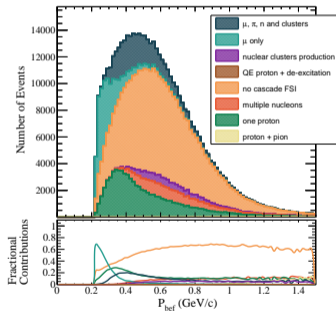


# Proton momentum before FSI

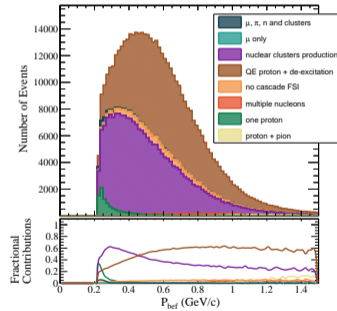
INCL cascade features a significant fraction of **events without a proton** in the final state. With de-excitation, we almost **do not have** events with no proton in the final state.



NuWro



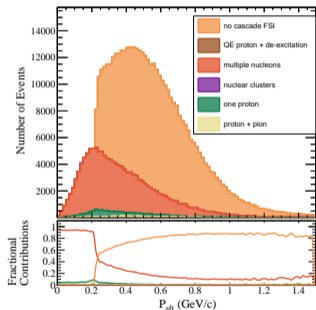
INCL



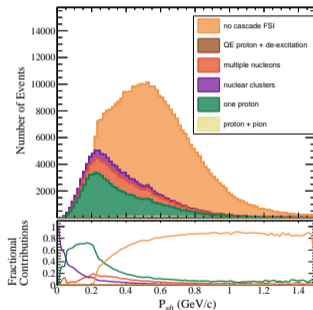
INCL + ABLA

# Proton momentum after FSI

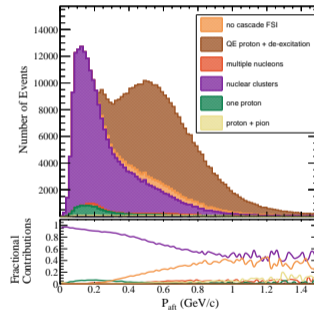
We "bring back" events from 0 proton channel, they **contribute to the low momentum** region of the distribution.



NuWro



INCL



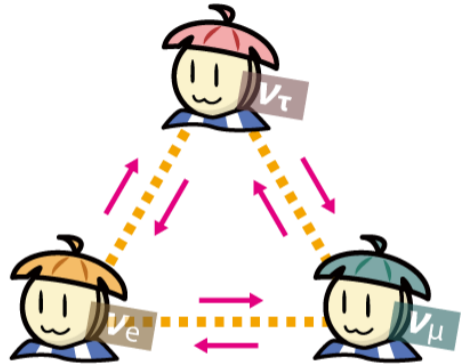
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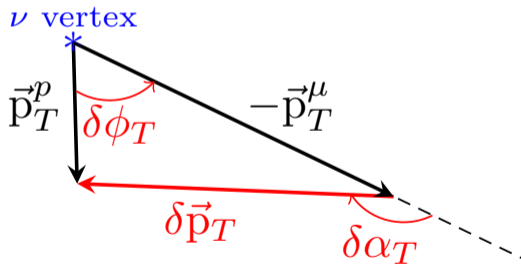
# Variables of interest

We use **Single Transverse Variables (STV)** that allow to disentangle different effects for better FSI estimation. STV are **observable** and **measurable**.

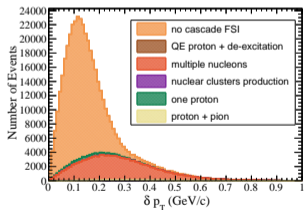
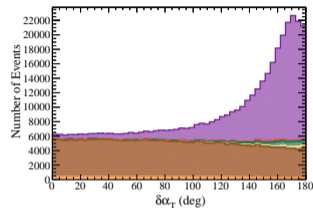
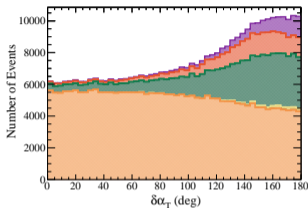
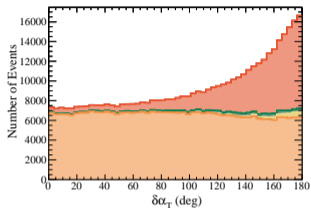
sensitive to FSI:  $\delta\alpha_T = \arccos \frac{-\vec{k}'_T \cdot \delta\vec{p}'_T}{k'_T \cdot \delta p'_T}$

sensitive to Fermi Motion:

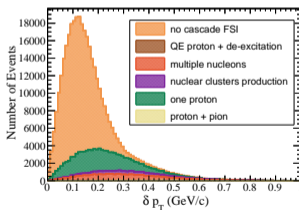
$$\delta\vec{p}_T = \vec{p}_T^{\vec{p}} + \vec{p}_T^{\vec{\mu}} = \vec{p}_T^{\vec{n}}$$



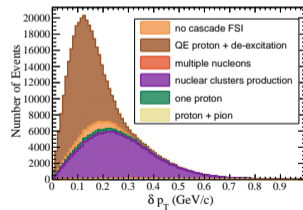
# STV



NuWro



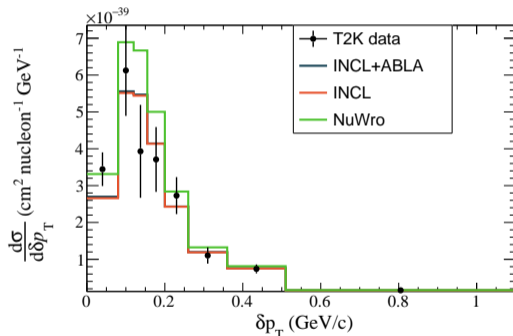
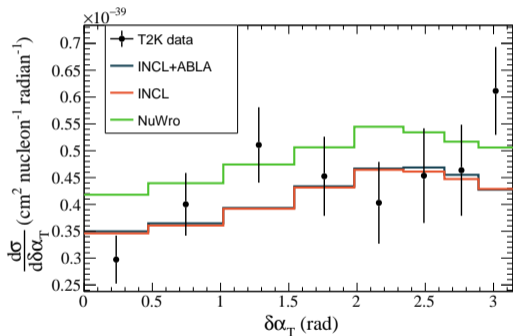
INCL



INCL + ABLA

# Comparison to T2K data: INCL + ABLA

Current detector **threshold is too large**, so we **cannot really see the effect of de-excitation**.

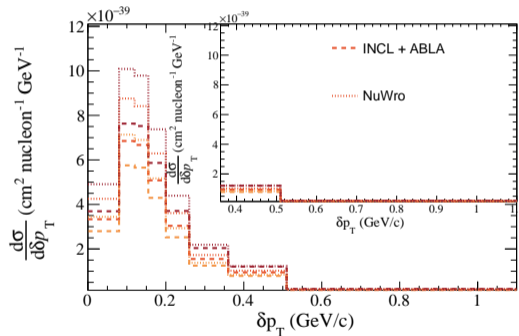
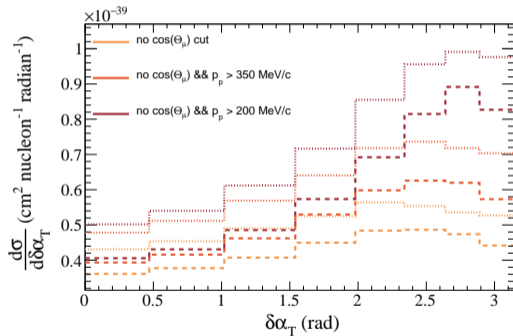


**Cuts (MeV):**  $p_\mu > 250$ ;  $450 < p_p < 1000$ ;  $\cos(\Theta_\mu) > -0.6$ ;  $\cos(\Theta_p) > 0.4$

T2K data taken from [Phys.Rev. D, 98 032003 \(2018\)](#)

# What if we change cuts to mimic better sensitivity?

We start to distinguish models from  $p_p > 200$  MeV/c



# Conclusion

We compared the simulation of the final-state interactions between the **NuWro** and **INCL** cascade models in CCQE events. We coupled INCL cascade to the ABLA de-excitation model.

- "transparent events" are **not** transparent: nuclear clusters may be produced
- INCL+ABLA simulation features **massive difference** in nucleon kinematics in comparison to NuWro (and the other similar generator used in neutrino scattering)
- An essential novelty of this study is the **simulation of nuclear cluster production** during cascade and de-excitation. It is important for the understanding of the **vertex activity** and calorimetric method of  $\nu$  **energy reconstruction**



# Prospects

**New generation** of detectors starts to use the **exclusive FSI**

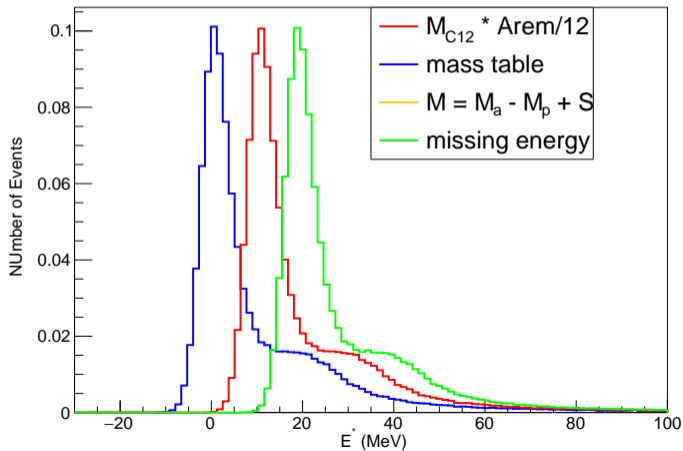
- ND280 upgrade of T2K to improve the detector threshold
- SK-Gd project: add gadolinium to SK to enhance the neutron detection efficiency
- The LAr program in USA is dedicated to measuring all the particles in the final state

The **de-excitation study** will be published soon. There is still plenty of work to be done: **neutron secondary interaction** studies,  $\bar{\nu}$  simulation and **pion FSI**.

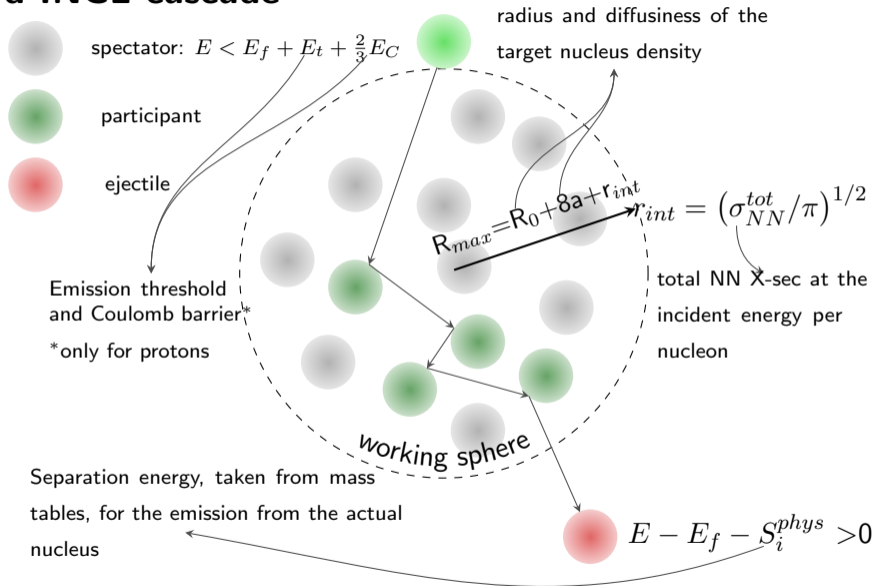
# BACKUP

# Why sometimes $E_{rec} > E_{\nu}$

NuWro, SF, excitation energy calculation

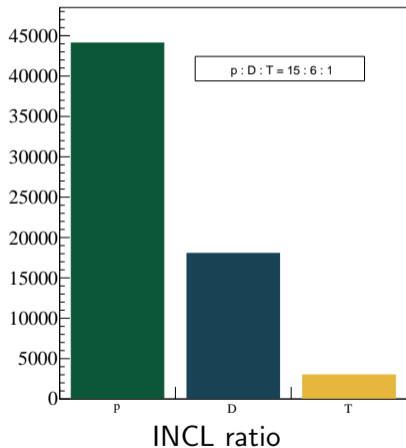


# Standard INCL cascade



# Nuclear clusters emission check

$^{12}\text{C}$  bombarded by 175 MeV neutrons



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## ARTICLE

### Production of protons, deuterons, and tritons from carbon bombarded by 175 MeV quasi mono-energetic neutrons

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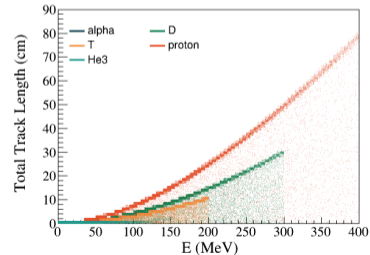
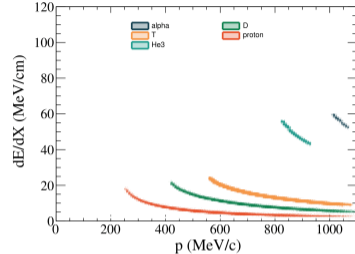
<sup>4</sup>Fast Neutron Research Facility, Chiang Mai University, P.O.Box 217, Chiang Mai 50200, Thailand

<sup>5</sup>Los Alamos National Laboratory, Los Alamos, NM 87545, USA

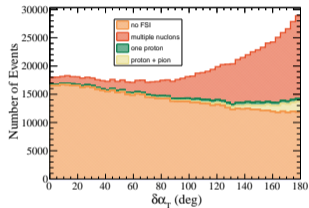
Paper ratio  $\approx 10 : 3 : 1$

# Vertex activity

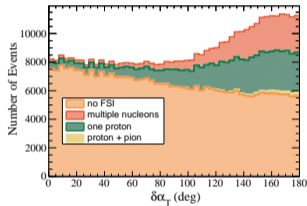
- take INCL and NuWro simulation **event by event**
- check the energy of the cluster/proton/muon **after FSI**
- try to use **ionisation curve** from the Geant4 simulation
- if there is no data, it means that particle **travels less** than 1(3) cm. Than we **use Birks law directly**, calculating path from the plot



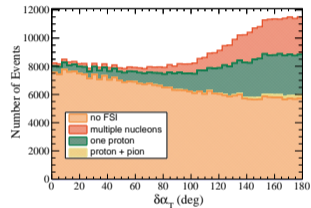
# Comparison to data: RW model



NuWro GFG

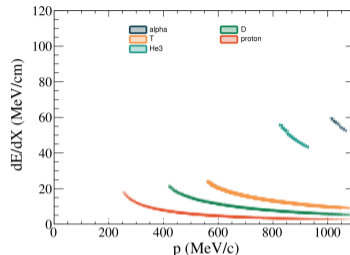
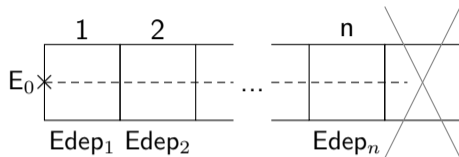


INCL RW



INCL + NuWro GFG

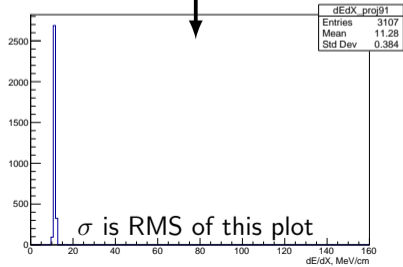
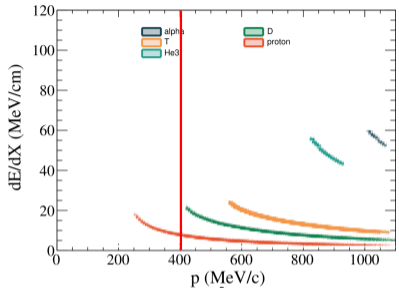
# Particle identification algorithm



- **initial kinetic energy**  $E_0$  is reconstructed as a sum of energy deposits along the whole track
- **momentum after passing 1 cm** is reconstructed using **5** mass hypotheses
- **for each momentum hypothesis**, the  $\frac{dE}{dX}_{rec}$  is calculated using the  $\frac{dE}{dX}$  dependence on momentum plot
- $\chi^2 = \sum \frac{\left(\frac{dE}{dX}_{sim} - \frac{dE}{dX}_{rec}\right)^2}{\sigma^2}$  is calculated for each hypothesis
- we choose hypothesis with the **lowest**  $\chi^2$



# $\sigma$ definition



To calculate  $\sigma$ , we need:

- take plot with  $dE/dX$  dependence on momentum
- find bin with the needed momentum
- to make a projection to  $dE/dX$  axis of this bin

