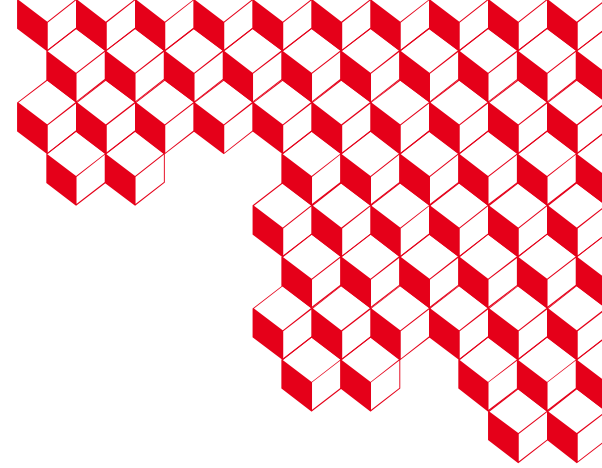




irfu



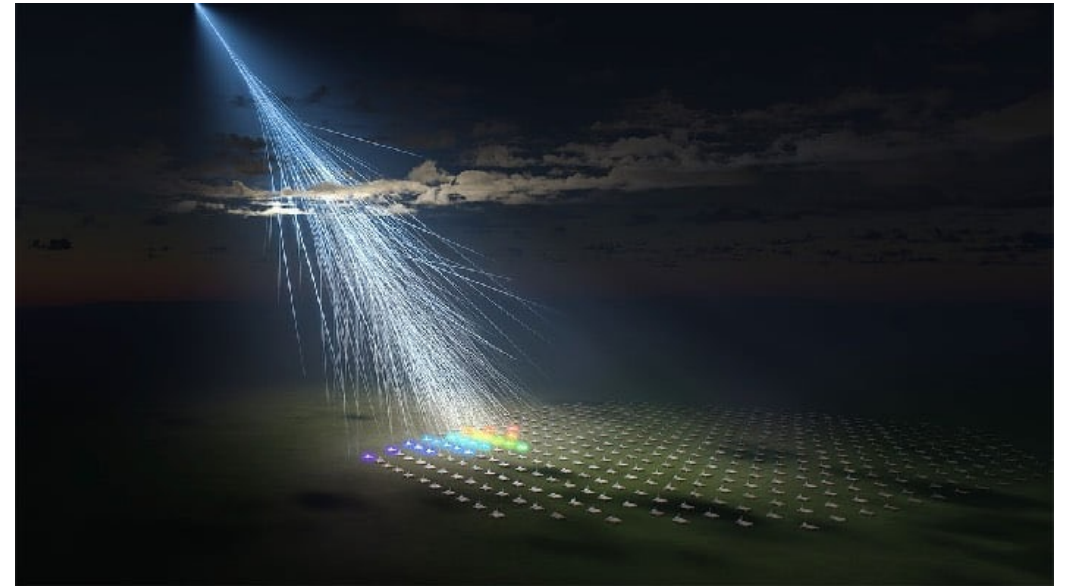
Rencontres de Moriond, VHEPU session Highlights

Mathieu de Bony

Summary

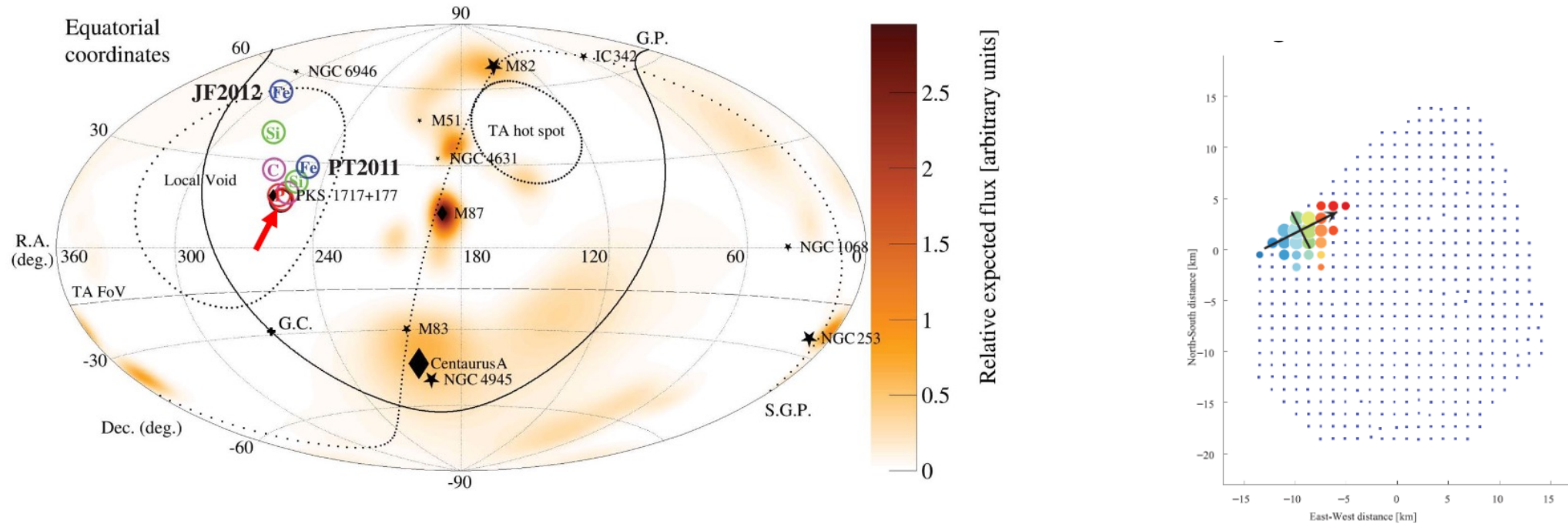
1. UHE cosmic-rays results
2. Future UHE cosmic-rays and neutrino experiments
3. VHE Neutrino astronomy results
4. Dark Matter search results
5. VHE gamma-ray results
6. Cosmology with Gravitational Waves

Ultra High Energy cosmic-rays with Pierre Auger Observatory and Telescope Array



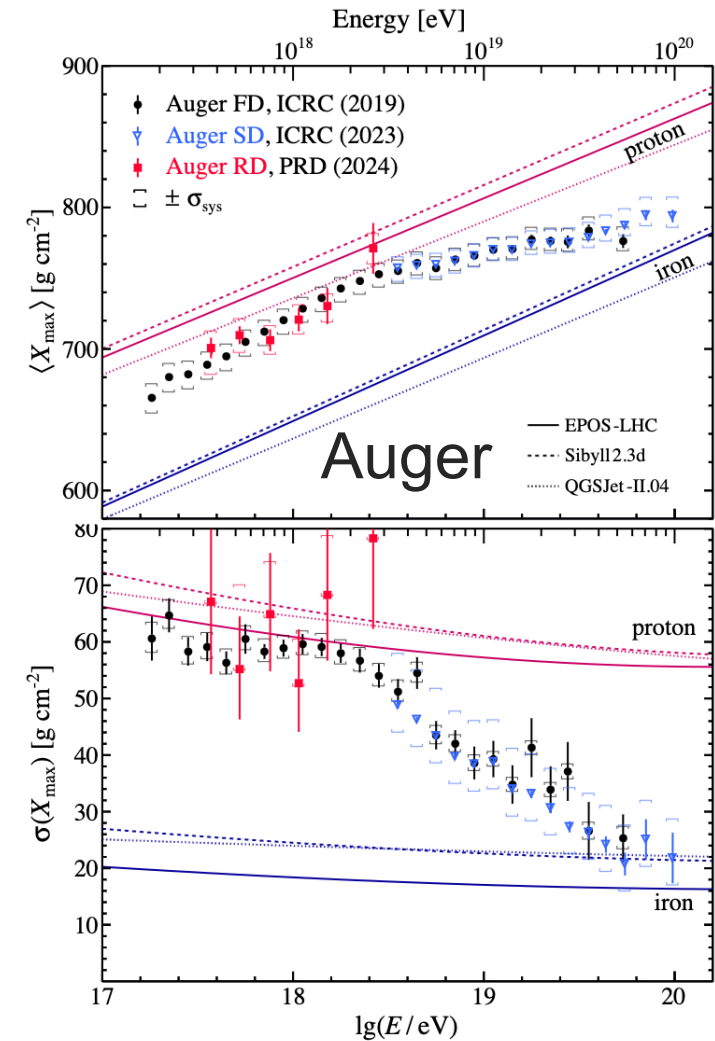
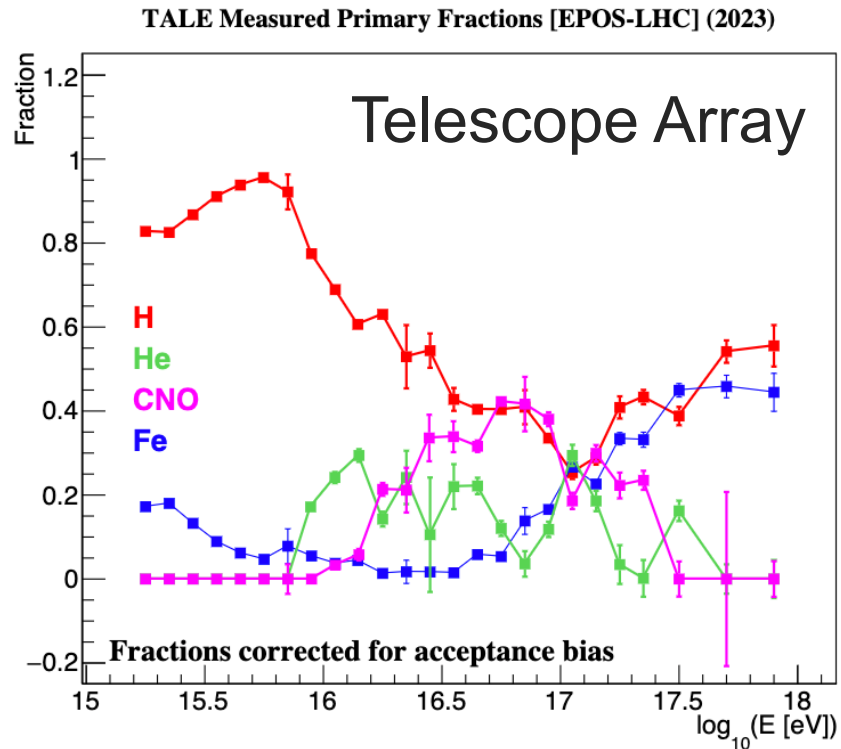
An extremely energetic cosmic ray event in Telescope Array

- Energy : 244 ± 29 (stat.) + 51-76(syst.)



Talk : Review of the Telescope Array results, E. Kido
Publication : TA Collaboration, Science, 382, 903 (2023)

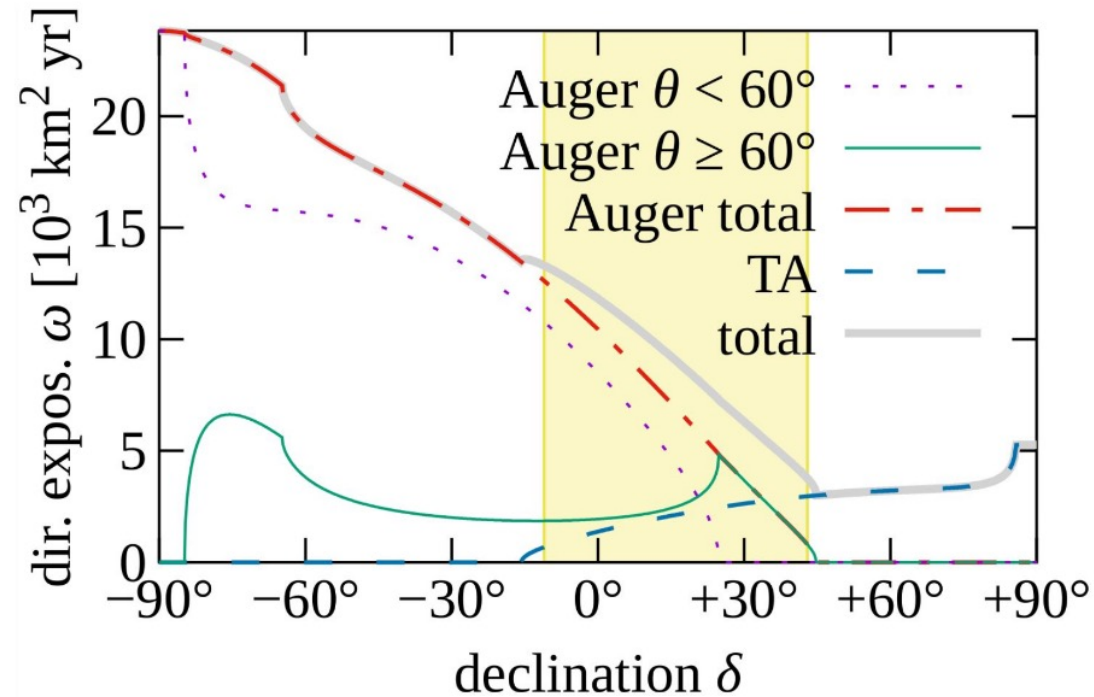
Energy dependant mass composition



- Talk :
- Review of the Telescope Array results, E. Kido
 - Ultra-high-energy cosmic rays with the Pierre Auger Observatory Current status and future perspectives, M. Roth and I. Maris

Anisotropy of UHE cosmic ray flux

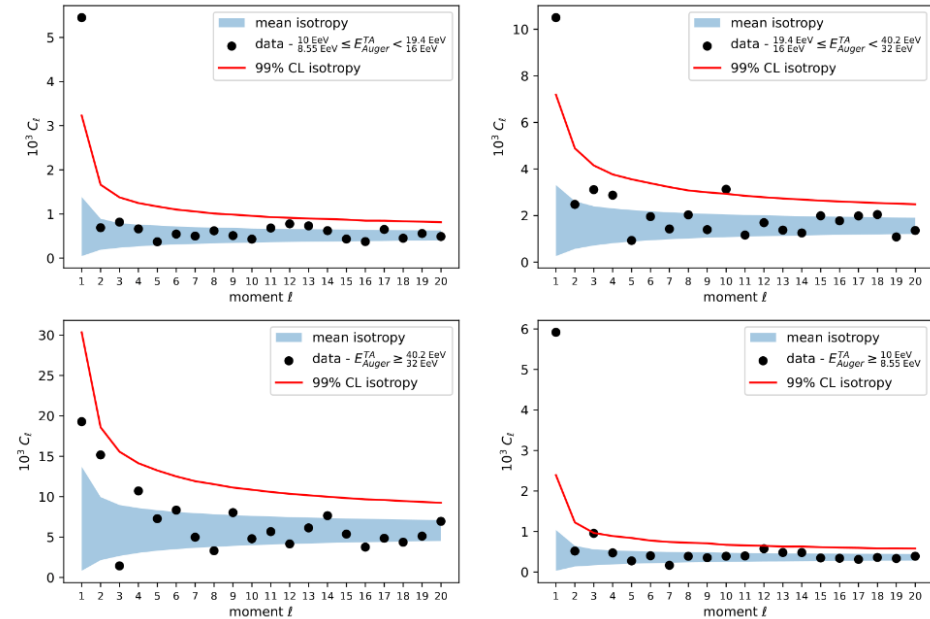
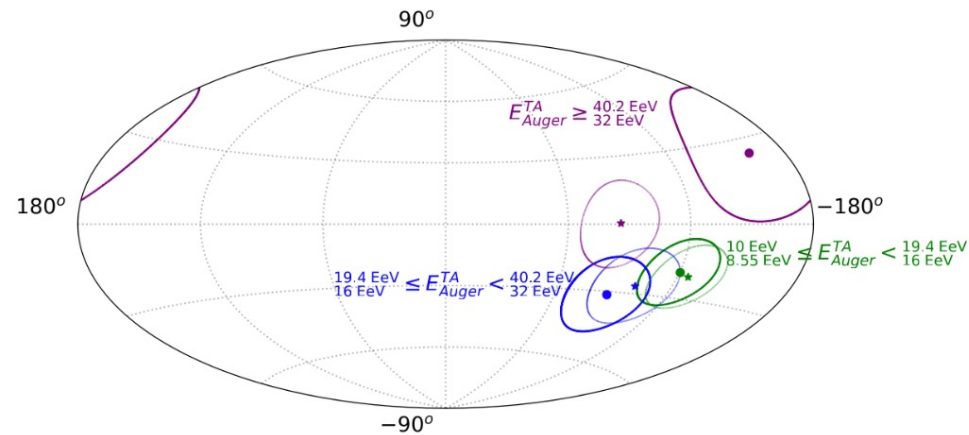
- Combination of data from Telescope Array and Pierre Auger Observatory



Talk : Results from the Auger-TA working group on UHECR arrival directions, F. Urban

Anysotropy of UHE cosmic ray flux

Dipoles and angular power spectrum

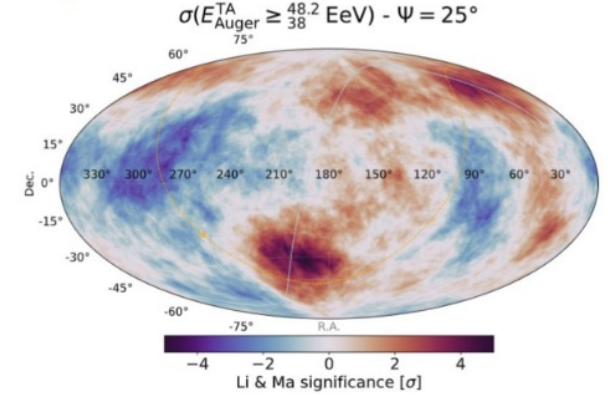
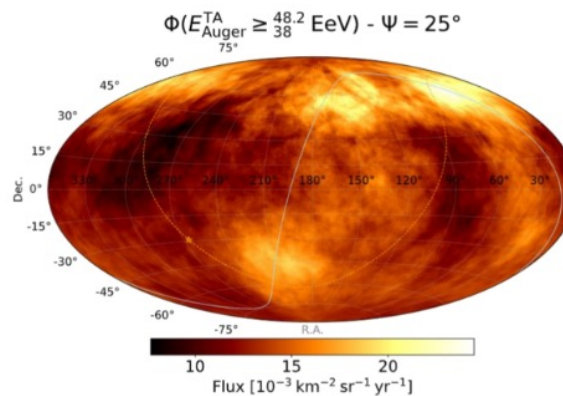
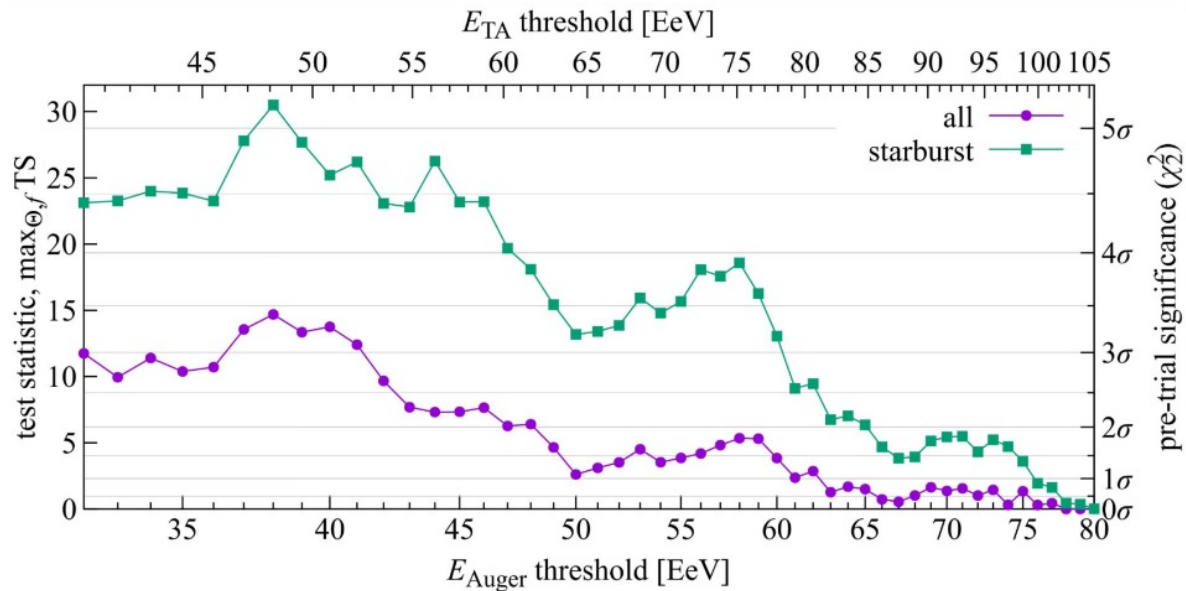


All energies : $l = 1$, 4.2σ deviation from isotropy

Talk : Results from the Auger-TA working group on UHECR arrival directions, F. Urban

Anisotropy of UHE cosmic ray flux

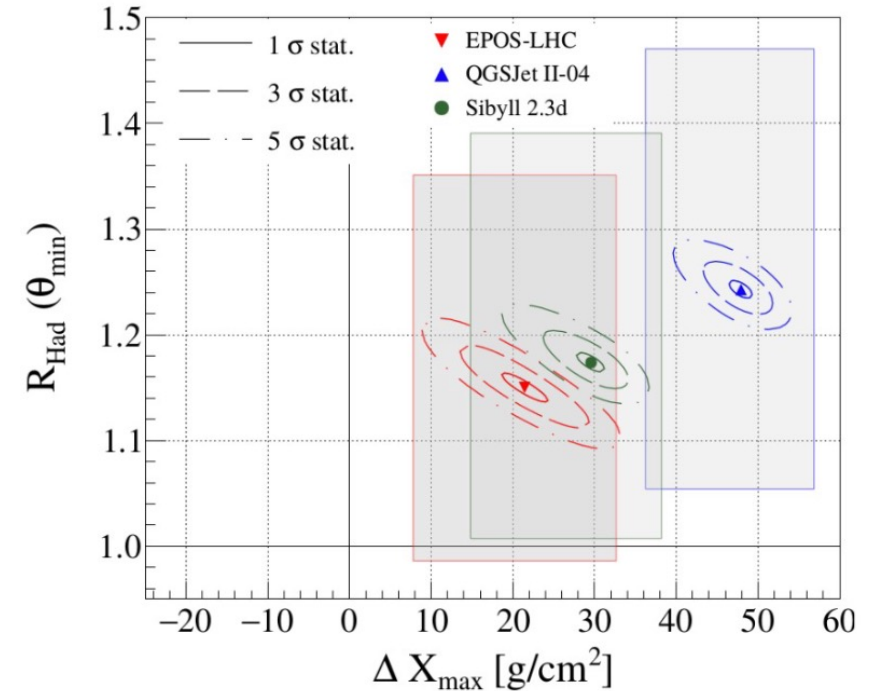
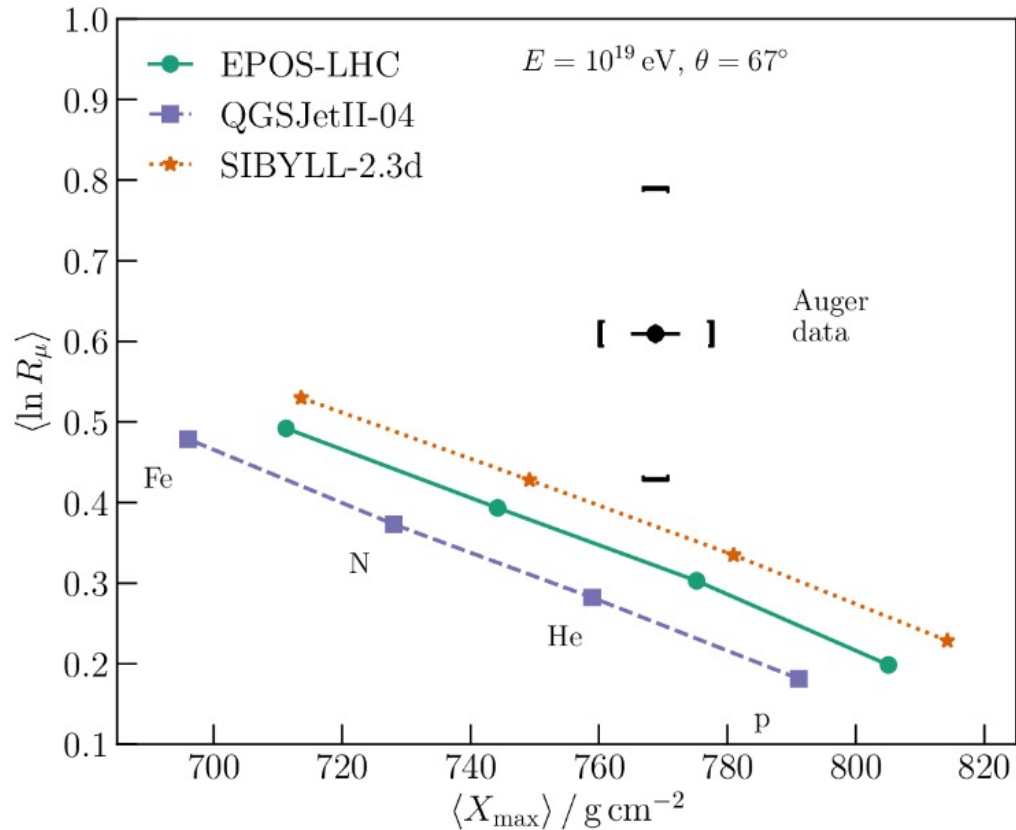
- 4.6 σ correlation with starburst galaxies



Talk : Results from the Auger-TA working group on UHECR arrival directions, F. Urban

Comparison of the new hadronic interaction models with Pierre Auger data

■ Too much muons in highly inclined showers



Discrepancy between X_{\max} prediction and measurement

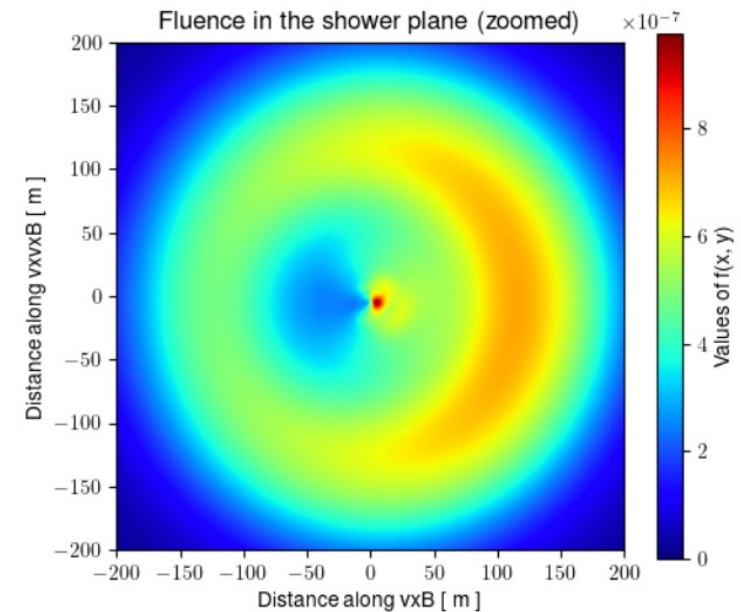
- Talk :
- Hadron interactions at ultra-high energy with the Pierre Auger Observatory, F. Riehn
 - EPOS LHC-R Up-to-date Hadronic Model for EAS Simulations, T. Pierog

New and future instruments UHE detector

■ Radio detection of atmospheric shower is the new trend

Allow large and cheap array

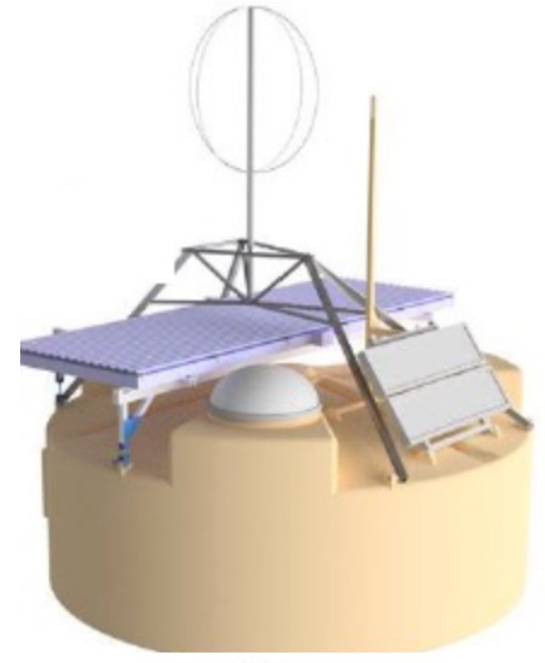
- For UHE cosmic-rays detector
- For UHE neutrino detector



Auger Prime

Objective : better and more details reconstruction of atmospheric showers

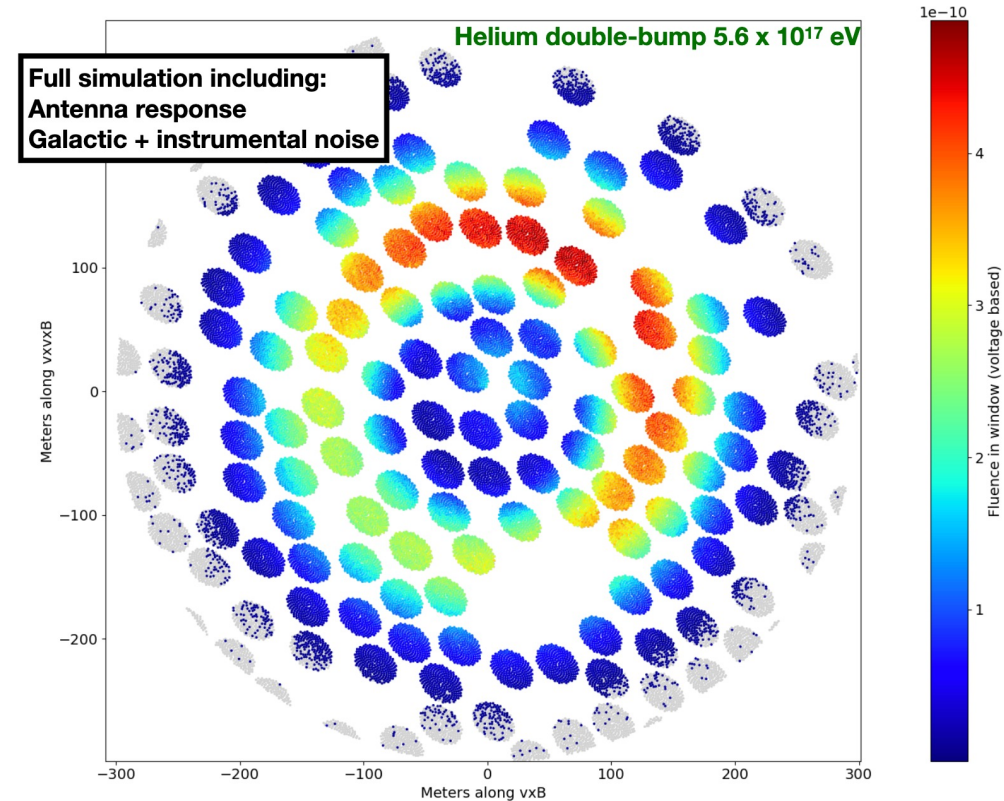
- Add **Plastic Scintillator Detector on top of each Water Cherenkov Detector**
 - Add smaller PM to WCD for better dynamic
 - New electronic for signal processing
 - **Radio detector on top of each WCD**
 - **Add an Underground Muon Detectors**
-
- **Measurement of the muonic content**
 - **Better mass reconstruction**



Talk :
- The AugerPrime upgrade of the Pierre Auger Observatory, T. Suomijärvi
- Measurements of UHE particles with the Radio Detector at Auge, C. Galea

Cosmic Ray detection with the Square Kilometre Array

- In continuation of a similar experiment at LOFAR
- The radio detectors is a the densest part of SKA-Low
 - Future most sensitive low frequency radio telescope
 - The whole telescope is composed of 131k antennas
- Trigger from scintillator detectors
- Precise measurement of the energy and X_{\max}



Talk : - High resolution air shower observations with the Square Kilometre Array, S. Buitink

GRAND 200k

- **Radio detection of highly inclined showers caused by Tau neutrino**
- In development phase :
 - GRANDProto300 : 300 antennas over 200 km²
 - GRAND@Auger : 10 antennas for cross calibration
 - GRAND@Nançay : 4 antennas for trigger testing
- Target :
 - **200k antennas over 20 sites**
 - Size of site : 10k km²
 - Number of antenna per site : 10k
- **Probe cosmogenic neutrino**
- **UHE neutrino astronomy**



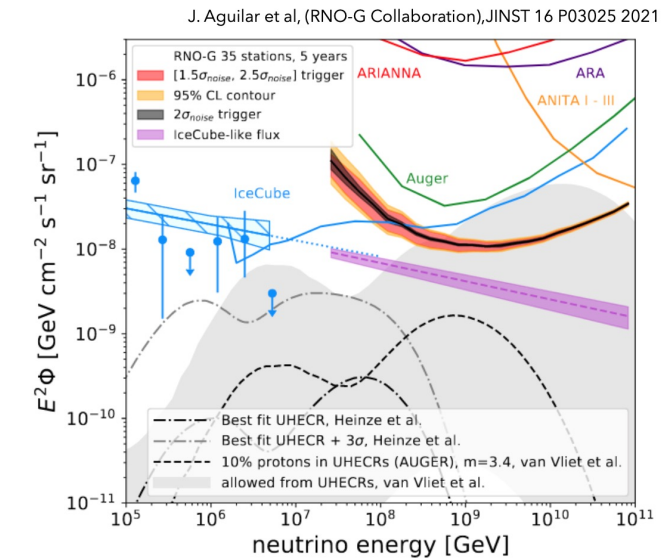
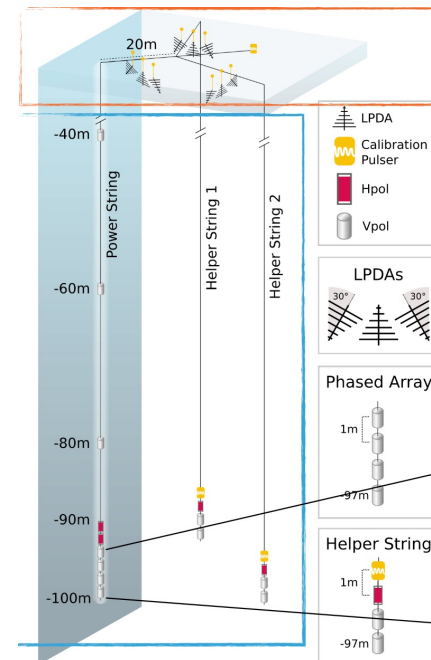
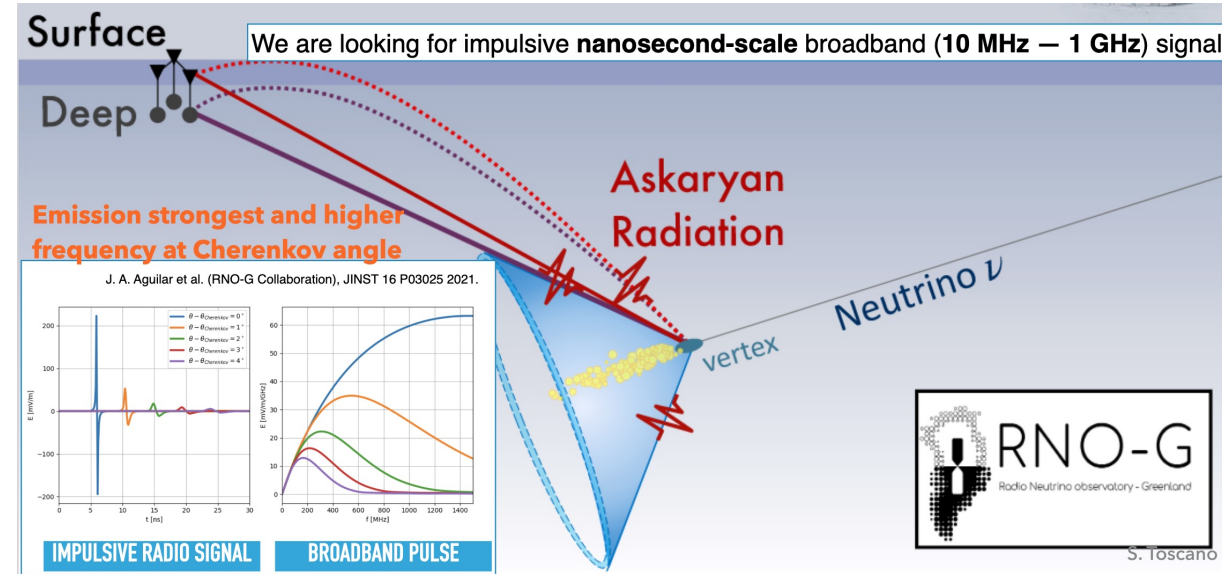
Talk : - The Giant Radio Array for Neutrino Detection and its prototype phases, M. Guelfand

RNO-G

- Radio detector in the ice to detect UHE neutrino
- Currently : 7 stations in Groenland
- Final : 35 stations of 50 km²

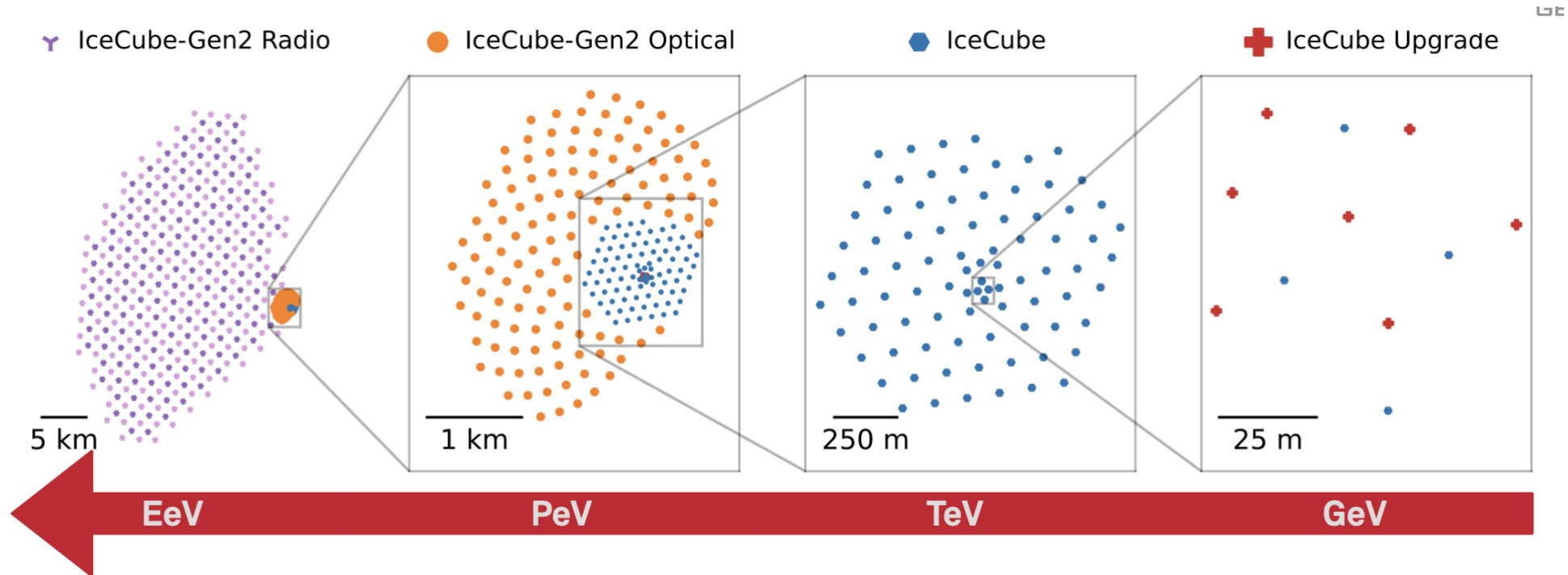
- Start exploring UHE neutrino astronomy
- R&D for IceCube-Gen 2

Talk : - RNO-G status and first results, S. Toscano



IceCube Gen2

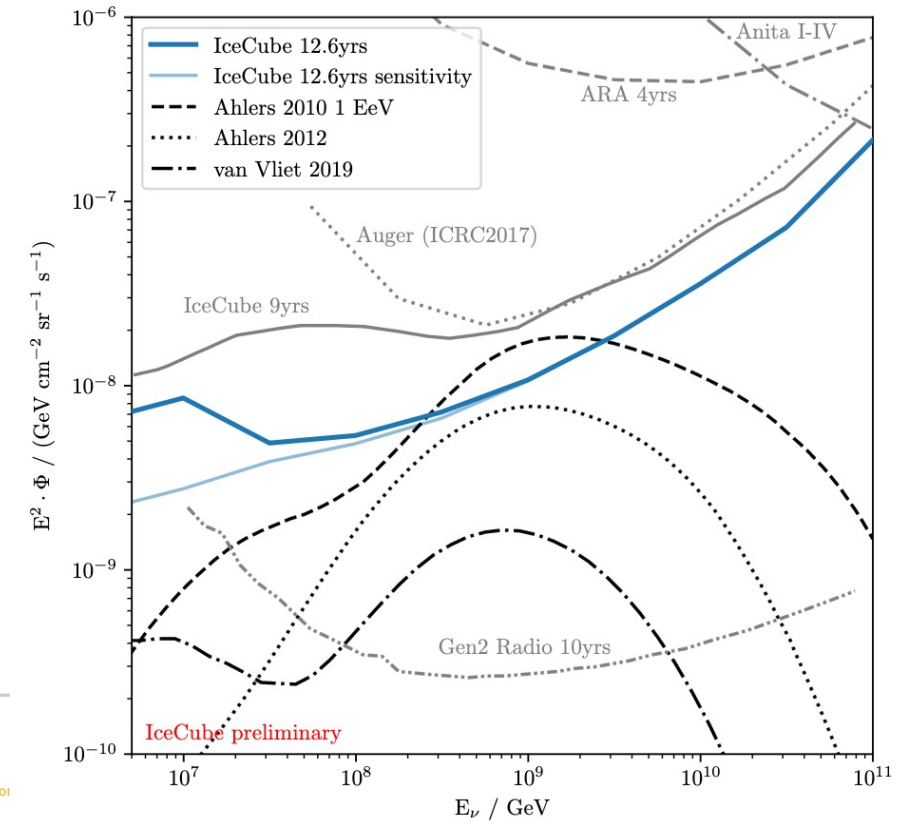
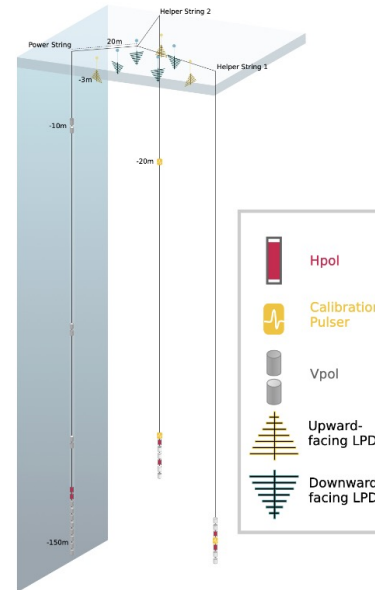
- 8 times larger optical array
- Add a radio array



Talk : - Recent cosmogenic neutrino search results with IceCube and prospects with IceCube-Gen2, M. Meier

IceCube Gen2

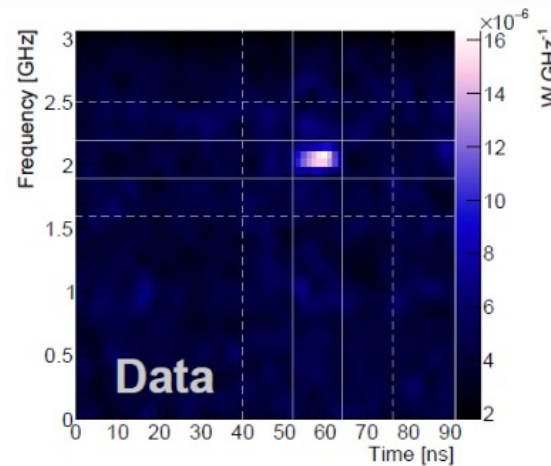
- 8 times larger optical array
- Add a radio array
 - 360 stations
 - Combination of shallow and deep antennas
 - 500 km² array
 - Will probe the cosmogenic neutrino flux



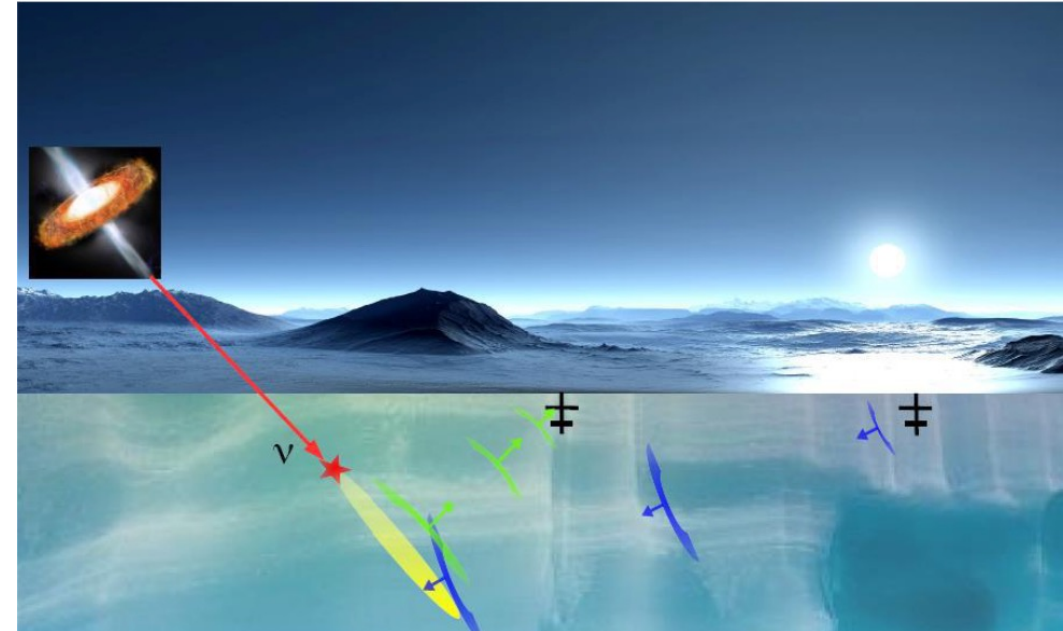
Talk : - Recent cosmogenic neutrino search results with IceCube and prospects with IceCube-Gen2, M. Meier

Radar detector for neutrino

- Detection with a radar system of the plasma created by showers in the ice
- First station deployed in 2023 in Groenland
 - Analysis ongoing



- Probe cosmogenic neutrino
- UHE neutrino astronomy



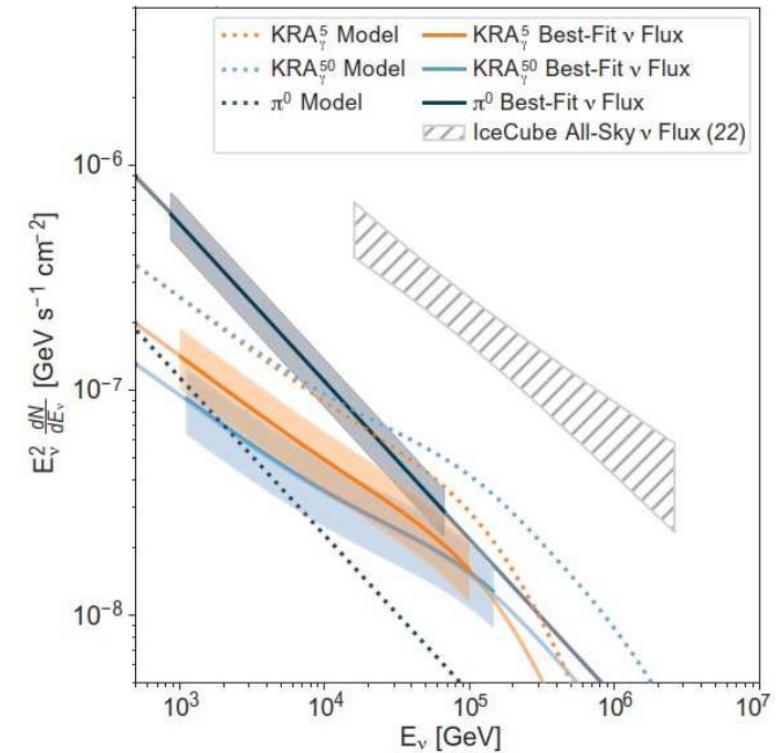
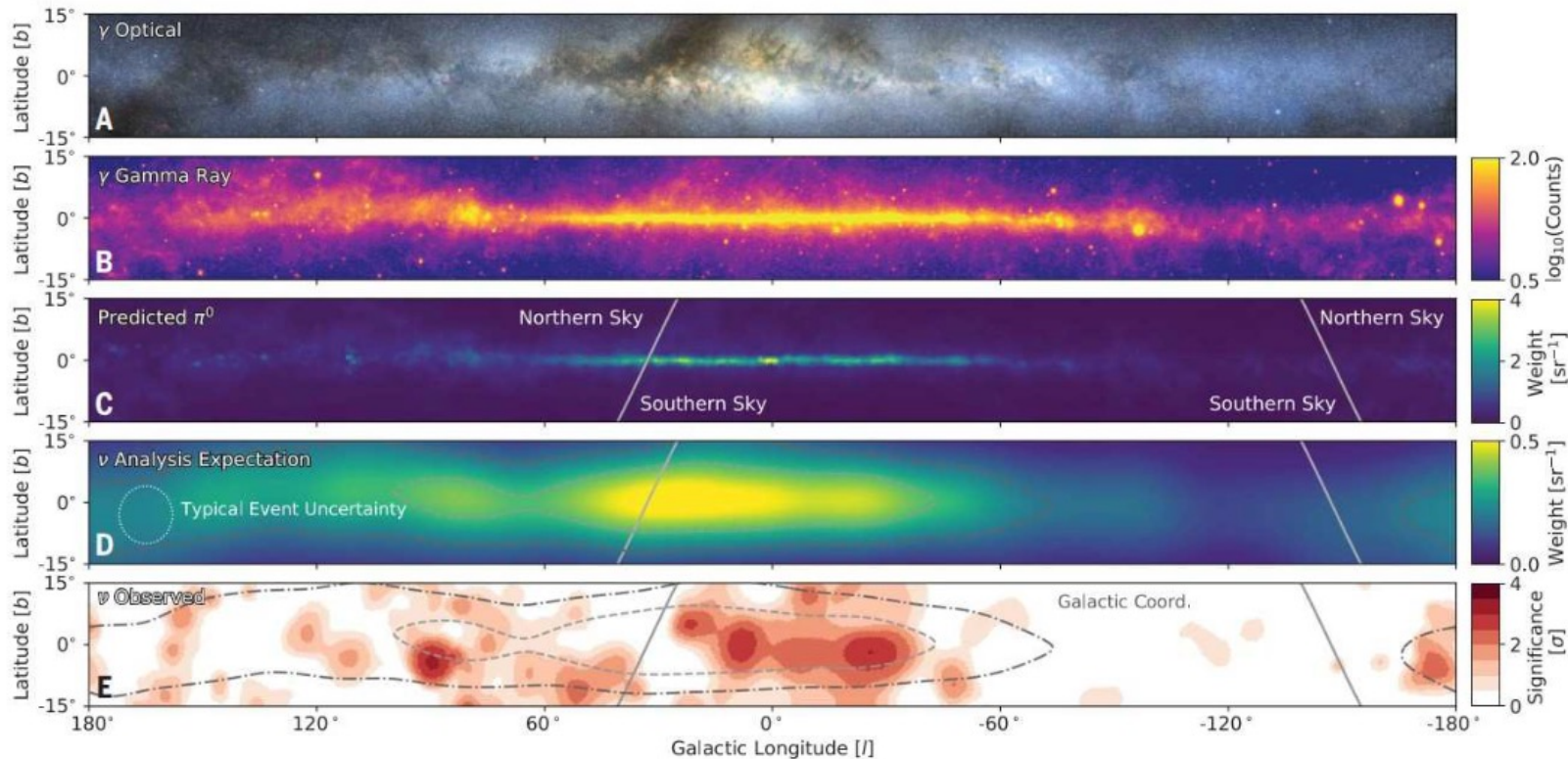
Talk : - The Radar Echo Telescope, K. de Vries

Recent neutrino astronomy results



Neutrino emission from the galactic plane

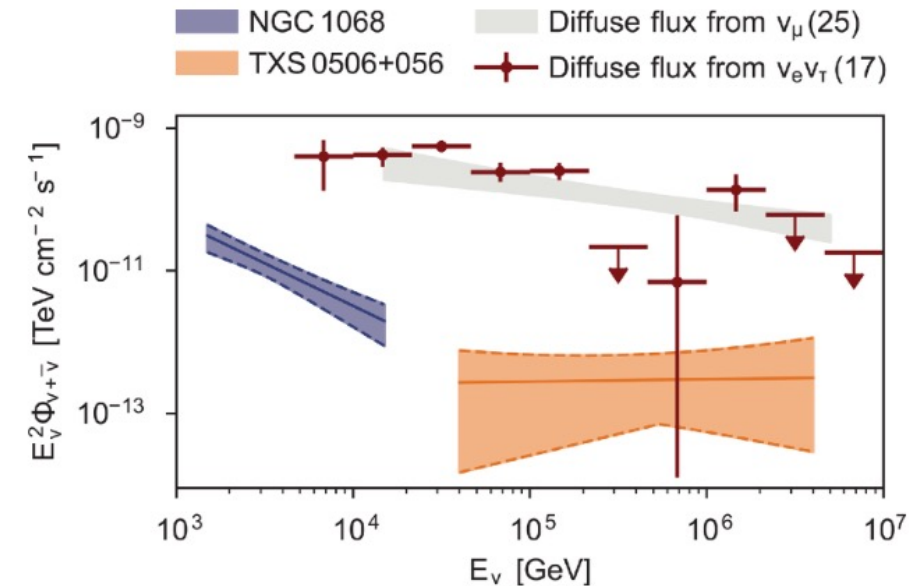
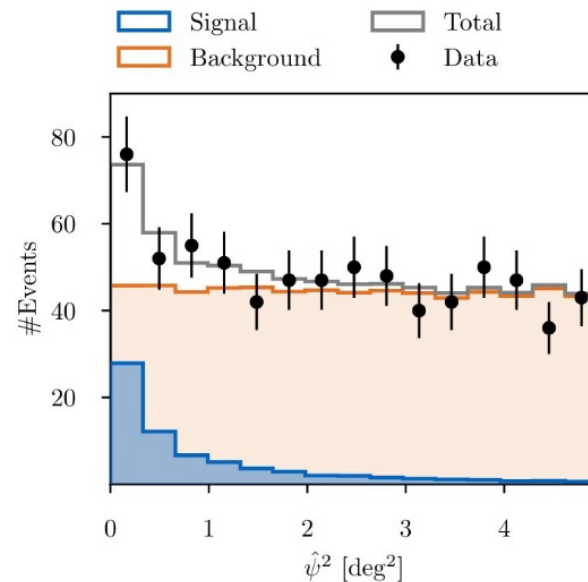
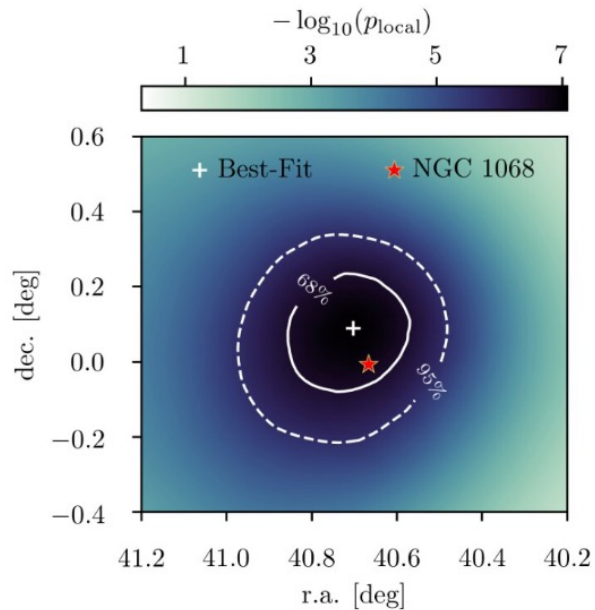
■ 4.48 σ excess for diffuse emission in IceCube



Talk : - Searches for Galactic Neutrinos with the IceCube Neutrino observatory, A. Sandrock

Neutrino emission from NGC 1068

■ 4.2 σ excess for emission from NGC 1068 (Seyfert Galaxy)



■ 3 σ excess for stacked Seyfert galaxies, 2.7 σ excess from 2 other Seyfert galaxies

■ Not able to explain diffuse emission with these source candidate (too soft spectra)

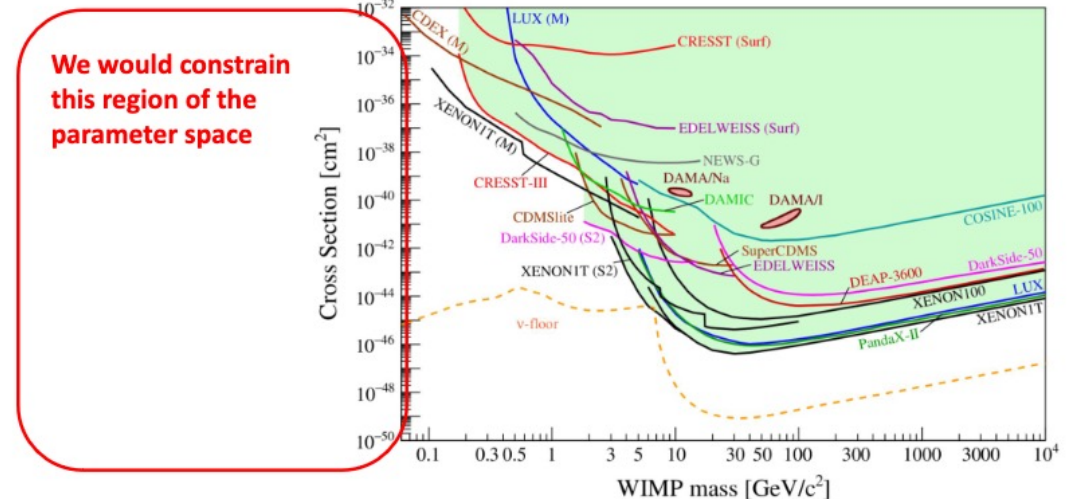
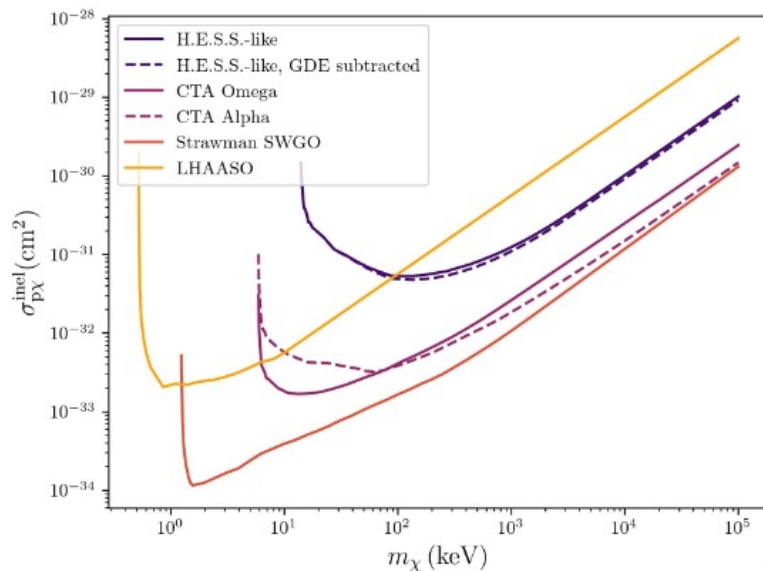
- Talk :
- Recent IceCube Results: Diffuse Flux, Point Sources, and Dark Matter, M. Jeong
 - NGC 1068 constraints on neutrino-dark matter scattering, J. Cline

Recent dark matter results



Inelastic scattering of cosmic rays on dark matter

- Explore potential for signal cause by inelastic scattering of cosmic-rays on dark matter
 - Need region with large DM content and large cosmic-ray flux
 - Observation of the Galactic Center with TeV gamma-rays
 - Could lead to limit on the cross section of interaction in an explore parameter space



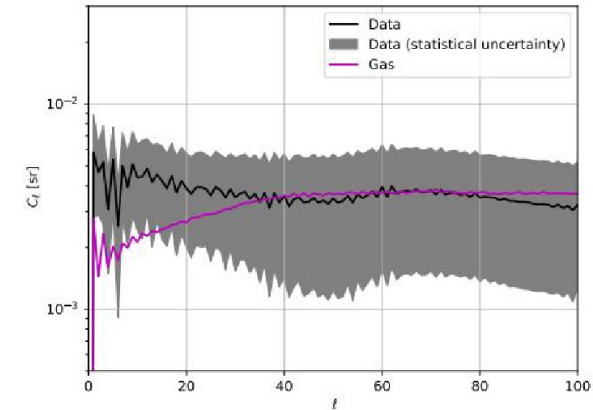
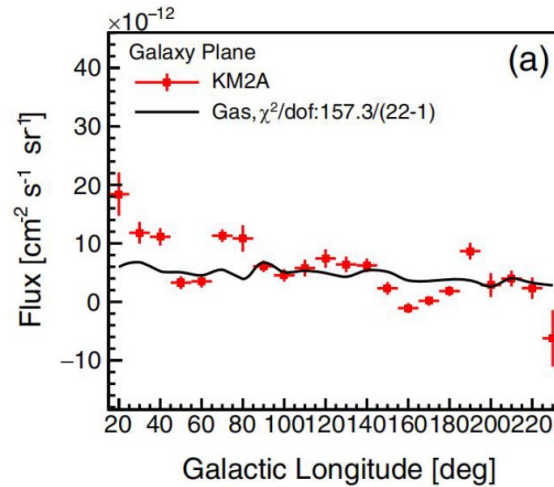
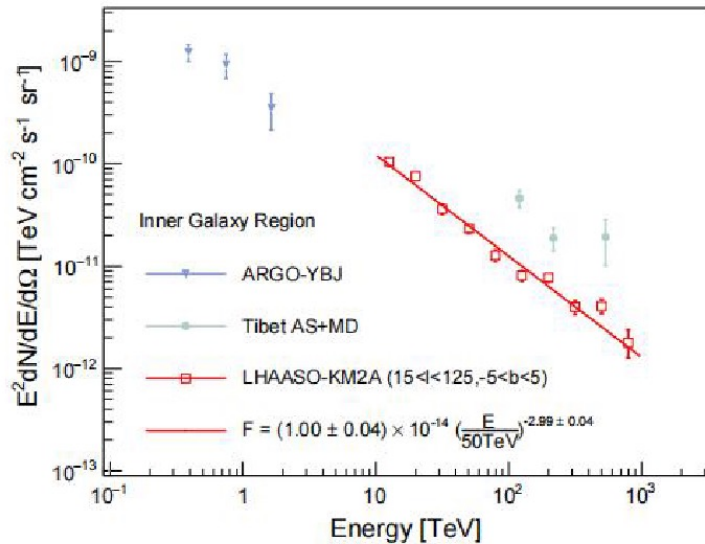
Talk : - Sub-GeV Dark Matter Indirect Detection From Cosmic-Ray Scattering, I. Reis

Last results from VHE Gamma-Ray astronomy



Diffuse galactic plane gamma-ray emission with LHAASO

■ Measurement of the diffuse gamma-rays galactic plane emission by LHAASO

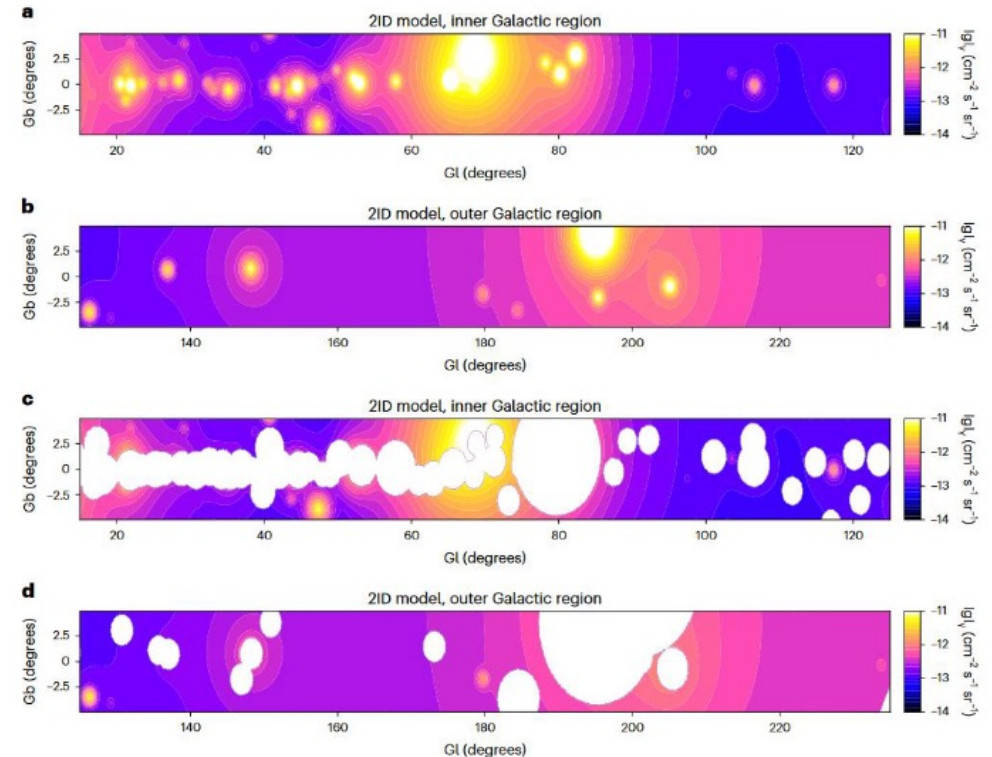
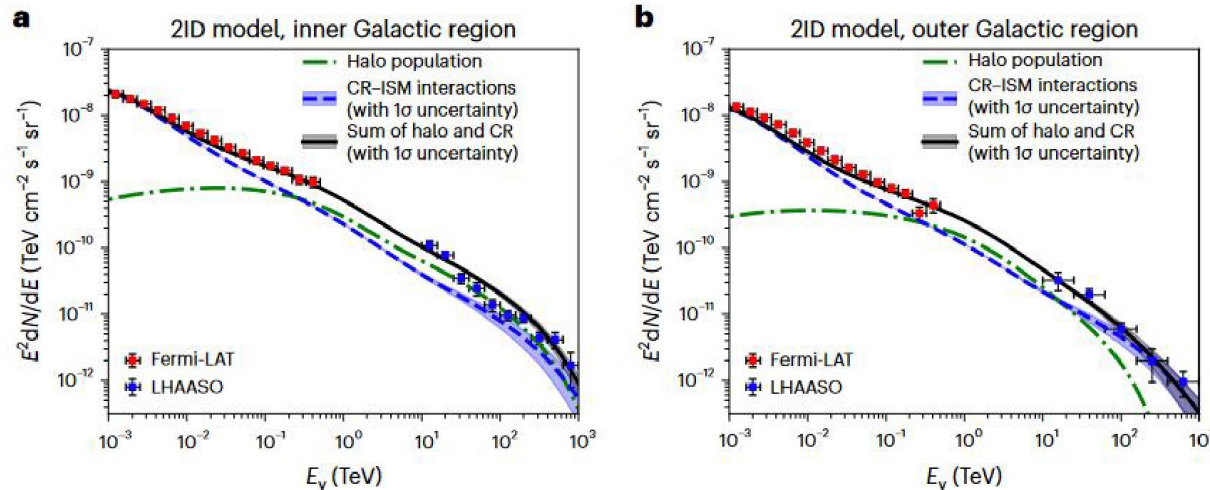


■ Some tension between gas and gamma-ray spatial distribution

Talk : - Probing cosmic ray origin and propagation with LHAASO, R. Zhang

Diffuse galactic plane gamma-ray emission with LHAASO

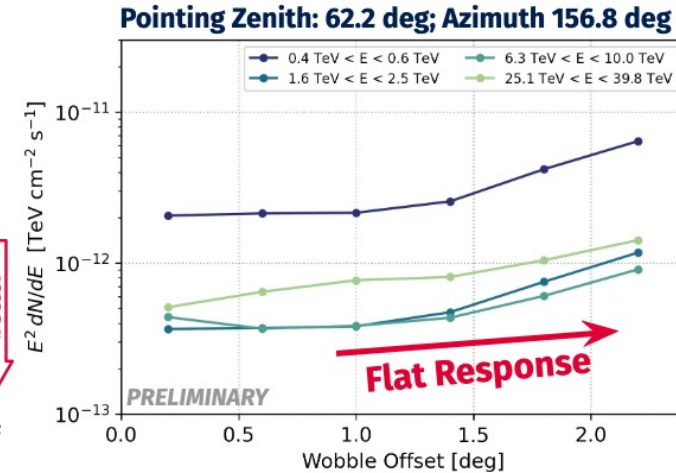
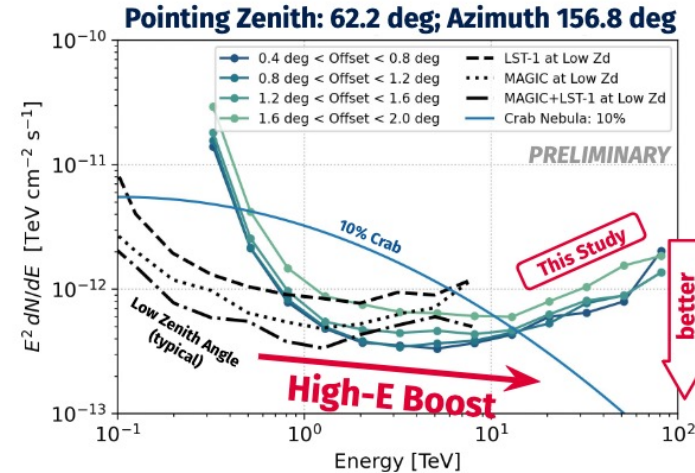
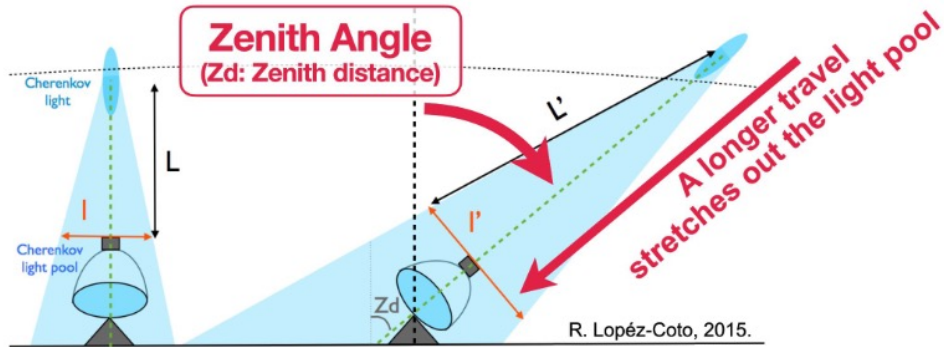
- Measured flux 2 to 3 time higher than predicted by models
 - Propagation effect ?
 - Unresolved population sources ?
 - Pulsar Halos ?
 - Not able to explain the inner galaxy region



Talk : - Probing cosmic ray origin and propagation with LHAASO, R. Zhang

Galactic center with LST-1

Observations of the galactic center at large zenith angle with LST-1

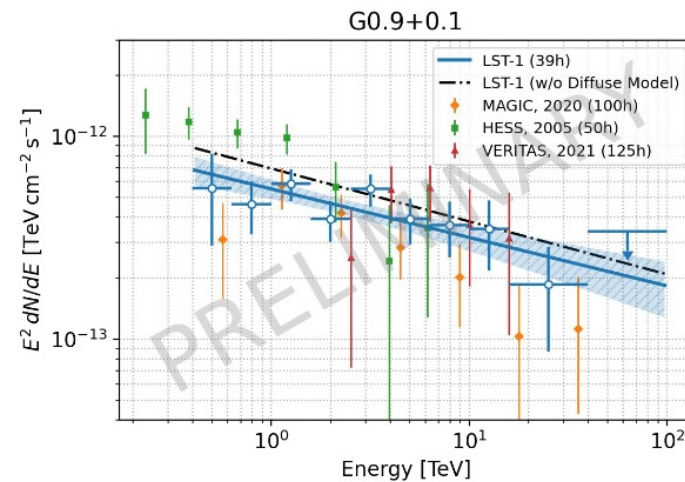
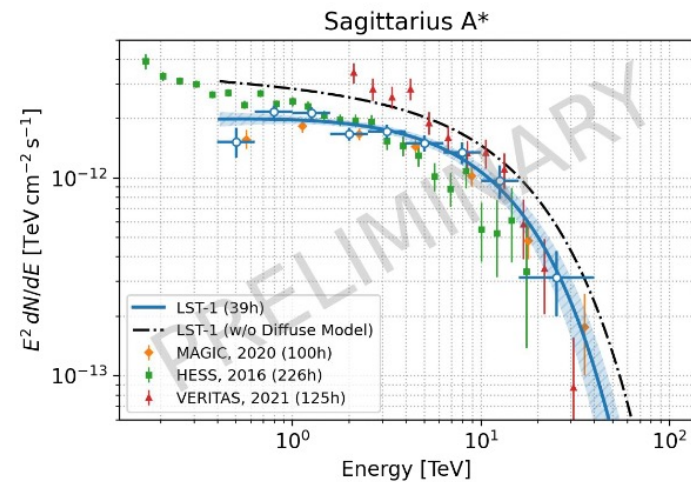


- Larger effective area at high energy
- Performance very sensitive to the zenith angle (even inside the FoV)
- Much higher energy threshold

Talk : - Galactic Center Observations with CTA-LST-1, S. Abe

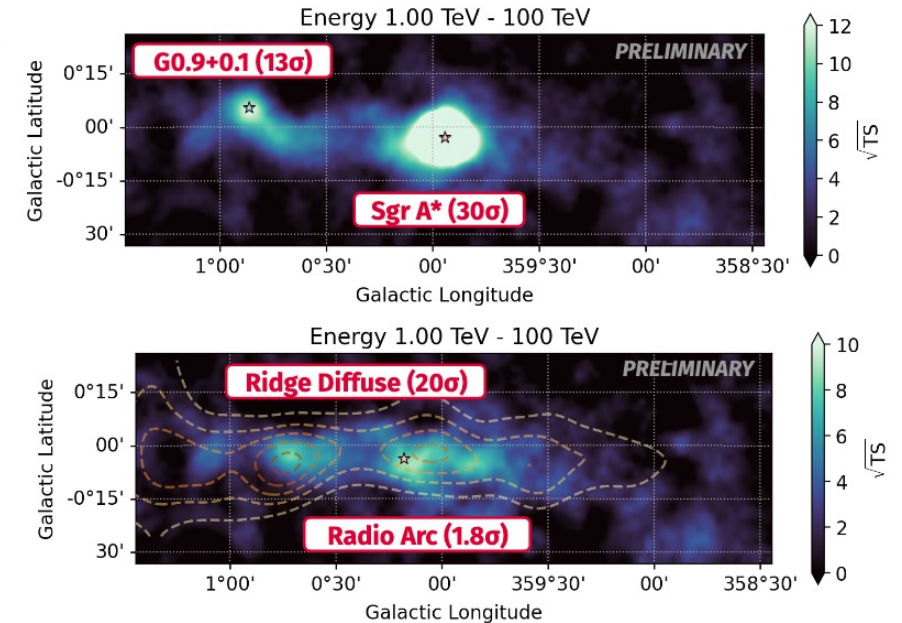
Galactic center with LST-1

Observations of the galactic center at large zenith angle with LST-1



TS Map

Sgr A* & G0.9+0.1 subtracted

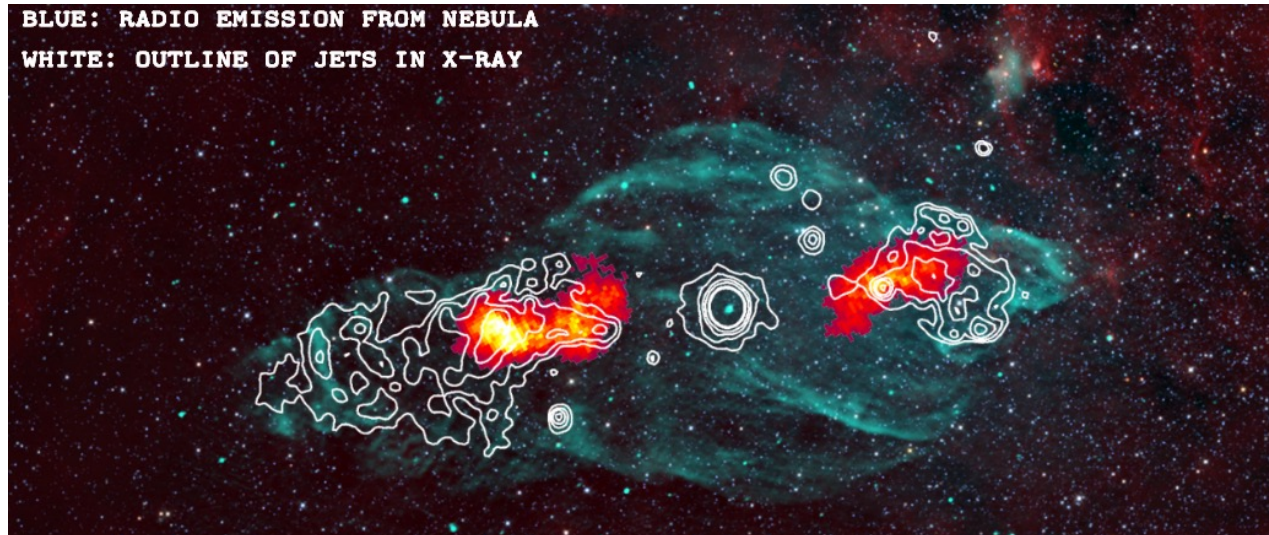


- Clear detection of the central source and the diffuse emission
- Results in agreement with previous experiments
- We could hope for high precision measurement once CTA is completed

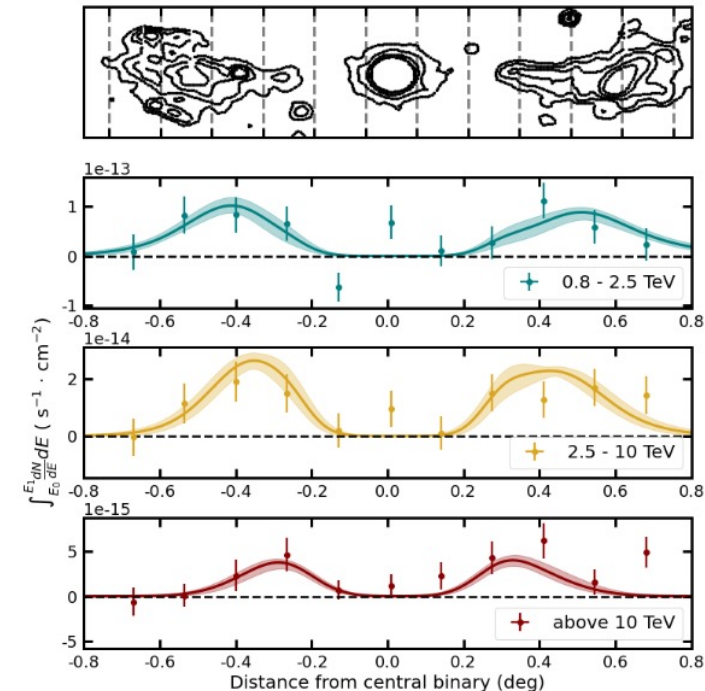
Talk : - Galactic Center Observations with CTA-LST-1, S. Abe

Detection of the jets of SS 443

■ Detection of the jets of SS 443 at VHE by H.E.S.S.



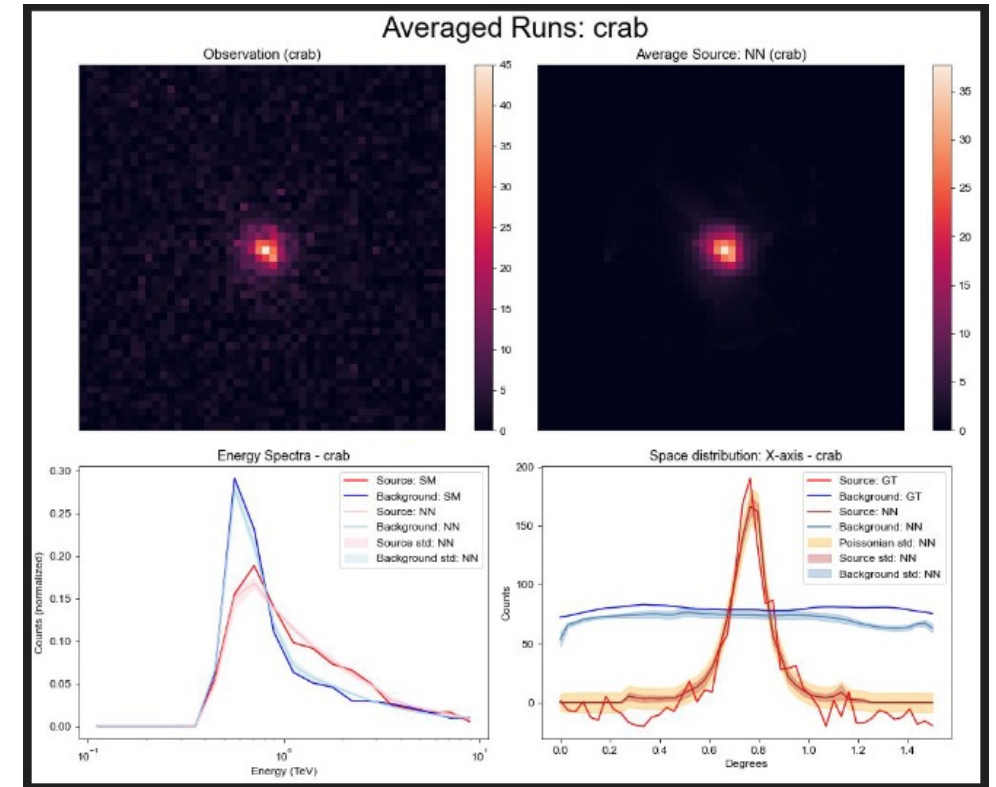
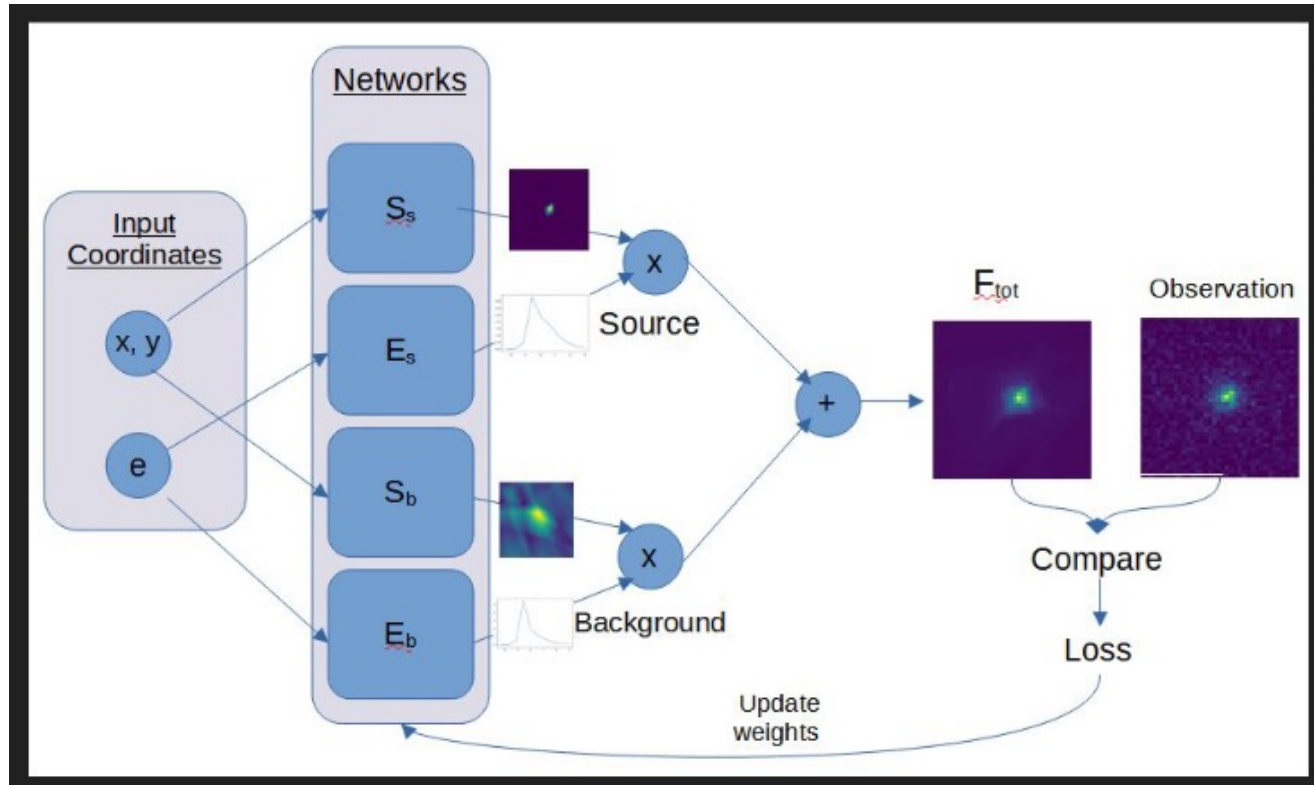
- Electrons accelerated due to shocks at the base of the outer jet
- Inverse Compton emission of VHE photons
- Origin of the shock region ?



Talk : - Acceleration and transport of relativistic electrons in the parsec-scale jets of the microquasar SS 433, L. Oliviera-Nieto

Separation of sources from background with ML

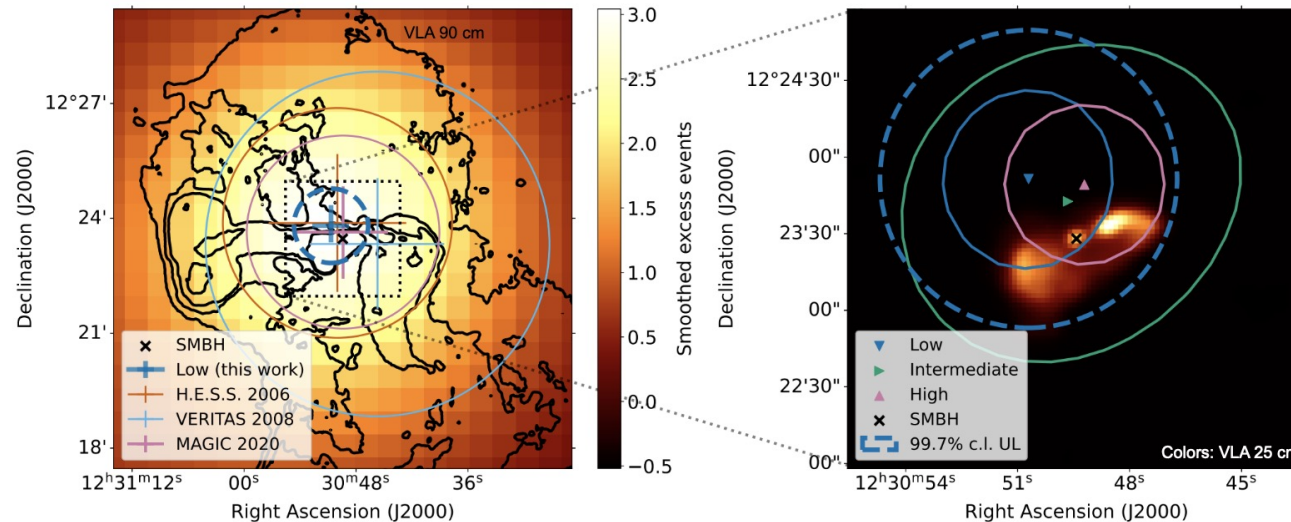
- Look at the potential of ML algorithm to separate source from background in IACT data



Talk : - Non-parametric Signal Separation with Probabilistic Machine Learning, M. Ullmo

Gamma-ray emission from M87

- Search for diffuse emission due to hadronic interaction in the low state of M87 by H.E.S.S.

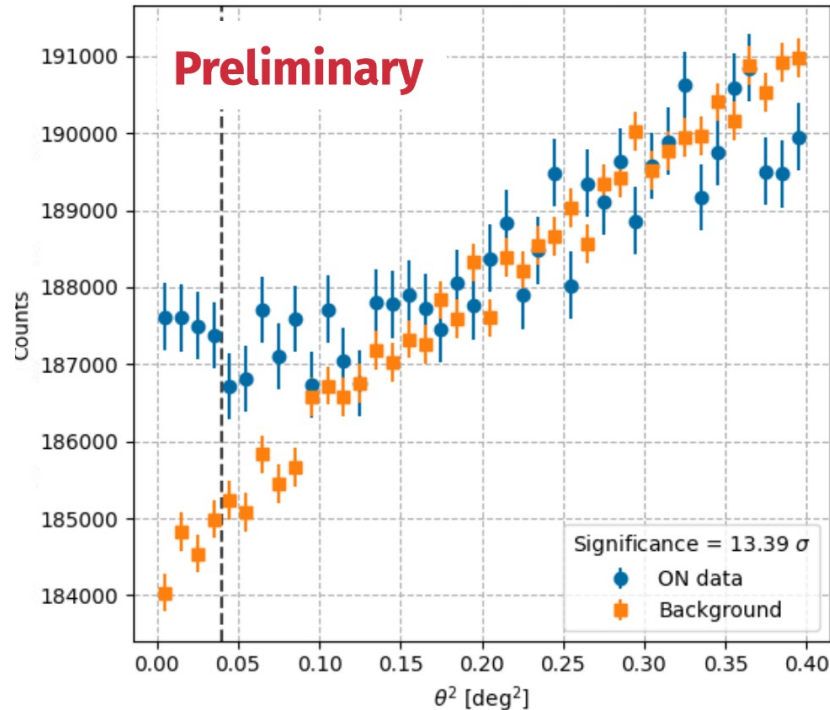


- Gamma-ray emission linked to hadronic process could not be more than 55% of the flux
- The leptonic emission could not come from radio lobes
- Unexplained curvature of the spectrum during flares

Talk : - Unveiling the gamma-ray mysteries of M87, V. Barbosa Martins

Furthest blazar ever detected at VHE by LST-1

- Detection at VHE of OP 313 ($z=0.997$) during a flare, 19h of data (LST-1)



- Average flux : 0.28 Crab
- Show the ability of LSTs and CTA at detecting high redshift sources

Talk : - AGN observations with LST-1: first results and discovery of the flat spectrum radio quasar OP 313, J. Otero-Santos

Astro-COLIBRI

- A tool to help at observations of transient phenomena
- astro-colibri.com

The screenshot displays the Astro-COLIBRI web interface. At the top, there are navigation buttons for 'Select action', 'Latest transients', and 'Cone search'. Below this is a filter bar for 'Observatories' (Swift, Fermi, HAWC, IceCube, AMON, Integral, GECAM, FlaapLUC, LVC, Catalogs, Other) and 'Event type' (FRB, Unclassified OT, Classified OT, SN, GRB, burst, neutrino, nuem, GW, 4FGL, TeVCAT, SGR/AXP, IceCat). A timeline at the top shows dates from 2023-12-01 to 2023-12-31 with various event icons. On the left, a list of transients is shown, including GRB 231214B (Gamma-ray burst), SN 2023zzi (Supernova), S231213ap (Gravitational wave), AT 2023aebz (Classified optical transient), and ZTF23abtniaf (Unclassified optical transient). The main panel features a 'Custom cone search' for S231213ap with RA/Dec: 170.95°, 29.83° and a 'Cone search' button. Below this is a sky map showing the search cone and various transients. On the right, 'Detailed info about selected source' for S231213ap is provided, including detection time (2023-12-13 11:14:17), RA/Dec, and classification (BBH: 1.09). A description explains that gravitational waves are distortions of space-time generated by accelerated masses. At the bottom, there are links for further details to GraceDB, TreasureMap, ALADIN, ESASKy, and TNS.

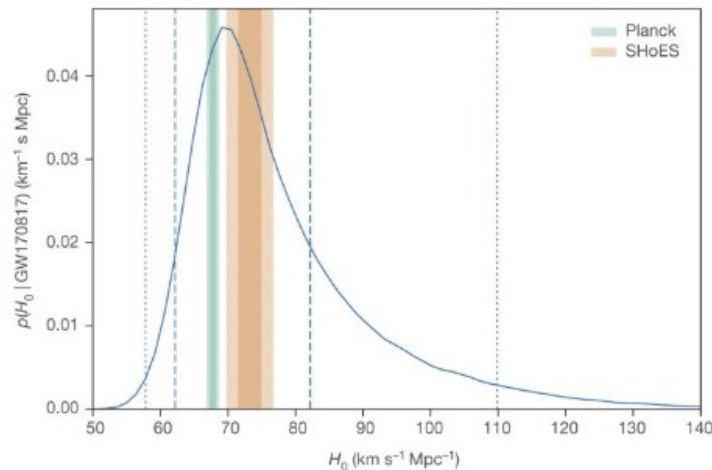
Talk : - Multi-messenger follow-up observations using Astro-COLIBRI, F. Schüssler

Cosmology with gravitational waves

- Attempt at new independent measurement of the Hubble constant

Hubble constant : Brighth sirens

- Comoving distance measurement from GW interferometer
- Redshift measurement by optical telescope

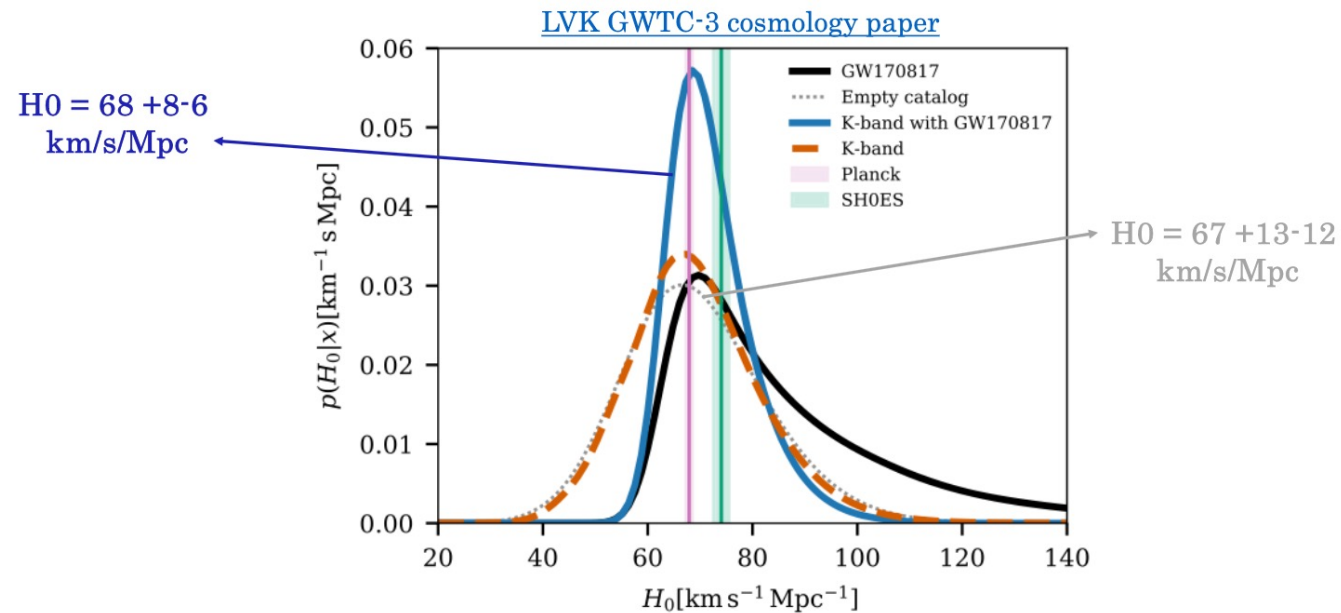


- $H_0 = 68+8-6 \text{ km/s/Mpc}$
- Detection of such events very rare (only one so far, GW 170817)
- Would need $O(100)$ events for precise measurement

Talk : - Results from multi-messenger investigations during the LIGO/Virgo/KAGRA O3 and O4a runs, R. Poggiani

Hubble constant : Dark Sirens

- Comoving distance and redshift measured by GW interferometer
- Degeneracy between mass and redshift

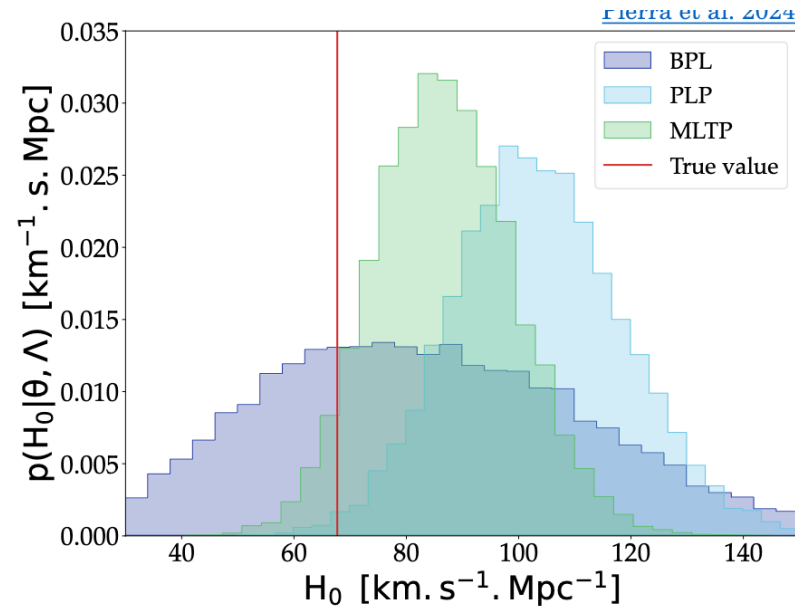


- $H_0 = 67+13-12$ km/s/Mpc

Talk : - Gravitational Wave Cosmology: Be Careful of the Black Hole Mass Spectrum, G. Pierra

Hubble constant : Dark Sirens

- Comoving distance and redshift measured by GW interferometer
- Degeneracy between mass and redshift

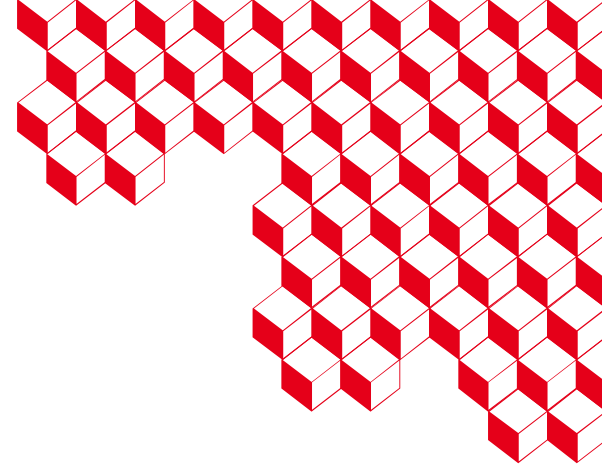


- Need a very good knowledge of the redshift dependant mass distributions of black holes
- Looking at how to improve robustness to systematic from the mass distribution

Talk : - Gravitational Wave Cosmology: Be Careful of the Black Hole Mass Spectrum, G. Pierra



irfu



Thanks for your attention