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SEARCHES FOR DARK MATTER PRODUCTION AT THE LHC





Fundamental Open Questions in Particle Physics

I. What is the origin of mass?

- Why are the vector bosons Z and W are massive whereas the photon is massless?
- Is there a Higgs boson or even more of them ?

Is there a new symmetry - Supersymmetry ?

- Can we get experimental evidence to support the Grand Unification of all fundamental forces?
- What is the origin of Dark Matter in the Universe?
 - \rightarrow Is a fundamental particle responsible for it?

III. What is the origin of the matter-anti-matter

asymmetry in our Universe?

- Does the answer lie in in CP violation?
- Neutrino masses and mixing how do they fit in the picture?



Fundamental Open Questions in Particle Physics

SUSY & DM Searches @ LHC 0. Buchmüller

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Searches for SUSY (& DM)



Searches for SUSY (& DM)



Supersymmetry

Extension of the Standard Model: Introduce a new symmetry Spin ½ matter particles (fermions) ⇔ Spin 1 force carriers (bosons) Standard Model particles SUSY particles



New Quantum number: R-parity:

- $R_p = (-1)^{B+L+2s} =$
 - +1 SM particles-1 SUSY particles

- <u>R-parity conservation</u>:
- SUSY particles are produced in pairs
- The lightest SUSY particle (LSP) is stable

What do we call a "SUSY search"?

The definition is purely derived from the experimental signature. Therefore, a "SUSY search signature" is characterized by Lots of missing energy, many jets, and possibly leptons in the final state



Missing Energy: • from LSP

<u>Multi-Jet:</u> • from cascade decay (gaugino)

Multi-Leptons:

from decay of charginos/neutralios

RP-Conserving SUSY is a very prominent example predicting this famous signature but ...

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What is its experimental signature?

. by no means is it the only New Physics model predicting this experimental pattern. Many other NP models predict this genuine signature



Missing Energy:

• Nwimp - end of the cascade

Multi-Jet:

• from decay of the Ns (possibly via heavy SM particles like top, W/Z)

Multi-Leptons:

• from decay of the N's

Model examples are Extra dimensions, Little Higgs, Technicolour, etc but a more generic definition for this signature is as follows.

Early SUSY Search Strategy at the LHC



Search Signatures

- SUSY-like decay chains range from short to long and simple to very complicated.
- All physics objects, MET, jets, leptons, photons, b's taus, tops, W, Z, etc are involved

> Comprehensive coverage of all possible signature requires a topology oriented search strategy:

References Analyses

0- leptons	1-lepton	OSDL	SSDL	≥3 leptons	2- photons	γ+lepton
Jets + MET	Single lepton + Jets + MET	Opposite- sign di- lepton + jets + MET	Same-sign di-lepton + jets + MET	Multi- lepton	Di-photon + jet + MET	Photon + lepton + MET

Already in less then two years of operation ATLAS & CMS managed to carry out the full list of these core "SUSY References Analyses"!

Inclusive SUSY Searches in 2013



Inclusive SUSY Searches in 2013



Inclusive SUSY Searches in 2013



CMSSM: Evolution with time



Global Fit to indirect and direct constraints on SUSY!

Other "fitter" groups find very similar results: e.g. SuperBayeS: arXiv:1212.2636 Fittino group: arXiv:1204.4199

X² increase from bluish to reddish

Source:



http://mastercode.web.cern.ch/mastercode/

Observable	Source Th./Ex.	Constraint	$\Delta \chi^2$ (CMSSM)	$\Delta \chi^2$ (NUHM1)	$\Delta \chi^2$ ("SM")
m_t GeV	43	173.2 ± 0.90	0.05	0.06	-
$\Delta \alpha_{had}^{(5)}(M_Z)$	42	0.02749 ± 0.00010	0.009	0.004	-
M _Z [GeV]	44	91.1875 ± 0.0021	2.7×10 ⁻⁵	0.26	-
Γ_Z [GeV]	26 / 44	$2.4952 \pm 0.0023 \pm 0.001_{SUSY}$	0.078	0.047	0.14
σ_{had} [nb]	26 / 44	41.540 ± 0.037	2.50	2.57	2.54
R_l	26 / 44	20.767 ± 0.025	1.05	1.08	1.08
$A_{\rm fb}(\ell)$	26 / 44	0.01714 ± 0.00095	0.72	0.69	0.81
$A_{\ell}(P_{\tau})$	26 / 44	0.1465 ± 0.0032	0.11	0.13	0.07
R _b	26 / 44	0.21629 ± 0.00066	0.26	0.29	0.27
R _c	26 / 44	0.1721 ± 0.0030	0.002	0.002	0.002
$A_{\rm fb}(b)$	26 / 44	0.0992 ± 0.0016	7.17	7.37	6.63
$A_{\rm fb}(c)$	26 / 44	0.0707 ± 0.0035	0.86	0.88	0.80
A_b	26 / 44	0.923 ± 0.020	0.36	0.36	0.35
Ac	26 / 44	0.670 ± 0.027	0.005	0.005	0.005
$A_{\ell}(SLD)$	26 / 44	0.1513 ± 0.0021	3.16	3.03	3.51
$\sin^2 \theta_w^{\epsilon}(Q_{fb})$	26 / 44	0.2324 ± 0.0012	0.63	0.64	0.59
M _W [GeV]	26 / 44	$80.399 \pm 0.023 \pm 0.010_{SUSY}$	1.77	1.39	2.08
$a_{\mu}^{nAP} - a_{\mu}^{om}$	53 / 42,54	$(30.2 \pm 8.8 \pm 2.0_{SUSY}) \times 10^{-10}$	4.35	1.82	11.19 (N/A)
M_h [GeV]	28 / 55,56	$> 114.4[\pm 1.5_{SUSY}]$	0.0	0.0	0.0
BR _{b→sy}	45 / 46	$1.117 \pm 0.076_{EXP}$ $\pm 0.082_{exp} \pm 0.050_{exerv}$	1.83	1.09	0.94
$BR(B_* \rightarrow \mu^+ \mu^-)$	29 / 41	CMS & LHCb	0.04	0.44	0.01
BREXP/SM	291 / 26	1.43 ± 0.43 ex p. mu	1.49	1.59	1.00
$BR(B_d \rightarrow \mu^+\mu^-)$	29 / 46	$< 4.6[\pm 0.01 \text{susy}] \times 10^{-9}$	0.0	0.0	0.0
BREXPISM	47/ 46	0.99 ± 0.32	0.02	≪ 0.01	≪ 0.01
BREAD/SM	291/148	1 008 ± 0 014 px p. ти	0.39	0.42	0.33
BR ^{EXV/SM}		< 4.5	0.0	0.0	0.0
$\Delta M_{\rm EXV/SM}^{\rm EXV/SM}$	202/ [22] [29] / [51] [59]	0.97 ± 0.01 even ± 0.27 even	0.02	0.02	0.01
	[29] / [46][51][52]	$1.00 \pm 0.01_{EXP} \pm 0.13_{SM}$	≪ 0.01	0.33	≪ 0.01
$\Delta \epsilon_{K}^{EXP/SM}$	49 / 51.52	$1.08 \pm 0.14_{EXP+TH}$	0.27	0.37	0.33
$\Omega_{CDM}h^2$	811 / 1131	$0.1120 \pm 0.0056 \pm 0.012$ stusy	8.4×10^{-4}	0.1	N/A
σ_{p}^{D1}	25	(m_{e0}, σ_p^{o1}) plane	0.13	0.13	N/A
iets + Br	18 20	(mo mo c) plane	1.55	2 20	N/A
$H/A H^{\pm}$	1211	$(M_A, \tan\beta)$ plane	0.0	0.0	N/A
Total x ² /d of	All	All	28.8/22	97 8/91	99 7/99 (91 5/99)
p-values	241	7411	15%	16%	9% (49%)

CMSSM: Evolution with time



SUSY Status – post 7 TeV LHC data

SUSY & DM Searches @ LHC 0. Buchmüller

- Constrained SUSY models like the CMSSM are severely put under pressure by the LHC limits!
- Experiments define new benchmarks and less complex SUSY models in order to present the interpretation of their searches.
- Aided by the discovery of a Higgs boson, the focus of the experimental search strategy and corresponding interpretation shifts towards other scenarios like "Natural SUSY" (i.e. 3rd generation squark searches).

Interpretation in Simplified Models



SMS: a few interesting features



How to summarize SMS limits?

Approach taken in the 2012 and 2013 Experimental SUSY PDG reviews [OB & Paul De Jong]:

http://pdg.lbl.gov/2012/reviews/rpp2012-rev-susy-2-experiment.pdf http://pdg.lbl.gov/2013/reviews/rpp2013-rev-susy-2-experiment.pdf

Model	Assumption	$m_{ ilde q}$	$m_{ ilde{g}}$
	$m_{ ilde{q}}pprox m_{ ilde{g}}$	1400	1400
CMSSM	all $m_{ ilde{q}}$	-	800
	all $m_{ ilde{g}}$	1300	-
Simplified model $\tilde{g}\tilde{g}$	$m_{ ilde{\chi}_1^0}=0$	-	900
	$m_{ ilde{\chi}_1^0} > 300$	-	no limit
Simplified model $\tilde{q}\tilde{q}$	$m_{ ilde{\chi}^0_1}=0$	750	-
	$m_{ ilde{\chi}_1^0} > 250$	no limit	-
Simplified model	$m_{ ilde{\chi}_1^0} = 0, m_{ ilde{q}} pprox m_{ ilde{g}}$	1500	1500
$ ilde{g} ilde{q}, ilde{g}ar{ ilde{q}}$	$m_{\tilde{\chi}_1^0} = 0$, all $m_{\tilde{g}}$	1400	-
	$m_{ ilde{\chi}_1^0}^{ ilde{\sim}1} = 0, ext{ all } m_{ ilde{q}}$	-	900



This was an appropriate approach for the rather limited amount of inclusive searches and corresponding SMS interpretations available in 2011 (7 TeV).

How to summarize SMS limits?

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Direct squark production – chosen limits



Signature: Jets + $E_t^{miss} + H_T$ Limit assumes all 1st & 2nd gen squarks to be mass degenerate or only one light squark!

Signature: 2 b-jets + E_{τ}^{mis}

Signature: 1Lepton + jets + E_{T}^{mis}





Gluino mediated squark production – limits chosen



Signature: Jets + H_T + E_t^{miss}

Signature: : Jets + b-tag + E_t^{miss}

Signature: 0/1 Leptons + $3 b-tag + E_t^{mis}$







SUSY & DM Searches @ LHC O. Buchmüller

Stop Searches Today



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Add pMSSM here



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Add pMSSM here


pMSSM Prediction & Present limits









Direct chargino/neutralino production









Dark Matter Searches: Direct Detection Experiments



Dark Matter Searches: Direct Detection Experiments



Direct Detection Landscape in a nutshell!



Direct Detection Landscape in a nutshell!



SUSY & Dark Matter: Evolution with time



SUSY & Dark Matter: Evolution with time



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Dark Matter Searches: Direct Detection vs Colliders



Direct Detection Experiments

DM-nucleus scattering



Collider Experiments

- Pair-production of DM
- missing energy signature



Mono-Mania (at the LHC)



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Mono-Mania (at the LHC)



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Monojet analyses better than direct detection?!



Claim [often made]: For low mass and the entire spin-dependent case monojet limits are stronger than direct detection limits!

Effective Field Theory (EFT) Interpretation

Example of considered operators:

$$O_V = \frac{(\bar{\chi}\gamma_\mu\chi)(\bar{q}\gamma_\mu q)}{\Lambda^2} \quad \text{Vector operator, s-channel}$$
$$(\bar{\chi}\gamma_\mu\gamma_5\chi)(\bar{q}\gamma_\mu\gamma_5q)$$

$$q$$

$$g_{q}$$

$$g_{q}$$

$$g_{\chi}$$

$$\bar{\chi}$$

_

 $O_{AV} = rac{(ar{\chi}\gamma_{\mu}\gamma_{5}\chi)(ar{q}\gamma_{\mu}\gamma_{5}q)}{\Lambda^{2}}$ Axial vector operator, s-channel

Assumption of EFT

If the operator (e.g. V or AV) mediator is suitably(!!) heavy it can be integrated out to obtain the effective V or AV contact operator. In this case (and only this case), the contact interaction scale Λ is related to the parameters entering the Lagrangian:

$$\Lambda = \frac{M_{mediator}}{\sqrt{g_q g_\chi}} \quad (4)$$

(relation in the full theory)

Validity of Effective Field Theory Limits





- Compare prediction of FT with EFT in $m_{med} m_{DM}$ plane. Three regions become visible:
 - Region I: EFT and FT agree better then 20% ➤ EFT is valid!
 - Region II: EFT yields significant weaker limits then FT
 > EFT limits are too conservative!
 - Region III: EFT yields significant stronger limits then FT
 ➢ EFT limits are too aggressive!

Validity of Effective Field Theory Limits



Minimal Simplified Dark Matter Model



MSDM: 4D Parameter Space

Mdм	M _{med}
gq	gdм

4-dimensional problem, projecting limits onto all 2-D plane:

- M_{DM} vs M_{med} assuming $g_q = g_{DM}$ and $g_q \neq g_{DM}$
- $M_{med} vs g_{q,}g_{DM}$ for fixed M_{DM}
- M_{DM} vs $g_{q,}g_{DM}$ for fixed M_{med}
- gq vs gdm for fixed Mdm, Mmed

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Collider vs Direct Detection





Collider vs Direct Detection



⁶⁵6

Projections for Future Experiments: M_{med} vs M_{DM}

arXiv:1409.4075 Limits from 8 TeV monojet search and projected limits for 3 LHC scenarios: - 13 TeV 30 fb⁻¹ - 14 TeV, 300 fb⁻¹ - 14 TeV, 3000 fb⁻¹

LUX 2013 limits and projected limits for LZ assuming 10 tonneyear exposure

Discovery reach accounting for coherent neutrino scattering



Projections for Future Experiments: σ vs M_{DM}



Direct Detection experiments and collider are complementary! They are probing different regions of the relevant parameter space!

Outlook: 8 TeV vs 14 TeV



Summary

SUSY & DM Searches @ LHC O. Buchmüller

> So far New Physics has not revealed itself!

- Even by 2010 the LHC has enter new territory for New Physics searches and since pushed e.g. the (coloured) SUSY mass scale to the ~1 TeV scale
- We were well prepared for an early discovery but we also knew that it could take more time and ingenuity before we can claim a discovery (if NP exist)
- The LHC experiments have established an impressive variety of very powerful direct searches for many different final states!
 - Based on these results we need to establish the "big picture" in order to understand find out if/where our search strategy might have weak spots or even holes!
 - This requires appropriate interpretations of the searches and a MEANIGFUL comparison with other experiments important example DM searches!
- The high energy running of the LHC starting 2015 will be our next very (as in VERY) real chance for discovery! The story continues ... stay tuned!

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BACKUP










ATLAS & CMS public results

All results presented in this talk (and many more) can be accessed via the public page of the ATLAS and CMS experiments:

ATLAS SUSY: https://twiki.cern.ch/twiki/bin/view/AtlasPublic/ SupersymmetryPublicResults

CMS SUSY :https://twiki.cern.ch/twiki/bin/view/CMSPublic/ PhysicsResultsSUS

ATLAS Summary

ATLAS Preliminary ATLAS SUSY Searches* - 95% CL Lower Limits Status: FPS 2013 $\int \mathcal{L} dt = (4.4 - 22.9) \text{ fb}^{-1}$ $\sqrt{s} = 7.8 \text{ TeV}$ e, μ , τ , γ Jets $\mathsf{E}_{\tau}^{\text{miss}}$ [\mathcal{L} dt[fb⁻¹] Model Mass limit Reference MSUGRA/CMSSM $m(\tilde{q})=m(\tilde{g})$ 2-6 iets Yes 20.3 ATLAS-CONF-2013-047 0 17 TeV MSUGRA/CMSSM 1 e, µ 3-6 jets Yes 20.3 1.2 TeV any $m(\tilde{q})$ ATLAS-CONF-2013-062 MSUGRA/CMSSM 7-10 jets any $m(\tilde{q})$ 0 Yes 20.3 1.1 TeV ATLAS-CONE-2013-054 Searches 2-6 jets 20.3 $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-047 $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ 0 Yes 740 GeV 0 2-6 iets Yes 20.3 1.3 TeV $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-047 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{\pm} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0}$ $\tilde{g}\tilde{g} \rightarrow qqqq\ell\ell(\ell\ell)\tilde{\chi}_{1}^{0}\tilde{\chi}_{1}^{0}$ 1 e, µ 3-6 jets Yes 20.3 1.18 TeV $m(\tilde{\chi}_{1}^{0}) < 200 \text{ GeV}, m(\tilde{\chi}_{1}^{\pm}) = 0.5(m(\tilde{\chi}_{1}^{0}) + m(\tilde{g}))$ ATLAS-CONF-2013-062 2 e, µ (SS) 3 jets Yes 20.7 1.1 TeV $m(\tilde{\chi}_{1}^{0}) < 650 \, \text{GeV}$ ATLAS-CONF-2013-007 Inclusive GMSB (Î NLSP) 2 e, µ 2-4 jets Yes 4.7 1.24 TeV $tan\beta < 15$ 1208 4688 $tan\beta > 18$ GMSB (*ℓ* NLSP) 1-2 τ 0-2 jets Yes 20.7 1.4 TeV ATLAS-CONF-2013-026 GGM (bino NLSP) $m(\tilde{\chi}_1^0) > 50 \text{ GeV}$ 2γ 0 Yes 4.8 1.07 TeV 1209.0753 GGM (wino NLSP) 619 GeV $m(\tilde{\chi}_1^0)$ >50 GeV $1 e, \mu + \gamma$ ATLAS-CONF-2012-144 0 Yes 4.8 GGM (higgsino-bino NLSP) 1 h Yes 4.8 900 GeV m(X10)>220 GeV 1211.1167 γ GGM (higgsino NLSP) 2 e, µ (Z) 0-3 jets Yes 5.8 690 GeV m(*H*)>200 GeV ATLAS-CONF-2012-152 mono-jet Yes 10.5 m(g)>10⁻⁴ eV ATLAS-CONF-2012-147 Gravitino LSP 0 645 GeV 20.1 $m(\tilde{\chi}_{1}^{0}) < 600 \, \text{GeV}$ gen. med. $\tilde{g} \rightarrow b \bar{b} \tilde{\chi}_{1}^{0}$ 0 3 b Yes 1.2 TeV ATLAS-CONE-2013-061 $\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_1^0$ 20.3 0 7-10 jets Yes 1.14 TeV $m(\tilde{\chi}_{1}^{0}) < 200 \, GeV$ ATLAS-CONE-2013-054 0-1 e,µ m(x10)<400 GeV 3 *b* Yes 20.1 1.34 TeV ATLAS-CONF-2013-061 $\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_1^0$ °ž a m($\tilde{\chi}_{1}^{0}$)<300 GeV $\tilde{g} \rightarrow b \bar{t} \tilde{\chi}_1^{\exists}$ 0-1 e, µ 3 b Yes 20.1 1.3 TeV ATLAS-CONF-2013-061 2 b $m(\tilde{\chi}_{1}^{0}) < 100 \, GeV$ $\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b \tilde{\chi}_1^0$ 0 Yes 20.1 100-630 GeV ATLAS-CONF-2013-053 $\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow t \tilde{\chi}_1^{\pm}$ 2 e. u (SS) 0-3 h Yes 20.7 430 GeV $m(\tilde{\chi}_1^{\pm})=2 m(\tilde{\chi}_1^0)$ ATLAS-CONE-2013-007 $\tilde{t}_1 \tilde{t}_1$ (light), $\tilde{t}_1 \rightarrow b \tilde{\chi}_1^{\pm}$ 1-2 e, µ 4.7 167 GeV $m(\tilde{\chi}_1^0)=55 \, GeV$ 1208.4305, 1209.2102 1-2 b Yes $\tilde{t}_1 \tilde{t}_1 (\text{light}), \tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0$ 2 e, µ 0-2 iets 20.3 $m(\tilde{\chi}_{1}^{0}) = m(\tilde{t}_{1}) - m(W) - 50 \text{ GeV}, m(\tilde{t}_{1}) < < m(\tilde{\chi}_{1}^{\pm})$ ATLAS-CONF-2013-048 Yes 220 GeV $\tilde{t}_1 \tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$ 2 e, µ 2 jets Yes 20.3 225-525 GeV $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-065 ť١ $\tilde{t}_1 \tilde{t}_1$ (medium), $\tilde{t}_1 \rightarrow b \tilde{\chi}_1^{\dagger}$ 0 2 b Yes 20.1 ťι 150-580 GeV $m(\tilde{\chi}_{1}^{0}) < 200 \text{ GeV}, m(\tilde{\chi}_{1}^{\pm}) - m(\tilde{\chi}_{1}^{0}) = 5 \text{ GeV}$ ATLAS-CONF-2013-053 1 e,μ Yes 20.7 $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-037 $\tilde{t}_1 \tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$ 1*b* 200-610 GeV ge $\tilde{t}_1 \tilde{t}_1$ (heavy), $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$ 0 2 b Yes 20.5 320-660 GeV $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-024 0 mono-jet/c-tag Yes 20.3 200 GeV $m(\tilde{t}_1)-m(\tilde{\chi}_1^0) < 85 \, GeV$ ATLAS-CONF-2013-068 $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow c \tilde{\chi}_1^0$ $\tilde{t}_1 \tilde{t}_1$ (natural GMSB) 2 e, µ (Z) Yes 20.7 500 GeV m(x10)>150 GeV ATLAS-CONF-2013-025 1 h $\tilde{t}_2 \tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$ 3 e, µ (Z) 1 *b* Yes 20.7 520 GeV $m(\tilde{t}_1)=m(\tilde{\chi}_1^0)+180 \text{ GeV}$ ATLAS-CONF-2013-025 $\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_1^0$ 20.3 85-315 GeV $m(\tilde{\chi}_1^0)=0 \text{ GeV}$ 2 e, µ 0 Yes ATLAS-CONF-2013-049 $\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow \tilde{\ell} \nu(\ell \tilde{\nu})$ 2 e, µ 0 Yes 20.3 $\mathfrak{m}(\tilde{\chi}_{1}^{0})=0 \text{ GeV}, \mathfrak{m}(\tilde{\ell}, \tilde{\nu})=0.5(\mathfrak{m}(\tilde{\chi}_{1}^{\pm})+\mathfrak{m}(\tilde{\chi}_{1}^{0}))$ ATLAS-CONF-2013-049 125-450 GeV $\begin{aligned} \tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \to \tilde{\tau} \nu (\tau \tilde{\nu}) \\ \tilde{\chi}_1^+ \tilde{\chi}_2^0 \to \tilde{\ell}_L \nu \tilde{\ell}_L \ell (\tilde{\nu} \nu), \ell \tilde{\nu} \tilde{\ell}_L \ell (\tilde{\nu} \nu) \end{aligned}$ 2τ Yes 20.7 180-330 GeV $m(\tilde{\chi}_{1}^{0})=0 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$ ATLAS-CONF-2013-028 目的 0 $m(\tilde{\chi}_{1}^{\pm})=m(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$ ATLAS-CONF-2013-035 3 e, µ 20.7 600 GeV 0 Yes $\tilde{\chi}_1^{\pm} \tilde{\chi}_2^{\bar{0}} \rightarrow W^* \tilde{\chi}_1^{\bar{0}} Z^* \tilde{\chi}_1^{\bar{0}}$ 3 e, µ 0 Yes 20.7 315 GeV $m(\tilde{\chi}_1^{\pm})=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$, sleptons decoupled ATLAS-CONF-2013-035 Direct $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$ $m(\tilde{\chi}_{1}^{\pm})-m(\tilde{\chi}_{1}^{0})=160 \text{ MeV}, \tau(\tilde{\chi}_{1}^{\pm})=0.2 \text{ ns}$ Disapp. trk 1 jet Yes 20.3 270 GeV ATLAS-CONF-2013-069 Stable, stopped g R-hadron 0 1-5 jets Yes 22.9 857 GeV $m(\tilde{\chi}_1^0)=100 \text{ GeV}, 10 \ \mu s < \tau(\tilde{g}) < 1000 \text{ s}$ ATLAS-CONF-2013-057 GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) \pm \tau(e, \mu)$ $1-2 \mu$ 0 15.9 475 GeV 10<tanβ<50 ATLAS-CONF-2013-058 GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}$, long-lived $\tilde{\chi}_1^0$ 2γ 0 Yes 4.7 230 GeV $0.4 < \tau(\tilde{\chi}_1^0) < 2 \text{ ns}$ 1304.6310 $1 \text{ mm} < c\tau < 1 \text{ m}, \tilde{g}$ decoupled 1210.7451 $\tilde{\chi}_1^0 \rightarrow qq\mu \text{ (RPV)}$ 1 *u* 0 Yes 44 700 GeV LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e + \mu$ 4.6 1.61 TeV $\lambda'_{311}=0.10, \lambda_{132}=0.05$ 1212.1272 2 e, µ 0 - $\lambda'_{311}=0.10, \lambda_{1(2)33}=0.05$ LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e(\mu) + \tau$ 1.1 TeV $1 e, \mu + \tau$ 0 4.6 1212.1272 Bilinear RPV CMSSM 1 e,µ 7 jets Yes 4.7 1.2 TeV $m(\tilde{q})=m(\tilde{g}), c\tau_{LSP} < 1 mm$ ATLAS-CONF-2012-140 PV $$\begin{split} &\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow e e \tilde{v}_{\mu}, e \mu \tilde{v}_{e} \\ &\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow \tau \tau \tilde{v}_{e}, e \tau \tilde{v}_{\tau} \end{split}$$ 4 e,μ 0 Yes 20.7 760 GeV $m(\tilde{\chi}_{1}^{0})>300 \text{ GeV}, \lambda_{121}>0$ ATLAS-CONF-2013-036 $3 e, \mu + \tau$ $m(\tilde{\chi}_{1}^{0})>80 \text{ GeV}, \lambda_{133}>0$ 0 Yes 20.7 350 GeV ATLAS-CONE-2013-036 66 GeV 0 6 jets 4.6 1210.4813 *φ̃*→aaa $\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$ 2 e, µ (SS) 0-3 b Yes 20.7 880 GeV ATLAS-CONF-2013-007 Scalar gluon 0 4 jets 4.6 100-287 GeV Other sqluon 1210,4826 incl. limit from 1110.2693 WIMP interaction (D5, Dirac χ) 0 mono-jet Yes 10.5 m(x)<80 GeV, limit of<687 GeV for D8 ATLAS-CONF-2012-147 704 GeV 10⁻¹ 1 $\sqrt{s} = 7 \text{ TeV}$ $\sqrt{s} = 8 \text{ TeV}$ $\sqrt{s} = 8 \text{ TeV}$ Mass scale [TeV] full data partial data full data

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 σ theoretical signal cross section uncertainty.

SUSY & DM Searches @ LHC O. Buchmüller

CMS Summary



Summary



- LHC&Experiments are on track for first collisions in 2008
 - Challenge: commissioning of machine and detectors of unprecedented complexity, technology and performance
- The LHC will discover (or exclude) the Higgs by ~2010
 - Electro Weak Symmetry Breaking
 - Large phase space can already be excluded with only ~1fb⁻¹
- The LHC will discover low energy SUSY (if it exists)
 - Could be easy; could also take more time and ingenuity before we can claim a discovery
 - First signals might emerge already in the first data
 - 1-3 TeV can be covered already with <10fb⁻¹
- The LHC will cover a new physics scale of 1-3 TeV
 - Many new physics models; Black hole, Extra Dimensions,Little Higgs, Split Susy, New Bosons, Technicolour, etc ...

In other words; the next five years will be an exciting time for particle physics ...

(Best) mass limits in a nutshell (RP conserving)

Direct squark		$\tilde{q} ightarrow q \chi_1^0$	$\tilde{u}_L o q \chi_1^0$	$\tilde{b} \to b \chi_1^0$	$\tilde{t} \to t \gamma$	χ^0_1	coloured sparticle		iclo
Best limit: [GeV]		~850	~500	~650	~650		production		
No limit for M _{LSP} [GeV]		~ 300	~120	~270	~26	60			
Direct squark		$\tilde{g} \to q \bar{q} \chi_1^0$	$\tilde{g} \to b \bar{b} \chi_1^0$	$\left \begin{array}{c} \tilde{g} \to t\bar{t}\chi \end{array} \right $	χ_1^0	M _{stop}	Stop M _{Lsp} < M _{top}	$\tilde{t} \to c \chi_1^0$	$\tilde{t} \to W b \chi_1^0$
Best limit [GeV]	imit: ~1200 ~1200 V]		~1200	~1400	I	В	est limit: [GeV]	~240	~320
No limit for M _{LSP} [GeV]		~480	~650	~700		Ne M	o limit for _{LSP} [GeV]	~210	~190

EWK sparticle production

Direct slepton	$\tilde{l}_L o l^{\pm} \chi_1^0$	$\tilde{l}_R \to l^{\pm} \chi_1^0$	$\chi_1^{\pm}\chi_2^0$	$light\; ilde{l}$	$heavy\; ilde{l}$		
Best limit: [GeV]	~300	~240	Best limit: [GeV]	~750	~300		
No limit for M _{LSP} [GeV]	~150	~90	No limit for M _{LSP} [GeV]	~350	~60		



Many Different Kinematic Variables



Multijets & missing energy search



2010: Entering New Territory at the LHC!



Simplified Model Spectra (SMS)



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Today: many more SMS and many more searches



(Minimal) Natural SUSY Spectrum



What do we call a "SUSY search"?

The definition is purely derived from the experimental signature. Therefore, a "SUSY search signature" is characterized by Lots of missing energy, many jets, and possibly leptons in the final state



Missing Energy: • from LSP

<u>Multi-Jet:</u> • from cascade decay (gaugino)

Multi-Leptons:

from decay of charginos/neutralios

RP-Conserving SUSY is a very prominent example predicting this famous signature but ...

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What is its experimental signature?

. by no means is it the only New Physics model predicting this experimental pattern. Many other NP models predict this genuine signature



Missing Energy:

• Nwimp - end of the cascade

Multi-Jet:

• from decay of the Ns (possibly via heavy SM particles like top, W/Z)

Multi-Leptons:

• from decay of the N's

Model examples are Extra dimensions, Little Higgs, Technicolour, etc but a more generic definition for this signature is as follows.

Rediscovery of the SM at a new energy frontier





Combination vs individual search



Natural SUSY: universal limits

If the gluino mass OR 3G mass lies in the red band, the point is excluded. If the gluino mass AND 3G mass lie in the yellow band the point may or may not be excluded. Otherwise the point is not excluded.



Combining with the latest published 8 TeV results:

Outlook: 8 TeV vs 14 TeV

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Use 30/fb for 2011/2012 for comparison

Higgs: pp \rightarrow H, H \rightarrow WW, ZZ and $\gamma\gamma$ mainly gg: factor ~2

SUSY – 3rd Generation: Mass scale ~ 500 GeV qq and gg: factor ~3 to 6

SUSY – Squarks/Gluino: Mass scale ~ 1.5 TeV qq,gg,qg: factor ~40 to 80

Z': Mass scale ~ 5 TeV qq: factor ~1000

Higgs: 15/fb@14 TeV to match 2011/2012 mainly gg: factor ~2

SUSY – 3rd Generation: 5/fb to 10/fb@14 TeV to match 2011/2012 qq and gg: factor ~3 to 6

SUSY – Squarks/Gluino: 0.4/fb to 0.8/fb@14 TeV to match 2011/2012 qq,gg,qg: factor ~40 to 80

Z' : O(1/pb) @14 TeV to match 2011/2012 qq: factor ~1000

RP-violating searches/interpretation



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RP violation searches: Summary



Like RP conserving searches, these searches are also probing the 1 TeV scale and even beyond!

Long-lived particle (SUSY) searches



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Dedicated searches for direct stop-pair production



Dedicated searches for direct stop-pair production



ATLAS-CONF-2013-065:

Scalar stop analysis with two leptons in the final state using a MVA technique.



Dedicated searches for direct stop-pair production



SMS limits: A word of caution!



Combining Searches = less model dependence



Combining Searches = less model dependence



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Validity of Effective Field Theory Limits



What those this imply on model-dependences of EFT limits?



This together with the observation that all DM theories for which the EFT is valid must have $m_{med} < \Gamma_{med}$ leads to the conclusion the the EFT only applies to a very (as in VERY) small class of DM models. EFT limits of monojet searches are therefore highly model-depended!

Why is SUSY so attractive?

 $\Delta m_H = f(m_B^2 - m_f^2)$

1. Quadratically divergent quantum corrections to the Higgs boson mass are avoided

(Hierarchy or naturalness problem)

- 2. Unification of coupling constants of the three interactions seems possible
- 3. SUSY provides a candidate for dark matter,



The lightest SUSY particle (LSP)

4. A SUSY extension is a small perturbation, consistent with the electroweak precision data



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A "typical" SUSY Spectrum

SUSY & DM Searches @ LHC O. Buchmüller Use the famous SPS1a benchmark point for illustration $[m_0=100, m_{12}=250, \tan\beta=10, A_0=-100, \mu>0]$ **CMSSM** $m_0, m_{1/2}, \tan\beta, A_0, sign(\mu)$ gluino/ 700 squarks GeV Mass / Advantage: 600 \tilde{t}_2 Only four free $ilde{b}_2 \\ ilde{b}_1$ charginos/ parameters (when 500 neutralinos $sign(\mu)$ fixed) One of the most \geq $H^0_{A^0}$ H^{\pm} 400 \tilde{t}_1 $\tilde{\chi}_4^0$ $\tilde{\chi}_3^0$ studied incarnations of the MSSM Higgs 300 sleptons sector Disadvantage: Not generally 200 \geq representative of SUSY (e.g. fixed h^0 LSP 100 mass relation between M_{aluion} and 0 M_{LSP})

Stop decay to charm and neutralino



ATLAS-CONF-2013-068:
Two different selections:
➤ Monojet-like selection
to cover region close to 'diagonal"
➤ MVA based c-tag selection
for remaining region



visible XS: εσ=0.7fb signal obs: 13 events signal exp: 14⁺⁵ events CL_B: 0.45 Monojet 95% excl. visible XS: $E\sigma = 136 fb$ signal obs: 2770 events signal exp: 2060+780-560 CL_R: 0.86

c-Tag:

95% excl.

Rediscovery of the SM at a new energy frontier




Dark Matter from invisible Higgs searches



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Dark Matter from invisible Higgs searches



What those this imply on model-dependences of EFT limits?



Look at EFT validity in m_{DM} – coupling* plane!



Look at EFT validity in m_{DM} – coupling* plane!

1. Region in which EFT is valid

For this we calculate the minimum coupling

$$\sqrt{g_q g_\chi} = m_{med} / \Lambda_{CMS}$$

that the simplified model must have for the EFT limits to apply. This is defined by region I (i.e. better then 20% agreement of FT and EFT).



Look at EFT validity in m_{DM} – coupling* plane!

1. Region in which EFT is valid (20%)

2. Require compatibility with relic density

When exclude the region in which relic abundance is larger then the observed value of $\Omega_{\chi\chi}h^2 = 0.119$ only mediator masses above a few hundred GeV fulfill this.



EFT limits of monojet searches only apply to a very (as in VERY) small class of DM models!



See discussion about equation 3.5 in arXiv:1308.6799 for further details.

What those this imply on model-dependences of EFT limits?



The observation that all DM theories for which the EFT is valid must have m_{med} < Γ_{med} and the small class to models it applies in any case leads to the conclusion the EFT only applies to a very small class of DM models. EFT limits of monojet searches are therefore highly model-depended!

Alternative Interpretation Ansatz: Simplified models



After three years of operation at the LHC the landscape for interpretation of searches has changed dramatically – new superior & modern approaches have replaced in many areas longstanding traditional ones (e.g. SUSY searches)



SUSY & DM Searches @ LHC O. Buchmüller









The proposal





The proposal



 $\bar{\chi}$

Beyond EFT limits: Simplified models

Working out the complementarity between direct DM detection experiments and collider based DM searches! 10^{-35} DM-proton cross section [cm²] Region I 10⁻³⁶ CMS limit 10⁻³⁷ 1000 Region II stronge mmed [GeV] 10⁻³⁸ COUPP Region III 10-39 100 $(\overline{\chi} \gamma^{\mu} \gamma^{5} \chi) (\overline{q} \gamma_{\mu} \gamma^{5} q)$ PICASSO, SIMPLE or 10⁻⁴⁰ CMS (EFT COUPP limit stronger 10^{-41} 101000 100 100 10 1000 10m_{DM} [GeV] $m_{\rm DM}$ [GeV]

EFT limits give the impression that monsjet searches outperform direct detection BUT EFT only applies a VERY small class of DM models. Simplified model limits give a much better Account of the REAL complementarity and thus seem superior for a comparison. Buchmüller

Beyond EFT limits: Simplified models

Working out the complementarity between direct DM detection experiments and collider based DM searches!



EFT limits give the impression that monsjet searches outperform direct detection BUT EFT only applies a VERY small class of DM models.

Simplified model limits give a much better Account of the REAL complementarity and thus seem superior for a comparison.

Monojet and Monophoto (plus E_T^{miss})



Monojet and Monophoto (plus E_T^{miss})



Dedicated searches for direct stop-pair production



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