A Facility for Hidden Sector exploration at SPS CERN



SHiP Collaboration as on 9th February 2015 (List of Institutes signing SHIP Technical Proposal)

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^cUniversità di Roma La Sapienza, Roma, Italy

Includes a contribution from 2 French groups signing the TP,

and acknowledgement to a personal contribution from Maxim Titov

Some history and current status

- ✓ Oct 2013: submitted our EOI: CERN-SPSC-2013-024 ; arXiv:1310.1762 ; SPSC-EOI-010
- ✓ January 2014: EOI discussed at SPSC

Encouraged to produce "an extended proposal with further developed physics goals, a more detailed technical design and a stronger collaboration."

✓ Work towards Technical Proposal in full swing

Extension of physics program Signal background studies and optimization Detector specification, simulation and even some detector R&D Optimization of Experimental Facility - beam line, target, and muon filter, RP, overall layout

- 1st SHiP Workshop in Zurich in June with a 100 experimentalists and theorists
 41 institutes from 14 countries expressed interest to contribute to the Technical Proposal
- 2nd SHiP Workshop/Collaboration meeting at CERN September 24-26 Revise progress in Working Groups towards Technical Proposal Extend physics for a general purpose facility: Tau neutrino, LFV and direct Dark Matter search
- ✓ 3rd SHiP Collaboration meeting at CERN December 15

Revise progress towards TP and Physics Proposal (PP). Formalize Collaboration as proposed by CERN management with 44 institutes (14 countries)

 4th SHiP Collaboration meeting in Naples, February 9-11 Finalize contents for TP and PP

✓ First drafts of TP and PP distributed for comments, end of February

Next steps: schedule of the SHIP facility

A few milestones:

✓ Technical proposal

 \rightarrow March 2015

 \rightarrow 2018 - 2022

 \rightarrow 2018

- CERN to decide on the strategy for the SHIP beam within a year after TP submission
- ✓ Technical Design Report
- Construction and installation
- Data taking and analysis of 2×10^{20} pot → 2023 - 2027++

CERN-SPSC-2015-XXX 23:59:59, 23 February 2015, v0.1

r A new general purpose fixed target facility is proposed at the CERN SPS accelerator which is aimed at exploring the domain of hidden particles and make measurements with tau neutrinos. Hidden particles are predicted by a large number of recently elaborated models of Hidden Sectors. The high 10 intensity of the SPS 400 GeV/c beam allows probing a wide variety of models with light long-lived ¹¹ exotic particles with masses below O(10) GeV/c², including feebly interacting low-energy phenomena 12 of SUSY. Moreover, the experimental programme of the facility may be extended in the future to The facility is serviced by a new dedicated beam line which is branching off the splitter section 15 in the North Area, followed by a new target station and a magnetic shield to suppress beam induced 18 background. In the initial phase, the facility will host a detector to search for hidden particles in 17 combination with a compact tau neutrino detector. The hidden particle detector consists of a long 18 evacuated decay volume ending with a magnetic spectrometer, calorimeters and muon detectors to 10 allow full reconstruction and particle identification of hidden particle decays. The neutrino detec-20 tor consists of an emulsion target equipped with tracking in a magnetic field followed by a muon The current SPS is capable of providing an integrated total of 2.10²⁰ protons on target in five 22 years of operation in nominal conditions. This allows accessing a significant fraction of the unexplored 24 parameter space for Hidden Sectors with sensitivities which are several orders of magnitude better than previous experiments. The associated neutrino detector allows performing a number of unique

EUROPEAN ORGANIZATION FOR NUCLEAR RESEARCH (CERN)

The SHiP facility is providing a unique experimental platform for physics at the Intensity Frontier

29 which is complementary to the searches for New Physics at the Energy Frontier.

¹Authors are listed on the following pages.

Introduction

- ✓ SHiP is a general purpose fixed target facility to explore the domain of hidden particles (HP) with masses below O(10) GeV. HP are predicted in many models explaining known shortcoming of the SM as described in Physics Proposal
- ✓ SHiP will use high intensity spills at 400 GeV similar to CNGS programme Full exploitation of the protons available at SPS after the demands of LHC and other fixed target experiments !
- ✓ SPS beam is ideal in terms of its intensity and energy in maximizing the sensitivity reach for HP in charm decays
- ✓ Also ideal for v_{τ} physics ($D_{S} \rightarrow \tau v_{\tau}$)

- ✓ Brief comparison with fixed target experiment at FNAL and KEK beams, and with colliding experiment at LHC running at √s = 14 TeV assuming a luminosity of 1000 fb⁻¹
- ✓ The scope of the proposed facility is wider than the physics objectives of SHiP, e.g. searches for LFV $\tau \rightarrow 3\mu$, direct search for DM require dedicated experiments

Introduction

- ✓ SHiP is complementary to the other running or planned projects searching for NP
 - direct searches by ATLAS and CMS at \sqrt{s} =14 TeV
 - (perhaps at FCC in pp-mode later)
 - indirect searches in flavour physics: LHCb, BELLE 2, NA62 (in kaon sector)
 - also muon anomalous magnetic moment, LFV searches, proton decay etc...
- ✓ Anticipate a recommendation by the SPSC committee and a decision by CERN on the strategy for SHiP within a year of the submission of TP
- ✓ The remaining choices of baseline technologies will be made in time to submit Technical Design Reports by 2018
- ✓ The detector construction, five years of data taking, and the data analysis of 2×10²⁰ p.o.t. can be achieved in ~10 years

Physics Motivation

SM is great but it is not a complete theory

Experimental facts of BSM physics

- Neutrino masses & oscillations
- The nature of non-baryonic Dark Matter
- Excess of matter over antimatter in the Universe
- Cosmic inflation of the Universe

Theoretical shortcomings

Gap between Fermi and Planck scales, Dark Energy, connection to gravity, resolution of the strong CP problem, the naturalness of the Higgs mass, the pattern of masses and mixings in the quark and lepton sectors, ...

No clear guidance on the scale of NP and on its coupling strength to the SM particles !

An example: See-saw generation of neutrino masses

Most general renormalisable Lagrangian of all SM particles (+3 singlets wrt the SM gauge group):

$$L_{\text{singlet}} = i\bar{N}_{I}\partial_{\mu}\gamma^{\mu}N_{I} - Y_{I\alpha}\bar{N}_{I}^{c}\tilde{H}L_{\alpha}^{c} - M_{I}\bar{N}_{I}^{c}N_{I} + \text{h.c.},$$

Yukawa term: mixing of N_I with active neutrinos to explain oscillations

Majorana term which carries no gauge charge

The scale of the active neutrino mass is given by the see-saw formula: $m_{\nu} \sim$ where $m_D \sim Y_{I\alpha}v$ - typical value of the Dirac mass term

Example:

For $M \sim 1$ GeV and $m_v \sim 0.05$ eV it results in $m_D \sim 10$ keV and Yukawa coupling $\sim 10^{-7}$

Smallness of the neutrino mass hints either on very large M or very small $Y_{l\alpha}$

BSM theories with a new energy scale (which may also have light particles)

SUSY is an example

- ✓ From naturalness expect SUSY masses comparable to the Higgs mass to avoid significant fine tuning
- ✓ Could still have light NP. SUSY breaking may be accompanied by s-goldstinos (P,S) with couplings ~ 1 / (SUSY breaking scale)
 → may have evaded detection

e.g. $D \to \pi X$, then $X \to l^+ l^-$ [Gorbunov, 2001, LHCb: Aaij et al., 2013] $B_s^0 \to P S$, then $P, S \to \mu^+ \mu^-$

✓ *R*-parity violating neutralinos in some SUSY models

e.g., [Dedes et al., 2001]

e.g. $D \to l \, \tilde{\chi}$, then $\tilde{\chi} \to l^+ \, l^- \, \nu$

Models with Dark sector: impressive list of ideas in the past (pioneered by Bjorken and Okun)

Holdom, Galison, Manohar, Arkani-Hamed, Weiner, Schuster, Essig, Pospelov, Toro, Batell, Ritz, Andreas, Goodsell, Abel, Khoze, Ringwald, Fayet, Cheung, Ruderman, Wang, Yavin, Morrissey, Poland, Zurek, Reece, Wang, ...

- ✓ Hidden particles are singlets with respect to the SM gauge group
- ✓ Very weak interactions with the SM particles through portals described by various operators with vector, Higgs, neutrino and axion forms
- ✓ Isolated Dark sector naturally provides DM candidates
- Dark Sector may have a rich structure of light hidden messengers between Dark sector and SM particles
- ✓ No theoretical input on the Dark Sector mass scale → may well happen to be accessible at future experiments

Vector portal

Mirror matter: to restore P, C and CP

Okun, Voloshin, Ellis, Schwarz, Tyupkin, Kolb, Seckel, Turner, Georgi, Ginsparg, Glashow, Foot, Volkas, Blinnikov, Khlopov, Gninenko, Ignatiev, Berezhiani,...

Dim 2: Hypercharge U(1) field, $B_{\mu\nu}$ New particle – massive vector photon (paraphoton, secluded photon, ...) renormalisable coupling – kinetic mixing $\rightarrow \varepsilon B_{\mu\nu}F^{\mu\nu}$ 10⁻²

Production: through a virtual photon: electron or proton fixed-target experiments, e^+e^- and hadron colliders, $\sigma \propto \epsilon^2$. Decay due to the mixing with photon to the pair of charged particles:

 e^+e^- , $\mu^+\mu^-$, $\pi^+\pi^-$, *etc*, *etc* or to invisible particles from the dark sector.

Constraints are coming from:

- SLAC and Fermilab beam dump experiments E137, E141, E774
- electron and muon anomalous magnetic moments
- KLOE, BaBar
- PS191, NOMAD, CHARM (CERN)

Higgs portal

Patt, Wilczek, Schabinger, Wells, No, Ramsey-Musolf, Walker, Khoze, Ro, Choi, Englert, Zerwas, Lebedev, Mambrini, Lee, Jaeckel, Everett, Djouadi, Falkowski, Schwetz, Zupan, Tytgat, Pospelov, Batell, Ritz,

Dim 2: Higgs field $H^{\dagger}H^{Bezrukov}$, Gorbunov, Gunion, Haber, Kane, Dawson,...

New particle – hidden (dark) scalar with renormalisable coupling

 $(\mu\chi+\lambda\chi^2)H^{\dagger}H$

$$\begin{pmatrix} H \\ h \end{pmatrix} = \begin{pmatrix} \cos \rho & -\sin \rho \\ \sin \rho & \cos \rho \end{pmatrix} \begin{pmatrix} \phi_0' \\ S' \end{pmatrix}$$

 10^{0}

- Convenient parameterization of an extended Higgs sector: two Higgs doublets, SUSY (e.g. light s-goldstino, scalar singlets, Higgs triplets, ...
- \checkmark Extra scalars may help in solving hierarchy problem, flavour problem, baryogenesis, Dark Matter, neutrino masses, inflaton, etc
- ✓ Production

- direct:
$$p+target \rightarrow \chi X$$

- in flavour decays: $B \rightarrow \chi K^*$
- $\tau H sin^2 \rho$

Neutrino portal

Minkowski, Yanagida, Gell-Mann, Ramond, Slansky, Glashow, Mohapatra, G. Senjanovic + too many names to write, the whole domain of neutrino physics

Dim 5/2 Higgs-lepton $H^T L$ renormalisable coupling $\rightarrow Y H^T \overline{N} L$

To set the scene (summary of the past results) :

from Atre et. al. (0901.3589)

Axion portal

Weinberg, Wilczek, Witten, Conlon, Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell, Cicoli, Goodsell, Ringwald, Lazarides, Shafi, Choi, Essig, Harnik, Kaplan, Toro, Gorbunov,...

Dim 4: Axion-like Particles, pseudoscalars Non-renormalisable couplings $\rightarrow \frac{a}{F}G_{\mu\nu}\tilde{G}^{\mu\nu}, \quad \frac{\partial_{\mu}a}{F}\bar{\psi}\gamma_{\mu}\gamma_{5}\psi$

- \checkmark Axions \rightarrow to solve strong CP-problem, string theory, extra dimensions
- ✓ Axion-Like Particles (or pseudo-Nambu-Goldstone bosons) → dark matter, SUSY …

Similar to the Higgs portal, from arXiv:1008.0636, Essig et al

BSM theories with no NP between Fermi and Planck scales (minimalistic approach)

vMSM (T.Asaka, M.Shaposhnikov PL B620 (2005) 17) explains all experimental short comings of the SM at once by adding 3 HNL: N₁, N₂ and N₃

N = Heavy Neutral Lepton - HNL

Role of N_1 with mass in keV region: dark matter Role of N_2 , N_3 with mass in 100 MeV – GeV region: "give" masses to neutrinos and produce baryon asymmetry of the Universe Role of the Higgs: give masses to quarks, leptons, Z and W and inflate the Universe.

Experimental and cosmological constraints

✓ Recent progress in cosmology

- ✓ The sensitivity of previous experiments did not probe the interesting region for HNL masses above the kaon mass
- ✓ Use heavy flavour and W,Z decays to extend mass sensitivity reach

Main goal of current and near future experiments: exploration of the Fermi scale

(and a bit beyond)

Experiments at the both energy and intensity frontiers are essential !

hidden sector:

HNL: baryon asymmetry of the Universe, dark matter, neutrino masses sgoldstino, light neutralino: SUSY paraphoton: mirror matter, dark matter

Energy Frontier: No sign of New Physics yet ! (or what has yet not been found)

ATLAS SUSY Searches* - 95% CL Lower Limits Status: ICHEP 2014

ATLAS Preliminary $\sqrt{s} = 7.8$ TeV

	Model	e,μ,τ,γ	Jets	$E_{ m T}^{ m miss}$	∫ <i>L dt</i> [fb	⁻¹] Mass limit	Reference
Inclusive Searches	$ \begin{array}{l} \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \overline{q}\bar{q}, \overline{q} \rightarrow q \overline{r}_{1}^{0} \\ \overline{g}\bar{s}, \overline{s} \rightarrow q q \overline{r}_{1}^{0} \\ \overline{g}\bar{s}, \overline{s} \rightarrow q q \overline{r}_{1}^{0} \\ \overline{g}\bar{s}, \overline{s} \rightarrow q q \overline{r}_{1}^{0} \rightarrow q q W^{\pm} \overline{\xi}_{1}^{0} \\ \overline{g}\bar{s}, \overline{s} \rightarrow q q (\ell \ell_{N}/w) \overline{\xi}_{1}^{0} \\ \overline{GMSB} (\tilde{\ell} \text{ NLSP}) \\ \overline{GMSB} (\tilde{\ell} \text{ NLSP}) \\ \overline{GGM} (bino \text{ NLSP}) \\ \overline{GGM} (hing \text{ NLSP}) \\ \overline{GGM} (hing \text{ NLSP}) \\ \overline{GGM} (hing \text{ Sino } \text{ NLSP}) \\ \overline{GGM} (hing \text{ Sino } \text{ NLSP}) \\ \overline{GGM} (hing \text{ Sino } \text{ NLSP}) \\ \overline{GaM} (hing \text{ NLSP} (hing \text{ NLSP} (hing \text{ NLSP}) \\ \overline{GaM} (hing \text{ NLSP} (hing \text{ NLSP} (hing \text{ NLSP} (hing \text{ NLSP}) \\ \overline{GaM} (hing \text{ NLSP} (hing$	$\begin{matrix} 0 \\ 1 \ e, \mu \\ 0 \\ 0 \\ 1 \ e, \mu \\ 2 \ e, \mu \\ 1 - 2 \ r, \mu - 1 - 1 \ \ell \\ 2 \ \gamma \\ 1 \ e, \mu + \gamma \\ \gamma \\ 2 \ e, \mu \ (Z) \\ 0 \end{matrix}$	2-6 jets 3-6 jets 7-10 jets 2-6 jets 2-6 jets 3-6 jets 0-3 jets 0-2 jets - 1 b 0-3 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 4.7 20.3 20.3 4.8 4.8 5.8 10.5	$q.k$ 1.7 TeV $m(\hat{q})=m(\hat{g})$ \hat{g} 1.2 TeV $any m(\hat{q})$ \hat{g} 1.1 TeV $any m(\hat{q})$ \hat{g} 1.1 TeV $any m(\hat{q})$ \hat{g} 1.1 TeV $any m(\hat{q})$ \hat{g} 1.33 TeV $m(\xi^n)=0 \text{ GeV} m(1^{ut} \text{ gen. } \hat{q})=m(2^{ut} \text{ gen. } \hat{q})$ \hat{g} 1.33 TeV $m(\xi^n)=0 \text{ GeV} m(\xi^n)=0.5(m(\tilde{\chi}^n)+m(\tilde{\chi}))$ \hat{g} 1.12 TeV $m(\xi^n)=0 \text{ GeV} m(\xi^n)=0.5(m(\tilde{\chi}^n)+m(\tilde{\chi}))$ \hat{g} 1.24 TeV $tan\beta > 20$ \hat{g} 619 GeV $m(\xi^n_1)>50 \text{ GeV}$ R^n 690 GeV $m(\chi^n_1)>220 \text{ GeV}$ R^n 690 GeV $m(\chi^n_1)>200 \text{ GeV}$ $g(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1))$ $m(\chi^n_1)>50 \text{ GeV}$ $m(\chi^n_1)>50 \text{ GeV}$ R^n 690 GeV $m(\chi^n_1)>220 \text{ GeV}$ $m(\chi^n_1)>200 \text{ GeV}$ $g(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m(\chi^n_1)+m(\chi^n_1)=0.5(m$	1405.7875 ATLAS-CONF-2013-062 1308.1841 1405.7875 1405.7875 ATLAS-CONF-2013-062 1208.4688 1407.0603 1208.4688 1407.0603 ATLAS-CONF-2014-001 ATLAS-CONF-2014-01 ATLAS-CONF-2012-144 1211.1167 ATLAS-CONF-2012-152
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EW direct	$ \begin{array}{c} \tilde{l}_{1,\mathbf{k}}\tilde{\ell}_{1,\mathbf{k}},\tilde{\ell} \rightarrow \ell\tilde{\kappa}_{1}^{0} \\ \tilde{k}_{1}^{*}\tilde{\kappa}_{1}^{*},\tilde{\kappa}_{1}^{*} \rightarrow \ell\nu(\tilde{r}) \\ \tilde{\kappa}_{1}^{*}\tilde{\kappa}_{1}^{*},\tilde{\kappa}_{1}^{*} \rightarrow \ell\nu(\tilde{r}) \\ \tilde{\kappa}_{1}^{*}\tilde{\kappa}_{2}^{*} \rightarrow \tilde{\ell}_{1}\nu_{1}^{\ell}\ell(\tilde{r})\nu, \ell\tilde{s}\tilde{\ell}_{1}\ell(\tilde{r}) \\ \tilde{\kappa}_{1}^{*}\tilde{\kappa}_{2}^{*} \rightarrow W\tilde{\kappa}_{1}^{0}\tilde{\kappa}_{1}^{*} \\ \tilde{\kappa}_{1}^{*}\tilde{\kappa}_{2}^{*} \rightarrow W\tilde{\kappa}_{1}^{0}\tilde{\kappa}_{1}^{*} \\ \tilde{\kappa}_{2}^{*}\tilde{\kappa}_{2}^{*} \rightarrow W\tilde{\kappa}_{1}^{0}\tilde{\kappa}_{1}^{*} \\ \tilde{\kappa}_{2}^{*}\tilde{\kappa}_{2}^{*} \rightarrow W\tilde{\kappa}_{1}^{0}\tilde{\kappa}_{1}^{*} \\ \tilde{\kappa}_{2}^{*}\tilde{\kappa}_{2}^{*} \rightarrow \tilde{\kappa}_{1}\ell \end{array} $	$\begin{array}{c} 2 \ e, \mu \\ 2 \ e, \mu \\ 2 \ \tau \\ 3 \ e, \mu \\ 2 \text{-} 3 \ e, \mu \\ 2 \text{-} 3 \ e, \mu \\ 1 \ e, \mu \\ 4 \ e, \mu \end{array}$	0 0 - 0 2 <i>b</i> 0	Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 20.3	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	1403.5294 1403.5294 1407.0350 1402.7029 1403.5294, 1402.7029 ATLAS-CONF-2013-093 1405.5086
Long-lived particles	$\begin{array}{l} \text{Direct} \tilde{\chi}_1^+ \tilde{\chi}_1^- \text{ prod., long-lived } \tilde{\chi}_1^\pm\\ \text{Stable, stopped } \tilde{g} \text{ R-hadron}\\ \text{GMSB, stable } \tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e,\\ \text{GMSB,} \tilde{\chi}_1^0 \rightarrow \gamma \tilde{G}, \text{ long-lived } \tilde{\chi}_1^0\\ \tilde{q}\tilde{q}, \tilde{\chi}_1^0 \rightarrow qq\mu \text{ (RPV)} \end{array}$	Disapp. trk 0 μ) 1-2 μ 2 γ 1 μ, displ. vtx	1 jet 1-5 jets - -	Yes Yes - Yes -	20.3 27.9 15.9 4.7 20.3	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	ATLAS-CONF-2013-069 1310.6584 ATLAS-CONF-2013-058 1304.6310 ATLAS-CONF-2013-092
NAR	$\begin{array}{l} LFV pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e + \mu \\ LFV pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e(\mu) + \tau \\ Bilinear RPV CMSSM \\ \tilde{X}_{1}^{\dagger} \tilde{x}_{1}^{\dagger}, \tilde{X}_{1}^{\dagger} \rightarrow WX_{0}^{\dagger}, \tilde{X}_{1}^{\dagger} \rightarrow ee\tilde{\nu}_{\mu}, e\mu\tilde{\nu}_{e} \\ \tilde{X}_{1}^{\dagger} \tilde{x}_{1}, \tilde{X}_{1}^{\dagger} \rightarrow WX_{0}^{\dagger}, \tilde{X}_{1}^{\dagger} \rightarrow ee\tilde{\nu}_{\mu}, e\mu\tilde{\nu}_{e} \\ \tilde{X}_{1}^{\dagger} \tilde{x}_{1}, \tilde{X}_{1}^{\dagger} \rightarrow WX_{0}^{\dagger}, \tilde{X}_{1}^{\dagger} \rightarrow \tau\tau\tilde{\nu}_{e}, e\tau\tilde{\nu}_{\tau} \\ \tilde{g} \rightarrow qq \\ \tilde{g} \rightarrow \tilde{q}q \\ \tilde{g} \rightarrow \tilde{t}_{1}t, \tilde{t}_{1} \rightarrow bs \end{array}$	$\begin{array}{c} 2 \ e, \mu \\ 1 \ e, \mu + \tau \\ 2 \ e, \mu \ (\text{SS}) \\ 4 \ e, \mu \\ 3 \ e, \mu + \tau \\ 0 \\ 2 \ e, \mu \ (\text{SS}) \end{array}$	- 0-3 b - - 6-7 jets 0-3 b	- Yes Yes - Yes	4.6 4.6 20.3 20.3 20.3 20.3 20.3 20.3	Fr 1.61 TeV X'_{11}=0.10, X_{132}=0.05 Fr 1.1 TeV X'_{11}=0.10, X_{132}=0.05 \$\vec{v}\$ 1.1 TeV X'_{11}=0.10, X_{132}=0.05 \$\vec{v}\$ 1.35 TeV m(\vec{v}	1212.1272 1212.1272 1404.2500 1405.5086 1405.5086 ATLAS-CONF-2013-091 1404.250
Other	Scalar gluon pair, sgluon $\rightarrow q\bar{q}$ Scalar gluon pair, sgluon $\rightarrow t\bar{t}$ WIMP interaction (D5, Dirac χ)	0 2 <i>e</i> , μ (SS) 0	4 jets 2 b mono-jet	- Yes Yes	4.6 14.3 10.5	sgluon 100-287 GeV sgluon 350-800 GeV incl. limit from 1110.2693 m(χ)<80 GeV, limit of <687 GeV for D8	1210.4826 ATLAS-CONF-2013-051 ATLAS-CONF-2012-147
	$\sqrt{s} = 7 \text{ TeV}$ full data p	√s = 8 TeV artial data	$\sqrt{s} = $ full	8 TeV data		10 ⁻¹ Mass scale [TeV]	

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 or theoretical signal cross section uncertainty.

Wait for new LHC data at \sqrt{s} = 13 *TeV*

Indirect bounds on the scale of New Physics

Most stringent limits come from observables in K⁰ & B⁰ mixing

Wait for new data from LHCb, NA-62 and later from BELLE-2 and LHCb Upgrade

Search for Heavy Majorana Neutral Leptons at accelerators (recent results and future perspectives)

M.Shaposhnikov NP B763 (2007) 45-59 A.Pilaftsis et.al. PR D72 (2005) 113001

From cosmology: $M_N < M_W$ or $M_N \ge \approx 300 \text{ GeV}$ (Sakharov condition \rightarrow CP has to be violated out of thermo-equilibrium)

- ✓ M_N < M_K
 Impressive limits exist from PS-191
 Will soon be validated by NA62
- ✓ M_N < M_{heavy flavour} LHCb, BELLE
 New beam-dump experiment at the SPS (SHIP) has the best sensitivity reach
- ✓ M_N < M_Z
 Can be best explored at Future Circular Collider in e+e- mode
- ✓ M_N > M_Z
 Prerogative of the ATLAS / CMS in the high luminosity phase of LHC

Intensity frontier: New fixed target facility is very timely to explore Hidden Sector and to search for HNL

Common experimental features:

✓ Production through meson decays (π , K, D, B, proton bremstrahlung, ...)

Final states	Models tested	
$\pi l, Kl, \rho l, l = (e, \mu, \nu)$	ν portal, HNL, SUSY neutralino	
$e^+e^-, \mu^+\mu^-$	V, S and A portals, SUSY s-goldstino	
$\pi^{+}\pi^{-}, K^{+}K^{-}$	V, S and A portals, SUSY s-goldstino	
$l^+l^-\nu$	HNL, SUSY neutralino	

- ✓ Full reconstruction and PID are essential to minimize model dependence
- ✓ Production and decay rates are strongly suppressed relative to SM
 - Production branching ratios O(10⁻¹⁰)
 - Long-lived objects

Decays

- Travel unperturbed through ordinary matter

✓ Challenge is background suppression → requires O(0.01) carefully estimated

Physics objectives of SHiP

 ✓ SHiP will directly search for weakly interacting New Physics Will exceed the sensitivity of previous experiments by a few orders of magnitude in the mass range O(10 GeV)

For example, probe HNL couplings close to the ultimate see-saw limit

- \checkmark Unique potential to explore physics of tau neutrinos
 - Observe \overline{v}_{τ} for the first time
 - Extract F_4 and F_5 structure functions never measured so far

General Purpose Fixed-Target SHIP Experimental Facility

SHIP Sensitivities to Selected Physics Channels

The Fixed-Target SHIP Facility at the SPS

- ✤ 85m long Junction cavern in the TDC2 line
- 170 m long machine Extraction Tunnel (4 m wide by 4m high similar to TDC2)
- 15 m long by 15 m wide Access building incl. a shaft to reach the Extraction Tunnel line

Proposed location by CERN beams and support departments (TT20 transfer line):

New Beam-Dump Experiment at the CERN SPS

New Intensity Frontier (SHIP Facility) at the SPS/CERN
 → Fixed-target ("beam-dump") is an ideal instrument to search for weakly interacting Hidden Particles (HP) in Heavy Flavor Decays

SPS can provide 2×10^{20} protons on target (*p.o.t.*) in <u>5 years</u> assuming the same operation as demonstrated during CNGS run

 → Large charm production cross section SPS: 4x10¹³ / 7s @ 400 GeV (√s = 27 GeV)
 → data sample of > 10¹⁷ D-mesons

→ Side benefit: Optimizing for heavy meson decays also optimizes facility for ν_τ (ν_e, ν_μ) physics: Br(Ds→τ+ν_τ)~ 5.6%
 → data sample of ~ 10¹⁵ τ-leptons

Comparison of SHIP with other projects:

- LHC (√s = 14 TeV): - 500 x σ_{cc} and 1 ab⁻¹ (i.e. 3-4 years): ~2 x 10¹⁶ c-hadrons in 4π → yield factor <100 smaller (acceptance) - FNAL 120 GeV pot: 10× smaller $\sigma_{cc'}$ 10× pot by 2025 for LBNE (?)

HNL operation not compatible with neutrino physics

◆ Incompatible with conventional neutrino facility
 ◆ SHIP is a very powerful general-purpose facility for <u>TODAY</u> and <u>FUTURE</u> !
 → Full exploitation and consolidation of the SPS complex after CNGS termination
 ◆ <u>Complementary physics program</u> to searches for new physics by <u>LHC</u> !

SHIP Experimental Requirements / Challenges

Use of neutrino and vector portals for development of experimental facility / sensitivity studies:

- ♦ HNLs produced in charm decays have significant P_T → large detector acceptance due to boost $\tau_{N2,3} \propto U^{-2}$, i.e. $ct \propto O(km)$
- Search for HNL from D-decays, i.e. M < 2 GeV
 B-decays: 20-100 smaller σ; B→ Dµν, i.e. limited to M ~ 5 GeV

- ◆ Place detector as close as possible to target (as background allows) to maximize geometrical acceptance → compromise between HP lifetime and production angle
- "Effective muon shield" (huge μ-flux of 5×10⁹ / spill) to reduce muoninduced bkg. from short-lived resonances accompanying charm production below neutrino-background; acceptable rate ~10⁵ μ /spill
- Decay vessel: "vacuum in detector volume" to reduce v-interactions
- Away from cavern walls to reduce v-interactions in detector proximity
- ✤ Magnetic spectrometer to reconstruct HNL mass.

threshold momentum p (GeV)

Low-p: from π/K -decay High-p: Ω/ρ decays to $\mu\mu$

SHIP Experimental Facility: Physics Objectives

Hidden particle decay volume

Spectrometer Particle ID

Tau neutrino detector

> 2013: originally designed to study HNL in vMSM

➢ Today:

- Search for wide range of weakly interacting exotic particles (incl. SUSY)

- Study physics of v_{τ} produced in D_S decays

Muon sweeper

Extension of SHIP Facility:

- Direct Dark Matter Searches (downstream SHIP)
- ► LFV Experiment $\tau \rightarrow 3\mu$ (dedicated detector)

Target / hadron absorber

CERN Task Force

Initiated by CERN Management after SPSC encouragement of the SHIP in January 2014

Detailed investigation aimed at overall feasibility, identifying options/issues & resource estimate:

- Physics motivation and requirements
- Experimental Area
- SPS configuration and beam time
- SPS beam extraction and delivery
- Target station
- Civil engineering
- Radioprotection

Document completed on July 2, 2014:

- Detailed cost estimate, manpower and schedule
- Compatible with commissioning runs in 2022, data taking 2023

CERN EN Working group responsible for providing design of facility for the SHIP TP

Schematic Principle of the SHIP Experimental Facility

Initial reduction of beam induced backgrounds:

- ★ Heavy material target to stop pions/kaons before they decay (to minimize neutrinos from π/K → μν)
 → blow up beam to dilute energy on target
- Slow (and uniform) beam extraction → reduce occupancy / combinatorial background
- Hadron Absorber
- ◆ Active Muon Filter → muon flux limit is driven by emulsion based neutrino detector and HP background
- Vacuum decay volume followed by a Hidden Sector detector measuring the HNL decay products

Hidden Sector detector optimization: beam energy / intensity, detector acceptance, background studies with full detector simulation \rightarrow design virtually background free experiment O (0.1 event)

SHIP Target and Target Complex

Active Muon Shield

Baseline: Active muon shield based entirely on magnet sweeper / passive absorber

Conceptial design:

- Need around 40 Tm of field to bend out the highest momenta muons (Eμ ~ 350 GeV)
- Return field of the magnets tends to bend low-energy muons back towards the detector
- Critical idea use a first magnet to separate µ+ and µaway from z-axis and then place the return field there
 → "wings" critical to design

♦ Realistic design of sweeper magnets in progress.
 → Challenges: flux leakage, constant field profile, modelling magnet shape

Prompt dose rates in the experimental hall 4x13 p.o.t. / 7s

SHIP Experimental Facility: Main Parameters

2 detector elements

of 5 m x 6 m in EOI

Hidden particle decay volume

Spectrometer Particle ID

Tau neutrino detector

- Total number of protons on target :
- Nominal P.O.T/year :
- Beam energy :
- Beam size:
- (Beam max sweep radius (spiral) :
- Spill intensity :
- Spill duration :
- Spill intensity variation/unit time :
- Cycle time :
- Target:
- Hadron stopper:
- Layout of experimental area

2 x 10²⁰ (5-10 years SPS operation *as-is*) 4 x 10¹⁹

400 GeV (\pm 5 GeV) 6mm RMS, possible increase 3cm, possible increase 4 x 10¹³

1 s

x2-3 variation, MD required 7.2 seconds Mo $4\lambda + W 6\lambda$ 7% water cooling 30 x 30 cm²

Iron 5m

Active muon shield + v-detector + 5m x 10m x 60m vessel + rear detectors plus space for future extension Operation 2023 and LS2 construction From optimization of active muon shield and acceptance: single detector element V (5m) and H(10m) \rightarrow baseline

(d) Ellipse 5×10

Geometric acceptance saturates for a given lifetime as a function of the detector length

Muon sweeper

Target / hadron absorber

Hidden Sector Detector: Objectives and Requirements

Direct detection of Hidden Sector Portals: → Full reconstruction / tracking and particle identification of final states with $e,\mu,\pi\pm,\gamma$ ($\pi0,\rho\pm$), (ν), and decays in flight

Cosmologically interesting and experimentally accessible $m_{HS} \sim 100 \text{ MeV} - 10 \text{ GeV}$:

- Sensitivity to as many modes as possible model independence
- Ultimately distinguish between models

Residual backgrounds sources:

- Neutrino inelastic scattering (e.g. ν/μ + p→X + K_L → μπν) → detector under vacuum, accompanying charged particles (tagging, timing), topological
- Muon inelastic scattering → accompanying charged particles (tagging, timing), topological
- Muon combinatorial (e.g. μμ with μ mis-ID) →
 Tagging, timing and topological
- ♦ Neutrons → Tagging, topological
- ♦ Cosmics → Tagging, timing and topological

Models tested	Final states
Neutrino portal, SUSY neutralino	$l\pi, lK, l\rho$ $l = (e, \mu, \nu)$
Vector, scalar, axion portals, SUSY sgoldstino	$e^+e^-,\mu^+\mu^-$
Vector, scalar, axion portals, SUSY sgoldstino	$\pi^+\pi^-, K^+K^-$
Neutrino portal ,SUSY neutralino, axino	$l^+l^-\nu$
Axion portal, SUSY sgoldstino	γγ
Axino	γ
SUSY sgoldstino	$\pi^{o}\pi^{o}$

X refers to fragmentation products which may help to tag these interactions:

Background source	Decay modes
$\nu \text{ or } \mu + \text{nucleon} \to X + K_L$	$K_L \rightarrow \pi e \nu, \pi \mu \nu, \pi^+ \pi^-, \pi^+ \pi^- \pi^0$
$\nu \text{ or } \mu + \text{nucleon} \to X + K_S$	$K_S ightarrow \pi^0 \pi^0, \pi^+ \pi^-$
$\nu \text{ or } \mu + \text{nucleon} \to X + n$	$n \to p e^- \bar{\nu}_e$
$\nu \text{ or } \mu + \text{nucleon} \to X + \lambda$	$\lambda ightarrow p\pi^-$
$n \text{ or } p+ \text{ nucleon} \rightarrow X + K_L, \text{ etc}$	As above
\	

Fraction of particles entering the vacuum vessel

Per ν_{μ} CC interaction:

Particle	Fraction _{Entering}
Neutron	1.98
٨	3.6×10^{-6}
K_S^0	3.6×10 ⁻⁶
K_L^0	0.5%

Combination of light and highly efficient taggers surrounding vaccum vessel (both the sides and the front face) are required

Vacuum Vessel ("Decay Volume")

Hidden Sector Detector Concept (based on existing technologies)

Reconstruction of HP decays in various final states:

- (Muon System of neutrino detector)

Hidden Sector Detector Layout : Overview of Technologies

TRACKER: NA62-like straw chambers in vacuum (10⁻³ mbar), 120 μm spatial resolution, 0.5% *X/X*0 for 4 stations

Timing Detector: < 100 ps resolution MPRC or Sci. bars with WLS-SiPM (ALICE, NA61, MINOS)

ECAL/HCAL: spiral shashlik CALO (HERA-B, LHCb)

- Magnetic spectrometer emphasize on lower
 - power < 1 MW
- design for modest 0.5 Tm with upgrade up to 1 Tm

MUON: Sci. bars / WLS fibres (e.g. MINOS) and SiPMs readout

SAMPIC Chip for the ECAL Calorimeter and Timing Detector

Collaborative effort between CEA and LAL:

SAMPIC chip can be used for:

- ➤ The TIMING detector requires ~50ps resolution
 → corresponds exactly to the initial target of SAMPIC. The expected rate is several orders of magnitude smaller than the limit of SAMPIC.
- ➤ The ECAL calorimeter using scintillators read by PMT → SAMPIC can be used in the low-speed mode with a timing resolution far better < 0.5ns required and allowing to fully capture pulses as long as 60 ns. Some optimization to the current SAMPIC chip for bi-gain operation is mandatory.

Potential interests of the French groups:

- Electronics developments: SAMPIC or others (Irfu, LAL, LPNHE)
- Micromegas detector for Tau Neutrino target tracker (Irfu)
- ✤ VETO photodetector (LPNHE)

Schematic Principle of the Experimental Facility

Initial reduction of beam induced backgrounds:

- ★ Heavy material target to stop pions/kaons before they decay (to minimize neutrinos from π/K → μν)
 → blow up beam to dilute energy on target
- Slow (and uniform) beam extraction → reduce occupancy / combinatorial background
- Hadron Absorber
- ◆ Active Muon Filter → muon flux limit is driven by emulsion based neutrino detector and HP background
- Tau neutrino detector located immediately downstream of active muon shield

Tau neutrino detector optimization: negligible occupancy, emulsion limit of 10^4 particles /mm², ideally suited for studying interactions of v_{τ} , observation of the production and decay of charm

Tau Neutrino Detector Layout

Neutrino target:

- > 1155 ECC bricks to be replaced 10 times
- 260 n-interactions integrated in 1 ECC brick (during 6 months exposure)
- Total emulsion surface: 8700 m2 (5% OPERA)

- ***** Emulsion Cloud Chamber (ECC)
 - →Passive material (Lead 1mm) 56 layers
 - \rightarrow High resolution (Nuclear emulsions) 57 films
- ***** Compact Emulsion Spectrometer:
 - \rightarrow 3 OPERA-like emulsion films
 - \rightarrow 2 Rohacell spacers (low density material)

Electronic Target Tracker for the Tau Neutrino Detector

Target Tracker Requirements:

- Maximum thickness of the plane is 5-6 cm with small dead space (< 1 mm)
- Spatial resolution (~ 100μm) and its angular dependence
- Capability of measuring the angle in each plane (efficiency versus the track angle: up to tg(θ) = 1)
- Performance in magnetic field (RD51 is currently using GOLIATH magnet in the test-beam area);

Target Tracker Layout:

- \rightarrow 12 planes with 2x1 m² surface
- Provide time stamp of the neutrino interaction in the brick"
- Matching between the electronic detectors and the emulsion tracker

Three possible technologies:

- Scintillating fiber tracker (250 μm scintillating fibres readout by SiPMs)
- ✤ GEM / Micromegas tracker

Micromegas for Target Tracker (based on ATLAS NSW Experience)

~ 2x1 m² MM can be built as a single module with min. dead space : ~ 1-strip pitch (300 μm) on each PCB

Measuring the arrival time of the signals opens a new dimension; in this case the MM functions like a TPC \Rightarrow Track vectors/plane for inclined tracks

σ < 100 um independently of track incident angle!

4-plet thickness total budget (ATLAS NSW -7.8 cm \rightarrow to be reduced by 20-30% or use 3-plet of MM chambers

Cost Effective Solution (compared to fiber tracker): ~40 kEUR for one MM module (i.e. for 4 layers of ~3 m² MM plane within a module) without electronics

MM spatial resolution: Strip pitch ~ 400-450 um Strip width ~ 300 um

SM Physics with SHIP: v_r-Factory With Lowest Background

- ✓ Expect O(10000) $v_{\tau} / \overline{v_{\tau}}$ interactions in 6 tons emulsion target with 2×10²⁰ pot **Prospects for** $v_{\tau} (v_{e}, v_{\mu})$ Physics : CC interacting
- > First observation of anti- v_{τ}
- Unique opportunity to measure ν_τ / anti-ν_τ cross-sections differentially
- Extraction of F₄, F₅ structure functions from CC-neutrino nucleon scattering (not accessible with lighter neutrinos)
- ➤ Charm physics with v and anti-v → anti-v highly sensitive to the s-quark (improve understanding of the strange quark content of nucleons)
- Study of v_e at high energies (E > 20 GeV)
- Exotic states searches (e.g. multi-quark)
- ➤ Measurement of the v_e production in charmed decays → normalization for longlived hidden particle searches

CC interacting v-fluxes and spectra:

	$\langle \mathrm{~E} \rangle \mathrm{(GeV)}$	# interactions
$ u_{\mu} $	29	$3.4 imes10^6$
ν_e	33	$2.3 imes10^5$
$\nu_{ au}$	48	$2.4 imes10^4$
$\bar{\nu}_{\mu}$	23	$1.1 imes 10^6$
$\bar{\nu}_e$	33	$8.5 imes10^4$
$\bar{\nu}_{ au}$	48	$1.2 imes10^4$

Neutrino-induced charm production:

	<e> (GeV)</e>	# interactions	
ν _μ	30.3	7.2 x 104	
Ve	42.5	3.3 x 104	
anti-v _µ	26.7	2.4×10^{4}	
anti-v _e	33.6	2.9×10^{4}	

Separate contributions of valence (absent in anti- v case) and sea quarks: 30-2000 x CHORUS statistics

Exotic state searches (e.g. pentaquark)

SHIP Sensitivities for selected physics channels

Physics Case for the General Beam Dump Facility

Different portals to Hidden sectors:

Portal	Particles	Operator(s)
"Vector"	Dark photons	$-\frac{\epsilon}{2\cos\theta_W}B_{\mu\nu}F'^{\mu\nu}$
"Axion"	Pseudoscalars	$\frac{a}{f_a}F_{\mu\nu}\widetilde{F}^{\mu\nu}, \frac{a}{f_a}G_{i\mu\nu}\widetilde{G}_i^{\mu\nu}, \frac{\partial_{\mu}a}{f_a}\overline{\psi}\gamma^{\mu}\gamma^5$
"Higgs"	Dark scalars	$(\mu \tilde{S} + \lambda S^2) H^{\dagger} \tilde{H}$
"Neutrino"	Sterile neutrinos	$y_N LHN$

Long lived weakly interacting particles:

XX.

- Heavy masses, O(100 GeV): ATLAS and CMS via missing ET
- Small masses, O(1-10 GeV): LHC-b, B-factories, direct observ., large couplings =short lifetimes

988

Small masses, O(1GeV), small couplings = long lifetimes: SHiP

Neutrino Portal: Sensitivity to Heavy Neutral Leptons (N2.3) in vMSM

◆ Ultimate see-saw limit is almost in reach → still would require increase of the SPS intensity by an order of magnitude (does not currently seem to be realistic)

Neutrino Portal: HNL Sensitivities in Collider Experiments

<u>SHIP complementarity</u>: Colliders are not very sensitive with low mass / long lifetimes:

- ► BELLE-2 using $B \rightarrow XlN$, where $N \rightarrow l\pi$ may go well below 10⁻⁴ in 0.5<M_N<5 GeV
- > TLEP using $Z \rightarrow Nv$ with $N \rightarrow$ lepton + 2 jets

Summary of past searches for HNL:

Depends on HNL decay length and efficiency:

♦ W → ℓN at LHC: extremely large BG, difficult triggering/ analysis.

♦ Z → Nv at e+e- collider: clean (expected sensitivity of FCC in e+e- mode assuming zero bkg.)

Minimal Vector Portal: Sensitivity to Dark Photons

SHiP,

protons

10

Higgs Portal: Sensitivity to Dark Scalars

SM Higgs + real singlet scalar (ϕ or h):

- > Could have mass $m_h < 5 \text{ GeV}$;
- Mixing with the SM Higgs with angle ρ:

Dark Scalar Production:

arXiv: 1310.6752 arXiv: 1310.8042

- direct: $p + target \rightarrow hX$
- flavour decays: $B \rightarrow hK^*$ (this study)

(D-CKM suppressed wrt B (5x10⁻¹⁰), while $\sigma(D)$ only 20k times larger $\sigma(B)$ at 27GeV)

Dark Scalar Decays: some uncertainty in the calculation of BR(h):

Dark scalar: $h \rightarrow \mu^+ \mu^-$

Sensitivity can be improved by SHIP:

Current sensitivity includes only B decays and only muon final states (need to add hadronic channels), cross section a bit conservative (/3)

Axion Portal: Sensitivity to Pseudo-Scalars, Axion-Like Particles

ALPs are well motivated from theory → typically pseudo-Goldstone bosons

ALP Coupling to two photons:

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</u>)² (<u>cr</u>)³ / (<u>r</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</sub>)² (<u>cr</u>)³ / (<u>r</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</sub>)² (<u>cr</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</u>)² (<u>cr</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</sub>)² (<u>cr</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</u>)² (<u>cr</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</u>)³

 ¹/_{μecry} ~ 4cm (<u>0.º GeV⁻¹</u>)⁴

 ¹/_{μecry} ~ 4cm </u></u></u></u></u></u></u>

SHIP can improve limits / can discover:

- Significant gain for fermion couplings
- Some gain for two-photon coupling
- Study of two-gluon coupling still to be done

- ALP decays to γγ, μμ, ee 10^{-10} Probe high 10^{0} BaBar 10^{-3} energy scales: 10^{-1} B→K+inv B 10^{-10} $\bowtie^{H} \mathbf{B} \rightarrow \mathbf{K} \mu^{+} \mu^{-}$ 10^{-2} GeV Interpretation of 10^{-5} 10-3 8 CHARM as 10^{-4} production 10^{-1} from neutral 10^{-5} pion mixing ! 10^{-} 10^{-6} BBN constrain Thermal equ. 10Ge 10^{-10} 10^{-1} 10^{0} 10^{-1} 10^{1} m_A [GeV] PRELIMINARY гл\GB: 🤶 e⁺e⁻ & μ⁺μ⁻ 🖉 10⁹ SHiP 10 beyond 1GeV 10 SHIP things are complicated 106 due to 10 dominance of excluded by past experiments hadronic decays 10-2 10⁻¹ $m_{PNGB}(GeV)^{1}$

✤ ALP Coupling to fermions:

- Production via ALP-pion mixing;

SHIP Physics Program: Direct SUSY Detection

- ► Light SUSY sgoldstinos (hep-ph/000735): Production/decay might be like HNL, i.e. $D \rightarrow \pi X$ followed by $X \rightarrow \pi + \pi -$
- Light R-parity violating SUSY neutralinos (hep-ph/0106199):
 B/D → Xχ₀, χ₀ → μ+μ−ν; χ₀ → Kν; K+e-LSP with R-parity "slightly" violated: τ < 0.1 s
- ► Light Pseudo-Dirac Gauginos: predicted in SUSY with U(1) R-symmetry $pp \rightarrow \Psi \Psi; \chi_2 \rightarrow ll \chi_1$
- Chern-Simons portal: new X vector boson coupled to SM bosons with Chern-Simons like interactions

10-4

 10^{-5}

 10^{-6}

 10^{-7}

 10^{-8}

 10^{-9}

 10^{-10}

 10^{-6}

 $c_{\mathbf{Z}}^{\mathbf{Z}}$

 10^{-5} 10^{-4} 0.001 0.01

Excluded by previous searches

Probed

by SHiP

Not enough events

in SHiP

Sensitivity

studies in

progress

0.1

SHIP Physics Program: SUSY Hidden Sector

Extension of SHIP Facility: Direct Dark Matter Search

Relativistic beam of light Dark Matter with $2x10^{20}$ pot

◆ The signature of DM is a neutral current scattering event → very similar to v-induced NC event

Tau neutrino detector could be used for DM detection (9 ton) → comparison with E613 FNAL → factor 100 more events in SHIP (but also more background (neutrino NC)

Tau neutrino detector is 9t for $DM \rightarrow$ a factor 10 less mass

Extension of SHIP Facility: LFV $\tau \rightarrow 3\mu$ Searches

SHIP Experimental Facility: use 1 mm heavy material (W) target

List of D-backgrounds:

Decay	\mathcal{B}
$D^+ \rightarrow \eta \mu^+ \nu_\mu$	$1.14 \cdot 10^{-3}$
$\eta \rightarrow \mu^+ \mu^-$	$5.8 \cdot 10^{-6}$
or	
$\eta \rightarrow \mu^+ \mu^- \gamma$	$3.1\cdot10^{-4}$
$D^+ \rightarrow \rho \mu^+ \nu_\mu$	$2.4 \cdot 10^{-3}$
$ ho ightarrow \mu^+ \mu^-$	$4.55 \cdot 10^{-5}$
$D^+ \rightarrow \eta' \mu^+ \nu_\mu$	$2.2 \cdot 10^{-4}$
$\eta\prime ightarrow \mu^+ \mu^- \gamma$	$1.8 \cdot 10^{-4}$
$D^+ \rightarrow \omega \mu^+ \nu_\mu$	$1.6 \cdot 10^{-3}$
$\omega ightarrow \mu^+ \mu^-$	$9.0 \cdot 10^{-5}$
$D_s^+ \to \eta \mu^+ \nu_\mu$	$2.7 \cdot 10^{-2}$
$\eta \rightarrow \mu^+ \mu^-$	$5.8 \cdot 10^{-6}$
or	
$\eta \rightarrow \mu^+ \mu^- \gamma$	$3.1\cdot10^{-4}$
$D_s^+ \rightarrow \eta' \mu^+ \nu_\mu$	$9.9 \cdot 10^{-3}$
$\eta\prime ightarrow \mu^+ \mu^- \gamma$	$1.8\cdot10^{-4}$

- the main handle:
 - decay vertex separation from the target
 - improving spatial/time resolution
- to be studied additionally
- imposes tight requirements on the trigger!

Sensitivity studies are in progress

Hadron Absorber

SHIP Experiment Progress in 1 Year

- SHIP is the universal tool to probe New Physics at the Intensity Fronterin the largely unexplored domain of new, very weakly interacting particles
- SHIP facility also opens unique opportunity for ντ- physics (largely extending original physics motivation of the SHIP proposal to the SM Physics)
- ➢ Future extension of the SHIP facility for DM and LFV searches is possible
- Major technological and engineering challenges of the SHIP Facility have been addressed during the last year:
 - → target and beam-line active muon shield (thanks to CERN EN), decay volume, background taggers, timing detectors ...

One year ago SHIP was just an idea → TODAY <u>SHIP IS THE GENERAL PURPOSE FACILITY</u> which could help:

b to increase diversity of the particle physics program world-wide

to explore the Fermi scale and to provide guidance on the scale of the New Physics or on the coupling strength of any new particles to the SM particles.