

# Résumé des conférences d'hiver 2014

## *Disclaimers:*

Résumé LHC-centrique

Beaucoup de résultats étaient déjà publics avant les conférences

## Résumé de:

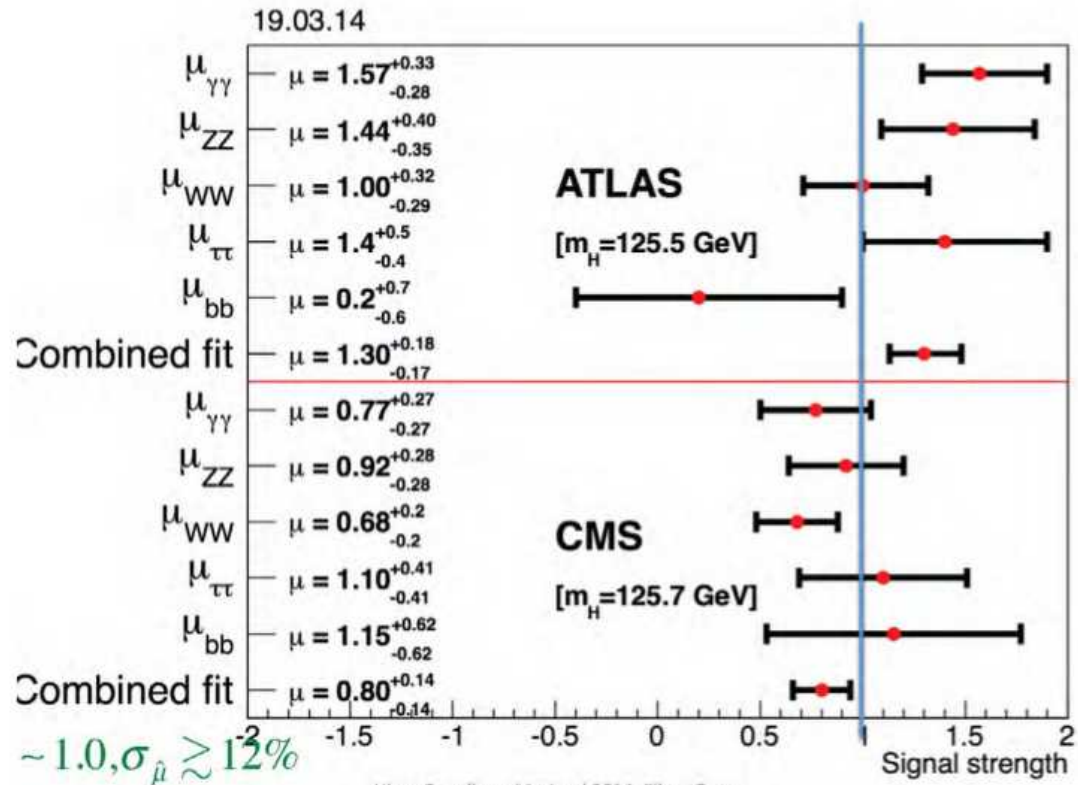
**Aspen:** <https://indico.cern.ch/event/276476/timetable/#20140119>

**Moriond EW:** <https://indico.in2p3.fr/conferenceOtherViews.py?view=standard&confId=9116>

# Higgs mass and couplings

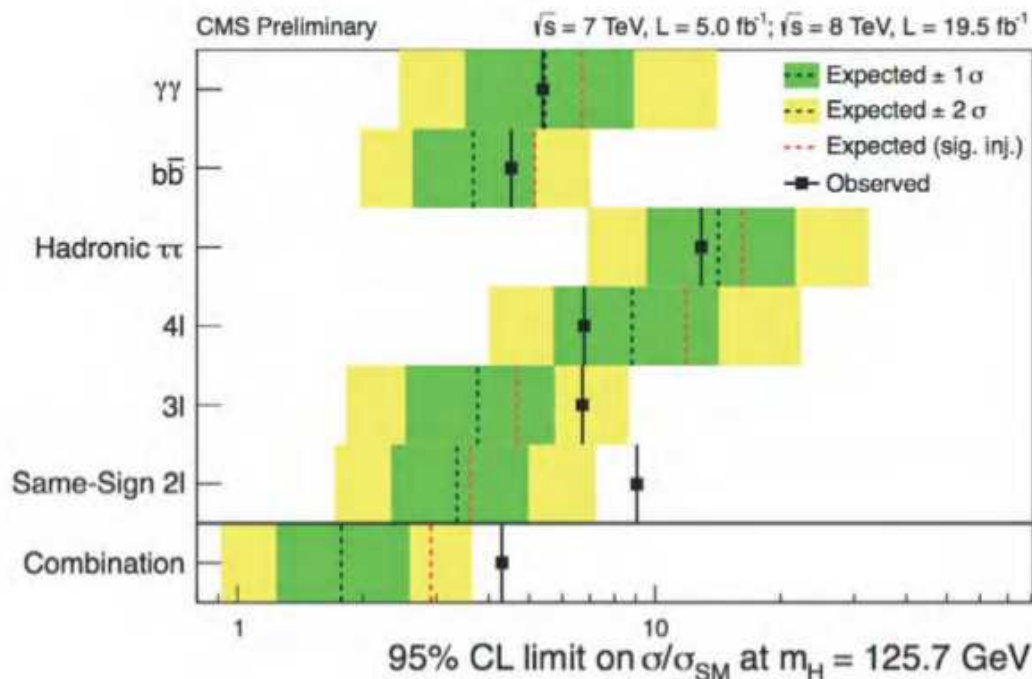
ATLAS	CMS (new ZZ(4l) not used)
$125.5 \pm 0.2 \text{ (stat)}^{+0.5}_{-0.6} \text{ (syst) GeV}$	$125.7 \pm 0.3 \text{ (stat)} \pm 0.3 \text{ (syst) GeV}$

- Overall comparison of all individual  $\mu$  values:



# Higgs: Production associée avec top-antitop (ttH)

- **CMS** has performed a search in  $t\bar{t}H$ ;  $H \rightarrow \gamma\gamma, b\bar{b}, \tau_h\tau_h$ , leptonic (2l SS, 3l, 4l), and combined (**Botta**).
- **Result is limit of  $\mu = 4.3$  (expected 2.9) 95% CL. Best fit value is  $\mu = 2.5 + 1.1 - 1.0$  (excess all in 2l SS).**



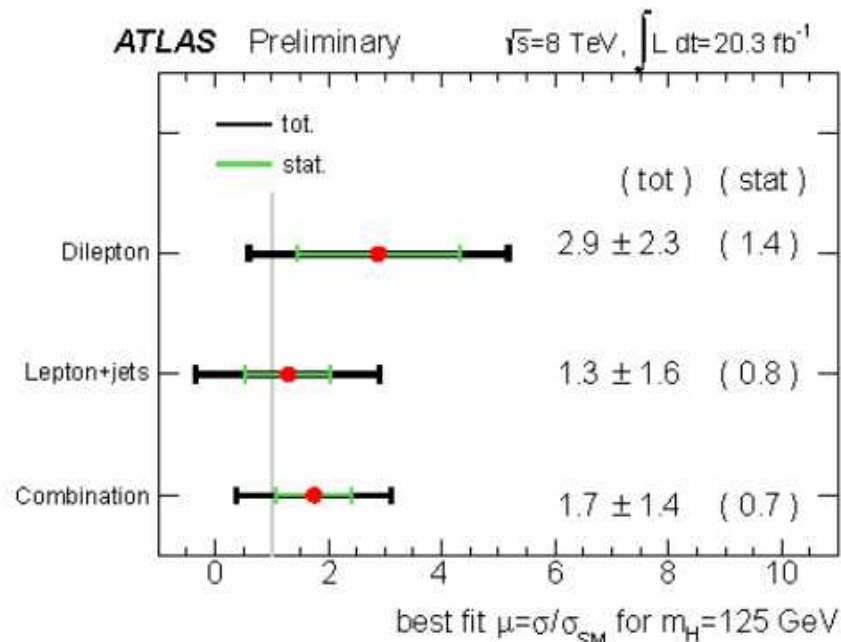
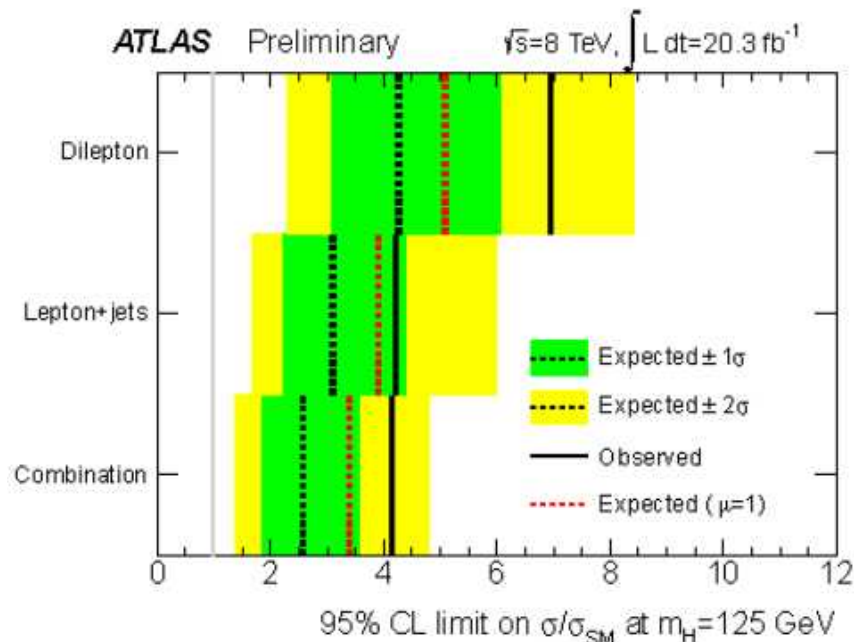
**Internal consistency for common value is 22%.**

**P-value for  $\mu = 1$  is  $1.6\sigma$**

**Note:  $t\bar{t}H$  increases by almost 5 for Run2 @ 13 TeV**

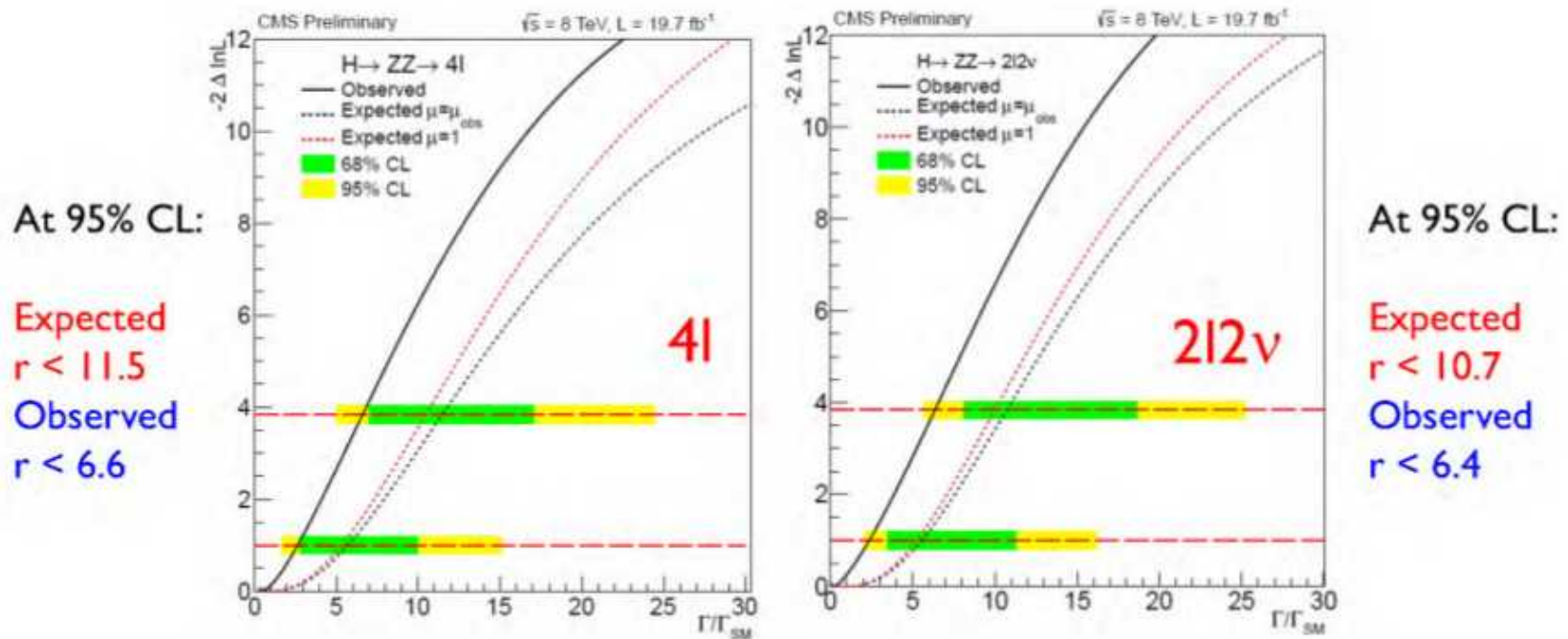
# Higgs: Production associée avec top-antitop (ttH)

- **ATLAS** has performed a search in  $t\bar{t}+H$ ;  $H \rightarrow \gamma\gamma$ . Observed limit of  $\mu = 4.7$  (5.4 expected). Leptonic (2l, 3l, 4l,  $\tau\tau$ ) modes in progress (*Le Menedeu*).
- New analysis in  $t\bar{t}+H$ ;  $H \rightarrow bb$ . Result is limit of  $\mu = 4.1$  (expected 2.6) 95% CL. Best fit value is  $\mu = 1.7 \pm 1.4$



# Higgs interferometry: Higgs width from on-shell/off-shell production ratio

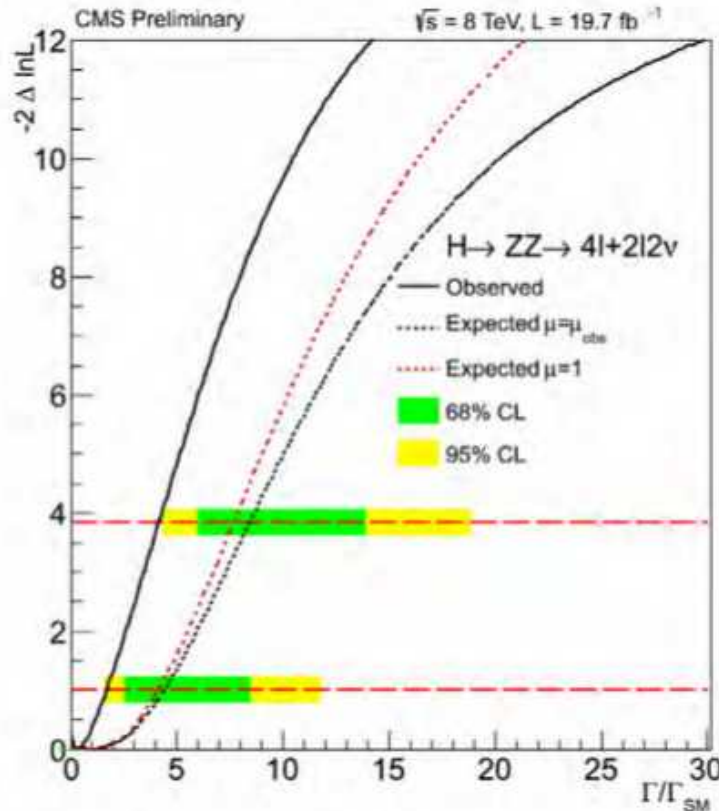
- At Moriond, CMS released first measurement of  $r = \Gamma/\Gamma_{SM}$ , using  $H \rightarrow ZZ$  decaying into  $4l$  and  $2l2\nu$  (PAS HIG-14-002) (*Covarelli*).
- They use their published  $H \rightarrow ZZ$  on-shell cross-section value  $\mu = 0.93^{+0.26}_{-0.24}$ , and also compare with  $\mu = 1.0$  for reference.
- They use a kinematic discriminant, similar to that of Campbell et al. to reduce the  $qq \rightarrow ZZ$  continuum relative to the  $gg$  signal.





# Higgs interferometry: Higgs width

- **Combination of two channels gives:**



- ▶ Combined **observed** (**expected**) values

- ▶  $r = \Gamma/\Gamma_{SM} < 4.2$  (**8.5**)  
@ 95% CL

(p-value = 0.02)

- ▶  $r = \Gamma/\Gamma_{SM} = 0.3^{+1.5}_{-0.3}$

- ▶ equivalent to:

- ▶  $\Gamma < 17.4$  (**35.3**) MeV  
@ 95% CL

- ▶  $\Gamma = (1.4^{+6.1}_{-1.4})$  MeV

**Very important result ! Observed limit is half of expected – data deficits in both channels ? Theory systs (LO+K<sub>f</sub>) under control ?**

# SUSY and Exotics at LHC: Summary of Run 1

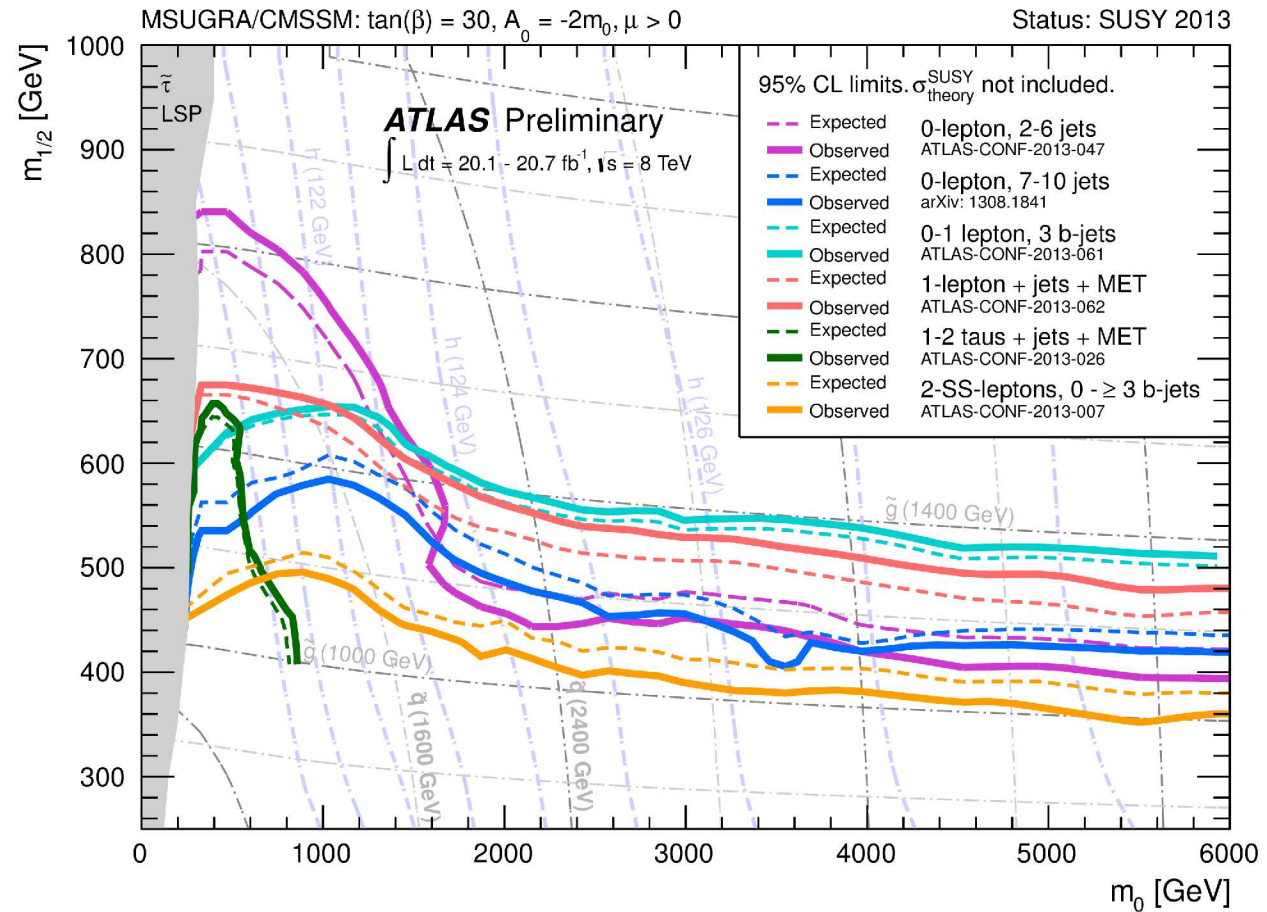
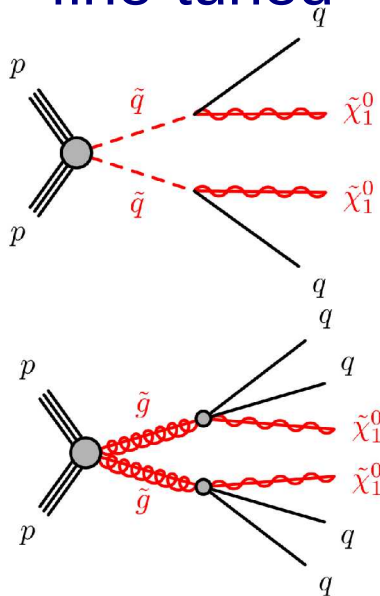
- The 8 TeV LHC data have been investigated extensively  
but still a lot of work in progress
- Unfortunately, still no hint of BSM physics in the LHC data...

	Approx. Lower Limit (95% C.L.)
VLQ T and stop ( $tt \rightarrow tt\chi\chi$ )	700 GeV
gluino	1.4 TeV
KK gluon $\rightarrow$ top-antitop res.	2 TeV
$Z' \rightarrow$ dilepton (SSM)	3 TeV
Excited quark $\rightarrow$ dijet	4 TeV

# Supersymmetry: Strong Production

- Summary of strong-production searches
- **cMSSM interpretation: squark and gluino mass  $> 1.4$  TeV (95% CL)**

- Conclusion: cMSSM is fine-tuned





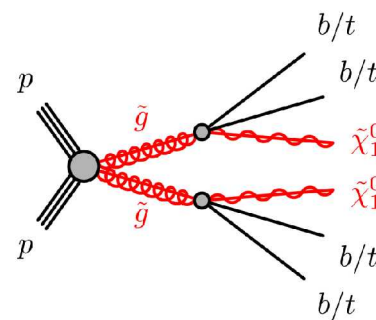
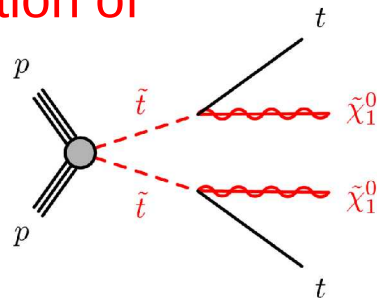
# Supersymmetry: 3rd generation

- Natural (i.e. not fine-tuned) SUSY requires:

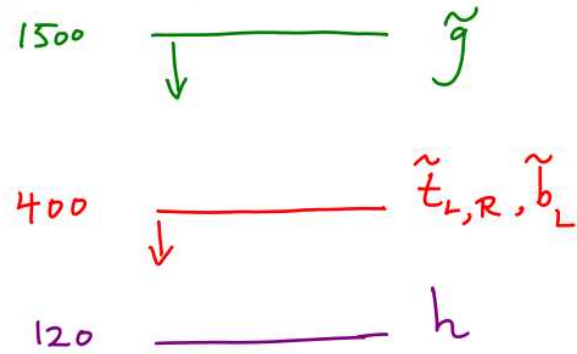
- stop/sbottom are light
- gluino somewhat light
- 1st and 2nd generation squarks are allowed to be very heavy

- 2 strategies:

- gluino production decaying to stop/sbottom
- direct production of stop/sbottom



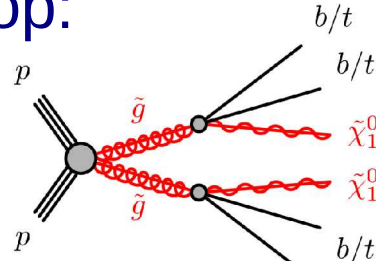
Compulsory Natural SUSY



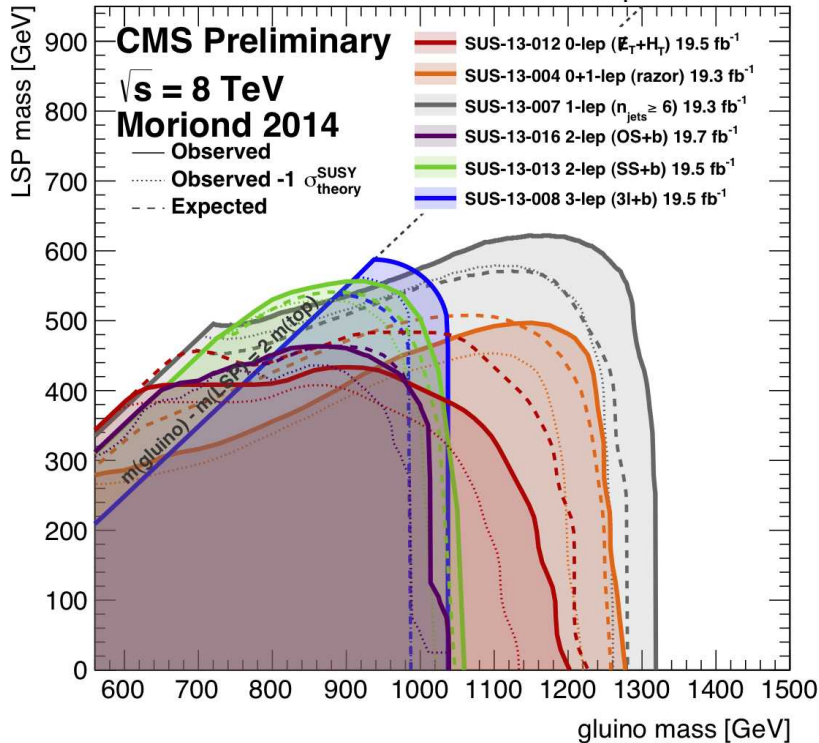
Unavoidable tunings:  $\left(\frac{400}{m_{\tilde{t}}}\right)^2$ ,  $\left(\frac{4m_{\tilde{t}}}{M_{\tilde{g}}}\right)^2$

# Supersymmetry: 3rd generation

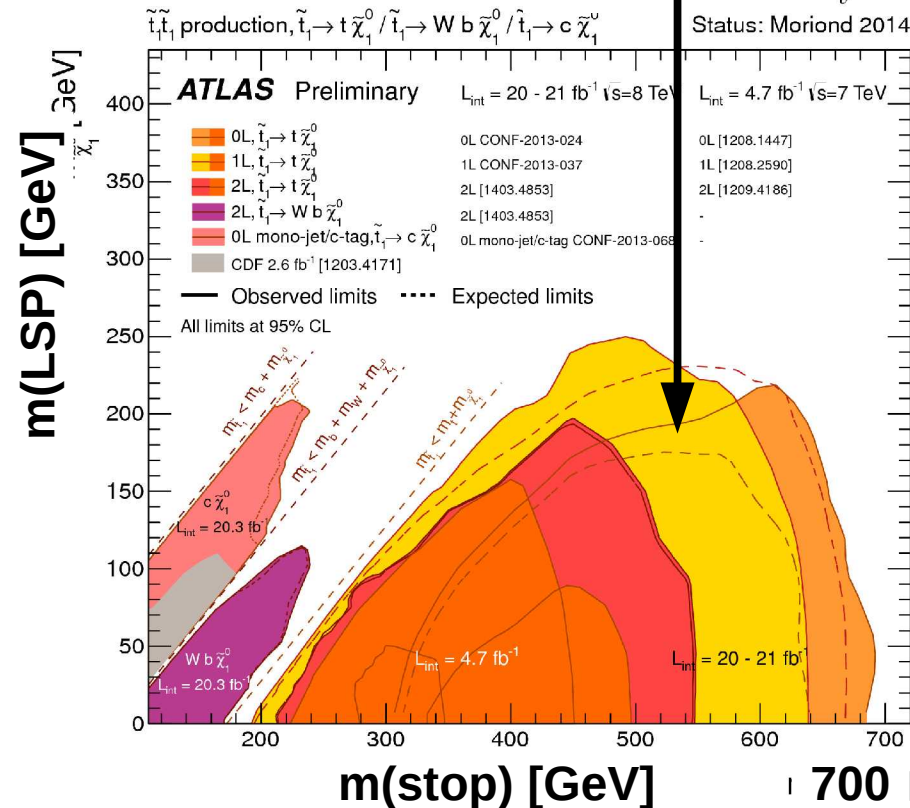
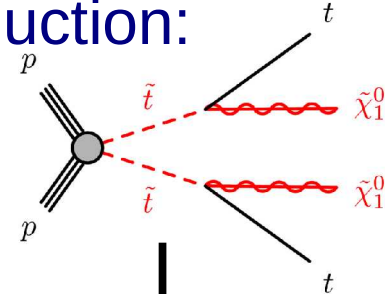
## ■ Gluino to stop:



$\tilde{g}\text{-}\tilde{g}$  production,  $\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_1^0$

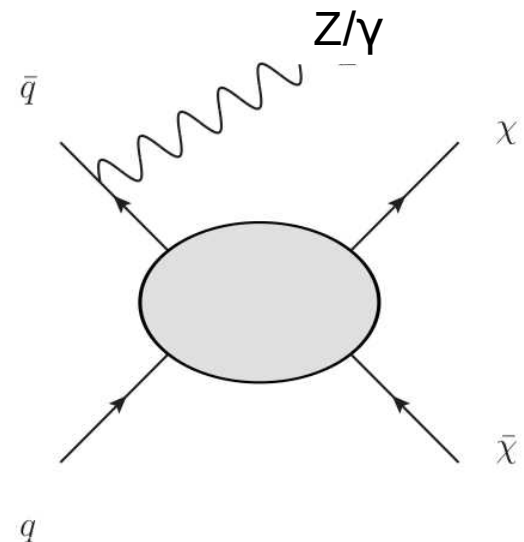
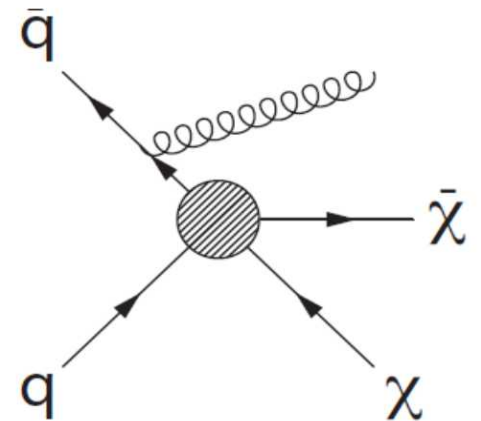
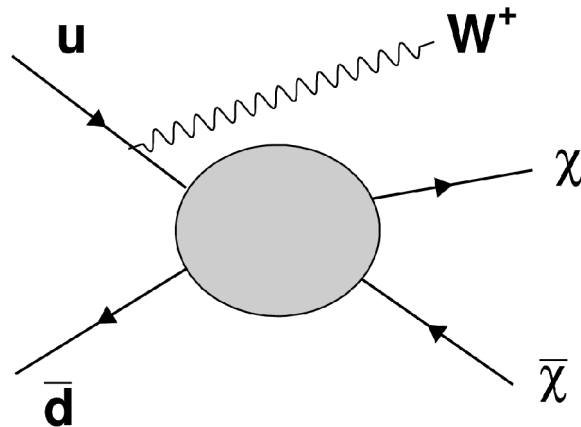


## ■ Direct stop production:



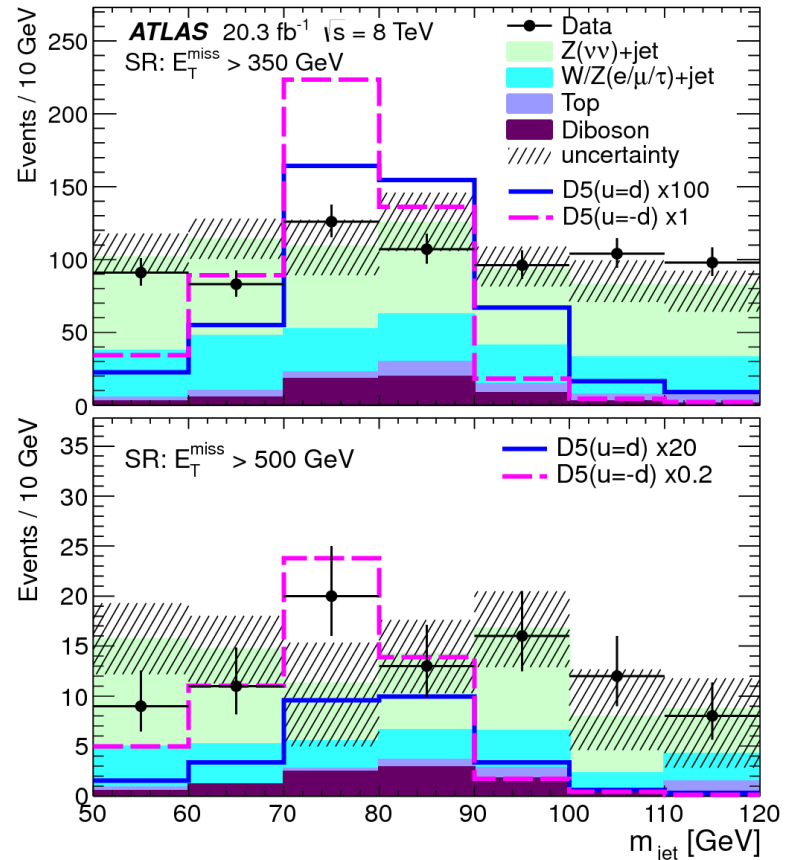
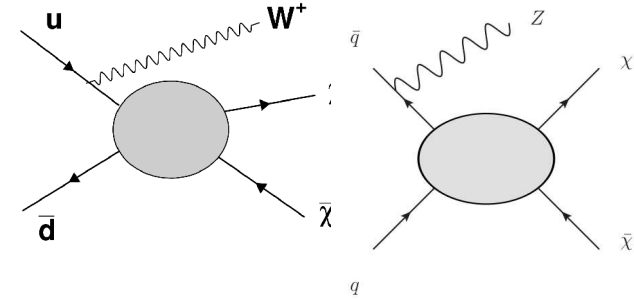
# Search for Dark Matter at LHC

- Dark matter pair-production
  - Observe only the Initial State Radiation
  - Consider  $g/\gamma/W/Z$  radiation:
    - mono-jet + missing ET
    - mono-photon + missing ET
    - $W$  or  $Z$  ( $lv/l\bar{l}$  or dijet + missing ET)
    - $Z \rightarrow$  dilepton + missing ET



# Search for Mono-W & Mono-Z (hadronic decay)

- High-momentum  $W \rightarrow qq$  or  $Z \rightarrow qq$  reconstructed with one large-radius jet
  - Cambridge–Aachen algorithm with a radius parameter of 1.2
- Select jets with mass consistent with  $W$  or  $Z$  hadronic decay
- Inclusive trigger MET  $> 150$  GeV
- Final selection:
  - MET  $> 350$  GeV
  - MET  $> 500$  GeV



# Search for Dark Matter

- Limits presented in several ways:

- Consider several operators of effective theory described in arXiv:1008.1783
- Loose constraint on EFT validity assumes WIMP production is near threshold:  $q^2 \sim 2m_\chi < 4\pi M_*$

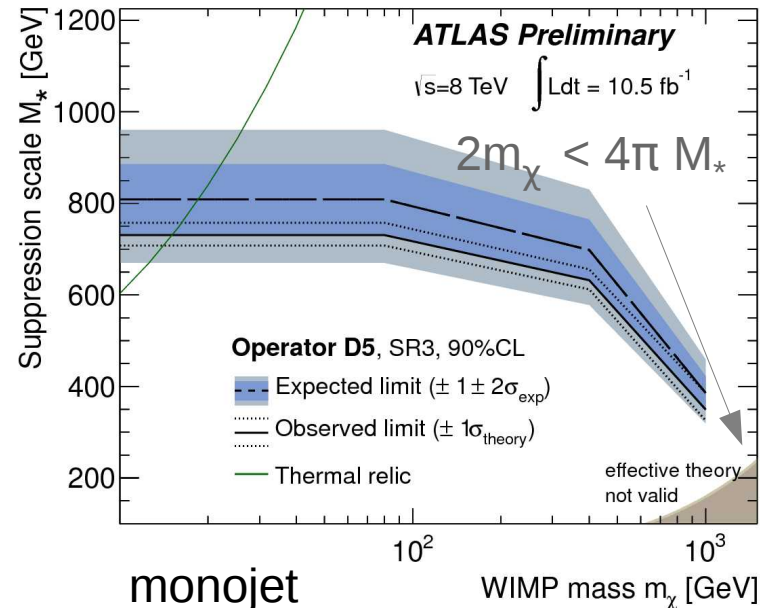
- $M = \text{mediator mass} = M_* \sqrt{g_q g_\chi}$

Perturbativity requires  $M < 4\pi M_*$

- Back-of-the-envelope:

if  $M_* < 400$  GeV, mediator is produced on-shell i.e. at much higher  $q^2$

Name	Initial state	Type	Operator
D1	$qq$	scalar	$\frac{m_q}{M_*^3} \bar{\chi} \chi \bar{q} q$
D5	$qq$	vector	$\frac{1}{M_*^2} \bar{\chi} \gamma^\mu \chi \bar{q} \gamma_\mu q$
D8	$qq$	axial-vector	$\frac{1}{M_*^2} \bar{\chi} \gamma^\mu \gamma^5 \chi \bar{q} \gamma_\mu \gamma^5 q$
D9	$qq$	tensor	$\frac{1}{M_*^2} \bar{\chi} \sigma^{\mu\nu} \chi \bar{q} \sigma_{\mu\nu} q$
D11	$gg$	scalar	$\frac{1}{4M_*^3} \bar{\chi} \chi \alpha_s (G_{\mu\nu}^a)^2$





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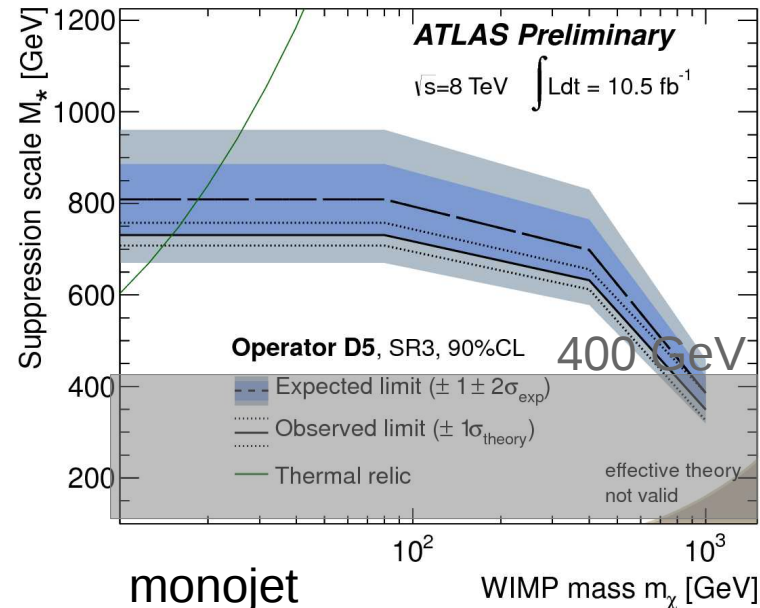
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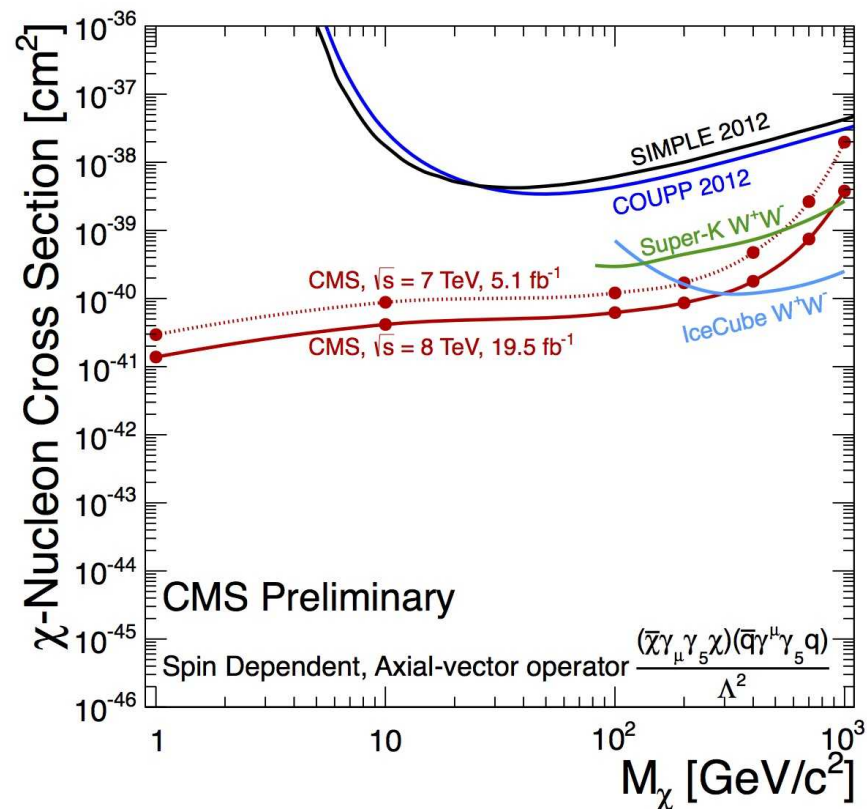
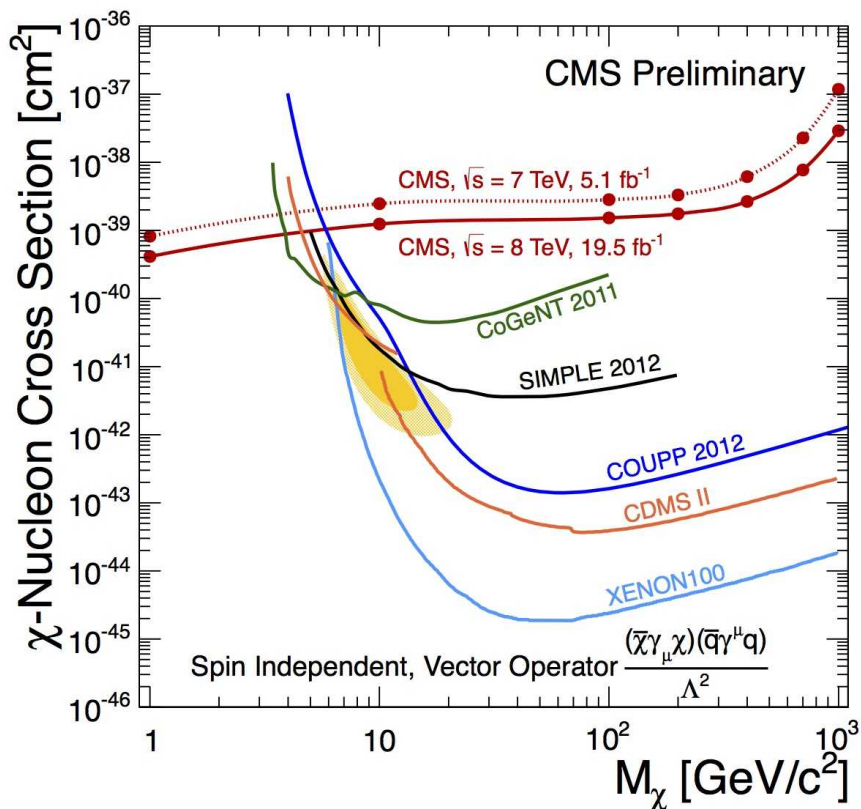
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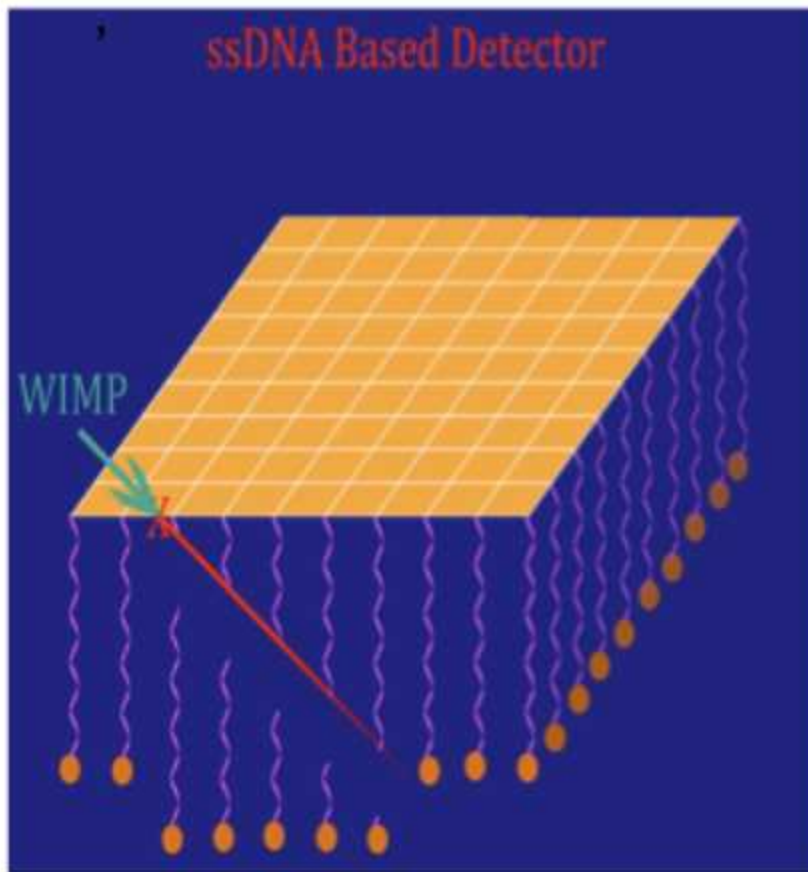


# Search for Dark Matter



# Direct Dark Matter Search with DNA

- 1 kg Gold, 1 kg ssDNA, identical sequences of bases with an order that is well known



**BEADED CURTAIN OF ssDNA**

WIMP from galaxy knocks out Au nucleus, which traverses DNA strings, severing the strand whenever it hits.

Katherine Freese (Aspen)

<https://indico.cern.ch/event/276476/session/8/contribution/22>

# Direct Dark Matter Search with DNA

## Diurnal Modulation (due to Earth's rotation)

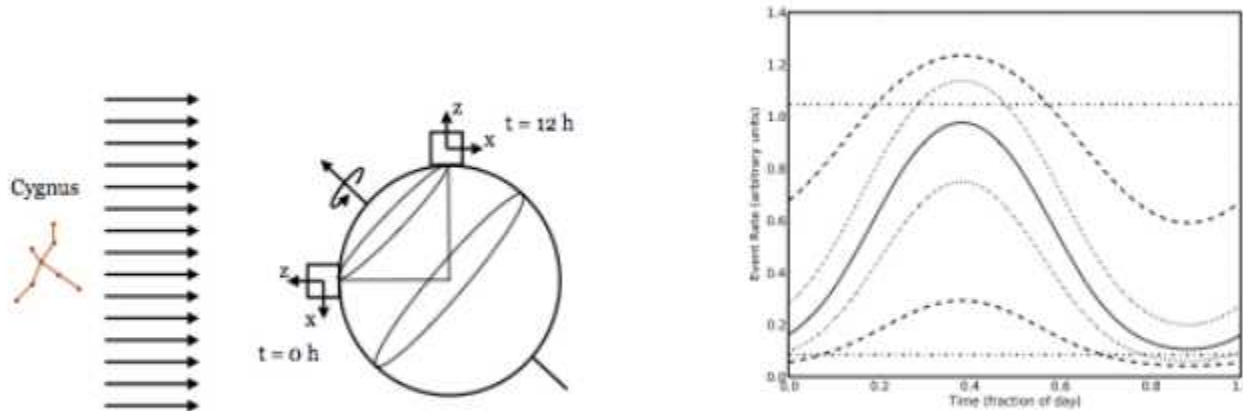
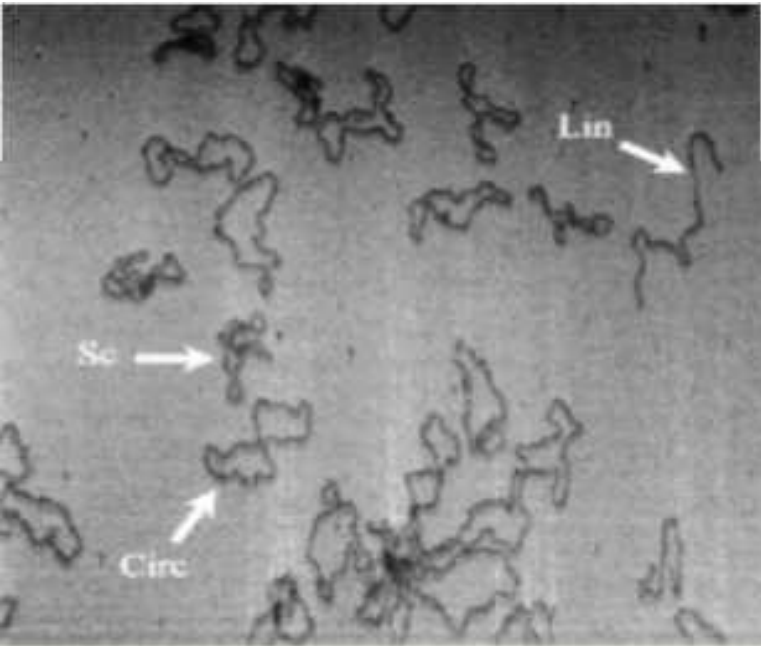


Fig. 2. (left) The daily rotation of the Earth introduces a modulation in recoil angle, as measured in the laboratory frame. (right) Magnitude of this daily modulation for seven lab-fixed directions, specified as angles with respect to the Earth's equatorial plane. The solid line corresponds to zero degrees, and the dotted, dashed, and dash-dot lines correspond to  $\pm 18^\circ$ ,  $\pm 54^\circ$  and  $\pm 90^\circ$ , with negative angles falling above the zero degree line and positive angles below. The  $\pm 90^\circ$  directions are co-aligned with the Earth's rotation axis and therefore exhibit no daily modulation. This calculation assumes a WIMP mass of 100 GeV and  $\text{CS}_2$  target gas. (from Ref. [13]).

Directional capability allows to reject background very efficiently

# Direct Dark Matter Search with DNA

Production de ségments  
d'ADN linéaires:



## Experimental issues

- How to keep ssDNA strands straight? Electric or magnetic field (Church)
- How to get severed strands to fall down: use electric or magnetic field?
- How to scoop the severed ssDNA (e.g. once per hour): use magnetizable rod?



# Direct Dark Matter Search with DNA

Gold target, 365 kg-day exposure

