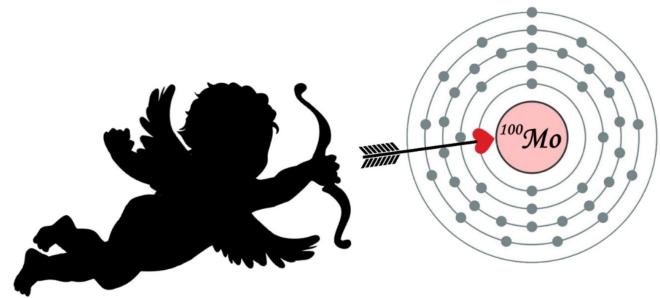


The CUPID-Mo experiment for $0\nu\beta\beta$ decay



Giovanni Benato

Séminaires du DPhP, December 7th, 2020

Matter producing $0\nu\beta\beta$ decay

What is matter?

- Everyday life: matter is made of atoms \Rightarrow Fails at high energy
- SM definition: matter is made of leptons and baryons (or quarks)

What is conserved in the SM?

- ~~B and L~~ \Rightarrow Non-perturbative effects at high energy
- $(B-L), (L_e - L_\mu), (L_\mu - L_\tau), (L_\tau - L_\nu)$ \Rightarrow Oscillation experiments

[S. Dell'Oro, S. Marcocci and F. Vissani,](#)
[PoS NEUTEL2017 \(2018\) 030](#)

M. Agostini, G. Benato, J. Detwiler,
J. Menendez and F. Vissani
Review paper coming out soon(ish)

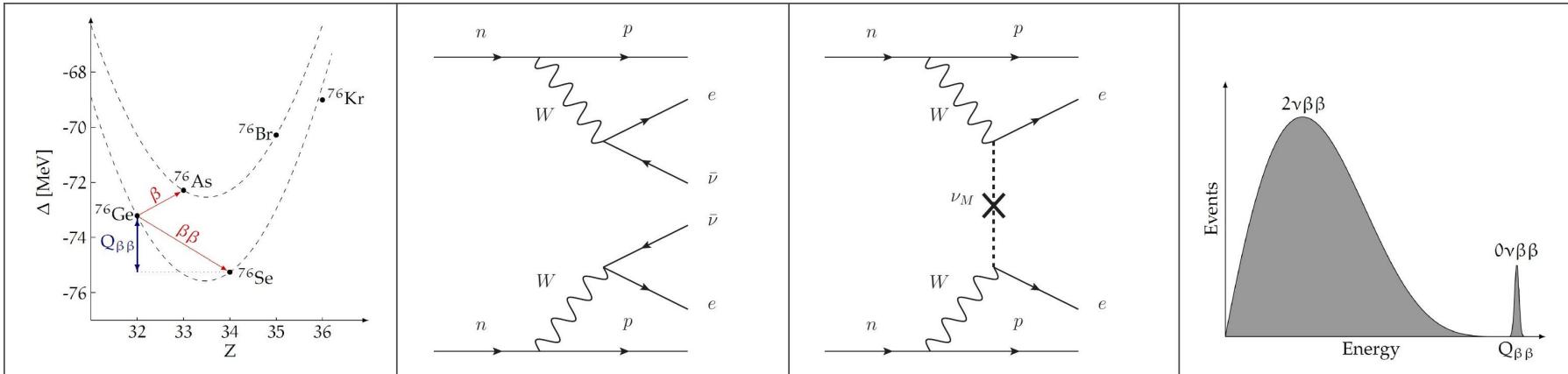
How can we test the conservation of B, L, (B-L)?

ΔL	ΔB	$\Delta(B-L)$	Process
-1	-1	0	$p \rightarrow e^+ + \pi^0$
+2	0	-2	$(A,Z) \rightarrow (A,Z+2) + 2e^-$



Matter creation and destruction

Expected $0\nu\beta\beta$ decay signature



$\beta\beta$ decay signature

- Continuum for $2\nu\beta\beta$ decay
- Peak at $Q_{\beta\beta}$ for $0\nu\beta\beta$ decay
⇒ Energy peak is the only necessary and sufficient signature to claim a discovery
- Additional signatures from signal topology, pulse shape discrimination, multiple channel readout, daughter tagging, ...

$0\nu\beta\beta$ decay rate

- $$(T_{1/2}^{0\nu})^{-1} = G_{0\nu} \cdot |M_{0\nu}|^2 \cdot |f|^2 / m_e^2$$
- $T_{1/2}^{0\nu}$ = $0\nu\beta\beta$ decay halflife
 - $G_{0\nu}$ = phase space (known)
 - $M_{0\nu}$ = nuclear matrix element (NME)
 - f = new physics term

Neutrino masses and $0\nu\beta\beta$ decay

- Neutrinos are purely left handed
⇒ Not true: experimentally $\Delta m_{ij} \neq 0$
- Majorana condition: in the rest frame, $v_{IR} = v_{IL}^c$
⇒ $v_i^c = \bar{v}$
⇒ v and \bar{v} are only distinguished by the spin
- Majorana condition is not Lorentz invariant
⇒ L must be violated at the (m_ν / p_ν) scale
- Majorana mass term:

$$\mathcal{L}_M = \frac{1}{2} \sum_{l,l'} \bar{v}_l m_{ll'} v_{l'}^c + h.c.$$

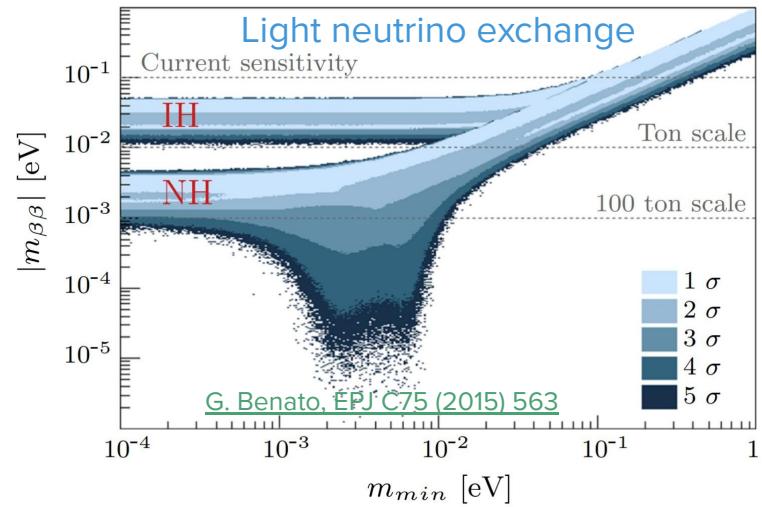
- What mechanism generates $m_{ll'}$?
- What predictions can we make on $|f|$?

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- What mechanism generates $m_{ll'}$?
- What predictions can we make on $T_{1/2}^{0\nu}$?



$$|f| = |m_{\beta\beta}| = \left| \sum_i U_{ei}^2 m_i \right|$$

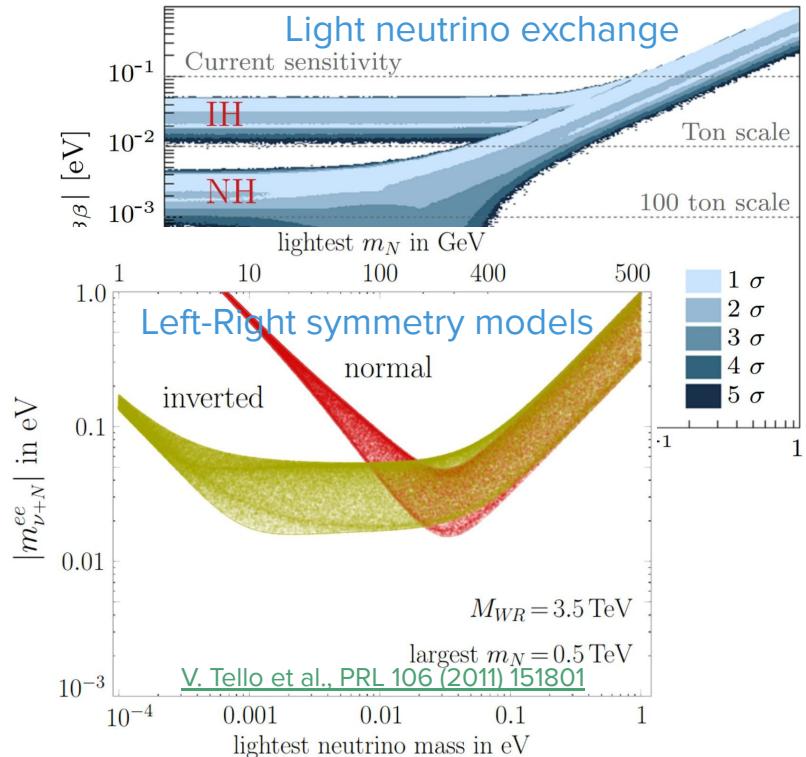
- “Minimal” model: add just one term to the Lagrangian
- “Simple” goal for future experiments
- “Easy” to compare results obtained with different isotopes

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- What predictions can we make on $T_{1/2}^{0\nu}$?

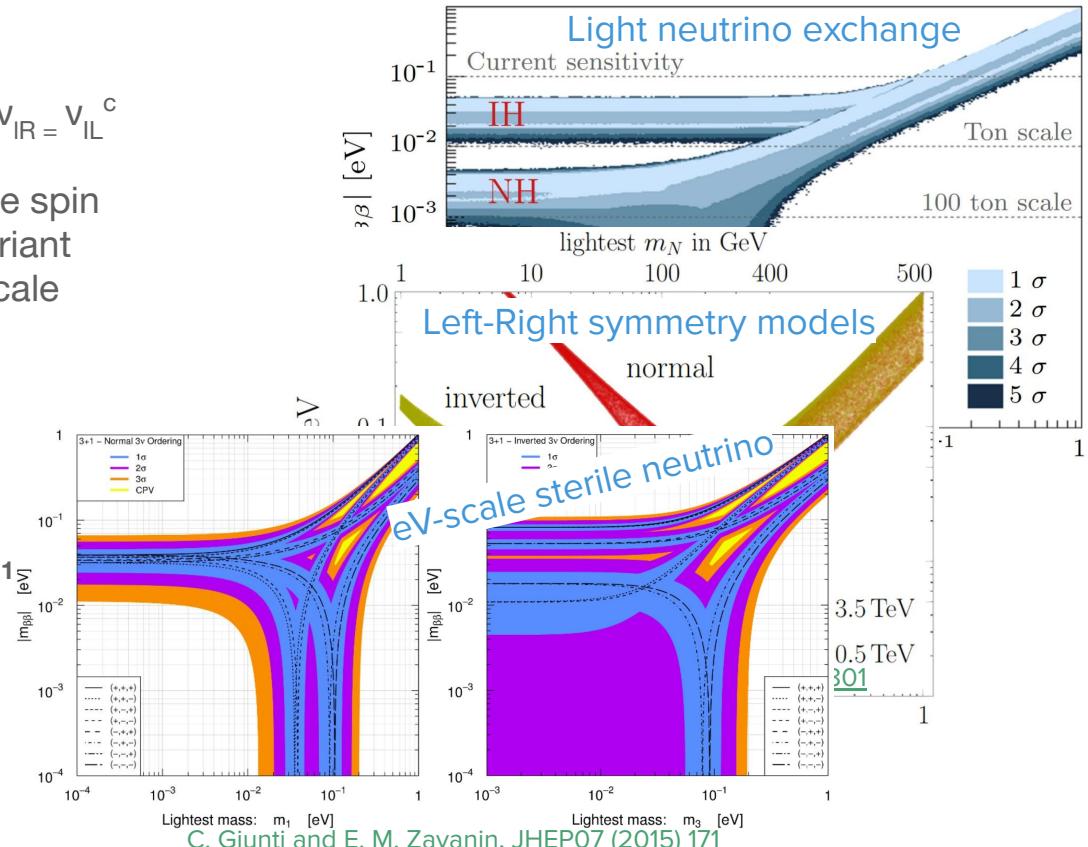


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- What mechanism generates $m_{II'}$?
- What predictions can we make on T_{13}

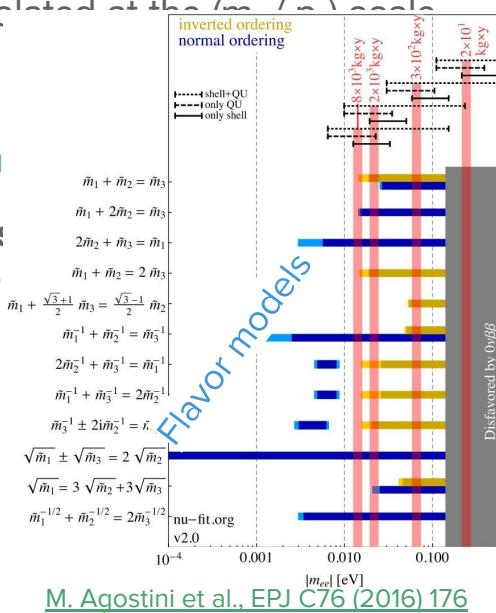


Neutrino masses and $0\nu\beta\beta$ decay

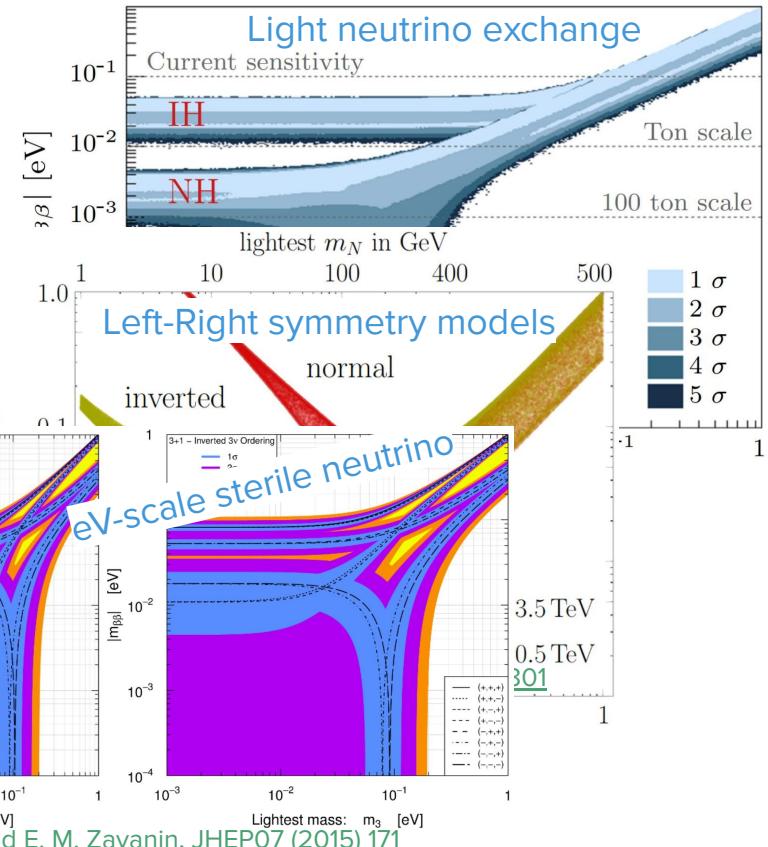
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⇒ L must be vice versa at the (m_1 / m_2) scale
- Majorana mass

$$\mathcal{L}_M = \frac{1}{2} \sum_{I,I'} \bar{v}_I m_I$$

- What mechanism?
- What predictions?



M. Agostini et al., EPJ C76 (2016) 176



Neutrino masses and $0\nu\beta\beta$ decay

- Neutrinos are purely left handed

\Rightarrow Not true: experimentally $\Delta m^2 \neq 0$

- Majorana condition:

$\Rightarrow v_L^c = v_L^-$

$\Rightarrow v$ and v are only c

- Majorana condition

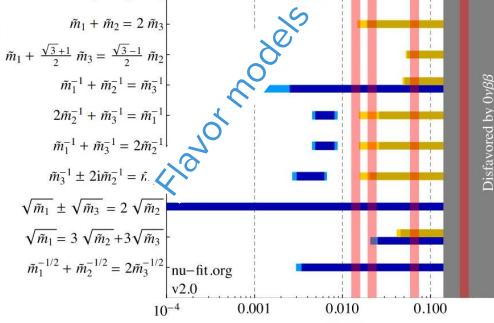
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- Majorana mass

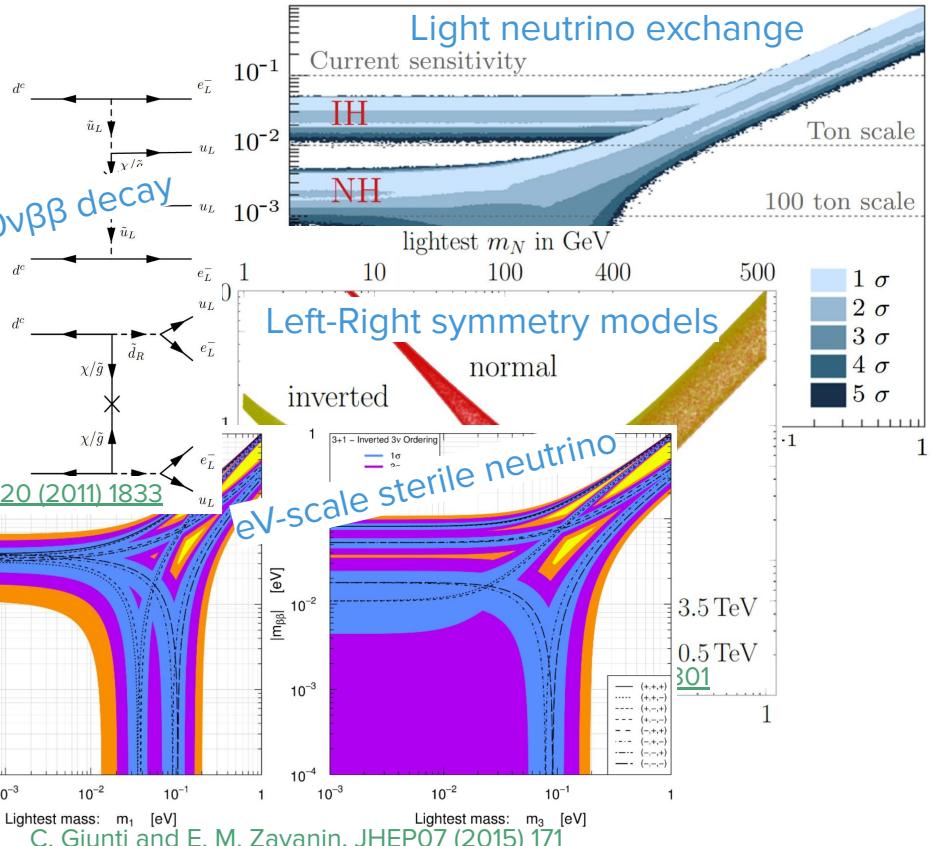
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M. Agostini et al., EPJ C76 (2016) 176



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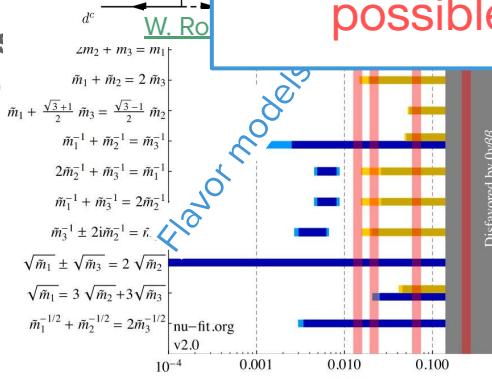
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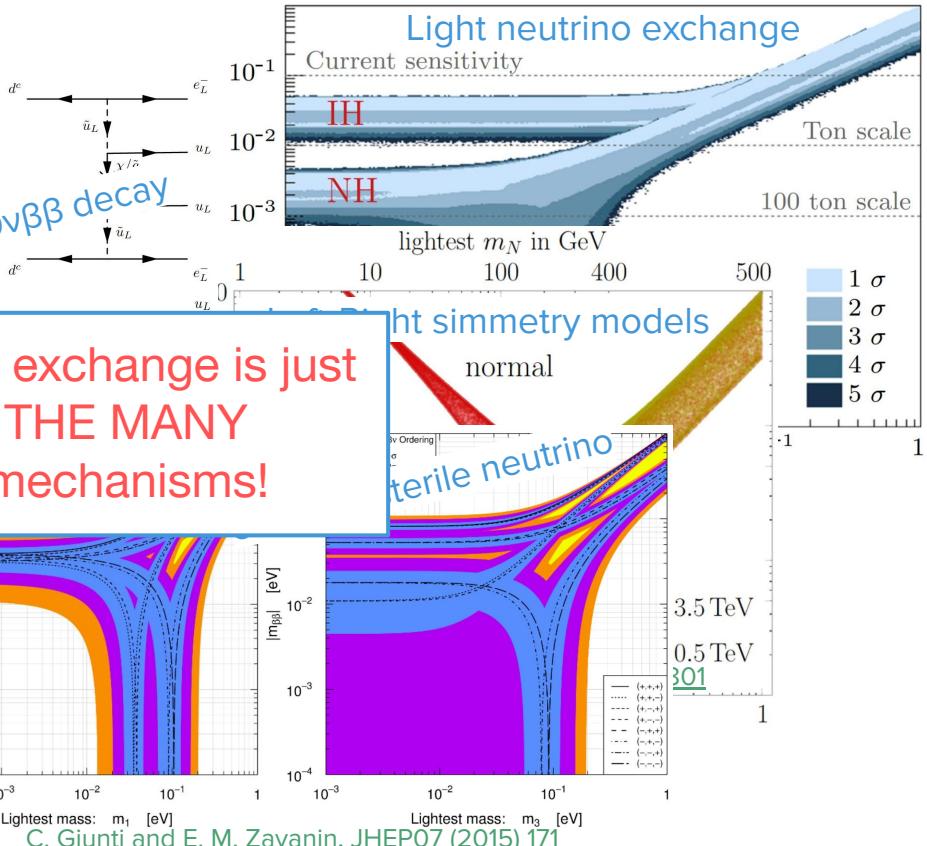
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- What mechanism?

- What predictions?



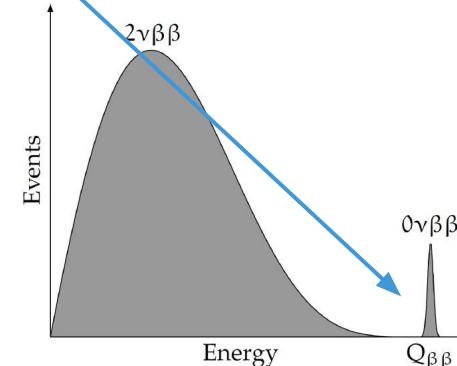
Light neutrino exchange is just
ONE OF THE MANY
possible mechanisms!



M. Agostini et al., EPJ C76 (2016) 176

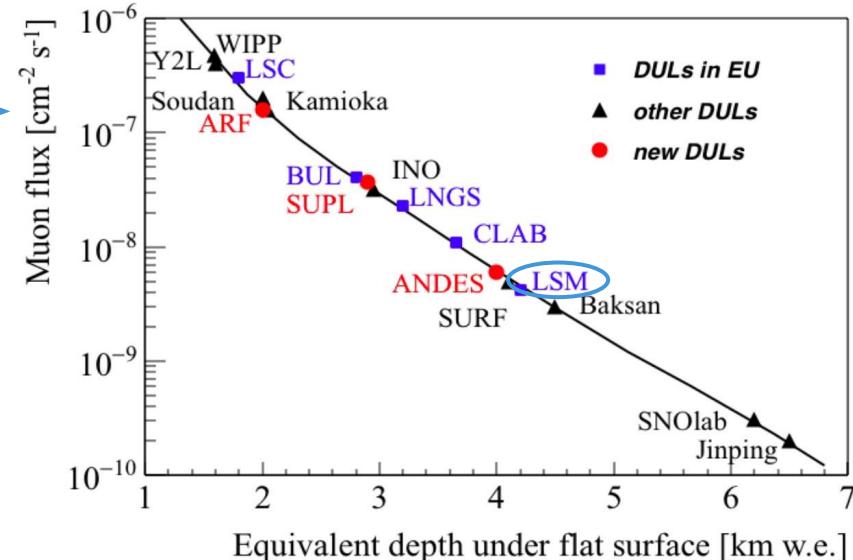
Backgrounds

- $Q_{\beta\beta}$ in the 2-3.5 MeV range for most used isotopes
- Cosmic muons
 - ⇒ Operate underground
- Neutrons (muon induced, fission, ...)
 - ⇒ Neutron absorbers (water, PE, borated PE, ...)
- Actinides (^{238}U and ^{232}Th) decay chains + Rn
 - α up to 8 MeV
 - β up to 3.3 MeV
 - γ up to 2.6 MeV
 - ⇒ Material selection
 - ⇒ Cleaning protocol
 - ⇒ Avoid recontamination
 - ⇒ Shielding and self-shielding
 - ⇒ Event topology
 - ⇒ Particle discrimination via pulse shape
- Irreducible $2\nu\beta\beta$ background
 - Tail of $2\nu\beta\beta$ spectrum
 - ⇒ Energy resolution
 - Pile-up of $2\nu\beta\beta$ events
 - ⇒ Time resolution



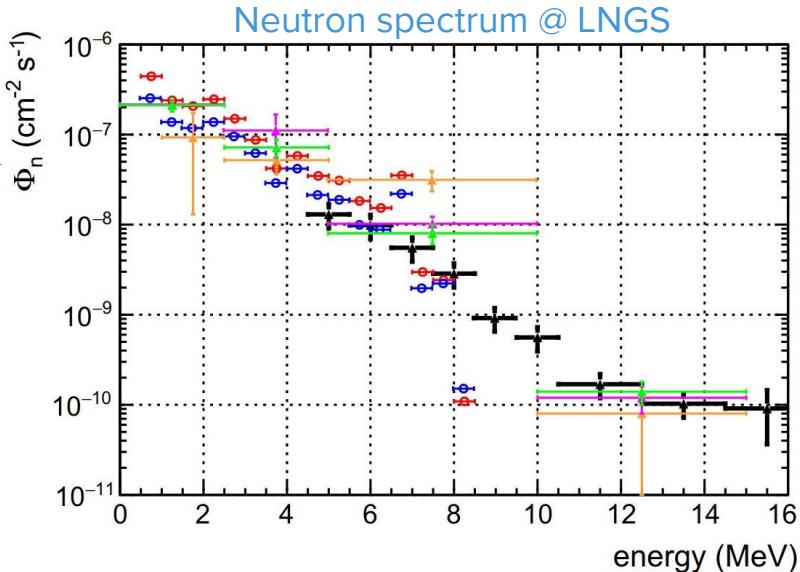
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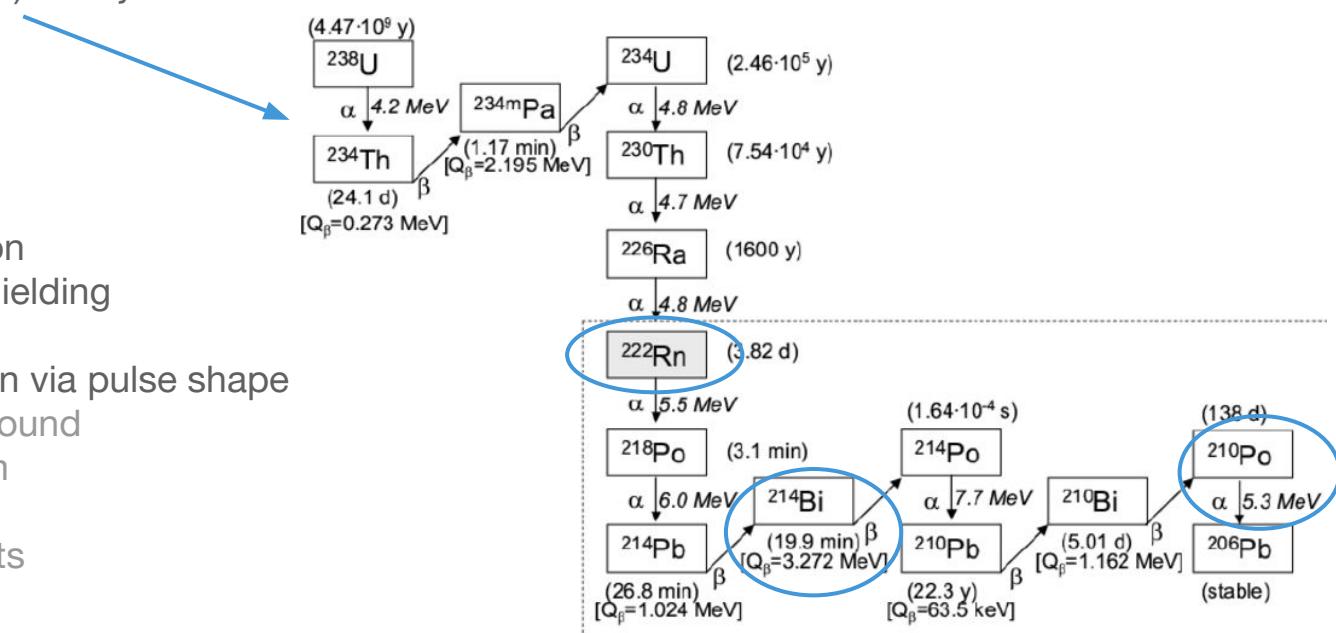


G. Bruno and B. Fulgione, EPJ C79 (2019) 747

Backgrounds

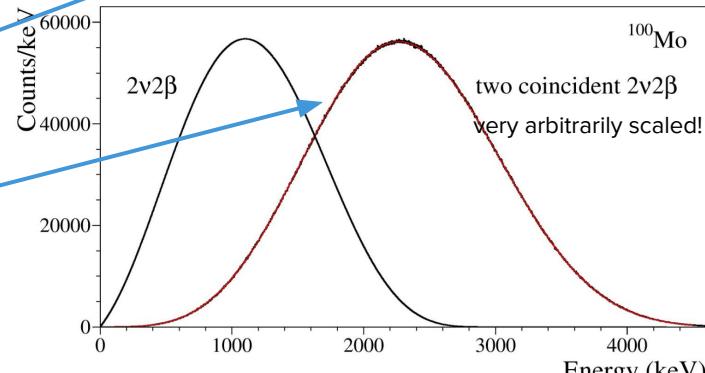
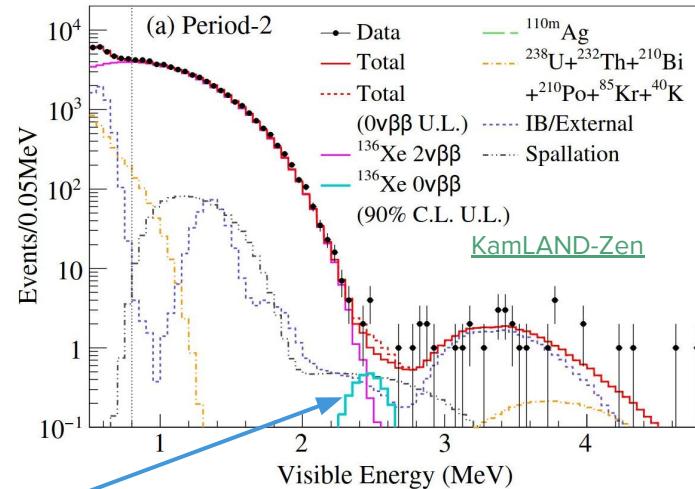
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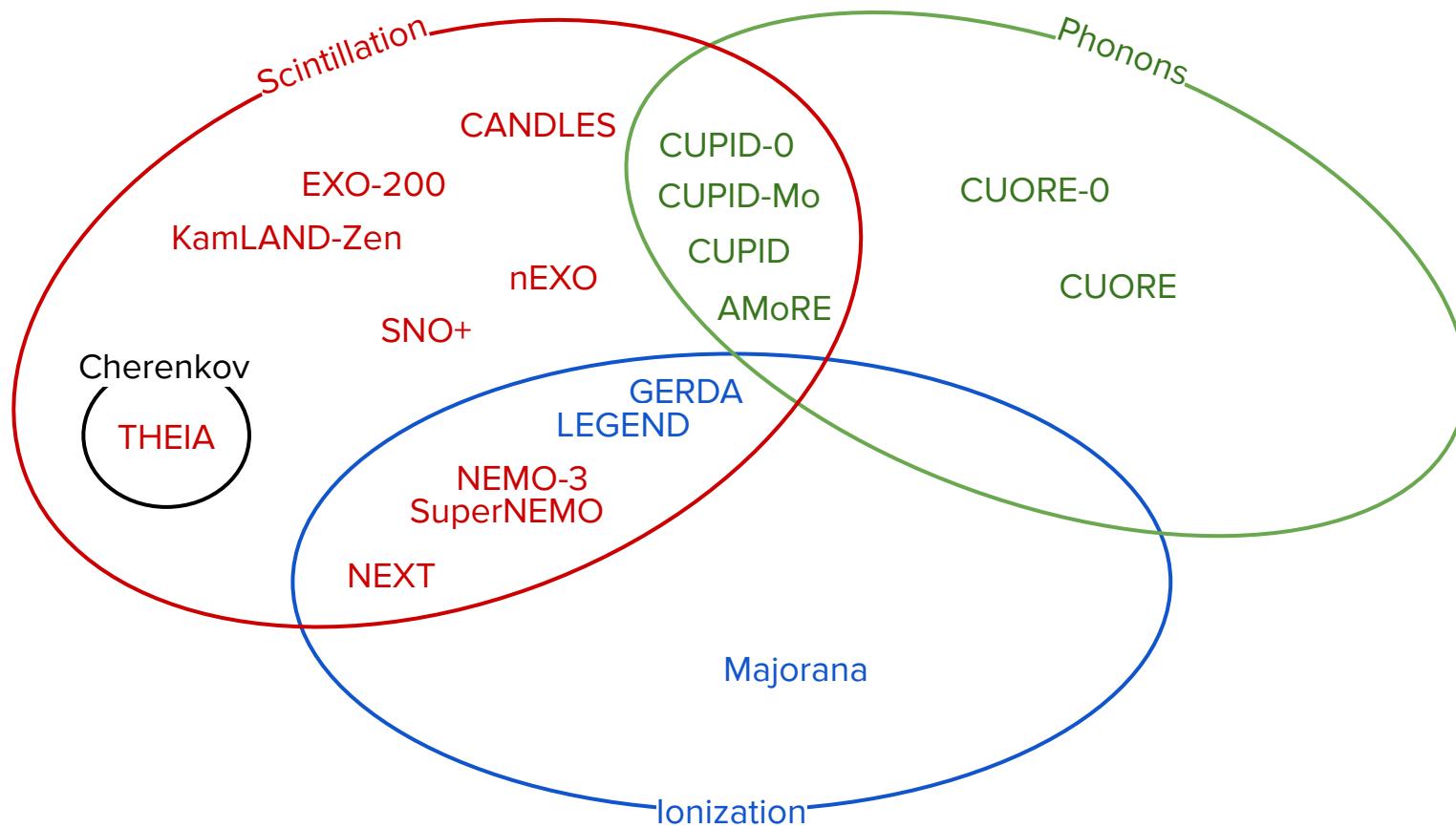
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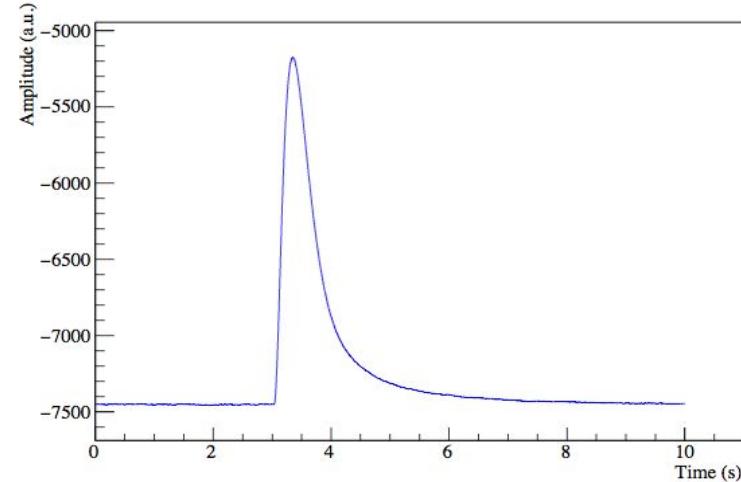
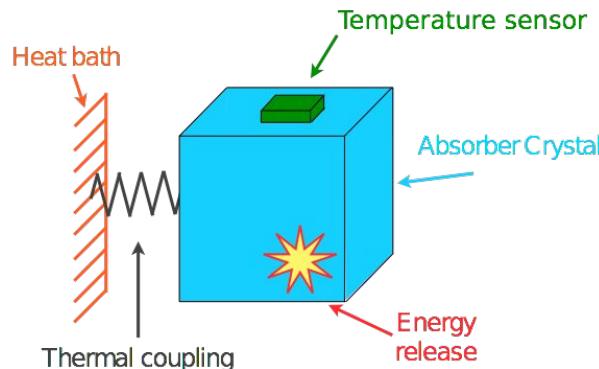
D. M. Cherniak et al., EPJ C72 (2012) 1989

$0\nu\beta\beta$ decay experimental fauna



Cryogenic calorimeters a.k.a bolometers

- Low heat capacity @ $T \sim 10$ mK
- Excellent energy resolution (**~0.2% FWHM**)
- Detector agnostic to origin of energy deposition
- Detector response of O(1) sec if readout with Neutron Transmutation Doped (NTD) Ge sensors



Simplified thermal model

- Crystal heat capacity: C
- Conductivity of coupling to thermal bath: G
- Signal amplitude $\propto \Delta T = E_{\text{dep}}/C$
- Decay constant: $\tau = G/C$

Scintillating crystals

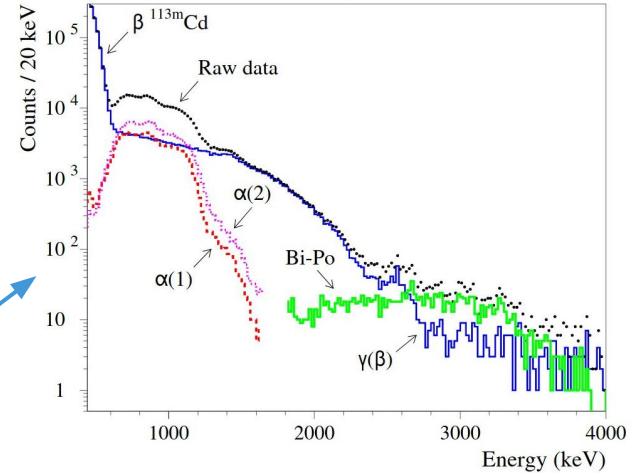
Scintillation light features

- Typical light yield (LY): O(10) photons/keV
→ Expected energy resolution: few %
- Amount of emitted light is particle dependent
- For some crystals, time profile of scintillation light is particle dependent

$^{116}\text{CdWO}_4$

light

heat

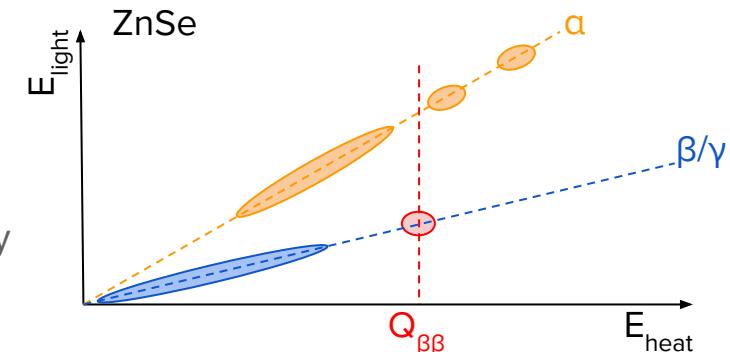
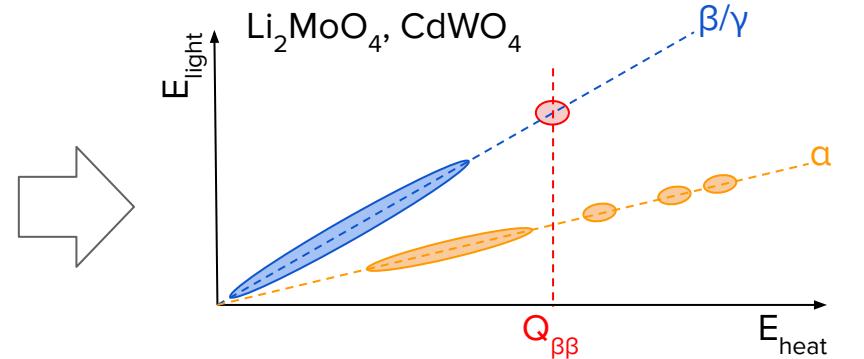
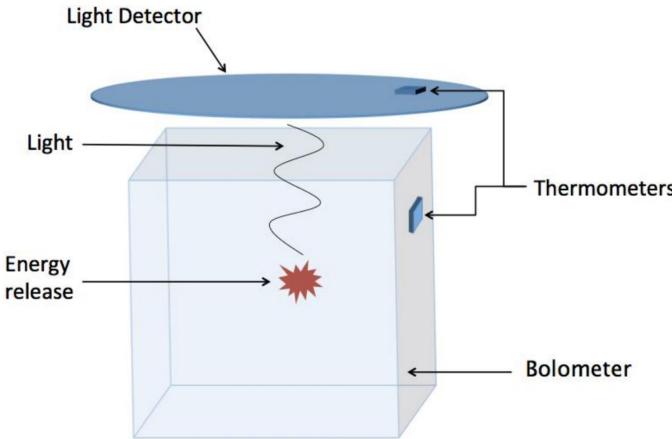


Scintillating crystals for $0\nu\beta\beta$ decay

- Heat to measure energy
- Scintillation light for particle identification (PID)

D. Helis et al., LTD 2018, 467
A. S. Barabash et al, Phys.Rev. D98 (2018) 092007

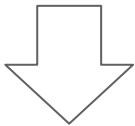
Scintillating bolometers



- Main background: surface α events
- Couple main crystal with secondary bolometer reading the scintillation (or Cherenkov) light
- Exploit different light yield (LY) of α vs β/γ to actively suppress background
- Typical light detector: thin Ge wafer coupled to thermometer (NTD, TES, KID, MMC)

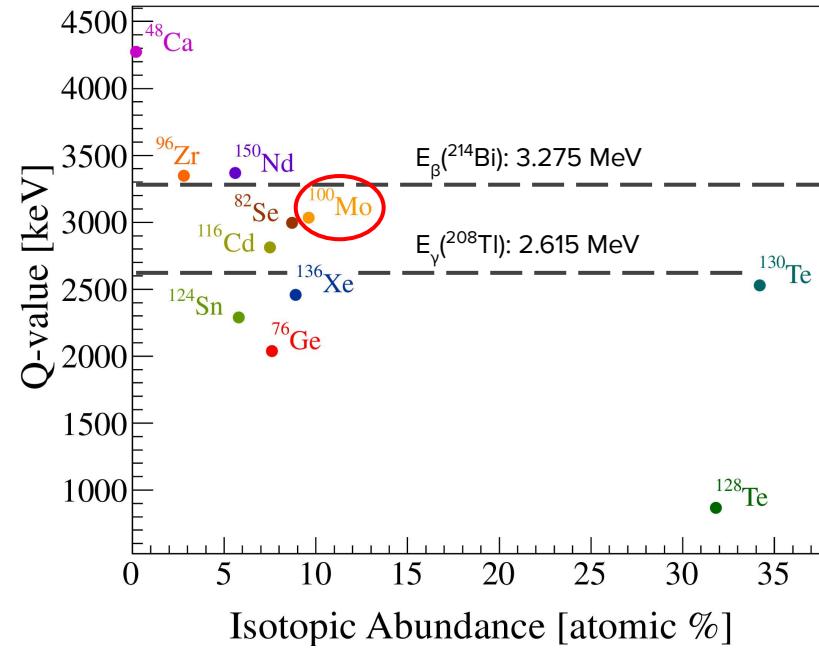
Isotope choice for bolometric experiments

- High isotopic abundance
- Enrichment possible at reasonable cost?
- $Q_{\beta\beta}$ above end point of β or γ radiation?
- Scintillating crystal available?
- Large scale crystal production possible?

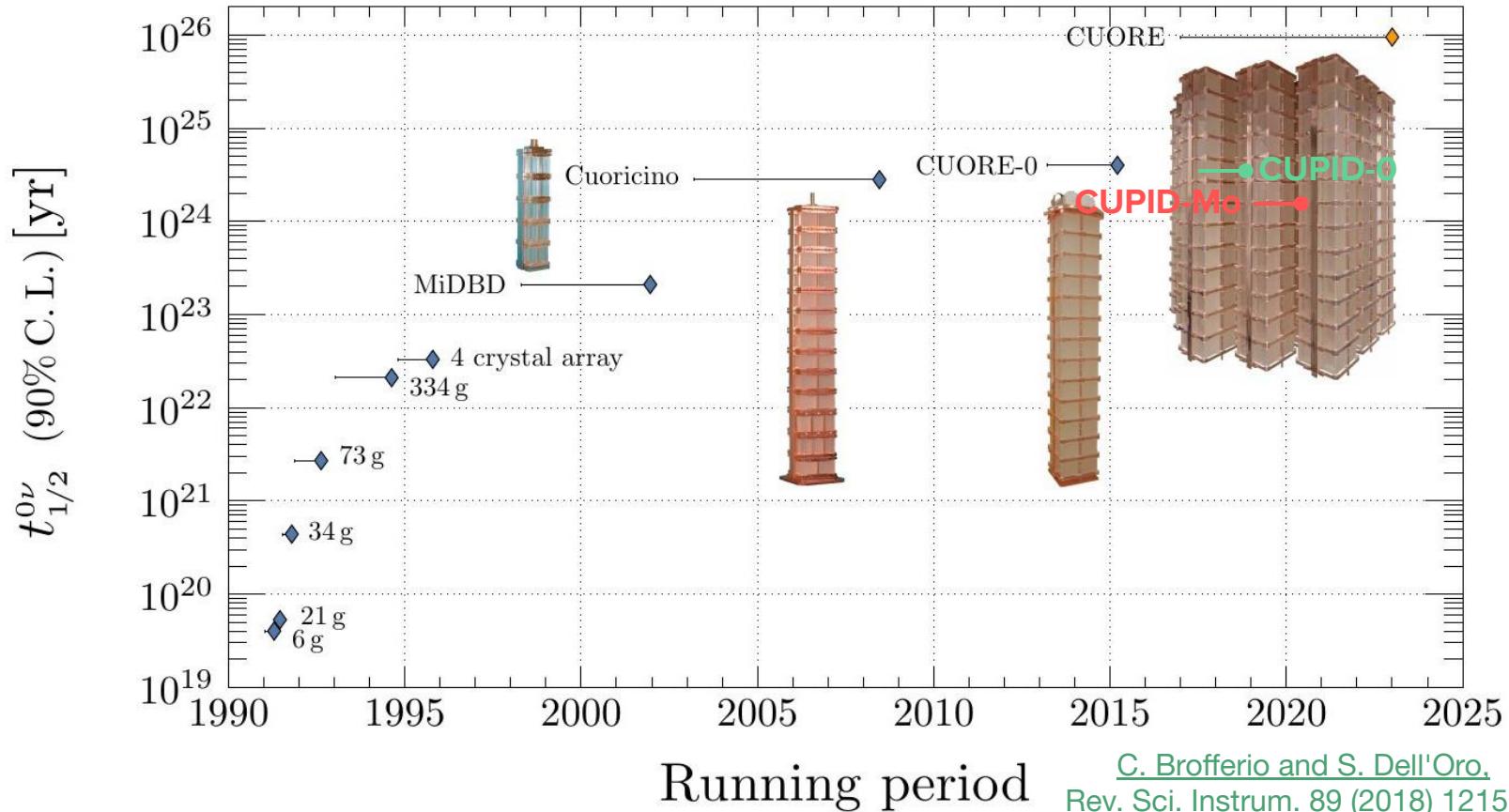


Advantages of bolometric approach

- Detectors and infrastructure are decoupled. Same cryogenic infrastructure re-usable with different isotopes and/or crystals
- Perfect for test of discovery or precision measurements



History of bolometric $0\nu\beta\beta$ decay searches



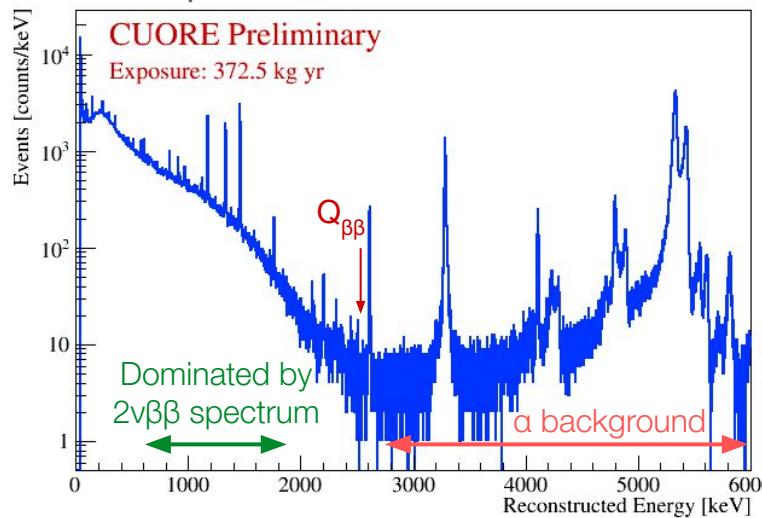
Running period

C. Brofferio and S. Dell'Oro,
Rev. Sci. Instrum. 89 (2018) 121502

CUORE: the Cryogenic Underground Observatory for Rare Events

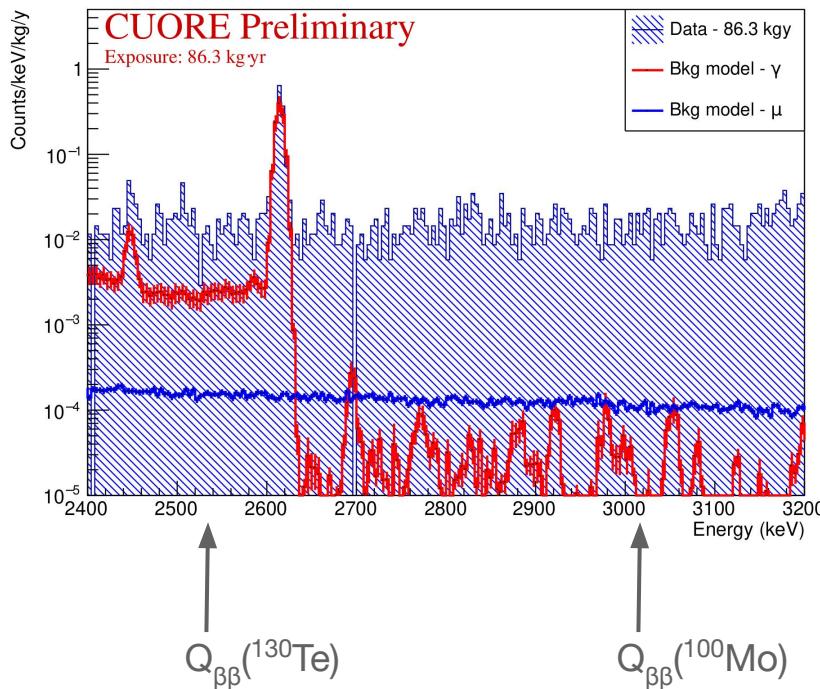


- 988 TeO₂ crystals with natural Te composition
→ **742 kg of total mass**, 206 kg of ¹³⁰Te mass
- Located in Hall A of the Gran Sasso National Lab
- Current limit: $T^{0\nu}_{1/2} > 3.2 \cdot 10^{25}$ yr @ 90% C.I.
- $Q_{\beta\beta}({}^{130}\text{Te}) = 2527.5$ keV
→ Above most γ background, below the ²⁰⁸Tl 2.6 MeV line
- TeO₂ crystals do not scintillate
→ no particle discrimination



Lessons learned from CUORE

ROI - External sources



- Most measured background is due to α particles (U/Th close to TeO_2 crystals)
→ α/β discrimination is required
- A $Q_{\beta\beta} > 2.6$ MeV would automatically reduce the remaining non- α background by >1 order of magnitude
- Muons are the dominant contribution after α 's
→ active muon veto

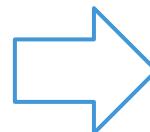
CUPID: the CUORE Upgrade with Particle IDentification

Goals:

- ~1500 $\text{Li}_2^{100}\text{MoO}_4$ scintillating crystals
→ ~250 kg of ^{100}Mo
- FWHM: 5 keV at $Q_{\beta\beta}$
- α rejection via PID with light detectors (LD)
- Background: 10^{-4} counts/keV/kg/yr
- Discovery sensitivity: $T_{1/2}^{0\nu} = 10^{27}$ yr

How do we get there?

- Large cryogenic infrastructure
→ Re-use CUORE cryostat
- Demonstrate LMO resolution
- Demonstrate PID performance of LDs
- Demonstrate reproducibility of performance
- Demonstrate low background

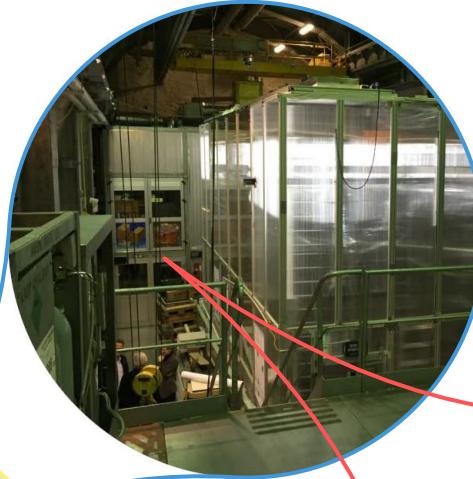
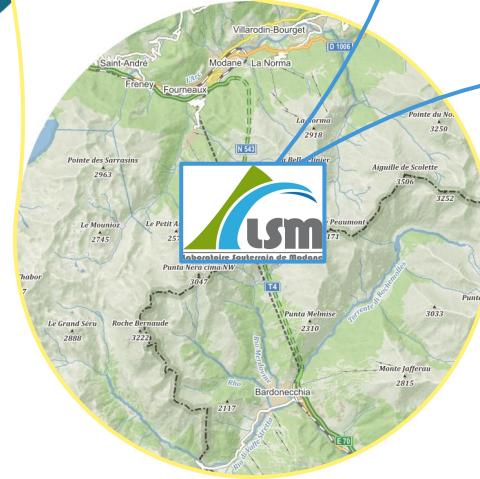


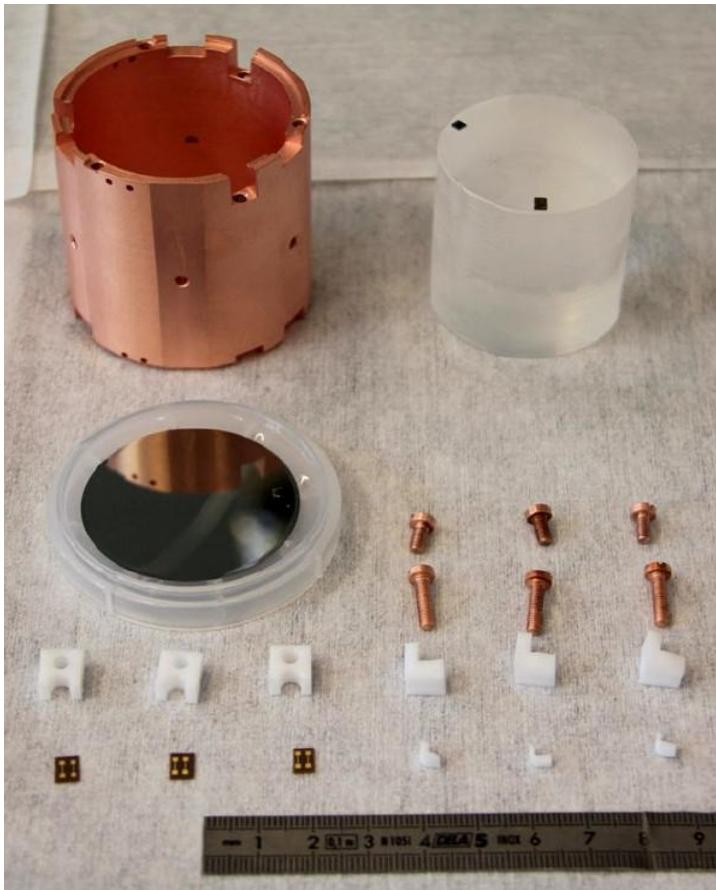
CUPID-Mo!

CUPID-Mo



CUPID-Mo

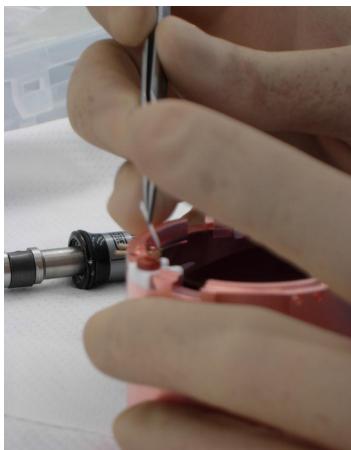
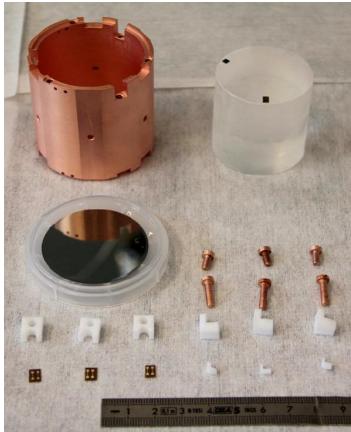




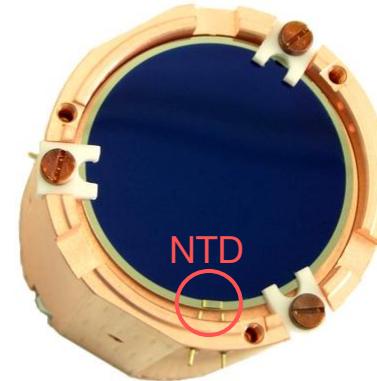
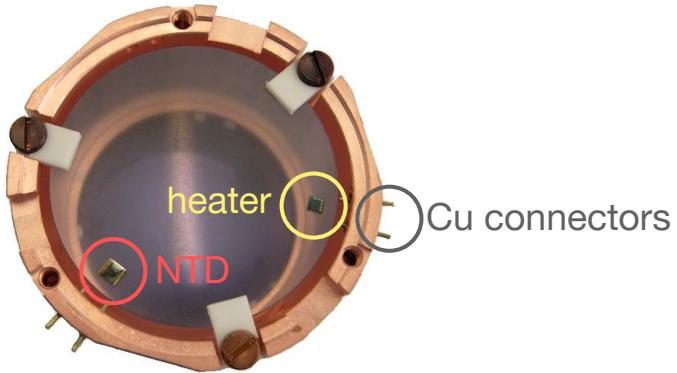
Ingredients

- 20 ^{enr}LMO crystals
 - Dimensions: \varnothing 44mm, h 45mm
 - Crystal mass: ~210 g each
 - ¹⁰⁰Mo isotopic fraction: $96.6 \pm 0.2\%$
- 20 Ge wafers
 - Dimensions: \varnothing 44.5mm, h 175 μ m
 - SiO coating on both sides to increase light collection
- Neutron transmutation doped (NTD) thermistors
 - 3x3x1mm³ for LMOs
 - 3x0.8x1mm³ for LDs
- Silicon-based resistors
 - To be used as heaters to inject pulser events

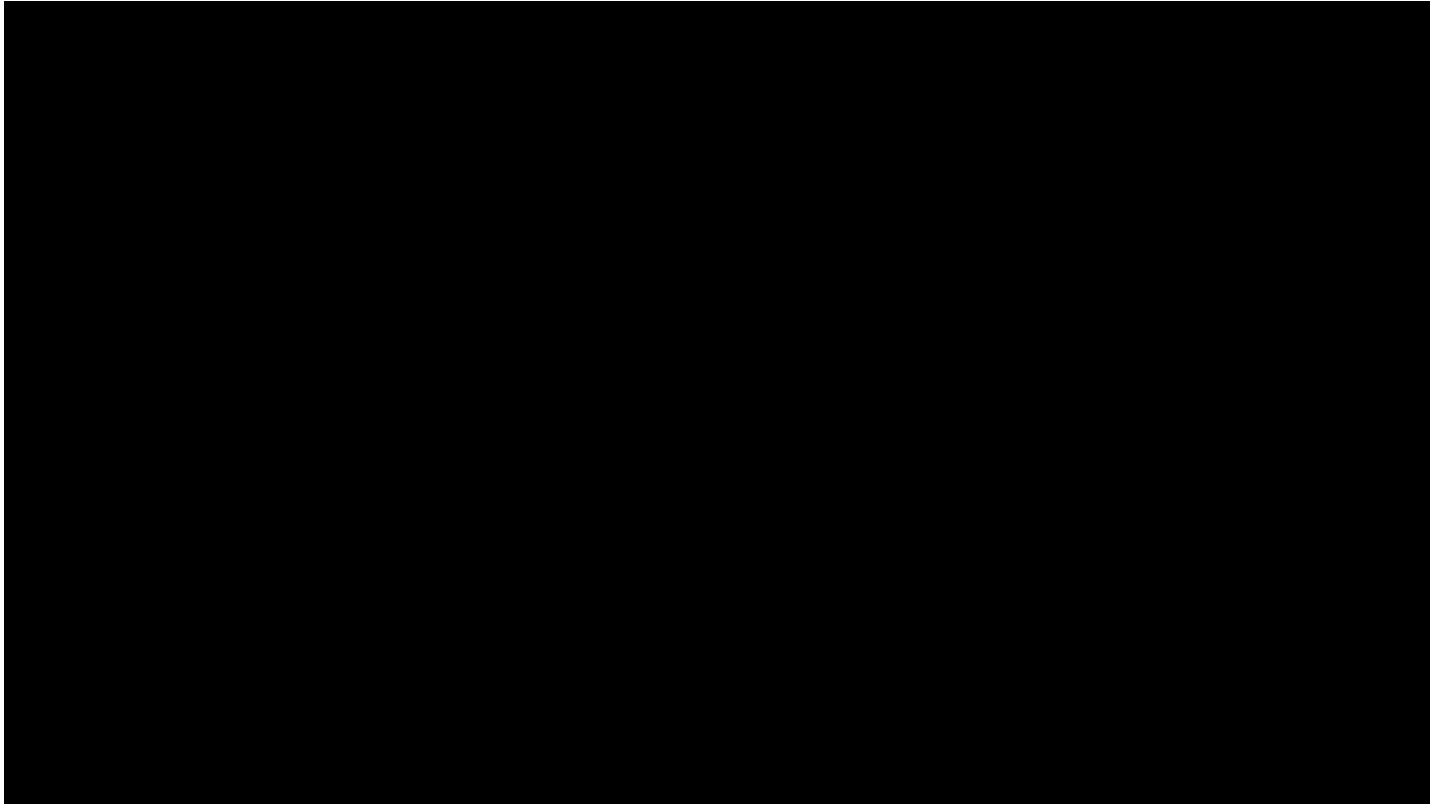
CUPID-Mo construction



CUPID-Mo construction

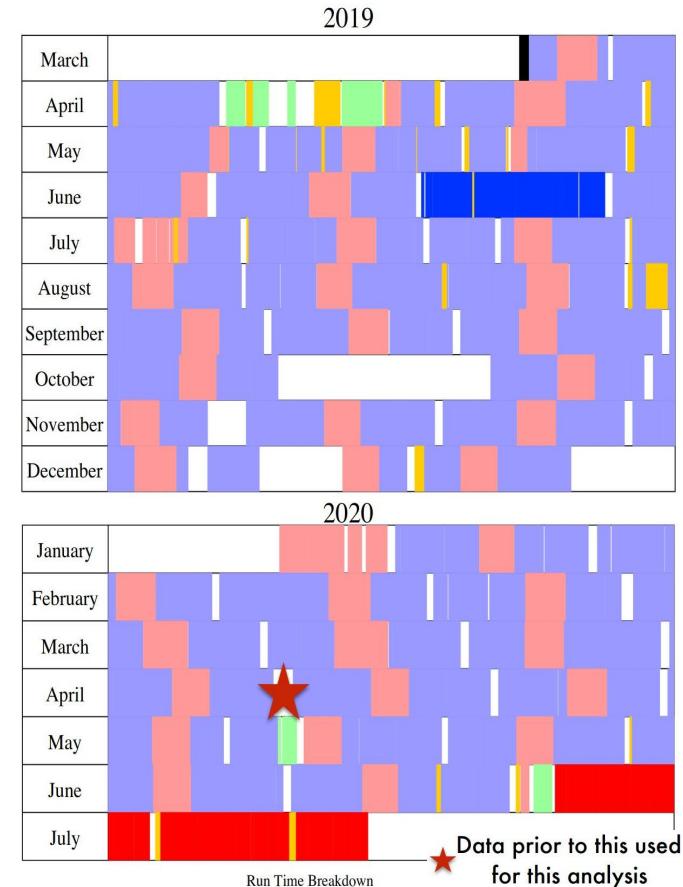
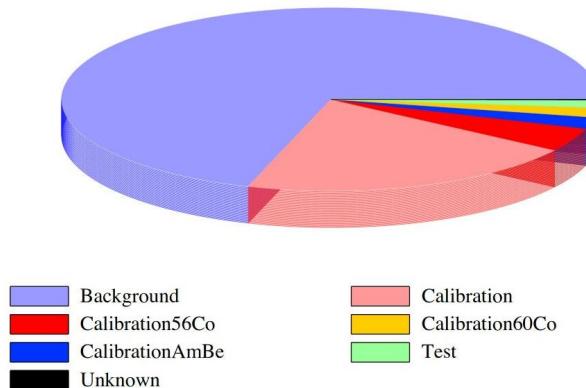


CUPID-Mo installation @ LSM

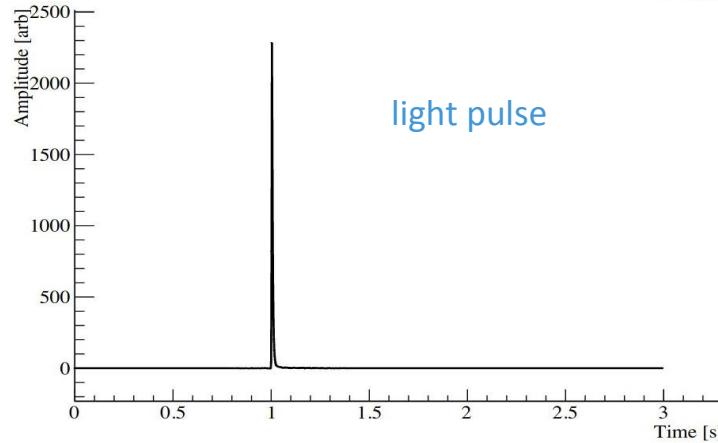
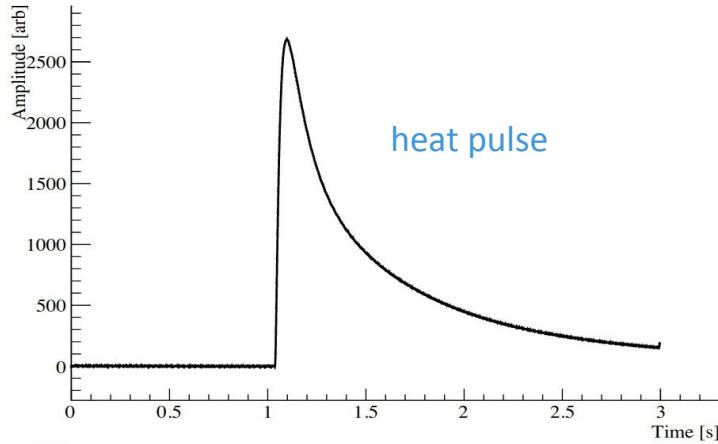


CUPID-Mo data taking

- Physics data: March 2019 - June 2020
- Analysed data: March-2019 - April 2020
 - 82% duty cycle
 - 73 days of calibration data
 - 240 days of physics ($0\nu\beta\beta$) data
 - 27 days excluded due to poor statistics in calibration
 - 213 days (7 datasets) of physics data good for analysis
 - Further 6% of data removed due to instabilities
 - One LMO removed due to poor performance
 - **Analyzed exposure: 2.16 kg·yr**



CUPID-Mo data processing

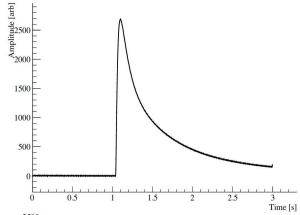


Data format

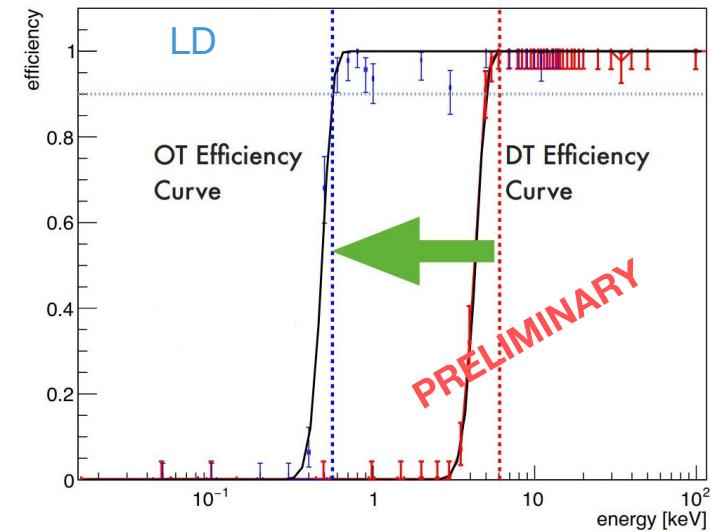
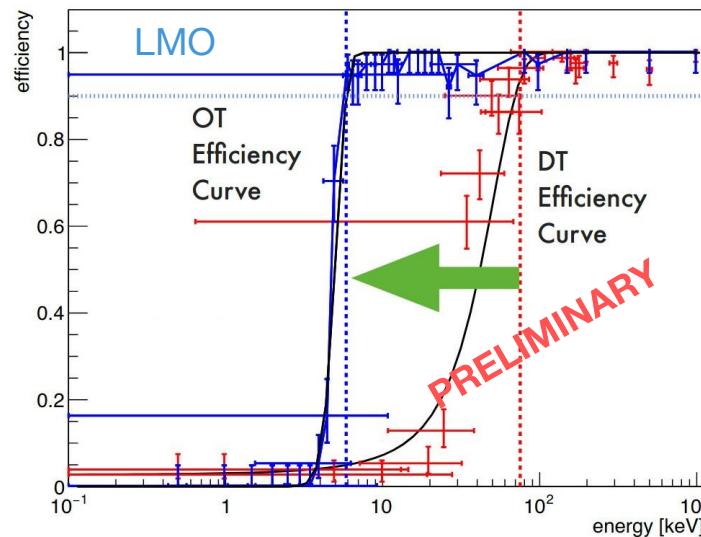
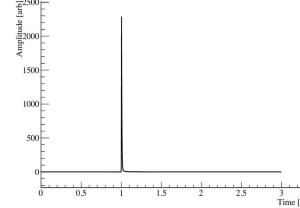
- 500 Hz sampling frequency
- Continuous data, no hardware trigger
- Data processed with Diana

[CUORE Collaboration, Phys. Rev. C93 \(2016\) 045503](#)
[CUPID-0 Collaboration, EPJ C78 \(2018\) 734](#)

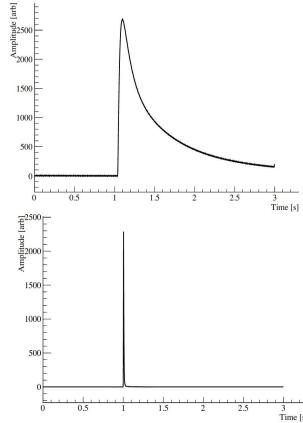
CUPID-Mo data processing



Optimum trigger

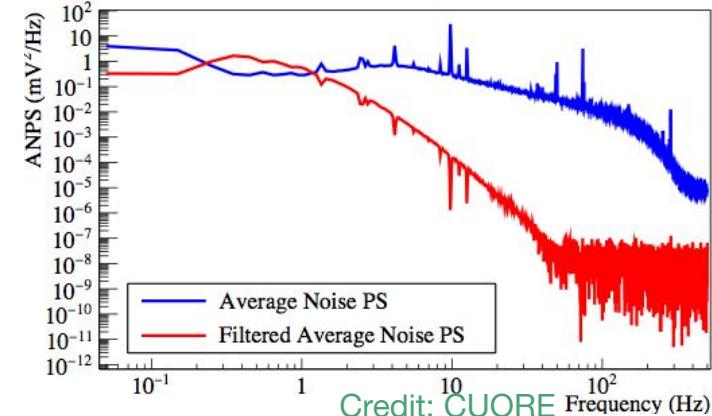
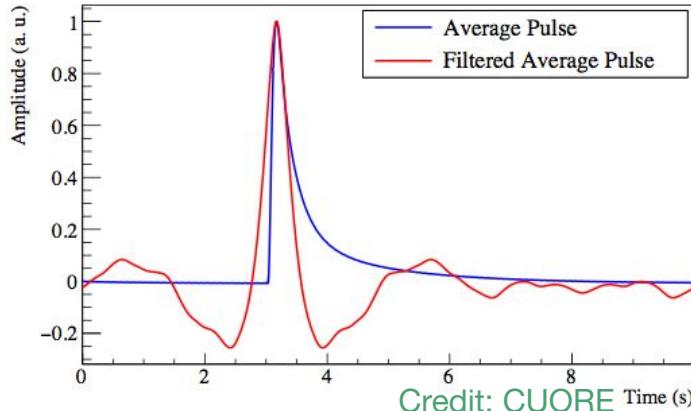


CUPID-Mo data processing

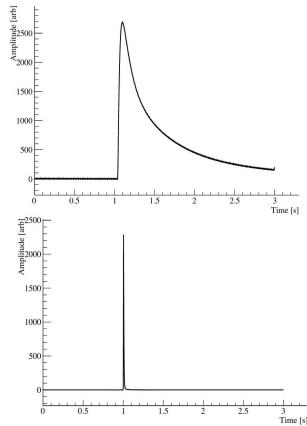


Optimum trigger Optimum filter

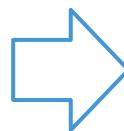
Optimum filter: matched filter that optimizes signal-to-noise ratio



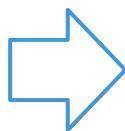
CUPID-Mo data processing



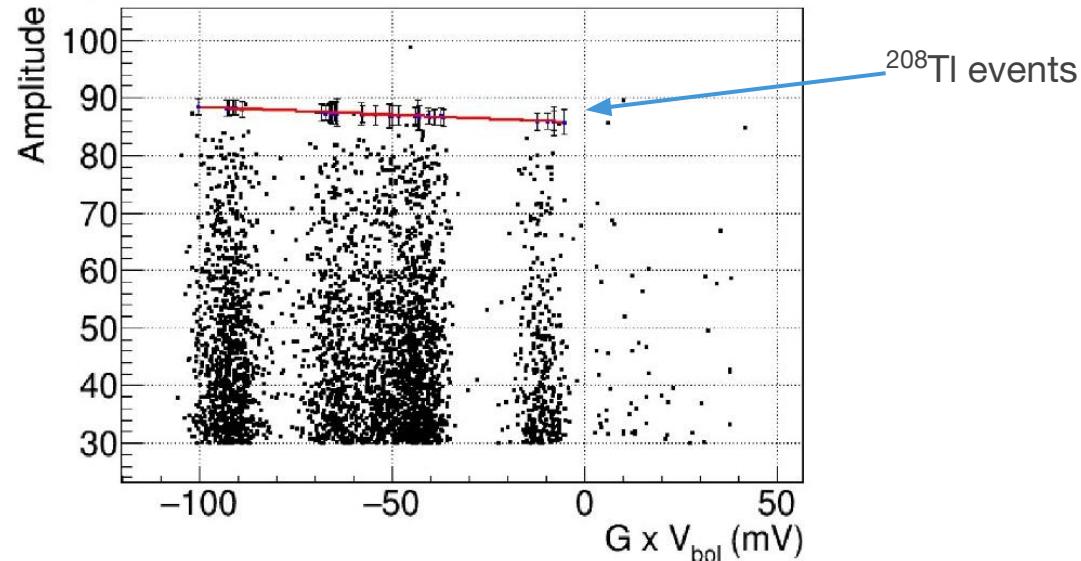
Optimum
trigger



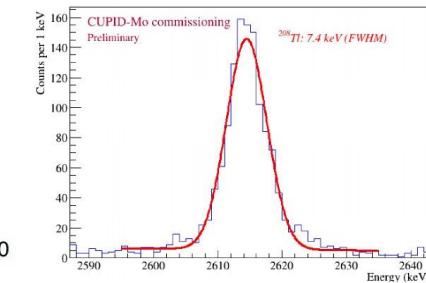
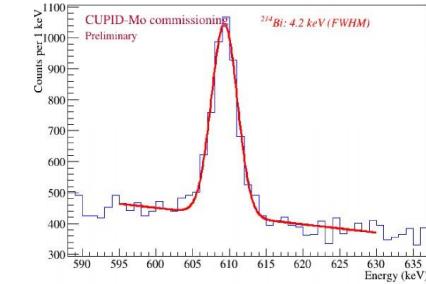
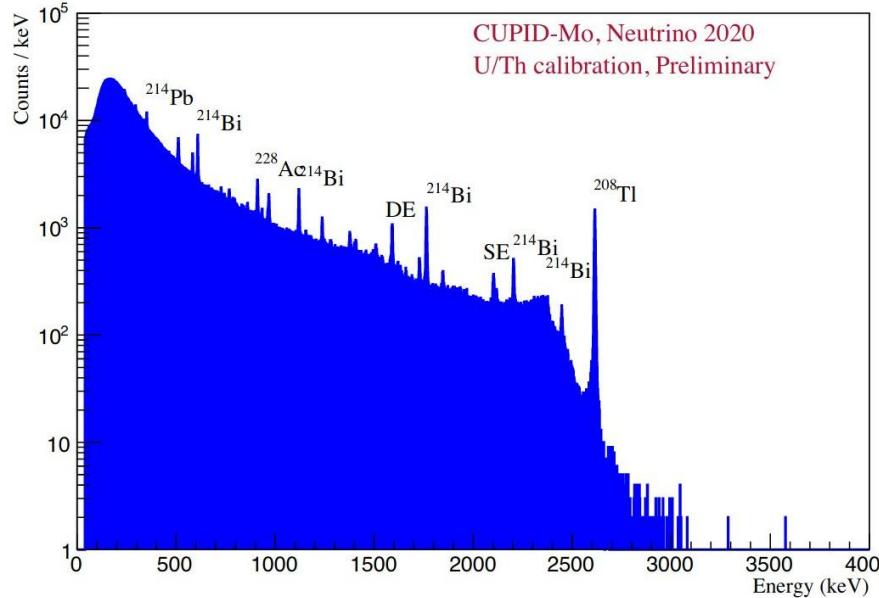
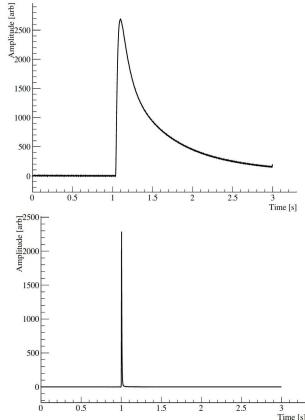
Optimum
filter



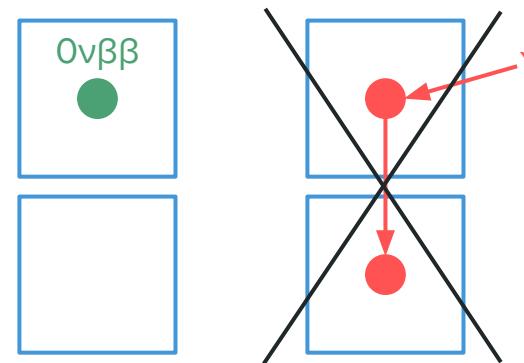
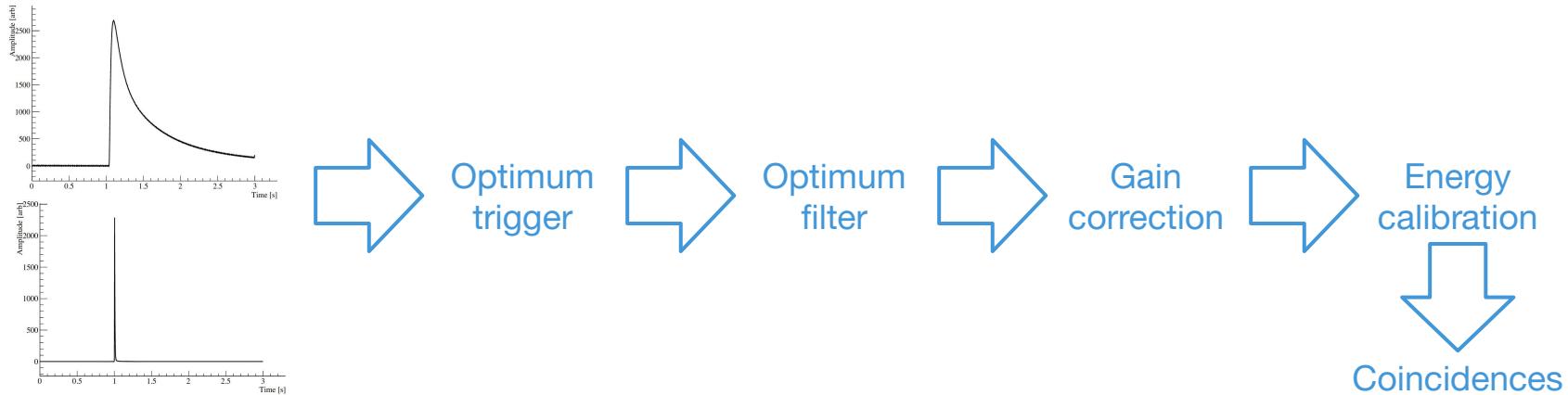
Gain
correction



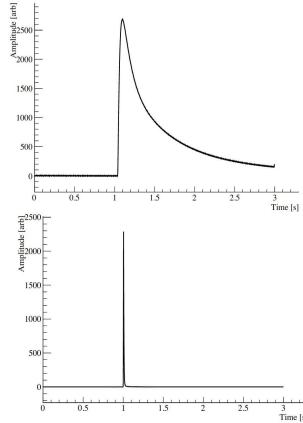
CUPID-Mo data processing



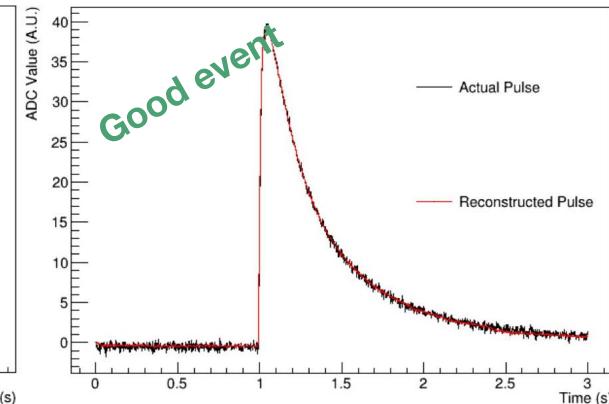
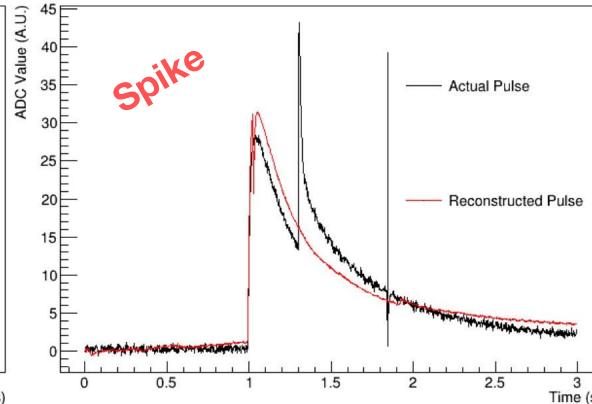
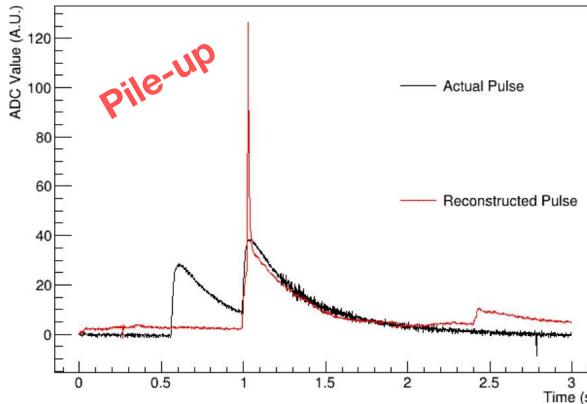
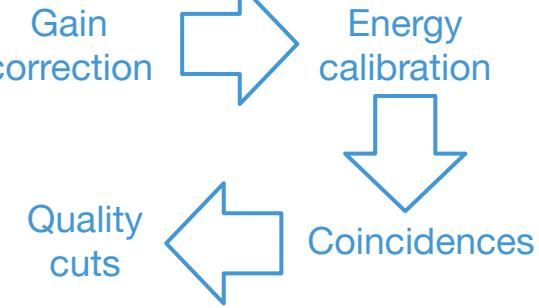
CUPID-Mo data processing



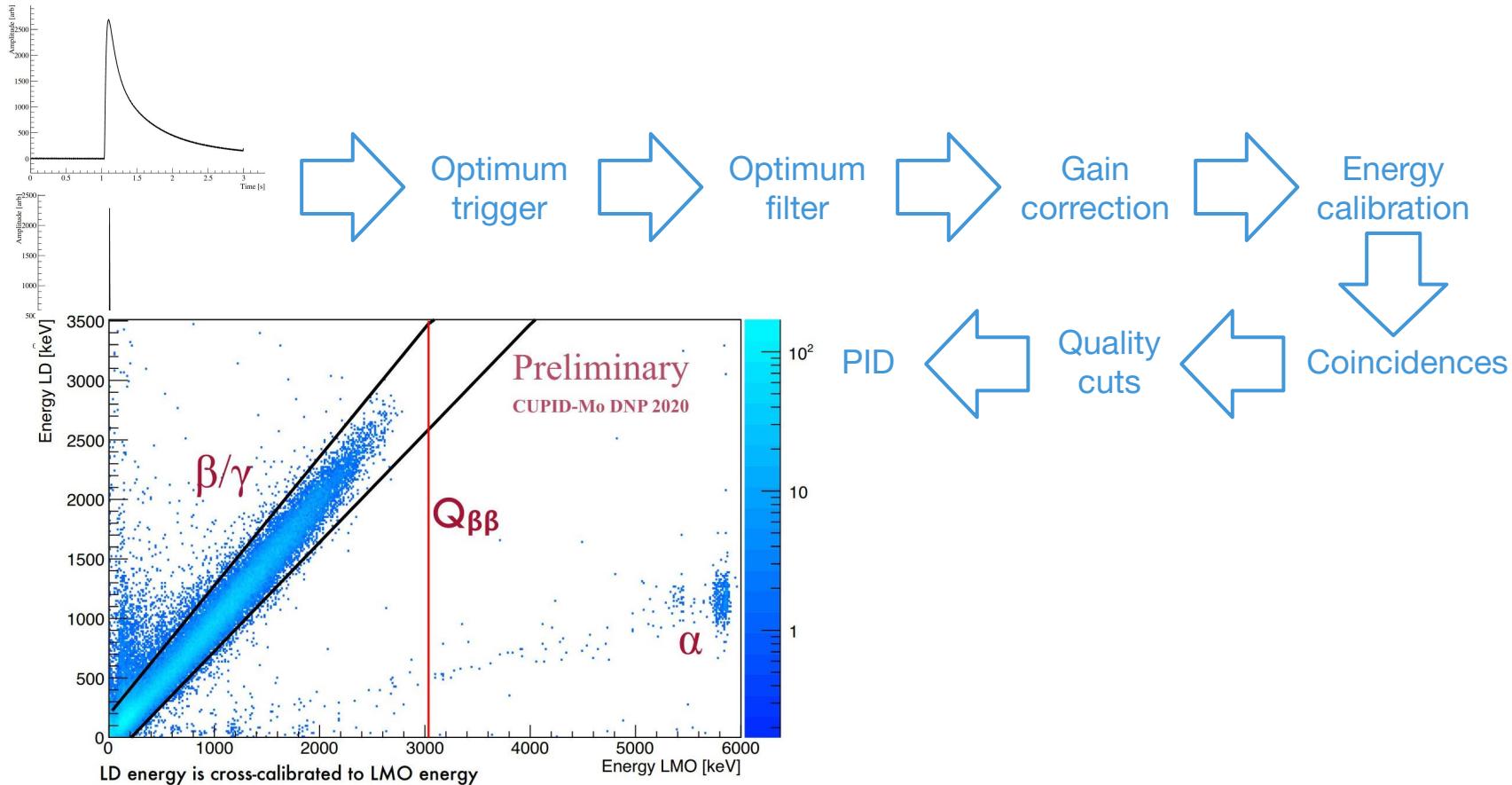
CUPID-Mo data processing



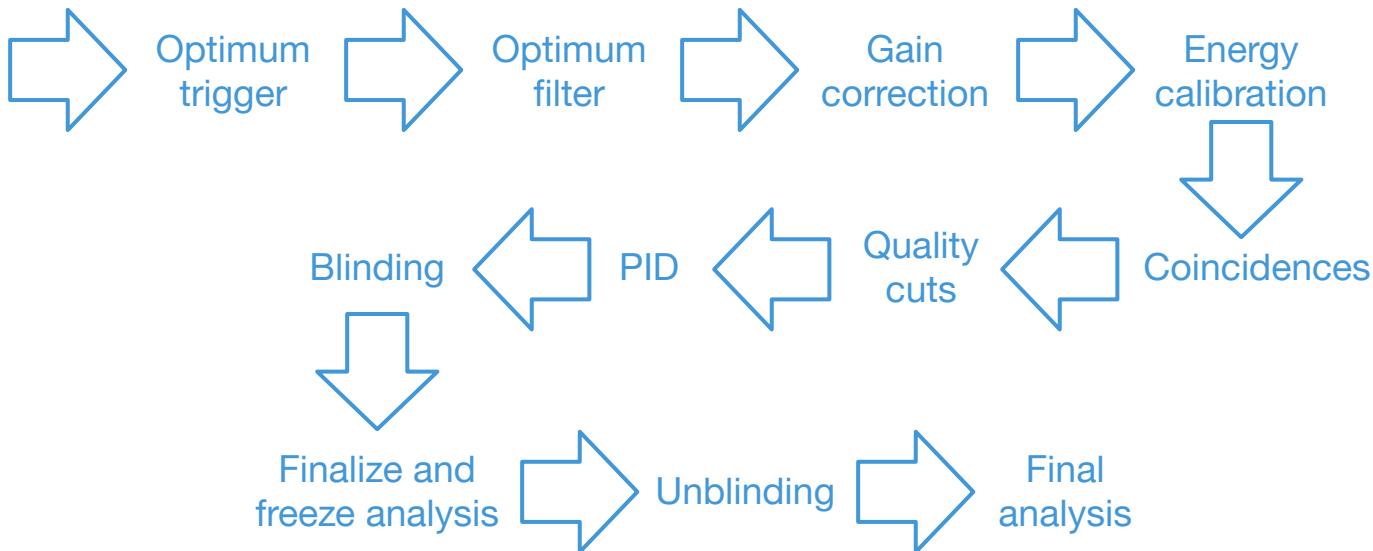
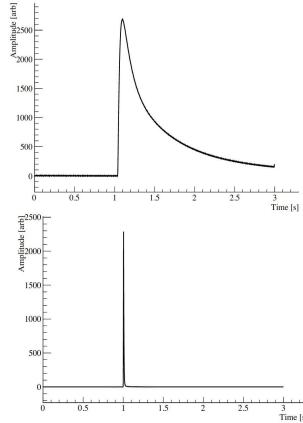
[CUPID-Mo, arXiv:2010.04033](https://arxiv.org/abs/2010.04033)
(submitted to JINST)



CUPID-Mo data processing

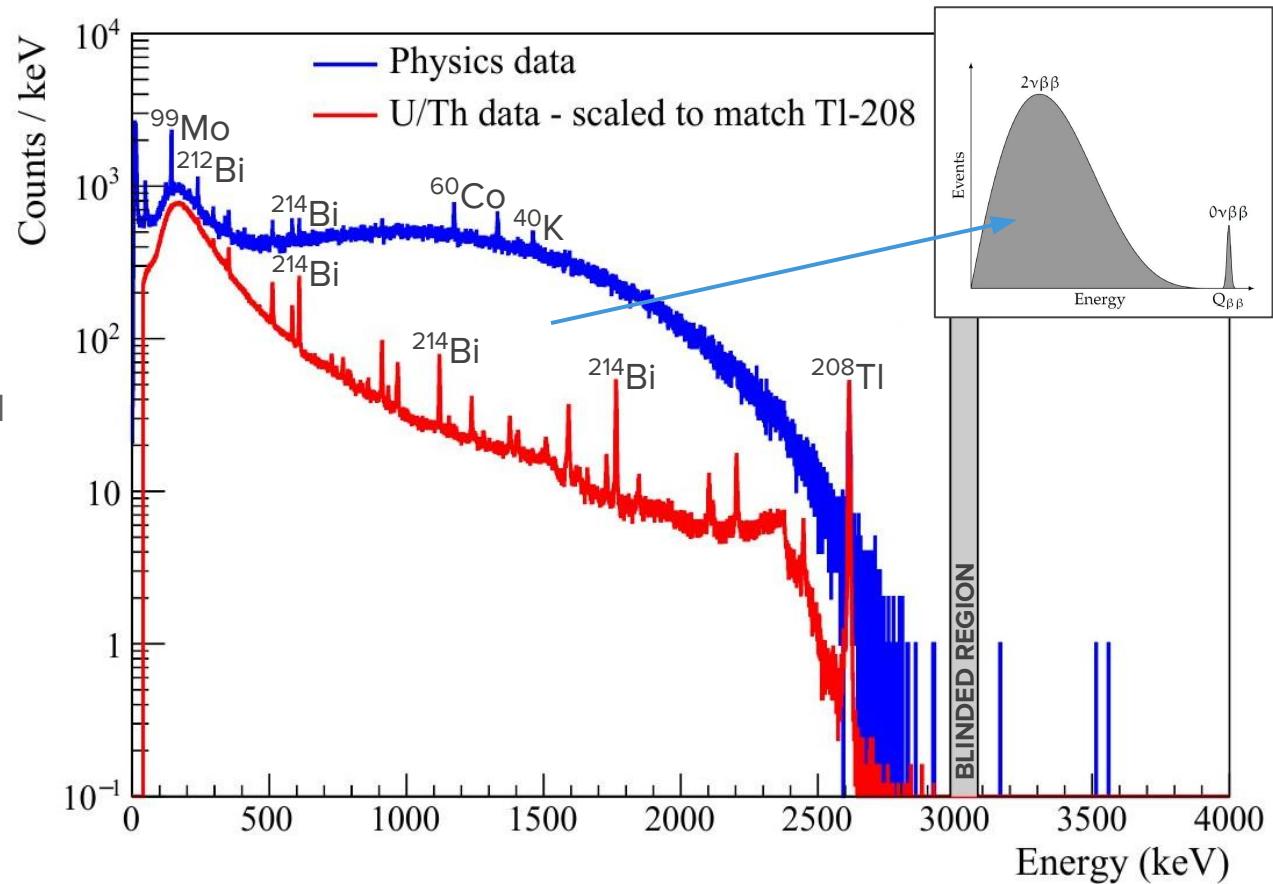


CUPID-Mo data processing

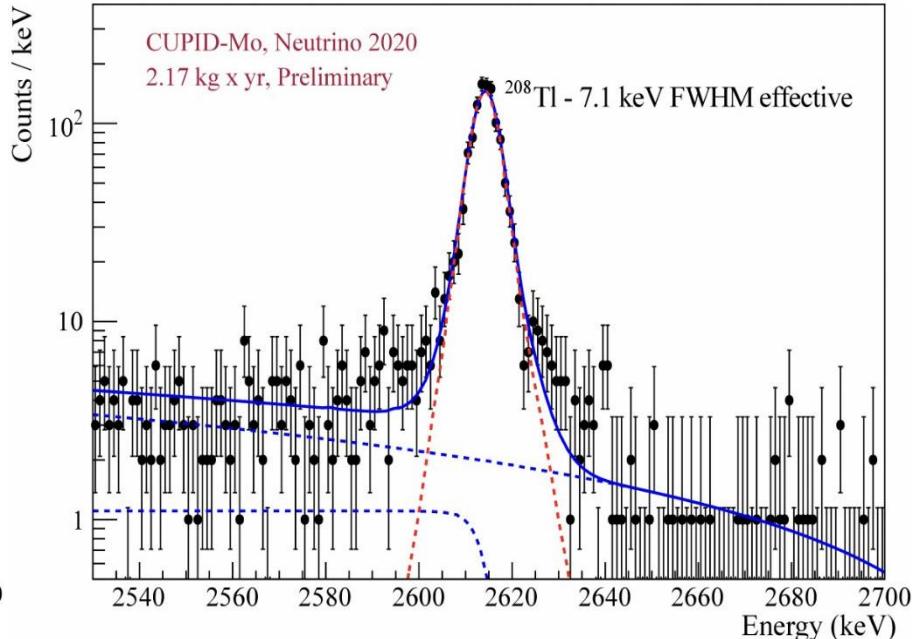
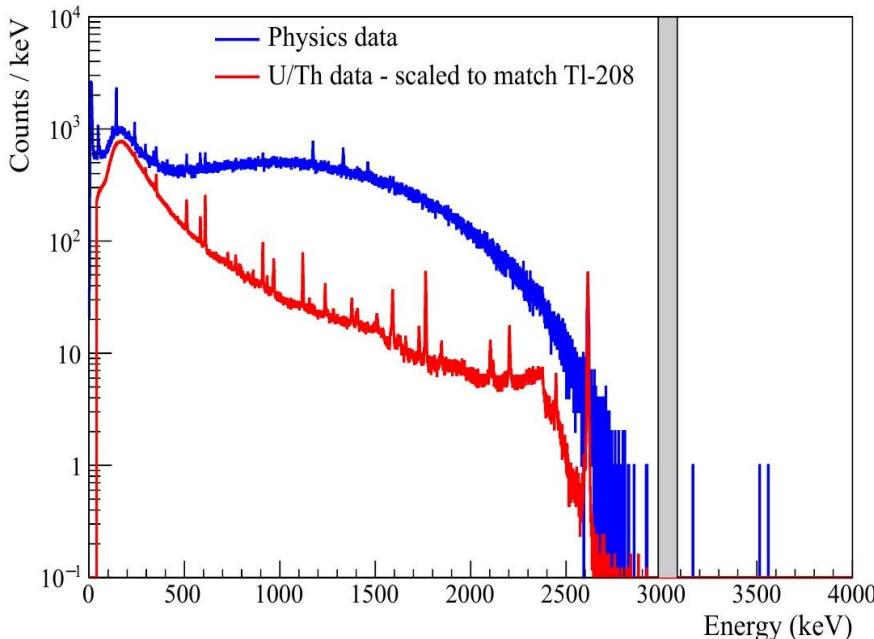


CUPID-Mo physics spectrum

- Blinded region:
 $Q_{\beta\beta} \pm 50$ keV
- Dominant $2\nu\beta\beta$ spectrum
- Most γ lines from external background sources
- Very few counts > 3 MeV after PID cut

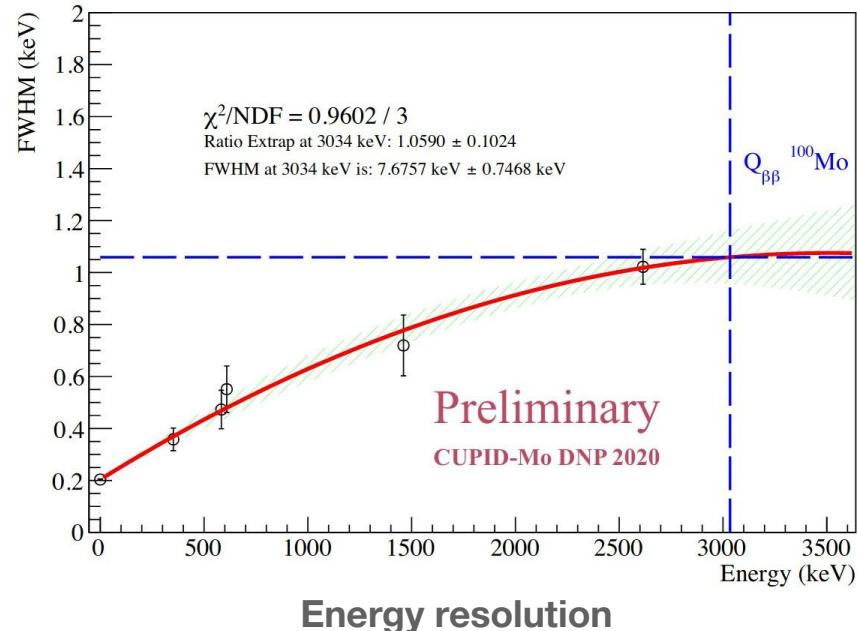
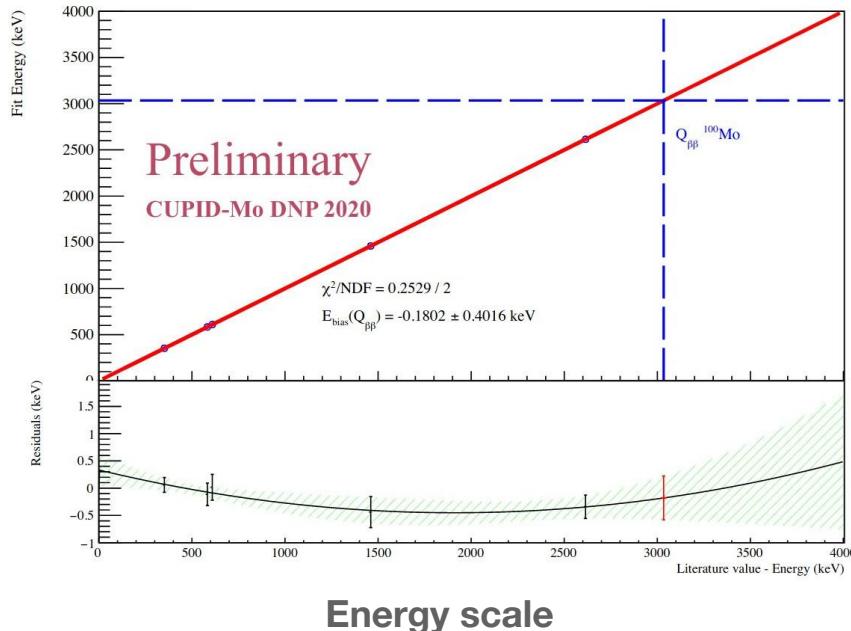


Peak-shape fit on physics data



- Fit visible peaks in physics data to study:
 - residual non-linearities
 - resolution on physics data
- Peak shape fit with Gaussian + smeared step (multi-Compton) + linear background

Energy scale and resolution in physics data



- Residuals fit with 2nd order polynomial
- Extrapolation to $Q_{\beta\beta}$ compatible with 0
- Fitted ratio of resolution in physics and calibration peaks
- Phenomenological fit to 2nd order polynomial
- **FWHM at $Q_{\beta\beta}$:** $7.6 \pm 0.7 \text{ keV}$
→ Sub-optimal noise conditions

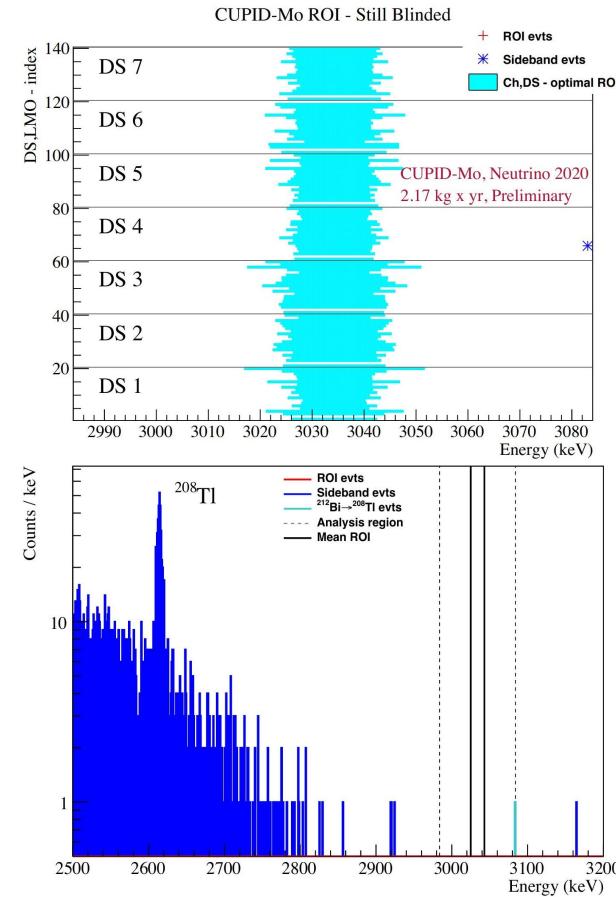
CUPID-Mo 0νββ decay analysis

0νββ decay analysis recipe

- Bayesian counting analysis in $Q_{\beta\beta} \pm 50$ keV
 - Side bands with background only
 - Central ROI with signal+background
- Two alternative background modelizations:
 - Flat+exponential (extrapolated from calibration spectrum)
 - Flat only component
- Resolution is channel-dataset dependent
 - ROI optimized independently
 - Mean ROI width: 17.9 keV

Unblinding

- First step: unblind 25 keV side-bands and check for surprises
 - One event found, so proceed
- Second step: unblind central 50 keV band
 - No events found



CUPID-Mo $0\nu\beta\beta$ decay analysis

Analysis recipe:

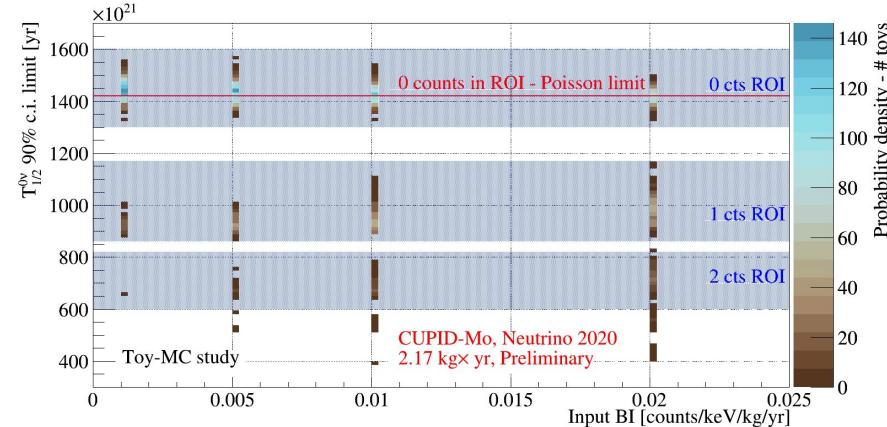
- Likelihood: $\mathcal{L} = \prod_i e^{-\lambda_i} \cdot \lambda_i^{n_i} / n_i!$
where i = band index = left side band / central ROI / right side band
- $\lambda_i = \Gamma_{0\nu} \cdot N_A \cdot f_{100} \cdot \epsilon_{\text{containment},i} \cdot \epsilon_{\text{analysis}} \cdot mt/m_a$ → Signal
 $+ BI \cdot mt \cdot \Delta E_i$ → Flat background
 $+ \tau \cdot A \cdot mt \cdot [\exp(-(E_{\min} - Q_{\beta\beta})/\tau) - \exp(-(E_{\max} - Q_{\beta\beta})/\tau)]$ → Exponential background (optional)
- Signal component in both ROI and side bands, but with very different containment
- Side bands used to evaluate background amplitude, central ROI to evaluate signal contribution
- Systematics implemented as additional nuisance parameters
- Gaussian prior on exponential amplitude (A) and decay constant (τ)
- Result on $0\nu\beta\beta$ decay rate extracted
marginalizing over the nuisance parameters

Systematic	Value	Prior
Isotopic fraction	0.966 ± 0.002	Gaussian
$0\nu\beta\beta$ containment MC	1.000 ± 0.015	Gaussian
$0\nu\beta\beta$ cont. detector response	0.95-1.00	Flat
Analysis efficiency	0.906 ± 0.004	Gaussian
Light yield selection	0.998-1.008	Flat

CUPID-Mo $0\nu\beta\beta$ decay analysis

Sensitivity study

- Toy spectra with various background levels:
→ BI = 10^{-3} , $5 \cdot 10^{-3}$, 10^{-2} , $2 \cdot 10^{-2}$ counts/keV/kg/yr
- Amplitude of exponential fixed to its best fit
- Poisson smearing on both flat and exponential background component
- 1000 toy-MC spectra generated for each BI



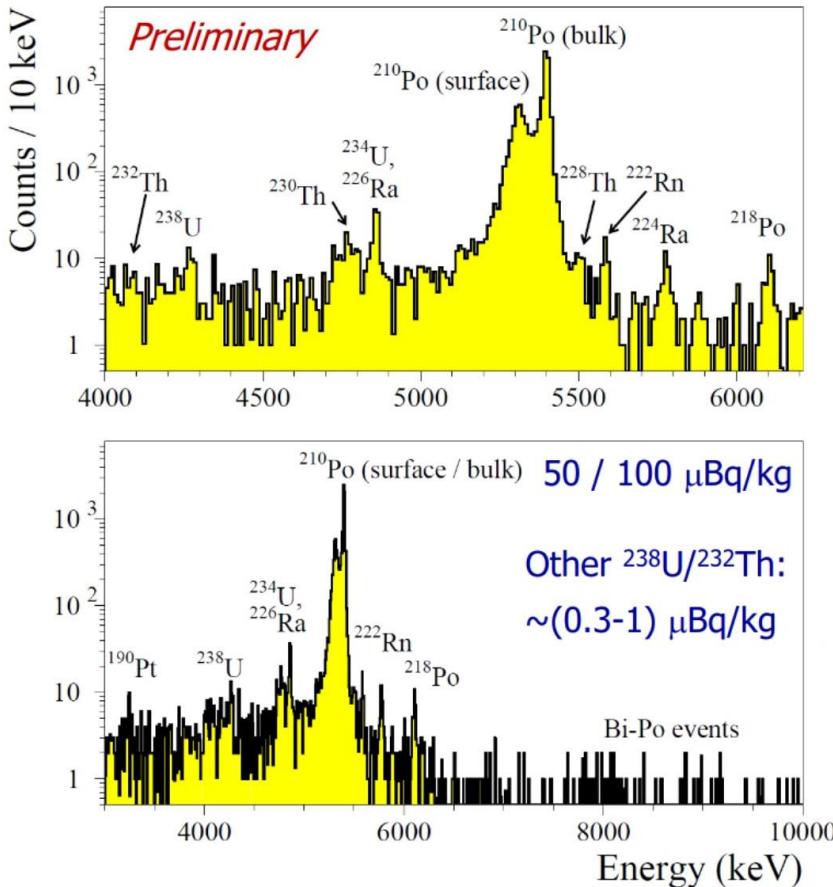
Results

- $T_{1/2}^{0\nu} > 1.5 \cdot 10^{24}$ yr @ 90% C.I.
 - $m_{\beta\beta} < 0.3\text{-}0.5$ eV (depending on NME)
 - BI $O(10^{-3})$ counts/keV/kg/yr
- Best result so far in ^{100}Mo !
- 4th most stringent limit with just 1.19 kg·yr of ^{100}Mo !
- Precise evaluation with background model ongoing

CUPID-Mo is a real experiment, not just a demonstrator!

[CUPID-Mo, arXiv:2011.13243](#)
[\(submitted to PRL\)](#)

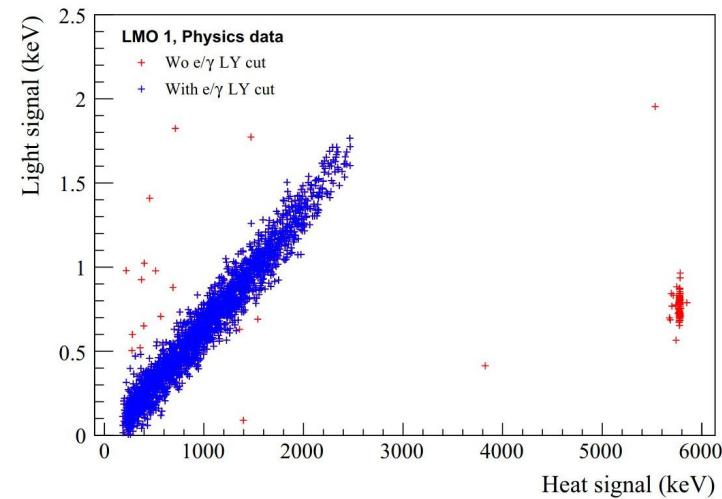
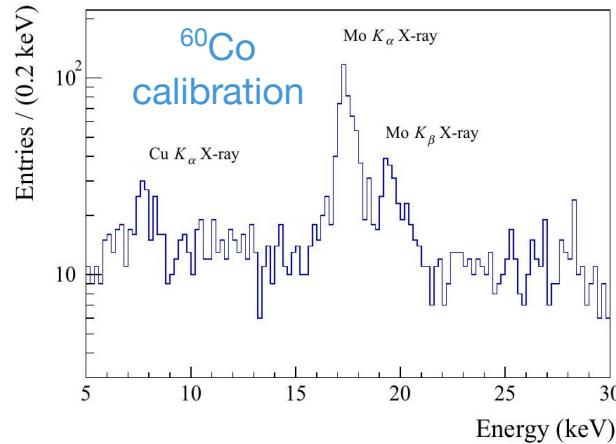
CUPID-Mo crystal purity



Chain	Nuclide	Activity [$\mu\text{Bq/kg}$]
^{232}Th	^{232}Th	0.22(9)
	^{228}Th	0.38(9)
	^{224}Ra	0.34(9)
	^{212}Bi	0.22(7)
^{238}U	^{238}U	0.35(10)
	$^{234}\text{U} + ^{226}\text{Ra}$	1.22(17)
	^{230}Rh	0.48(12)
	^{222}Rn	0.47(10)
	^{218}Po	0.35(9)
	^{210}Po	98(6)
	^{190}Pt	0.19(8)

[D. Poda, Poster @ Neutrino 2020](#)

CUPID-Mo PID performance



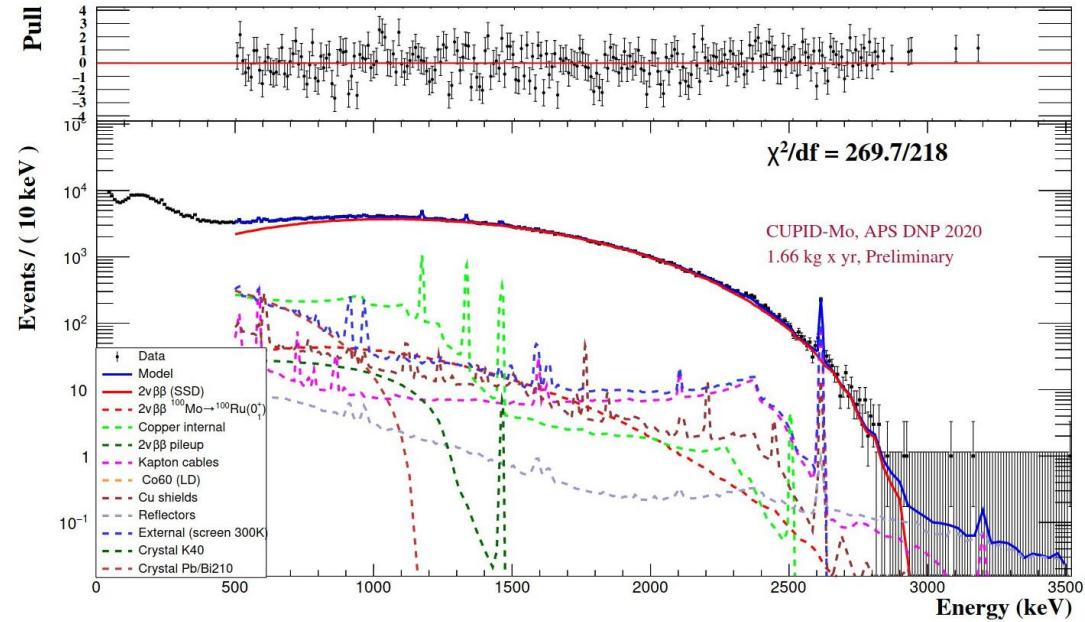
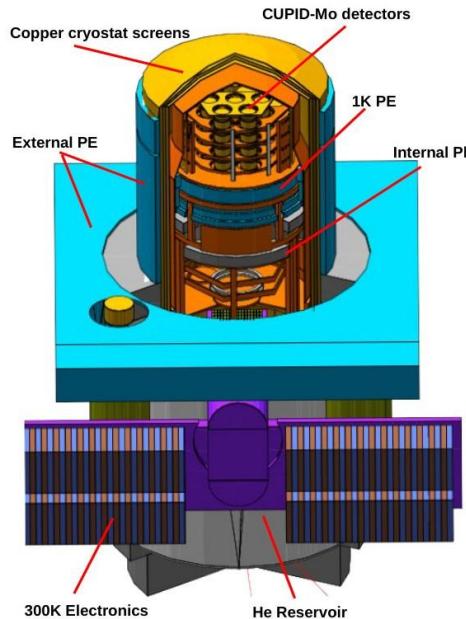
- PID performance quantified in terms of discrimination power:

$$DP = \frac{\mu_{\gamma} - \mu_{\alpha}}{\sigma_{\gamma}^2 + \sigma_{\alpha}^2}$$

- Median DP=15, worst detector DP=6.3
 - a rejection >99.9% for all detectors
 - β/γ efficiency > 99.9% for all detectors

CUPID-Mo Collaboration, EPJ C80 (2020) 44

CUPID-Mo work-in-progress

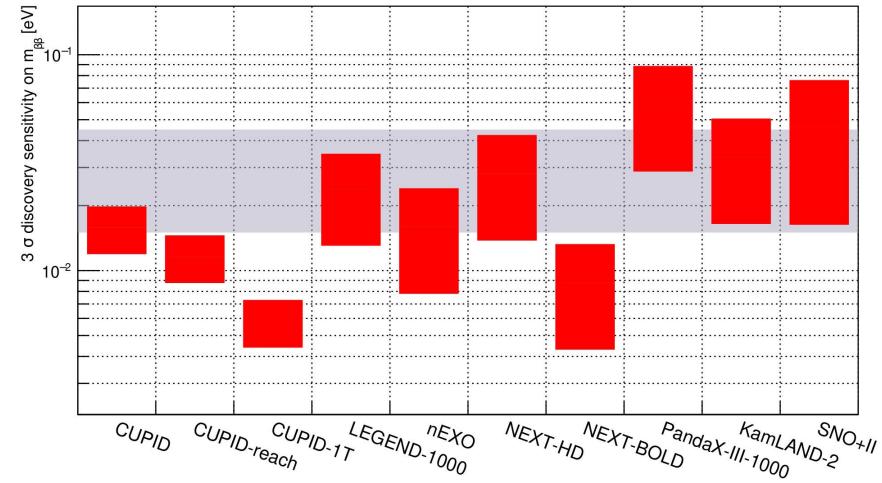


Full background-model fit will allow:

- a precise measurement of the LMO bulk and surface contamination
- a precise measurement of ^{100}Mo $2\nu\beta\beta$ decay half life to ground and excited states
- a test of nuclear physics models for the prediction of the $2\nu\beta\beta$ decay spectrum

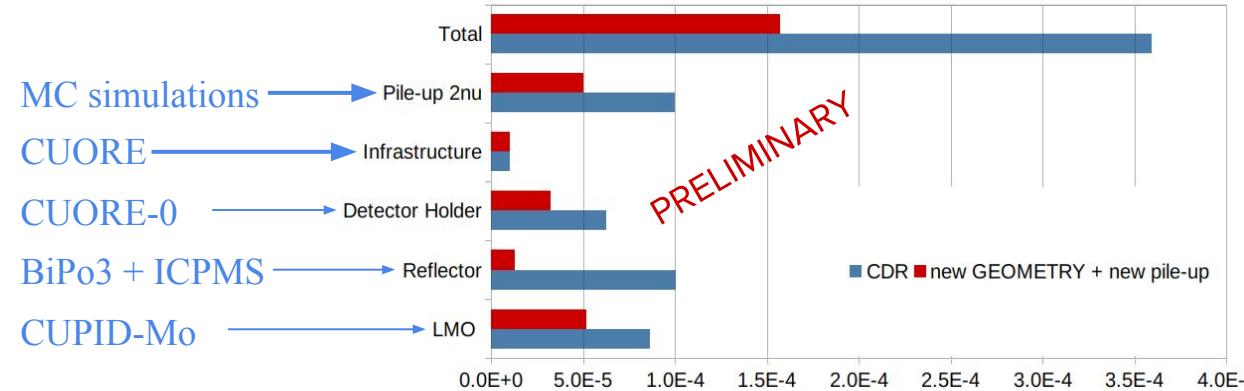
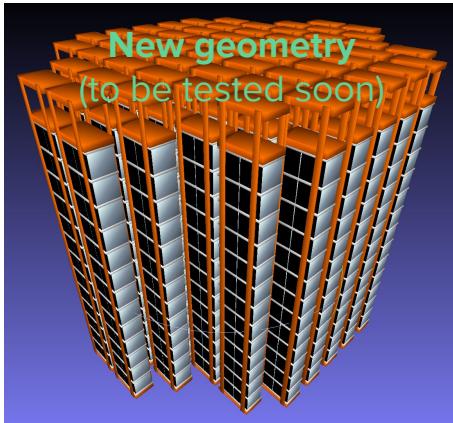
CUPID

- CUPID pre-CDR: very conservative
 - Exactly what we could start building today.
 - **No improvement assumed!**
- CUPID reach: assume improvement at reach before construction
 - Improved signal timing with TES on LD
 - Improved radiopurity
 - Zero-bkg condition: $2 \cdot 10^{-5}$ counts/keV/kg/yr
- CUPID 1 ton: new 4x larger cryostat
 - 1 ton of ^{100}Mo , plus possibly other isotopes
 - BI: $5 \cdot 10^{-6}$ counts/keV/kg/yr

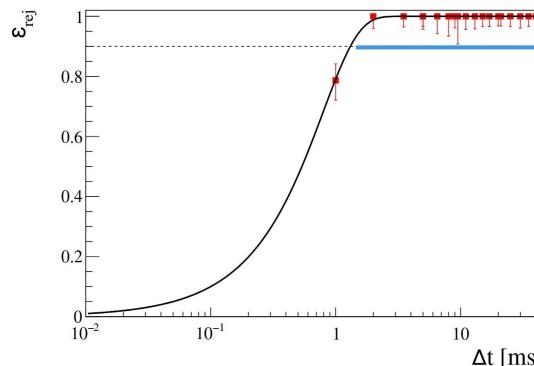


Parameter	CUPID Baseline	CUPID-reach	CUPID-1T
Crystal	$\text{Li}_2^{100}\text{MoO}_4$	$\text{Li}_2^{100}\text{MoO}_4$	$\text{Li}_2^{100}\text{MoO}_4$
Detector mass (kg)	472	472	1871
^{100}Mo mass (kg)	253	253	1000
Energy resolution FWHM (keV)	5	5	5
Background index (counts/(keV·kg·yr))	10^{-4}	2×10^{-5}	5×10^{-6}
Containment efficiency	79%	79%	79%
Selection efficiency	90%	90%	90%
Livetime (years)	10	10	10
Half-life exclusion sensitivity (90% C.L.)	1.5×10^{27} y	2.3×10^{27} y	9.2×10^{27} y
Half-life discovery sensitivity (3 σ)	1.1×10^{27} y	2×10^{27} y	8×10^{27} y
$m_{\beta\beta}$ exclusion sensitivity (90% C.L.)	10–17 meV	8.2–14 meV	4.1–6.8 MeV
$m_{\beta\beta}$ discovery sensitivity (3 σ)	12–20 meV	8.8–15 meV	4.4–7.3 meV

To-do list for CUPID



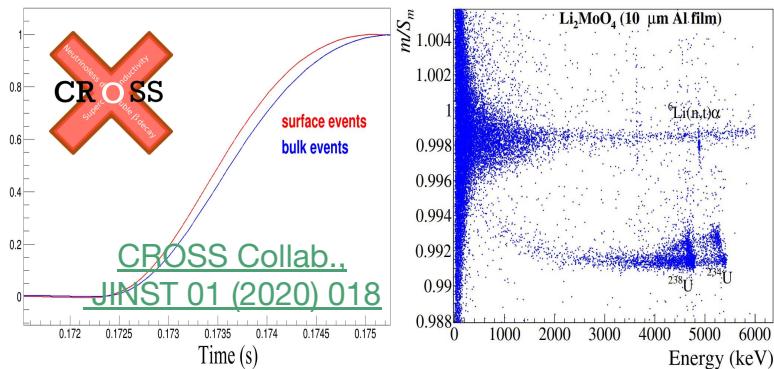
Suppression of $2\nu\beta\beta$ background



< 2 ms resolution on LMO
with space for improvement
→ Higher sampling freq.
→ LD are faster!

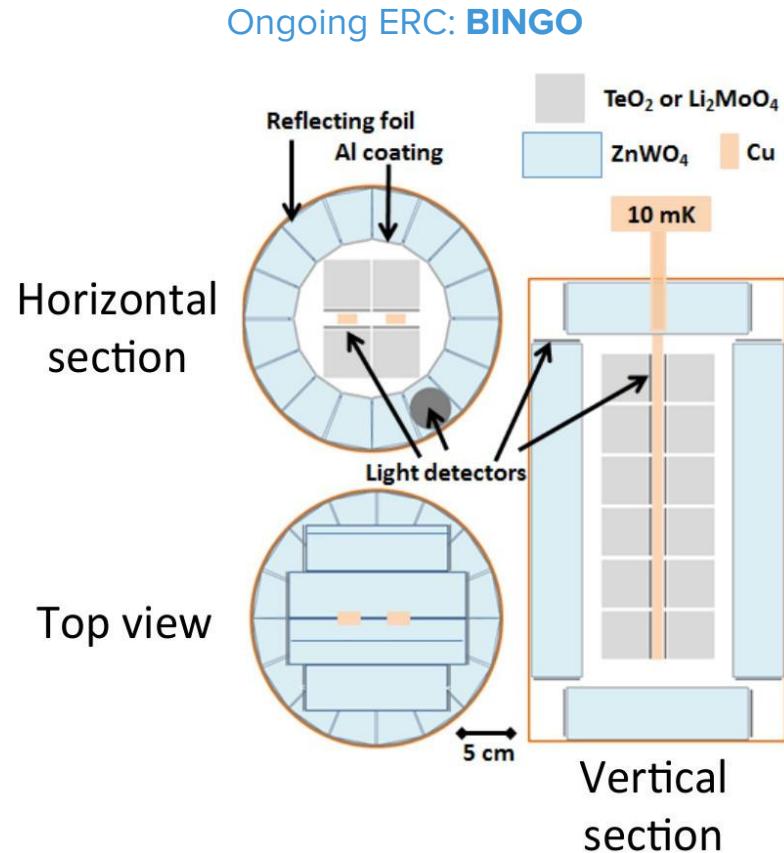
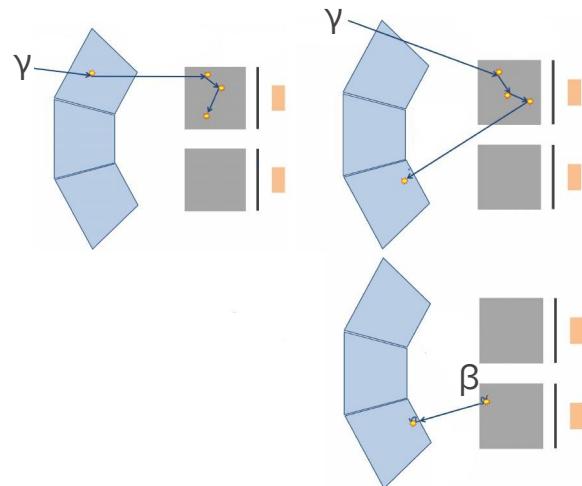
[CUPID Collaboration.](#)
[arXiv:2011.11726](#)

CROSS: pulse shape on heat channel



How do we proceed beyond CUPID?

- Increase mass
→ Easy, just need to find money 😊
- Reduce background
→ Active shield, active crystal mounting
→ Faster and more sensitive LDs,
e.g. TES or Neganov-Luke assisted LDs
- Multi-isotope approach allows confirmation of discovery with same setup



THANK YOU!